## The Generalized Gibbs Ensemble (GGE) via the Hilbert space Monte Carlo sampling approach

Vincenzo Alba<sup>1</sup> and Maurizio Fagotti<sup>2</sup>

<sup>1</sup>International School for Advanced Studies (SISSA), Via Bonomea 265, 34136, Trieste, Italy, INFN, Sezione di Trieste <sup>2</sup>Département de Physique, Ecole normale superieure, CNRS, 24 rue Lhomond, 75005 Paris, France (Dated: July 7, 2015)

By combining classical Monte Carlo simulations and Bethe ansatz techniques we numerically construct the Truncated Generalized Gibbs Ensemble (TGGE) for the spin- $\frac{1}{2}$  isotropic Heisenberg (XXX) chain. The key idea is to sample the Hilbert space,i.e., the eigenstates, of the model with the appropriate GGE probability measure. The method can be trivially extended to other integrable systems, such as the Lieb-Liniger model. We benchmark our approach focusing on GGE expectation values of several local observables. We compare our results with the Generalized Thermodynamic Bethe Ansatz (GTBA), finding spectacular agreement. Although the method is devised for finite-size chains, finite-size effects decay exponentially with system size, and moderately large chains are sufficient to extract thermodynamic quantities. Remarkably, the it is possible to extract in a simple way the steady-state root distributions, which, in the Bethe ansatz formalism, encode complete information about the GGE expectation values in the thermodynamic limit. Finally, it is straightforward to modify the method to simulate GGE extensions, in which, instead of the conserved charges, one includes arbitrary functions of the Bethe roots.

Introduction.— The issue of how statistical ensembles arise from the out-of-equilibrium dynamics in isolated quantum many-body system is still a fundamental, yet challenging, problem. The main motivation of the renewed interest in this topic is the high degree of control reached in out-of-equilibrium experiments with cold atomic gases<sup>50–63</sup>. The paradigm experiment is the so-called gloabal quantum quench<sup>4</sup>, in which a system is initially prepared in an eigenstate  $|\Psi_0\rangle$  of a many-body Hamiltonian  $\mathcal{H}$ . Then a global parameter of  $\mathcal{H}$  is suddenly changed, and the system is let to evolve unitarily under the new Hamiltonian  $\mathcal{H}'$ . At long times after the quench the system is expected to equilibrate, due to dephasing 10 more, Eisert?? la citazione e' a cazzo, as confirmed by experiments cite experiments. On the other hand, in integrable models the presence of non-trivial local conserved quantities, besides the energy, strongly affects the dynamics and the nature of the steady state. As for now, it is still unclear wheather such steady-state can be described by a statistical ensemble, and how to construct it.

Recently, it has been proposed that the stationary/equilibrium value of a generic local operator  $\hat{\mathcal{O}}$  is described by a Generalized Gibbs Ensemble<sup>3</sup> (GGE) as  $\langle \hat{\mathcal{O}} \rangle \equiv \text{Tr}(\mathcal{O}\rho^{GGE})$ . Here  $\rho^{GGE}$  extends the Gibbs density matrix by including all the local conserved quantities  $\hat{\mathcal{I}}_j$  (charges) as

$$\rho^{GGE} = Z^{-1} \exp\left(-\lambda_j \hat{\mathcal{I}}_j\right),\tag{1}$$

where we assume summation over the repeated index j, and Z is a normalization factor. In (1) the  $\lambda_j$  are Lagrange multipliers to be fixed by imposing  $\langle \Psi_0 | \hat{\mathcal{I}}_j | \Psi_0 \rangle = \langle \hat{\mathcal{I}}_j \rangle$ , and  $\hat{\mathcal{I}}_2 = \mathcal{H}'$  is the post-quench Hamiltonian. In realistic situations one deals with the truncated GGE (TGGE), i.e., considering only a finite subset of the charges **citations**.

While the validity of the GGE has been largely confirmed in non-interacting field theories **cite papers**, in interacting ones the scenario is far less clear. For Bethe ansatz solvable models the so-called quench-action method<sup>41</sup> **cite more** allows for an exact treatement of the steady state, provided that the overlap

between the initial state  $|\Psi_0\rangle$  and the eigenstates of  $\mathcal{H}'$  are known. Surprisingly, in several cases the quench-action is in disagreement with the TGGE **citations**, whereas it seems to be supported by numerical simulations<sup>119</sup>. While one might argue that the GGE can be "repaired" by enlarging the set of charges included in (1), no quantitative study in this direction has been conducted yet. One intriguing possibility is that local charges are not sufficient, and one has to include quasi-local ones<sup>122</sup> **more citations**.

On the other hand, numerical methods, such as the time dependent density matrix renormalization group 123,124 (tDMRG), have been mostly used to simulate the post-quench dynamics in microscopic models. However, no numerical attempt to explore the GGE itself has been undertaken yet. The aim of this work is to provide a Monte-Carlo-based framework for studying the GGE, and its possible extensions, in Bethe ansatz solvable models. Although we restrict ourselves to finite-size systems, thermodynamic quantities can be extracted by a standard finite-size scaling analysis. Moreover, since we observe that finite-size corrections typically decay exponentially with system size, moderately large systems are sufficient to access the thermodynamic limit. The method relies on the detailed knowledge of the Hilbert space structure provided by the Bethe ansatz formalism, and on the Bethe-Takahashi (B-T) equations cite Takahashi book, revies. The key idea is to sample the model Hilbert space according to the GGE probability measure given in (1). We should mention that for the Gibbs ensemble the same idea has been explored in Ref. 121. The method allows one to obtain GGE expectation value for generic observables, provided that their expression in terms of the roots of the B-T equations (rapidities) are known. Remarkably, it is also possible to extract the steadystate rapidity distributions, which encode the complete information about ensemble averages in the thermodynamic limit. It is also straightforward to extend the GGE including arbitrary functions of the rapidities. This could be useful, for instance, to explore the effect of quasi-local charges We should check how their expressions in terms of the rapidities look like. Finally, we should mention that GGE averages of local observables can be computed using exact diagonalization or Quantum Monte Carlo. However, both these methods require the operatorial expression of the conserved charges (see Re. 89 for the *XXX* chain), while the approach presented here relies only on their expression (typically simple) in terms of the B-T roots.

Here we perform a detailed benchmark of the approach focusing on the spin- $\frac{1}{2}$  isotropic Heisenberg chain (XXX) chain) that is the venerable prototype of integrable models cite Bethe and more. We restrict ourselves to several TGGEs (cf. (1)) constructed including  $\hat{\mathcal{I}}_2, \hat{\mathcal{I}}_3, \hat{\mathcal{I}}_4$ , and varying the associated Lagrange multipliers  $\lambda_i$ . We focus on the GGE expectation values  $\langle \hat{\mathcal{I}}_j \rangle$  and the variance of their ensemble fluctuations  $\sigma^2(\hat{\mathcal{I}}_j) \equiv \langle \hat{\mathcal{I}}_j^2 \rangle - \langle \hat{\mathcal{I}}_j \rangle^2$ . Both these quantities are related to well-known physical observables, such as the energy density, the energy current, the specific heat, and the thermal Drude weight Cite something since it is non-trivial quantity. We also compute the average magnetization and the spin susceptibility. We find that the Monte Carlo data for moderately large chains perfectly agree with the thermodynamic limit results obtained using the Generalized Thermodynamic Bethe Ansatz (GTBA) approach. This is the first direct numerical check of the GTBA for the XXX chain. Finally, we extract the first few rapidity distributions corresponding to the the Gibbs ensemble at several temperatures, and the GGE. In both cases finite-size effects are negligible, especially for small rapidities, which are the relevant ones to describe the long-wavelength physics I didn't check it properly but it is **trivial**. For the Gibbs ensemble we compare our results with the standard finite-temperature Thermodynamic Bethe Ansatz (TBA), finding perfect agreement.

The Heisenberg spin chain.— The isotropic spin- $\frac{1}{2}$  Heisenberg (XXX) chain with L sites is defined by the Hamiltonian

$$\mathcal{H} \equiv J \sum_{i=1}^{L} \left[ \frac{1}{2} (S_i^+ S_{i+1}^- + S_i^- S_{i+1}^+) + S_i^z S_{i+1}^z \right], \quad (2)$$

where  $S_i^\pm \equiv (\sigma_i^x \pm i\sigma_i^y)/2$  are spin operators acting on the site  $i, S_i^z \equiv \sigma_i^z/2$ , and  $\sigma_i^{x,y,z}$  the Pauli matrices. We fix J=1 and use periodic boundary conditions, identifying sites L+1 and 1. The total magnetization  $S_T^z \equiv \sum_i S_i^z = L/2 - M$ , with M number of down spins (particles), commutes with (2), and it is used to label its eigenstates.

In the Bethe ansatz each eigenstate of (2) is univocally identified by M parameters  $\{x_{\alpha} \in \mathbb{C}\}_{\alpha=1}^{M}$ . In the thermodynamic limit  $L \to \infty$  they form "string" patterns along the imaginary direction of the complex plane (string hypothesis)**cita Bethe Takahashi**. Strings of length  $1 \le n \le M$  (n-strings) are parametrized as  $x_{n;\gamma}^{j} = x_{n;\gamma} - i(n-1-2j)$ , with  $j=0,1,\ldots,n-1$ . Here  $x_{n;\gamma} \in \mathbb{R}$  is the real part of the string (string center), j labels the different string components, and  $\gamma$  denotes strings of the same length but with different centers. The string hypothesis is not correct for finite chains, although deviations typically decay exponentially with L. Physically, the n-strings correspond to eigenstate components containing bound states of n particles. The string centers  $x_{n;\gamma}$  are solu-

tions of the Bethe-Takahashi equations

$$L\vartheta_n(x_{n;\gamma}) = 2\pi I_{n;\gamma} + \sum_{(m,\beta)\neq(n,\gamma)} \Theta_{m,n}(x_{n;\gamma} - x_{m;\beta}).$$
 (3)

Here  $\vartheta_n(x) \equiv 2\arctan(x/n), \ \Theta_{m,n}(x)$  is the scattering phase between different rapidities, and  $I_{n;\gamma} \in \frac{1}{2}\mathbb{Z}$  are the so-called Bethe-Takahashi quantum numbers. The  $I_{n;\gamma}$  obey the upper bound  $|I_{n;\gamma}| \leq I_{\text{MAX}}(n,L,M)$ , with  $I_{\text{MAX}}$  a known function of n,M,L. Every choice of  $I_{n;\gamma}$  identifies an eigenstate of (2). Notice that each eigenstate contains strings of different lengths. We define the "string content" of an eigenstate as  $\mathcal{S} \equiv \{s_1,\ldots,s_M\}$ , with  $0 \leq s_n \leq \lfloor M/n \rfloor$  the number of n-strings. By definition one has  $\sum_j j s_j = M$ .

Besides the total magnetization and the momentum, the XXX chain has non-trivial *local* conserved charges  $\mathcal{I}_j$ , with  $[\mathcal{I}_j, \mathcal{I}_k] = 0 \,\forall j, k$ . These are obtained as

$$\mathcal{I}_{j+1} \equiv \frac{i}{(j-1)!} \frac{d^j}{dy^j} \log(\Lambda(\lbrace x_{n;\gamma} \rbrace, y)) \bigg|_{y=i}.$$
 (4)

Here  $\Lambda$  in the Algebraic Bethe ansatz is the eigenvalue of the quantum transfer matrix T(y), with y the spectral parameter. The dependence of  $\Lambda$  on the rapidities  $x_{n;\gamma}$  is known. One can check that  $\mathcal{I}_2 = \mathcal{H}$ . The range of  $\mathcal{I}_j$  increases linearly with j, i.e., larger j correspond to less local charges. Remarkably, the eigenvalues of  $\mathcal{I}_j$  over a generic eigenstate are obtained by summing the contributions of the different string sectors independently. For instance, the energy is obtained as  $E = \sum_n E_n$ , where  $E_n = 2\sum_\gamma n/(n^2 + x_{n;\gamma}^2)$ .

Hilbert space Monte Carlo sampling.— For a finite chain the GGE ensemble can be generated by sampling the eigenstates of (2) with the probability (1). This can be done efficiently using Monte Carlo. One starts with an initial M particle eigenstate of (2), with string content  $S = \{s_1, \ldots, s_M\}$ , identified by Bethe-Takahashi quantum number configuration  $C = \{I_{n;\gamma}\}_{n=1}^{M}$  ( $\gamma = 1, \ldots, s_n$ ). Let us denote the expectation values of the conserved charges as  $\{\mathcal{I}_j\}$ . The basic idea is to generate a new eigenstate with a Metropolis update. Specifically, each Monte Carlo step (mcs) consists of three moves:

- 1. Choose a new particle number sector M' and a string content S' with probability P(M', S').
- 2. Generate a quantum number configuration C' compatible with the S' obtained in step 1 and solve the Bethe-Takahashi equations (3) to extract the new rapidities  $\{x_{n;\gamma}\}$ .
- 3. After calculating the expectation values of the charges  $\mathcal{I}'_j$  accept the new eigenstate with the Metropolis probability:

$$\operatorname{Min}\left(1, \frac{L - 2M' + 1}{L - 2M + 1}e^{-\sum_{j}\lambda_{j}(\mathcal{I}'_{j} - \mathcal{I}_{j})}\right). \tag{5}$$

In (5) the factor in front of the exponential takes into account that eigenstates in the same SU(2) multiplet have the same charges expectation value, i.e., the  $\mathcal{I}_j$  are SU(2) scalars. Crucially, the steps 1 and 2 are necessary in order to account for

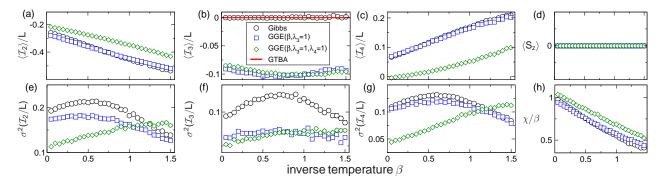


FIG. 1. The Generalized Gibbs Ensemble (GGE) for the Heisenberg spin chain with L=16 sites: numerical results obtained using the Hilbert space Monte Carlo sampling. Only the first three conserved charges  $\mathcal{I}_n$  (n=1,2,3), with associated Lagrange multipliers  $\lambda_n$ , are included in the GGE. Here  $\mathcal{I}_2$  is the Hamiltonian and  $\lambda_2 \equiv \beta$  the inverse temperature. In all the panels different symbols correspond to different values of  $\lambda_3, \lambda_4$ . The circles correspond to the Gibbs ensemble, i.e.,  $\lambda_3 = \lambda_4 = 0$ . (a) The GGE average  $\langle \mathcal{I}_2/L \rangle$  plotted as a function of  $\beta$ . (b) Variance of the GGE fluctuations  $\sigma^2(\mathcal{I}_2/L) \equiv \langle (\mathcal{I}_2/L)^2 \rangle - \langle \mathcal{I}_2/L \rangle^2$  as a function of  $\beta$ . (c)(d) and (e)(f): Same as in (a)(b) for  $\mathcal{I}_3$  and  $\mathcal{I}_4$ , respectively. In all panels the dash-dotted lines are the analytical results obtained using the Generalized Thermodynamic Bethe Ansatz (GTBA). (g) The GGE expectation value of the total magnetization  $\langle S_z \rangle$ . Notice that  $\langle S_z \rangle = 0$  due to the SU(2) invariance of the conserved charges. (h)  $\chi/\beta$  plotted versus  $\beta$ , with  $\chi$  being the magnetic susceptibility per site.

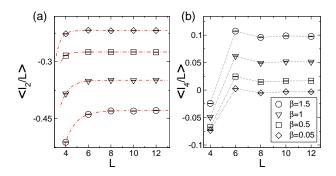


FIG. 2. Finite-size scaling of the GGE averages in the Heisenberg chain: Numerical results obtained from the Hilbert space Monte Carlo sampling. Here the GGE is constructed including  $\mathcal{I}_2, \mathcal{I}_3, \mathcal{I}_4$ , with Lagrange multipliers  $\lambda_2 = \beta, \lambda_3 = \lambda_4 = 1$ . (a)  $\langle \mathcal{I}_2/L \rangle$  plotted versus the chain size L for several values of  $\beta$ . The dash-dotted lines are exponential fits. (b) Same as in (a) for  $\mathcal{I}_4$ .

the density of states of the Heisenberg spin chain. The steps 1-3 define a Markov chain, which, after a thermalization, generates eigenstates sampled according to (1). Notice that it is straighforward to extend the algorithm to consider fixed particle number M. More interestingly, by trivially modifying (5) it is possible to simulate more exotic ensembles in which instead of the charges  $\mathcal{I}_j$ , one conseiders arbitrary functions of the rapidities. For instance, this would be useful in order to explore the effect of quasi-local charge on the GGE.

We should mention that a similar method has been developed in Ref... to construct the Gibbs ensemble in the Heisenberg spin chain.

The GGE expectation values  $\langle \mathcal{O} \rangle$  are as the average of the expectation values of  $\mathcal{O}$  over the eigenstates  $|\{x_{n;\gamma}\}\rangle$  generated by the Monte Carlo as

$$\langle \mathcal{O} \rangle = \lim_{N_{\text{mes}} \to \infty} \frac{1}{N_{\text{mes}}} \sum_{\{x_{n;\gamma}\}} \langle \{x_{n;\gamma}\} | \mathcal{O} | \{x_{n;\gamma}\} \rangle,$$
 (6)

where  $N_{
m mcs}$  is the number of Monte Carlo steps, i.e. the number of eigenstates in the sum.

Local observables.— The validity of the Monte Carlo method is illustrated in Fig. 3 considering the GGE expectation values of the charge densities  $\langle \mathcal{I}_i/L \rangle$  (panels (a)-(c) in the Figure) and the variance of their ensemble fluctuations  $\sigma^2(\mathcal{I}_j) \equiv \langle \mathcal{I}_j^2 \rangle - \langle \mathcal{I}_j \rangle^2$  (panels (d)-(f)). Finally, panels (g)(h) plot the total magnetization  $\langle S_z \rangle$  (i.e., the particle number) and the spin susceptibility  $\chi$  (particle number fluctutations). Notice that  $\sigma^2(\mathcal{I}_2)$  is related to the specific heat,  $\mathcal{I}_3 \equiv \sum_{\alpha\beta\gamma} \epsilon_{\alpha\beta\gamma} \sigma_i^{\beta} \sigma_{i+1}^{\gamma} \sigma_{i+2}^{\alpha}$  is the energy current, and  $\sigma^2(\mathcal{I}_3)$  is related to the energy Drude weight. Here the data are for the truncated TGGE constructed with the first three charges  $\mathcal{I}_2, \mathcal{I}_3, \mathcal{I}_4$ . We consider several values of the Lagrange multipliers, namely  $\lambda_3 = \lambda_4 = 0$  (Gibbs ensemble, circles in the Figure),  $\lambda_3=1$  and  $\lambda_4=0$  (squares), and  $\lambda_3 = \lambda_4 = 1$  (rhombi). All our results are plotted versus the inverse temperature  $\lambda_2=\beta$ . The data are Monte Carlo results for  $N_{\rm mcs}=5\cdot 10^5$ . In most of the cases, especially for small  $\beta$  the Monte Carlo error bars are small than the symbols. As expected, the different ensemble give different expectation values, implying that the local observables we consider are able to distinguish different GGEs. Notice that in panel (b)  $\langle \mathcal{I}_3 \rangle = 0$  for the Gibbs ensemble due to the parity invariance of  $\mathcal{I}_i$  with even j, while in (d)  $\langle S_z \rangle = 0$  due to the SU(2) symmetry of (2). In all the panels in Fig. 3 the continuous lines are the analyttic results obtained in the thermodynamic limit by solving the GTBA equations. These which fully match the Monte Carlo data, which signals that the finite-size effects are negligible already for L=16, at least for the values of the  $\lambda_i$  considered.

The finite-size corrections are more carefully investigated in Fig. 2. Fig. 2 plots  $\langle \mathcal{I}_2 \rangle$  and  $\langle \mathcal{I}_4 \rangle$  (panels (a) and (b), respectively) versus  $\beta$ . Here we focus on the TGGE with  $\lambda_2=\beta,\lambda_3=0$  and  $\lambda_4=1$ . Panel (a) demonstrates that finite-size effects decay exponentially with L for any  $\beta$ . Clearly, corrections are larger at lower temperature, as ex-

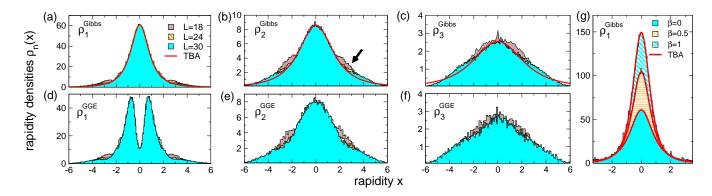


FIG. 3. The rapidity densities  $\rho_n(x)$  (for n=1,2,3) for the infinite temperature Gibbs (panels (a)-(c)) and the GGE equilibrium states (panels (d)-(f)): Numerical results for the Heisenberg spin chain obtained using the Hilbert space Monte Carlo sampling. Here the GGE is constructed including only  $\mathcal{I}_2$  and  $\mathcal{I}_4$  with fixed Lagrange multipliers  $\lambda_2=0$  and  $\lambda_4=1$ . In all the panels the data are the histograms of the n-strings rapidities sampled in the Monte Carlo. The width of the histogram bins is  $\Delta x=2/L$ . In each panel different histograms correspond to different chain sizes L. All the histograms are divided by  $10^3$  for convenience. In (b) the arrow is to highlight the finite-size effects. In panels (a)-(c) the lines are the Thermodynamic Bethe Ansatz (TBA) results. (g) Finite-temperature effects: Monte Carlo data for  $\rho_1^{\text{Gibbs}}$  for different values of the inverse temperature  $\beta$ .

pected. Moreover, they increase with the range of the operator as shown in panel (b), although the behavior remains exponential.

Extracting the rapidity densities.— In the thermodynamic limit in each n-string sector the roots of (3) become dense. Thus, each eigenstate is characterized by the root distribution  $\boldsymbol{\rho} \equiv \{\rho_n\}_{n=1}^\infty$ . Formally, the  $\rho_n$  are defined as  $\rho_n = \lim_{L \to \infty} [L(x_{n;\gamma+1} - x_{n;\gamma})]^{-1}$ . For a generic observable  $\mathcal{O}$ , the GGE average becomes a functional integral as

$$\operatorname{Tr}\left\{\exp\left(\lambda_{j}\mathcal{I}_{j}\right)\mathcal{O}\right\} \to \int \mathcal{D}\boldsymbol{\rho} \exp\left(S[\boldsymbol{\rho}] + \lambda_{j}\mathcal{I}_{j}[\boldsymbol{\rho}]\right)\mathcal{O}[\boldsymbol{\rho}].$$
 (7)

Here  $S[\rho]$  is the Yang-Yang entropy.  $S[\rho]$  counts the number of eigenstates leading to the same  $\rho$ , and it is extensive. In (7)  $\mathcal O$  is assumed that  $\mathcal O$  becomes a smooth function of  $\rho$  in the thermodynamic limit. Since both S and  $\mathcal I_j$  are extensive, the integral in (7) is dominated by the saddle point  $\rho^{sp}$ , with  $\delta(S+\lambda_j\mathcal I_j)/\delta\rho|_{\rho=\rho^{sp}}=0$ . Here  $\rho^{sp}$  acts as a representative state for the ensemble, and it contains the full information about the GGE equilibrium steady state. Eq. (6) suggests/implies that the representative state root densities  $\rho^{sp}_n$  can be obtained the histograms of the roots  $x_{n;\gamma}$  sampled in the Monte Carlo history, in the limit  $L\to\infty$ .

This is supported in Fig. 3 considering several GGEs. Panels (a)-(c) plot the root densities  $\rho_n^{sp}(x)$  for n=1,2,3 as a function of x for the representative state (saddle point) of the infinite-temperature Gibbs ensemble. In each panel the different histograms correspond to different chain sizes  $18 \le L \le 30$ . The data are obtained from Monte Carlo histories with  $4 \cdot 10^5$  Monte Carlo steps. The width of the histogram bins is varied with che chain size as 2/L. In all the panels the full lines are the analytic results obtained from the Thermodynamic Bethe Ansatz (TBA) (cf. (10)). Remarkably, the Monte Carlo data are in good agreement with the TBA results. This agreement is perfect for n=1, whereas it become progressively worse upon considering larger n>1 (see panels (b)(c)). Clearly, the deviations from the TBA result vanish

upon increasing the system size (see for instance the arrow in panel (b)). These finite-size effects are larger on the tails of the distributions. This is expected since large rapidities correspond to large quasi-momenta, which are more sensitive to the lattice effects. Finally, finite-size effects increase with n, i.e., with the bound state sizes. The finite-temperature Gibbs ensemble is discuss in Fig. 3 (g), focusing on  $\beta = 1/2$  and  $\beta = 1$  (the different histograms in the panel). Only results for  $\rho_1(x)$ , for a chain with L=30 are presented. The infinite temperature histogram is reported for comparison. The continuous lines are now the analytic results obtained by solving the finite-temperature TBA equations and perfectly agree with the Monte Carlo data. Upon lowering the temperature the height of the peak at x=0 increases. This reflects that at  $\beta = \infty$  the tail of the root distributions vanish exponentially, whereas for  $\beta = 0$  they are  $\sim 1/x^4$ .

Finally, panels (d)-(f) plot  $\rho_n(x)$  for the GGE ensemble. Specifically, we focus on the TGGE with the two charges  $\mathcal{I}_2,\mathcal{I}_4$  with  $\lambda_2=0$  and  $\lambda_4=1$ . In contrast with the thermal case (see (a))  $\rho_1$  exhibits a double peak structure. Similar to the infite-temperature Gibbs ensemble ((a)-(c) in the Figure), the data suggest that for L=30 finite-size effects are negligible, at least for  $-2 \le x \le 2$ .

## I. CONCLUSIONS

## II. THE STRING ROOT DENSITIES AT INFINITE TEMPERATURE

For infinite temperature the densities  $\rho_n$  are given as

$$\rho_n(x) = \frac{2}{\pi} \frac{1}{(n^2 + x^2)(x^2 + (2+n)^2)}$$
 (8)

Notice that

$$\int_{-\infty}^{+\infty} \rho_n(x) dx = \frac{1}{n(n+1)(n+2)}$$
 (9)

Icluding the first order correction to the infinite temperature result one obtains

$$\rho_n(x) = \frac{2}{\pi} \frac{1}{(n^2 + x^2)(x^2 + (2+n)^2)} - \frac{8}{\pi} \frac{n(n+2)}{(n^2 + x^2)^2 (x^2 + (2+n)^2)^2} J\beta + \mathcal{O}(J^2 \beta^2)$$
(10)

- <sup>1</sup> M. Rigol, V. Dunjko, V. Yurovsky, and M. Olshanii, Phys. Rev. Lett. **98**, 050405 (2007).
- <sup>2</sup> S. Popescu, A. J. Short, and A. Winter, Nature Physics **2**, 754 (2006).
- <sup>3</sup> M. Rigol, V. Dunjko, and M. Olshanii, Nature **452**, 854 (2008).
- <sup>4</sup> A. Polkovnikov, K. Sengupta, and M. Vengalattore, Rev. Mod. Phys. 83, 863 (2011).
- <sup>5</sup> J. Eisert., M. Friesdorf, and C. Gogolin, arXiv:1408.5148.
- <sup>6</sup> C. Kollath, A. M. Läuchli, and E. Altman, Phys. Rev. Lett. 98, 180601 (2007).
- <sup>7</sup> S. R. Manmana, S. Wessel, R. M. Noack, and A. Muramatsu, Phys. Rev. Lett. **98**, 210405 (2007).
- <sup>125</sup> P. Calabrese and J. Cardy, J. Stat. Mech. P06008 (2007).
- <sup>9</sup> M. Cramer, C. M. Dawson, J. Eisert, and T. J. Osborne, Phys. Rev. Lett. **100**, 030602 (2008).
- <sup>10</sup> T. Barthel and U. Schollwöck, Phys. Rev. Lett. **100**, 100601 (2008).
- <sup>11</sup> M. Cramer, A. Flesch, I. P. McCulloch, U. Schollwöck, and J. Eisert, Phys. Rev. Lett. **101**, 063001 (2008).
- <sup>12</sup> M. Kollar and M. Eckstein, Phys. Rev. A **78**, 013626 (2008).
- <sup>126</sup> A. Iucci and M. A. Cazalilla, Phys. Rev. A **80**, 063619 (2009).
- <sup>14</sup> S. Sotiriadis, P. Calabrese, and J. Cardy, EPL **87**, 20002 (2009).
- <sup>15</sup> G. Roux, Phys. Rev. A **79**, 021608 (2009).
- <sup>16</sup> M. Rigol, Phys. Rev. Lett. **103**, 100403 (2009).
- <sup>17</sup> M. Rigol, Phys. Rev. A **80**, 053607 (2009).
- <sup>18</sup> P. Barmettler, M. Punk, V. Gritsev, E. Demler, and E. Altman, Phys. Rev. Lett. **102**, 130603 (2009).
- <sup>19</sup> P. Barmettler, M. Punk, V. Gritsev, E. Demler, and E. Altman, New J. Phys. **12**, 055017 (2010).
- <sup>20</sup> M. Cramer and J. Eisert, New J. Phys. **12**, 055020 (2010).
- A. Flesch, M. Cramer, I. P. McCulloch, U. Schollwöck, and J. Eisert, Phys. Rev. A 78, 033608 (2008).
- <sup>22</sup> G. Roux, Phys. Rev. A **81**, 053604 (2010).
- <sup>23</sup> D. Fioretto and G. Mussardo, New J. Phys. **12**, 055015 (2010).
- <sup>24</sup> G. Biroli, C. Kollath, and A. M. Läuchli, Phys. Rev. Lett. **105**, 250401 (2010).
- <sup>25</sup> L. F. Santos and M. Rigol, Phys. Rev. E **82**, 031130 (2010).
- <sup>26</sup> M. C. Bañuls, J. I. Cirac, and M. B. Hastings, Phys. Rev. Lett. 106, 050405 (2011).
- P. Calabrese, F. H. L. Essler, and M. Fagotti, Phys. Rev. Lett. 106, 227203 (2011).
- <sup>28</sup> C. Gogolin, M. P. Mueller, and J. Eisert, Phys. Rev. Lett. **106**, 040401 (2011).
- <sup>29</sup> M. Rigol and M. Fitzpatrick, Phys. Rev. A **84**, 033640 (2011).
- <sup>30</sup> T. Caneva, E. Canovi, D. Rossini, G. E. Santoro, and A. Silva, J. Stat. Mech. (2011) P07015.
- <sup>31</sup> L. Santos, A. Polkovnikov, and M. Rigol, Phys. Rev. Lett. **107**, 040601 (2011).
- <sup>32</sup> A. C. Cassidy, C. W. Clark, and M. Rigol, Phys. Rev. Lett. **106**, 140405 (2011).

- <sup>33</sup> F. H. L. Essler, S. Evangelisti, and M. Fagotti, Phys. Rev. Lett. 109, 247206 (2012).
- <sup>34</sup> M. A. Cazalilla, A. Iucci, and M.-C. Chung, Phys. Rev. E 85, 011133 (2012).
- <sup>35</sup> J. Mossel and J.-S. Caux, New J. Phys. **14** 075006 (2012).
- <sup>36</sup> M. Rigol and M. Srednicki, Phys. Rev. Lett. **108**, 110601 (2012).
- <sup>37</sup> J. Mossel and J.-S. Caux, J. Phys. A: Math. Theor. **45**, 255001 (2012).
- <sup>130</sup> M. Fagotti and F. H. L. Essler, Phys. Rev. B **87**, 245107 (2013).
- <sup>39</sup> M. Fagotti, Phys. Rev. B **87**, 165106 (2013).
- <sup>131</sup> M. Collura, S. Sotiriadis, and P. Calabrese, Phys. Rev. Lett. 110, 245301 (2013).
- <sup>41</sup> J.-S. Caux and F. H. L. Essler, Phys. Rev. Lett. **110**, 257203 (2013).
- <sup>133</sup> M. Kormos, A. Shashi, Y.-Z. Chou, J.-S. Caux, and A. Imambekov, Phys. Rev. B 88, 205131 (2013).
- <sup>43</sup> B. Bertini, D. Schuricht, and F. H. L. Essler, arXiv:1405.4813 (2014).
- <sup>44</sup> S. Sotiriadis and P. Calabrese, J. Stat. Mech. (2014) P07024.
- <sup>45</sup> F. H. L. Essler, S. Kehrein, S. R. Manmana, and N. J. Robinson, Phys. Rev. B **89**, 165104 (2014).
- <sup>46</sup> M. Fagotti, M. Collura, F. H. L. Essler, and P. Calabrese, Phys. Rev. B 89, 125101 (2014).
- <sup>47</sup> M. Fagotti, J. Stat. Mech. (2014) P03016.
- B. Wouters, J. De Nardis, M. Brockmann, D. Fioretto, M. Rigol, and J.-S. Caux, Phys. Rev. Lett. 113, 117202 (2014).
- B. Pozsgay, M. Mestyán, M. A. Werner, M. Kormos, G. Zarànd, and G. Takács, Phys. Rev. Lett. 113, 117203 (2014).
- M. Greiner, O. Mandel, T. Hänsch, and I. Bloch, Nature (London) 419, 51 (2002).
- <sup>51</sup> T. Kinoshita, T. Wenger, and D. S. Weiss, Nature (London) **440**, 900 (2008).
- <sup>52</sup> S. Hofferberth, I. Lesanovsky, B. Fischer, T. Schumm, and J. Schiedmayer, Nature (London) 449, 324 (2007).
- <sup>53</sup> I. Bloch, J. Dalibard, and W. Zwerger, Rev. Mod. Phys. **80**, 885 (2008).
- <sup>54</sup> S. Trotzky, Y.-A. Chen, A. Flesch, I. P. McCulloch, U. Schollwöck, J. Eisert, and I. Bloch, Nature Phys. 8, 325 (2012).
- M. Gring, M. Kuhnert, T. Langen, T. Kitagawa, B. Rauer, M. Schreitl, I. Mazets, D. A. Smith, E. Demler, and J. Schmiedmayer, Science 337, 6100 (2012).
- M. Cheneau, P. Barmettler, D. Poletti, M. Endres, P. Schaua, T. Fukuhara, C. Gross, I. Bloch, C. Kollath, and S. Kuhr, Nature (London) 481, 484 (2012).
- <sup>57</sup> U. Schneider, L. Hackeruller, J. P. Ronzheimer, S. Will, S. Braun, T. Best, I. Bloch, E. Demler, S. Mandt, D. Rasch, and A. Rosch, Nature Phys. 8, 213 (2012).
- <sup>58</sup> M. Kuhnert, R. Geiger, T. Langen, M. Gring, B. Rauer, T. Kitagawa, E. Demler, D. Adu Smith, and J. Schmiedmayer, Phys. Rev. Lett. **110**, 090405 (2013).

- <sup>59</sup> T. Langen, R. Geiger, M. Kuhnert, B. Rauer, and J. Schmiedmayer, Nature Phys. 9, 640 (2013).
- <sup>60</sup> F. Meinert, M. J. Mark, E. Kirilov, K. Lauber, P. Weinmann, A. J. Daley, and H.-C. Nagerl, Phys. Rev. Lett. **111**, 053003 (2013).
- <sup>61</sup> T. Fukuhara, A. Kantian, M. Endres, M. Cheneau, P. Schaua, S. Hild, C. Gross, U. Schollwöck, T. Giamarchi, I. Bloch, and S. Kuhr, Nature Phys. 9, 235 (2013).
- <sup>62</sup> J. P. Ronzheimer, M. Schreiber, S. Braun, S. S. Hodgman, S. Langer, I. P. McCulloch, F. Heidrich-Meisner, I. Bloch, and U. Schneider, Phys. Rev. Lett. **110**, 205301 (2013).
- <sup>63</sup> S. Braun, M. Friesdorf, S. Hodgman, M. Schreiber, J. Ronzheimer, A. Riera, M. del Rey, I. Bloch, J. Eisert, and U. Schneider, arXiv:1403.7199.
- 64 J. M. Deutsch, Phys. Rev. A 43, 2046 (1991).
- 65 M. Srednicki, Phys. Rev. E **50**, 888 (1994).
- 66 M. Srednicki, J. Phys. A 29, L75 (1996).
- <sup>67</sup> M. Srednicki, J. Phys. A **32**, 1163 (1999).
- <sup>68</sup> S. Goldstein, J. L. Lebowitz, R. Tumulka, and N. Zanghí, Phys. Rev. Lett. **96**, 050403 (2006).
- <sup>69</sup> S. Goldstein, J. L. Lebowitz, C. Mastrodonato, R. Tumulka, and N. Zanghi, Proc. R. Soc. A 466, 3203 (2010).
- <sup>70</sup> S. Goldstein, J. L. Lebowitz, R. Tumulka, and N. Zanghi, Eur. Phys. J. H **35**, 173 (2010).
- <sup>71</sup> T. N. Ikeda, Y. Watanabe, and M. Ueda, Phys. Rev. E **84**, 021130 (2011).
- <sup>72</sup> T. N. Ikeda, Y. Watanabe, and M. Ueda, Phys. Rev. E **87**, 012125 (2013).
- <sup>73</sup> R. Steinigeweg, J. Herbrych, and P. Prelovšek, Phys. Rev. E 87, 012118 (2013).
- <sup>74</sup> W. Beugeling, R. Moessner, and M. Haque, Phys. Rev. E 89, 042112 (2014).
- <sup>75</sup> R. Steinigeweg, A. Khodja, H. Niemeyer, C. Gogolin, and J. Gemmer, Phys. Rev. Lett. **112**, 130403 (2014).
- <sup>76</sup> S. Sorg, L. Vidmar, L. Pollet, and F. Heidrich-Meisner, arXiv:1405.5404v2.
- W. Beugeling, R. Moessner, and M. Haque, arXiv:1407.2043.
- <sup>78</sup> V. Khemani, A. Chandran, H. Kim, and S. L. Sondhi, arXiv:1406.4863.
- <sup>79</sup> H. Kim, T. N. Ikeda, and D. Huse, arXiv:1408.0535.
- <sup>80</sup> L. Bonnes, F. H. L. Essler, and A. M. Läuchli, arXiv:1404.4062 (2014).
- <sup>81</sup> J.-S. Caux and J. Mossel, J. Stat. Mech. (2011) P02023.
- <sup>82</sup> V. Alba, M. Fagotti, and P. Calabrese, J. Stat. Mech. (2009) P10020.
- <sup>83</sup> N. Kitanine, J. M. Maillet, and V. Terras, Nucl. Phys. B **554**, 647 (1999).
- <sup>84</sup> N. Kitanine, J. M. Maillet, and V. Terras, Nucl. Phys. B **567**, 554 (2000).
- <sup>85</sup> L. Amico, R. Fazio, A. Osterloh, and V. Vedral, Rev. Mod. Phys. 80, 517 (2008).
- <sup>86</sup> M. Takahashi, *Thermodynamics of one-dimensional solvable models*, Cambridge University Press 1999.
- <sup>87</sup> C. N. Yang and C. P. Yang, J. Math. Phys. **10**, 1115 (1969).
- <sup>88</sup> M. Takahashi, Prog. Theor. Phys. **46**, 401 (1971).
- <sup>89</sup> M. P. Grabowski and P. Mathieu, Ann. Phys. N.Y. **243**, 299 (1995).
- <sup>90</sup> J. Eisert, M. Cramer, and M. B. Plenio, Rev. Mod. Phys. **82**, 277 (2009).
- <sup>91</sup> P. Calabrese, J. Cardy, and B. Doyon Eds., Special issue: Entanglement entropy in extended systems, J. Phys. A 42, 50 (2009).
- <sup>92</sup> P. Calabrese and J. Cardy, J. Phys. A **42** 504005 (2009).
- <sup>93</sup> V. E. Korepin, N. M. Bogoliubov, and A. G. Izergin, *Quantum Inverse Scattering Methods and Correlation Functions*, Cambridge

- University Press 1997.
- <sup>94</sup> X. Zotos and P. Prelovšek, Phys. Rev. B 53, 983 (1996).
- <sup>95</sup> H. Castella and X. Zotos, Phys. Rev. B **54**, 4375 (1996).
- <sup>96</sup> X. Zotos, F. Naef, and P. Prelovšek, Phys. Rev. B **55**, 11029 (1997)
- <sup>97</sup> F. C. Alcaraz, M. I. Berganza, and G. Sierra, Phys. Rev. Lett. **106**, 201601 (2011).
- 98 I. Pizorn, arXiv:1202.3336.
- <sup>99</sup> M. I. Berganza, F. C. Alcaraz, and G. Sierra, J. Stat. Mech. (2012) P01016.
- <sup>100</sup> G. Wong, I. Klich, L. A. P. Zayas, and D. Vaman, JHEP **12** (2013) 020
- <sup>101</sup> M. Storms, and R. R. P. Singh, Phys. Rev. E **89**, 012125 (2014).
- <sup>102</sup> R. Berkovits, Phys. Rev. B **87**, 075141 (2013).
- <sup>103</sup> F. H. L. Essler, A. M. Läuchli, and P. Calabrese, Phys. Rev. Lett. 110, 115701 (2013).
- <sup>104</sup> M. Nozaki, T. Numasawa, T. Takayanagi, Phys. Rev. Lett. **112**, 111602 (2014).
- <sup>105</sup> G. Ramirez, J. Rodriguez-Laguna, and G. Sierra, arXiv:1402.5015.
- F. Ares, J. G. Esteve, F. Falceto, and E. Sánchez-Burillo, arXiv:1401.5922.
- <sup>107</sup> Y. Huang, and J. Moore, arXiv:1405.1817.
- <sup>108</sup> T. Pálmai, arXiv:1406.3182.
- J. Mölter, T. Barthel, U. Schollwöck, and V. Alba, arXiv:1407.0066.
- 110 H.-H. Lai and K. Yang, arXiv:1409:1224
- J. Sato, B. Aufgebauer, H. Boos, F. Göhmann, A. Klümper, M. Takahashi, and C. Trippe, Phys. Rev. Lett. 106, 257201 (2011).
- <sup>112</sup> M. Fagotti and P. Calabrese, Phys. Rev. A **78**, 010306 (2008).
- <sup>113</sup> V. Gurarie, J. Stat. Mech. (2014) P02014.
- <sup>114</sup> M. Collura, M. Kormos, and P. Calabrese, J. Stat. Mech. (2014) P01009.
- <sup>115</sup> M. Kormos, L. Bucciantini, and P. Calabrese, EPL **107**, 40002 (2014).
- <sup>116</sup> J.-S. Caux and J.-M. Maillet, Phys. Rev. Lett. **95**, 077201 (2005).
- <sup>117</sup> J.-S. Caux, R. Hagemans and J.-M. Maillet, J. Stat. Mech. P09003 (2005).
- <sup>118</sup> J.-S. Caux, J. Math. Phys. **50**, 095214 (2009).
- B. Pozsgay, M. Mestyán, M. A. Werner, M. Kormos, G. Zaránd, and G. Takács, Phys. Rev. Lett. 113, 117203 (2014).
- B. Wouters, M. Brockmann, J. De Nardis, D. Fioretto, M. Rigol, and J.-S. Caux, Phys. Rev. Lett. 113, 117202 (2014).
- <sup>121</sup> S.-J. Gu, N. M. R. Peres, Y.-Q. Li, Eur. Phys. J. B **48**, 157 (2005).
- <sup>122</sup> E. Ilievski, M. Medejak, and T. Prosen, arXiv:1506.05049.
- <sup>123</sup> S. R. White and A. E. Feiguin, Phys. Rev. Lett. **93**, 076401 (2004).
- 124 A. J. Daley, C. Kollath, U. Schollock, and G. Vidal, J. Stat. Mech. (2004) P04005.
- <sup>125</sup> P. Calabrese and J. Cardy, J. Stat. Mech. (2007) P06008.
- <sup>126</sup> A. Iucci and M. A. Cazalilla, Phys. Rev. A **80**, 063619 (2009).
- <sup>127</sup> P. Calabrese, F. H. L. Essler, and M. Fagotti, Phys. Rev. Lett. **106**, 227203 (2011).
- <sup>128</sup> P. Calabrese, F. H. L. Essler, and M. Fagotti, J. Stat. Mech. (2012) P07016.
- P. Calabrese, F. H. L. Essler, and M. Fagotti, J. Stat. Mech. (2012) P07022.
- <sup>130</sup> M. Fagotti and F. H. L. Essler, Phys. Rev. B **87**, 245107 (2013).
- <sup>131</sup> M. Collura, S. Sotiriadis and P. Calabrese, Phys. Rev. Lett. **110**, 245301 (2013)
- <sup>132</sup> M. Collura, S. Sotiriadis and P. Calabrese, J. Stat. Mech. (2013) P09025.
- <sup>133</sup> M. Kormos, M. Collura and P. Calabrese, arXiv:1307.2142.

<sup>134</sup> G. Mussardo, Phys. Rev. Lett. **111**, 100401 (2013).