

# Constructing the GGE in integrable models: A Hilbert space Monte Carlo approach

Vincenzo Alba<sup>1</sup> and Maurizio Fagotti<sup>2</sup>

<sup>1</sup>*International School for Advanced Studies (SISSA), Via Bonomea 265, 34136, Trieste, Italy, INFN, Sezione di Trieste*

<sup>2</sup>*Département de Physique, Ecole normale supérieure, CNRS, 24 rue Lhomond, 75005 Paris, France*

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## I. INTRODUCTION

The issue of how statistical ensembles emerge from the out-of-equilibrium unitary dynamics in isolated quantum many-body system is a fundamental, yet still challenging, problem. Much of the motivation for the renewed interest in this topic originated from the possibility of simulating efficiently the out-of-equilibrium dynamics in cold atom experiments. The paradigm of out-of-equilibrium experiment is the quantum quench, in which

In a generic *global* quench in an isolated quantum many-body system let us consider the initial state  $|\Psi_0\rangle$ . In the thermodynamic limit for generic  $|\Psi_0\rangle$  the system is expected to equilibrate. The most relevant question is whether expectation values of observables can be computed in a thermodynamic ensemble.

Moreover, in integrable models the out-of-equilibrium dynamics is strongly affected by the presence of non-trivial local integral of motion.

It is believed that the equilibrium expectation values of local operators at long times are described by a Generalized Gibbs Ensemble (GGE). More quantitatively,

$$\lim_{t \rightarrow \infty} \lim_{L \rightarrow \infty} \langle \Psi(t) | \mathcal{O} | \Psi(t) \rangle = \langle \mathcal{O} \rangle_{GGE}, \quad (1)$$

with

$$\langle \mathcal{O} \rangle_{GGE} \equiv \lim_{L \rightarrow \infty} \frac{\text{Tr}(\mathcal{O} \rho^{GGE})}{\text{Tr}(\rho^{GGE})}. \quad (2)$$

with  $L$  being the system size. The GGE density matrix  $\rho^{GGE}$  is defined as

$$\rho^{GGE} = \lim_{\mathcal{N} \rightarrow \infty} \frac{1}{Z} \exp \left( - \sum_{j=1}^{\mathcal{N}} \lambda_j \mathcal{I}_j \right) \quad (3)$$

where  $Z$  is the normalization,  $\mathcal{I}_j$  the integral of motion,  $\lambda_j$  the associated Lagrange multipliers. These are determined by imposing that  $\langle \Psi_0 | \mathcal{I}_j | \Psi_0 \rangle = \langle \mathcal{O} \rangle_{GGE}$ .

We provide a Monte Carlo method to simulate the GGE statistical ensemble.

Remarkably, this allows to extract the rapidity densities defining the ensemble representative state in the thermodynamic limit.

Finite-size effects are under control and for moderately large chain sizes the Monte Carlo results agree very well with the thermodynamic limit results, especially for small rapidities. Higher rapidities are more sensitive to finite size effects as they reflect short length scales.

The method allows to address local observables for which the expression in terms of the rapidity is known as well.

An alternative numerical method would be Quantum Monte Carlo. However, this would require the implementation of the higher conserved charges, other than the Hamiltonian. As the range of higher charges becomes large this is not easily doable in practice.

We provide the first numerical verification of the validity of the GTBA equations for the Heisenberg spin chain.

Interestingly, finite-size corrections are generically small, i.e., exponentially decaying with the chain size.

Our Monte Carlo approach can be trivially extended to include other conserved charges, both local and non-local, and arbitrary functions of the rapidities. While the former is related to the GGE average of the particle number, the latter measures its fluctuations. Notice that due to the  $SU(2)$  symmetry of the conserved charges one has that  $\langle S_z \rangle = 0$  (panel (g)).

## II. THE MODEL AND THE METHOD

### A. The Heisenberg spin chain

The isotropic spin- $\frac{1}{2}$  Heisenberg ( $XXX$ ) chain is defined by the Hamiltonian

$$\mathcal{H} \equiv J \sum_{i=1}^L \left[ \frac{1}{2} (S_i^+ S_{i+1}^- + S_i^- S_{i+1}^+) + S_i^z S_{i+1}^z \right], \quad (4)$$

where  $S_i^\pm \equiv (\sigma_i^x \pm i\sigma_i^y)/2$  are spin operators acting on the site  $i$  of the chain,  $S_i^z \equiv \sigma_i^z/2$ , and  $\sigma_i^{x,y,z}$  the Pauli matrices. We fix  $J = 1$  and use periodic boundary conditions identifying sites  $L + 1$  and  $1$ . The total magnetization  $S_T^z \equiv \sum_i S_i^z = L/2 - M$ , with  $M$  number of down spins (particles), commute with (4). Thus, the eigenstates of (4) can be labelled by  $M$ .

Each eigenstate of (4) is univocally identified by a set of  $M$  complex parameters (so-called rapidities)  $\{x_\alpha \in \mathbb{C}\}_{\alpha=1}^M$ . In the thermodynamic limit  $L \rightarrow \infty$  the rapidities  $x_\alpha$  form “string” patterns along the imaginary direction in the complex plane (string hypothesis). The rapidities forming a string of length  $1 \leq n \leq M$  (so-called  $n$ -string) are parametrized as  $x_\gamma^{(n,j)} = x_\gamma^{(n)} - i(n-1-2j)$ ,  $j = 0, 1, \dots, n-1$ , where  $x_\gamma^{(n)} \in \mathbb{R}$  is the real part of the string (string center),  $j$  labels the different rapidities in the same  $n$ -string, and  $\gamma$  denotes strings of the same length but with different centers. Although the string hypothesis is not correct for finite chains, deviations typically decay exponentially with system size. The string centers  $x_\gamma^{(n)}$  are solutions of the Bethe-Takahashi equations

$$L\vartheta(x_\gamma^{(n)}/n) = 2\pi I_\gamma^{(n)} + \sum_{(m,\beta) \neq (n,\gamma)} \Theta_{m,n}(x_\gamma^{(n)} - x_\beta^{(m)}). \quad (5)$$

Here  $\vartheta(x) \equiv 2 \arctan(x)$ ,  $\Theta_{m,n}(x)$  is the scattering phase between different rapidities, and  $I_\gamma^{(n)} \in \frac{1}{2}\mathbb{Z}$  are the so-called Bethe-Takahashi quantum numbers. The  $I_{n;\gamma}$  obey the upper bound  $I_{\text{MAX}}(n, L, M) \in \frac{1}{2}\mathbb{Z}$ , with  $I_{\text{MAX}}$  a known function of  $n, M, L$ . Every choice of  $I_{n;\gamma}$  identifies an eigenstate of (4). Notice that each eigenstate contains strings of different lengths. It is useful to define the string content of an eigenstate as  $\mathcal{S} \equiv \{s_1, \dots, s_n\}$ , with  $s_n$  the number of  $n$ -strings.

Besides  $S_T$  and the momentum, the  $XXX$  chain has non-trivial *local* conserved charges  $\mathcal{I}_j$ , with  $[\mathcal{I}_j, \mathcal{I}_k] = 0 \forall j, k$ , and

$\mathcal{I}_2 = \mathcal{H}$ . These can be obtained as

$$\mathcal{I}_{j+1} \equiv \frac{i}{(j-1)!} \frac{d^j}{dx^j} \log(\Lambda(\{x_{n;\gamma}\}, x)) \Big|_{x=i}. \quad (6)$$

Here  $\Lambda$  is the eigenvalue of the quantum transfer matrix  $T(x)$ , with  $x$  the spectral parameter. The dependence of  $\Lambda$  on the rapidities  $x_{n;\gamma}$  is known.

## B. The string hypothesis

An important feature of the Bethe equations (??) is that Here the scattering phases  $\Theta_{m,n}$  are defined as

$$\Theta_{m,n}(x) \equiv \begin{cases} \vartheta\left(\frac{x}{|n-m|}\right) + \sum_{r=1}^{(n+m-|n-m|-1)/2} 2\vartheta\left(\frac{x}{|n-m|+2r}\right) + \vartheta\left(\frac{x}{n+m}\right) & \text{if } n \neq m \\ \sum_{r=1}^{n-1} 2\vartheta\left(\frac{x}{2r}\right) + \vartheta\left(\frac{x}{2n}\right) & \text{if } n = m \end{cases},$$

where  $\vartheta(x) \equiv 2 \arctan(x)$ . Similar to the Bethe quantum numbers the Bethe-Takahashi quantum numbers  $I_\gamma^{(n)}$  identify the solutions of the Bethe-Takahashi equations. It can be shown that  $I_\gamma^{(n)}$  are integers (half integers) if  $L - \alpha_n$  is odd (even). Moreover,  $I_\gamma^{(n)}$  obey the constraint

$$|I_\gamma^{(n)}| \leq \frac{1}{2}(L-1 - \sum_{m=1}^M t_{mn}\alpha_m), \quad (7)$$

where  $t_{mn} \equiv 2\min(m, n) - \delta_{mn}$ . Given a set of solutions of (5), and correspondingly an eigenstate of (4), we define its “string content” as  $\mathcal{S} \equiv \{\mu_1, \mu_2, \dots, \mu_M\}$ , with  $\mu_n$  the number of  $n$ -strings. Clearly, one has that  $\sum_{n=1}^M \mu_n = M$ . The energy and the total momentum associated to a given solution of the Bethe-Takahashi equations are given as

$$E(\{I_\gamma^{(n)}\}) = -L/4 + \sum_{n,\gamma} \frac{2n}{(x_\gamma^{(n)})^2 + n^2} \quad (8)$$

$$K(\{I_\gamma^{(n)}\}) = \sum_{n,\gamma} \frac{2\pi I_\gamma^{(n)}}{L} \quad (9)$$

## C. The thermodynamic limit

In the thermodynamic limit, i.e., for  $L \rightarrow \infty$  the Bethe eigenstates are described by the distributions of the string centers  $\rho_n$ . These are defined such that  $L\rho_n(x)$  gives the number of  $n$ -strings in the interval  $[x, x+dx]$ . More formally  $\rho_n^p(x)$  can be defined as

$$\rho_n^p(x) \equiv \frac{1}{L(x_{\gamma+1}^{(n)} - x_\gamma^{(n)})} \quad (10)$$

The Bethe-Takahashi equations (5) are replaced by the integral equations

$$\rho_n^p(x) + \rho_n^h(x) = a_n(x) + \sum_{m=1}^{\infty} a_{nm}\rho_m^p(x) \quad (11)$$

Moreover, sum over eigenstates can be replaced by

$$\sum_{\text{states}} \rightarrow \int \mathcal{D}\rho^p e^{S_{YY}[\rho^p]}, \quad (12)$$

where  $\rho^p \equiv \{\rho_n^p\}_{n=1}^{\infty}$ , and  $S_{YY}[\rho]$  is the Yang-Yang entropy

$$S_{YY}[\rho] \equiv N \sum_{n=1}^{\infty} \int_{-\infty}^{+\infty} \rho_n(x) \log \left( 1 + \frac{\rho_n^h(x)}{\rho_n(x)} \right) + \rho_n^h \log \left( 1 + \frac{\rho_n(x)}{\rho_n^h(x)} \right). \quad (13)$$

The Yang-Yang entropy counts the number of eigenstates that in the thermodynamic limit lead to the same  $\rho$ .

## D. Hilbert space structure

For the  $XXX$  chain with  $L$  sites the total number of eigenstates with fixed number of particles  $M$  is given as  $C_M^L - C_{M-1}^L$ , with  $C_y^x \equiv x!/(y!(x-y)!)$  the binomial coefficient. Thus, the probability of an eigenstate of the  $XXX$  chain with  $M$  particles is given as  $(C_M^L - C_{M-1}^L)/C_{L/2}^L$ .

For fixed  $M$  the total number of eigenstates  $D(\mathcal{S})$  corresponding to a string content  $\mathcal{S} \equiv \{\mu_1, \mu, \dots, \mu_M\}$  is given as

$$D(\mathcal{S}) = \prod_{i=1}^M C_{\mu_i}^{L - \sum_{j=1}^M t_{ij}\mu_j}. \quad (14)$$

### E. The conserved charges

Notice that  $\Lambda(x)$  depends on the solutions of the Bethe equations (??)  $\{x_\alpha\}$  and it is given as

$$\Lambda \equiv \left(\frac{x+i}{2}\right)^L \prod_{\alpha} \frac{x-x_{\alpha}-2i}{x-x_{\alpha}} + \left(\frac{x-i}{2}\right)^L \prod_{\alpha} \frac{x-x_{\alpha}+2i}{x-x_{\alpha}}. \quad (15)$$

Notice that the second term in (15) does not contribute to the charges (cf. (6)), as long as  $j < L$ . It is trivial to check that  $\mathcal{I}_1 \equiv K$  is the total momentum, while  $\mathcal{I}_2 \equiv \mathcal{H}$ . Moreover, the range of the conserved charge  $\mathcal{I}_j$  is  $j$ , meaning that  $\mathcal{I}_j$  acts non-trivially only on a block of  $j$  adjacent spins.

### III. THE HILBERT SPACE MONTE CARLO SAMPLING

Here we describe the Hilbert space Monte Carlo approach to simulate the GGE ensemble. The method allows to calculate correlation functions and to extract the rapidity densities  $\rho_n(x)$  of the ensemble representative state.

The basic idea of the method is to sample the Hilbert space, i.e., the eigenstates, of the  $XXX$  chain using a Metropolis algorithm.

Here we focus on a chain of length  $L$ . Given an eigenstate of the  $XXX$  chain in the sector with  $M$  particles and identified by a Bethe-Takahashi quantum number configuration  $\mathcal{C}$ , the basic Monte Carlo update generates another configuration  $\mathcal{C}'$ , and consists of five steps:

1. Select a particle number sector  $0 \leq M' \leq L/2$  with probability

$$P_{M'} = \frac{C_{M'}^L - C_{M'-1}^L}{C_{L/2}^L}. \quad (16)$$

2. For fixed  $M'$  select the string content  $\mathcal{S} = \{\mu_1, \mu_2, \dots, \mu_{M'}\}$  with probability

$$P_{\mathcal{S}} = \frac{D(\mathcal{S})}{C_{L/2}^L}, \quad (17)$$

where  $D(\mathcal{S})$  given in (14).

3. Generate randomly a Bethe-Takahashi quantum number configuration  $\mathcal{C}'$  consistent with the string content  $\mathcal{S}$  obtained in step 2.
4. Solve the Bethe-Takahashi equations (5) for using the quantum number configuration  $\mathcal{C}'$  obtained in step 3. Obtain the value of the conserved charges  $\mathcal{I}'_j$  with  $j = 2, m$ .
5. Accept the new configuration with the Metropolis probability  $T(\mathcal{C} \rightarrow \mathcal{C}')$

$$T = \text{Min} \left( 1, \frac{L-2M'+1}{L-2M+1} e^{-\sum_j \lambda_j (\mathcal{I}'_j - \mathcal{I}_j)} \right). \quad (18)$$

Here  $\lambda_j$  are the Lagrange multipliers, and  $\mathcal{I}_j$  is the expectation value of the conserved charge on the eigenstate identified by the Bethe-Takahashi quantum numbers  $\mathcal{C}$ .

In (18) the factor  $L-2M+1$  corresponds to the  $S_z$  degeneracy in the sector with  $M$  particles. Crucially, the Metropolis acceptance probability (18) relies on the fact that the conserved charges  $\mathcal{I}_j$  are  $SU(2)$  scalars.

The procedure outlined above defines a Markov chain. After a thermalization time the configurations generated are distributed according to the  $GGE$  probability (3).

Notice that step 1 and 2 take into account the density of the states of the  $XXX$  chain, while step 5 generates the correct  $GGE$  probability (3).

### IV. EXTRACTING THE RAPIDITY DENSITIES

The rapidity densities  $\rho_n^p$  of the ensemble representative state can be measured in Monte Carlo from the histograms of the roots generated in the Monte Carlo history.

This is illustrated in Fig. 2 for several ensembles. In the Figure panels (a)-(c) plot the first three root densities  $\rho_n(x)$  for  $n = 1, 2, 3$  as a function of  $x$  for the representative state of the infinite-temperature Gibbs ensemble. We restrict ourselves to the interval  $-6 \leq x \leq 6$ . In each panel the different histograms correspond to different chain sizes  $18 \leq L \leq 30$ . The histograms are obtained from  $4 \cdot 10^5$  Monte Carlo steps. The width of the histogram bins is given as  $2/L$ . In all the panels the full lines are the analytic results obtained from the Thermodynamic Bethe Ansatz (TBA) (cf. (21)). Remarkably, although finite-size effects are present, the Monte Carlo data are in good agreement with the TBA results. This is especially true for  $\rho_1$ , whereas finite-size corrections become progressively worse upon considering larger  $n > 1$  (see panels (b)(c)). Clearly, these corrections are vanishing upon increasing the system size (see for instance the arrow in panel (b)). Moreover, the finite size effects become larger upon increasing the rapidity  $x$ . This is expected as larger rapidity correspond to larger quasi-momenta, which are more sensitive to the lattice effects. Finally, finite-size effects increase with  $n$ . This is somewhat expected since larger  $n$  correspond to more extended many-spin bound states.

Finite-temperature ensemble can be also considered. This is illustrated in panel (g) discussing the Gibbs ensemble at inverse temperature  $\beta = 1/2$  and  $\beta = 1$  (different histograms in the panel). We focus on  $\rho_1(x)$ . The data for infinite temperature are reported for comparison. All the histograms are obtained for  $L = 30$ . The full lines are now the analytic results obtained by solving the TBA equations and perfectly agree with the Monte Carlo data. Notice that upon lowering the temperature the height of the zero rapidity peak increases. This reflects that at  $T = 0$  the ground state of the antiferromagnetic Heisenberg chain are packed around  $x = 0$ , and the contribution of nonzero rapidity is vanishing exponentially.

Finally, in panels (d)(f) we present the rapidity densities  $\rho_n(x)$  for the  $GGE$  ensemble. Specifically, here we focus on

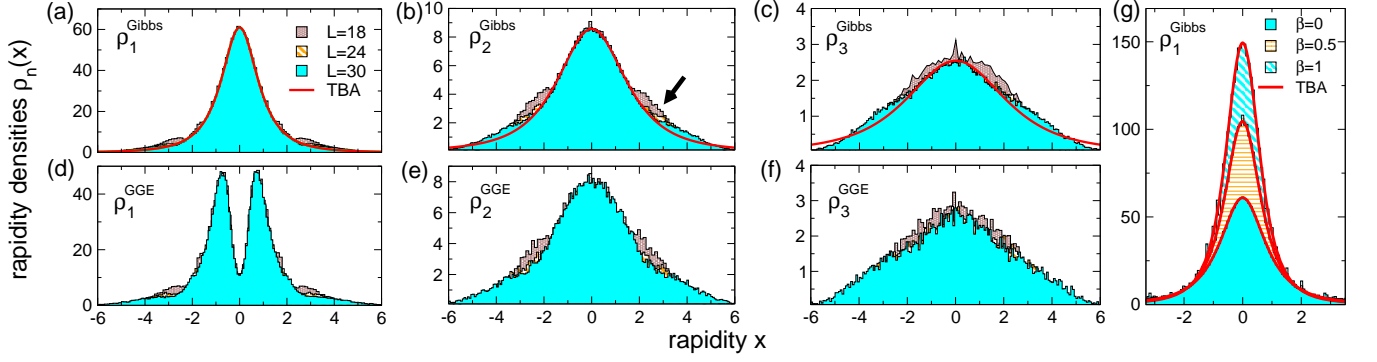


FIG. 1. The rapidity densities  $\rho_n(x)$  (for  $n = 1, 2, 3$ ) for the infinite temperature Gibbs (panels (a)-(c)) and the GGE equilibrium states (panels (d)-(f)): Numerical results for the Heisenberg spin chain obtained using the Hilbert space Monte Carlo sampling. Here the GGE is constructed including only  $\mathcal{I}_2$  and  $\mathcal{I}_4$  with fixed Lagrange multipliers  $\lambda_2 = 0$  and  $\lambda_4 = 1$ . In all the panels the data are the histograms of the  $n$ -strings rapidities sampled in the Monte Carlo. The width of the histogram bins is  $\Delta x = 2/L$ . In each panel different histograms correspond to different chain sizes  $L$ . All the histograms are divided by  $10^3$  for convenience. In (b) the arrow is to highlight the finite-size effects. In panels (a)-(c) the lines are the Thermodynamic Bethe Ansatz (TBA) results. (g) Finite-temperature effects: Monte Carlo data for  $\rho_1^{\text{Gibbs}}$  for different values of the inverse temperature  $\beta$ .

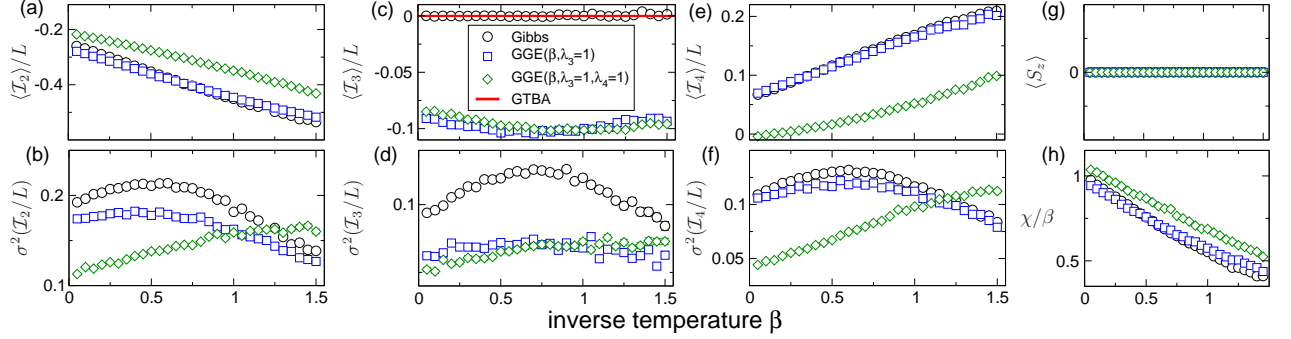


FIG. 2. The Generalized Gibbs Ensemble (GGE) for the Heisenberg spin chain with  $L = 16$  sites: numerical results obtained using the Hilbert space Monte Carlo sampling. Only the first three conserved charges  $\mathcal{I}_n$  ( $n = 1, 2, 3$ ), with associated Lagrange multipliers  $\lambda_n$ , are included in the GGE. Here  $\mathcal{I}_2$  is the Hamiltonian and  $\lambda_2 \equiv \beta$  the inverse temperature. In all the panels different symbols correspond to different values of  $\lambda_3, \lambda_4$ . The circles correspond to the Gibbs ensemble, i.e.,  $\lambda_3 = \lambda_4 = 0$ . (a) The GGE average  $\langle \mathcal{I}_2 \rangle / L$  plotted as a function of  $\beta$ . (b) Variance of the GGE fluctuations  $\sigma^2(\mathcal{I}_2/L) \equiv \langle (\mathcal{I}_2/L)^2 \rangle - \langle \mathcal{I}_2/L \rangle^2$  as a function of  $\beta$ . (c)(d) and (e)(f): Same as in (a)(b) for  $\mathcal{I}_3$  and  $\mathcal{I}_4$ , respectively. In all panels the dash-dotted lines are the analytical results obtained using the Generalized Thermodynamic Bethe Ansatz (GTBA). (g) The GGE expectation value of the total magnetization  $\langle S_z \rangle$ . Notice that  $\langle S_z \rangle = 0$  due to the  $SU(2)$  invariance of the conserved charges. (h)  $\chi/\beta$  plotted versus  $\beta$ , with  $\chi$  being the magnetic susceptibility per site.

the GGE obtained including the two charges  $\mathcal{I}_2$  and  $\mathcal{I}_4$ . We fix the associated Lagrange multipliers to  $\lambda_2 = 0$  and  $\lambda_4 = 1$ . Similar to the Gibbs ensemble, panels (a)(c) suggest that for  $L = 30$  the finite size effects are negligible, at least in the interval  $-2 \leq x \leq 2$ .

## V. LOCAL OBSERVABLES

We now turn to discuss the behavior of some *local* observables. This is discussed in Fig. 2. Here we focus on GGE expectation values of local observables. Here the GGE is constructed including the conserved charges  $\mathcal{I}_j$  with  $j = 2, 3, 4$ . In the Figure we consider several values of the associated Lagrange multipliers, namely  $\lambda_3 = 0$  and  $\lambda_4 = 0$  (Gibbs ensemble, circles in the Figure),  $\lambda_3 = 1$  and  $\lambda_4 = 0$  (squares

in the Figure), and  $\lambda_3 = \lambda_4 = 1$  (rhombi). All the data are obtained using the Monte Carlo sampling approach for a chain with  $L = 16$  sites. The expectation values are the Monte Carlo averages of the local observables over the eigenstates sampled by the Monte Carlo. All the expectation values are plotted versus the inverse temperature  $\lambda = \beta$ . Panel (a) focuses on the GGE expectation value of energy density  $\langle \mathcal{I}_2 \rangle / L$ . The variance of the associated ensemble fluctuations  $\sigma^2(\mathcal{I}_2/L) \equiv \langle (\mathcal{I}_2/L)^2 \rangle - \langle \mathcal{I}_2/L \rangle^2$  are shown in panel (b). The latter are related to the specific heat of the chain.

As expected, different GGEs (different symbols in the Figure) implies different behaviors of  $\langle \mathcal{I}_2 \rangle / L$  as a function of  $\beta$ . The continuous lines are the analytic results obtained by solving the GTBA equations. Remarkably, the latter fully match the Monte Carlo data, i.e., finite-size effects are negligible already at  $L = 16$ .

**How do I reconcile this with Fig. 1 where finite size ef-**

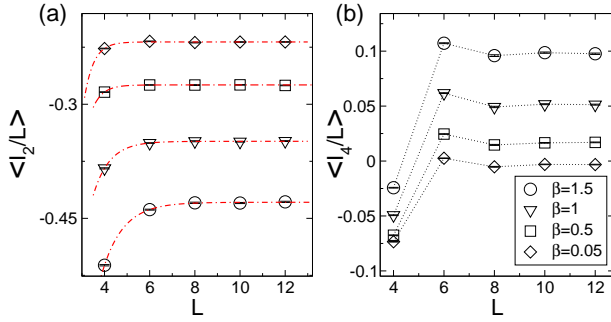


FIG. 3. Finite-size scaling of the GGE averages in the Heisenberg chain: Numerical results obtained from the Hilbert space Monte Carlo sampling. Here the GGE is constructed including  $\mathcal{I}_2, \mathcal{I}_3, \mathcal{I}_4$ , with Lagrange multipliers  $\lambda_2 = \beta, \lambda_3 = \lambda_4 = 1$ . (a)  $\langle \mathcal{I}_2 \rangle / L$  plotted versus the chain size  $L$  for several values of  $\beta$ . The dash-dotted lines are exponential fits. (b) Same as in (a) for  $\mathcal{I}_4$ .

**fects are visible even for  $L = 30$ . Is that true that shorter correlations are less sensitive to higher momenta?**

This remains true if one considers the GGE expectation values of operators with larger supports.

For instance, this is confirmed for the energy current  $\mathcal{I}_3 \equiv \sum_{\alpha\beta\gamma} \epsilon_{\alpha\beta\gamma} \sigma_i^\beta \sigma_{i+1}^\gamma \sigma_{i+2}^\alpha$  in panels (c)(d). Notice that for the Gibbs ensemble one has  $\langle \mathcal{I}_3 \rangle = 0$ , due to the parity invariance. Similar behavior is observed for  $\mathcal{I}_4$  in panels (e)(f).

Finally, in panels (g)(h) we focus on the total magnetization  $\langle S_z \rangle$  and the spin susceptibility  $\chi$ .

The finite-size corrections for local observables are more carefully investigated in Fig. ??, focusing on  $\mathcal{I}_3$  and  $\mathcal{I}_4$  (shown in panel (a) and (b), respectively). The Figure plots

the GGE expectation values  $\langle \mathcal{I}_2 \rangle$  and  $\langle \mathcal{I}_4 \rangle$  versus  $\beta$ . Here we focus on the GGE with Lagrange multipliers  $\lambda_2 = \beta, \lambda_3 = 0$  and  $\lambda_4 = 1$ . Panel (a) demonstrates that finite-size effects decay exponentially with  $L$  for any  $\beta$ . Notice that these corrections are larger at lower temperature, as expected since the correlation length increases with  $\beta$ . Finally, finite-size corrections increase with the range of the operator considered as shown in panel (b), although the behavior remains exponential.

## VI. THE STRING ROOT DENSITIES AT INFINITE TEMPERATURE

For infinite temperature the densities  $\rho_n$  are given as

$$\rho_n(x) = \frac{2}{\pi} \frac{1}{(n^2 + x^2)(x^2 + (2+n)^2)} \quad (19)$$

Notice that

$$\int_{-\infty}^{+\infty} \rho_n(x) dx = \frac{1}{n(n+1)(n+2)} \quad (20)$$

Including the first order correction to the infinite temperature result one obtains

$$\rho_n(x) = \frac{2}{\pi} \frac{1}{(n^2 + x^2)(x^2 + (2+n)^2)} - \frac{8}{\pi} \frac{n(n+2)}{(n^2 + x^2)^2(x^2 + (2+n)^2)^2} J\beta + \mathcal{O}(J^2\beta^2) \quad (21)$$

- <sup>1</sup> M. Rigol, V. Dunjko, V. Yurovsky, and M. Olshanii, Phys. Rev. Lett. **98**, 050405 (2007).
- <sup>2</sup> S. Popescu, A. J. Short, and A. Winter, Nature Physics **2**, 754 (2006).
- <sup>3</sup> M. Rigol, V. Dunjko, and M. Olshanii, Nature **452**, 854 (2008).
- <sup>4</sup> A. Polkovnikov, K. Sengupta, and M. Vengalattore, Rev. Mod. Phys. **83**, 863 (2011).
- <sup>5</sup> J. Eisert, M. Friesdorf, and C. Gogolin, arXiv:1408.5148.
- <sup>6</sup> C. Kollath, A. M. Läuchli, and E. Altman, Phys. Rev. Lett. **98**, 180601 (2007).
- <sup>7</sup> S. R. Manmana, S. Wessel, R. M. Noack, and A. Muramatsu, Phys. Rev. Lett. **98**, 210405 (2007).
- <sup>8</sup> P. Calabrese and J. Cardy, J. Stat. Mech. P06008 (2007).
- <sup>9</sup> M. Cramer, C. M. Dawson, J. Eisert, and T. J. Osborne, Phys. Rev. Lett. **100**, 030602 (2008).
- <sup>10</sup> T. Barthel and U. Schollwöck, Phys. Rev. Lett. **100**, 100601 (2008).
- <sup>11</sup> M. Cramer, A. Flesch, I. P. McCulloch, U. Schollwöck, and J. Eisert, Phys. Rev. Lett. **101**, 063001 (2008).
- <sup>12</sup> M. Kollar and M. Eckstein, Phys. Rev. A **78**, 013626 (2008).
- <sup>13</sup> A. Iucci and M. A. Cazalilla, Phys. Rev. A **80**, 063619 (2009).
- <sup>14</sup> S. Sotiriadis, P. Calabrese, and J. Cardy, EPL **87**, 20002 (2009).
- <sup>15</sup> G. Roux, Phys. Rev. A **79**, 021608 (2009).
- <sup>16</sup> M. Rigol, Phys. Rev. Lett. **103**, 100403 (2009).
- <sup>17</sup> M. Rigol, Phys. Rev. A **80**, 053607 (2009).

- <sup>18</sup> P. Barmettler, M. Punk, V. Gritsev, E. Demler, and E. Altman, Phys. Rev. Lett. **102**, 130603 (2009).
- <sup>19</sup> P. Barmettler, M. Punk, V. Gritsev, E. Demler, and E. Altman, New J. Phys. **12**, 055017 (2010).
- <sup>20</sup> M. Cramer and J. Eisert, New J. Phys. **12**, 055020 (2010).
- <sup>21</sup> A. Flesch, M. Cramer, I. P. McCulloch, U. Schollwöck, and J. Eisert, Phys. Rev. A **78**, 033608 (2008).
- <sup>22</sup> G. Roux, Phys. Rev. A **81**, 053604 (2010).
- <sup>23</sup> D. Fioretto and G. Mussardo, New J. Phys. **12**, 055015 (2010).
- <sup>24</sup> G. Biroli, C. Kollath, and A. M. Läuchli, Phys. Rev. Lett. **105**, 250401 (2010).
- <sup>25</sup> L. F. Santos and M. Rigol, Phys. Rev. E **82**, 031130 (2010).
- <sup>26</sup> M. C. Bañuls, J. I. Cirac, and M. B. Hastings, Phys. Rev. Lett. **106**, 050405 (2011).
- <sup>27</sup> P. Calabrese, F. H. L. Essler, and M. Fagotti, Phys. Rev. Lett. **106**, 227203 (2011).
- <sup>28</sup> C. Gogolin, M. P. Mueller, and J. Eisert, Phys. Rev. Lett. **106**, 040401 (2011).
- <sup>29</sup> M. Rigol and M. Fitzpatrick, Phys. Rev. A **84**, 033640 (2011).
- <sup>30</sup> T. Caneva, E. Canovi, D. Rossini, G. E. Santoro, and A. Silva, J. Stat. Mech. (2011) P07015.
- <sup>31</sup> L. Santos, A. Polkovnikov, and M. Rigol, Phys. Rev. Lett. **107**, 040601 (2011).
- <sup>32</sup> A. C. Cassidy, C. W. Clark, and M. Rigol, Phys. Rev. Lett. **106**, 140405 (2011).



- <sup>33</sup> F. H. L. Essler, S. Evangelisti, and M. Fagotti, Phys. Rev. Lett. **109**, 247206 (2012).
- <sup>34</sup> M. A. Cazalilla, A. Iucci, and M.-C. Chung, Phys. Rev. E **85**, 011133 (2012).
- <sup>35</sup> J. Mossel and J.-S. Caux, New J. Phys. **14** 075006 (2012).
- <sup>36</sup> M. Rigol and M. Srednicki, Phys. Rev. Lett. **108**, 110601 (2012).
- <sup>37</sup> J. Mossel and J.-S. Caux, J. Phys. A: Math. Theor. **45**, 255001 (2012).
- <sup>38</sup> M. Fagotti and F. H. L. Essler, Phys. Rev. B **87**, 245107 (2013).
- <sup>39</sup> M. Fagotti, Phys. Rev. B **87**, 165106 (2013).
- <sup>40</sup> M. Collura, S. Sotiriadis, and P. Calabrese, Phys. Rev. Lett. **110**, 245301 (2013).
- <sup>41</sup> J.-S. Caux and F. H. L. Essler, Phys. Rev. Lett. **110**, 257203 (2013).
- <sup>42</sup> M. Kormos, A. Shashi, Y.-Z. Chou, J.-S. Caux, and A. Imambekov, Phys. Rev. B **88**, 205131 (2013).
- <sup>43</sup> B. Bertini, D. Schuricht, and F. H. L. Essler, arXiv:1405.4813 (2014).
- <sup>44</sup> S. Sotiriadis and P. Calabrese, J. Stat. Mech. (2014) P07024.
- <sup>45</sup> F. H. L. Essler, S. Kehrein, S. R. Manmana, and N. J. Robinson, Phys. Rev. B **89**, 165104 (2014).
- <sup>46</sup> M. Fagotti, M. Collura, F. H. L. Essler, and P. Calabrese, Phys. Rev. B **89**, 125101 (2014).
- <sup>47</sup> M. Fagotti, J. Stat. Mech. (2014) P03016.
- <sup>48</sup> B. Wouters, J. De Nardis, M. Brockmann, D. Fioretto, M. Rigol, and J.-S. Caux, Phys. Rev. Lett. **113**, 117202 (2014).
- <sup>49</sup> B. Pozsgay, M. Mestyán, M. A. Werner, M. Kormos, G. Zarànd, and G. Takács, Phys. Rev. Lett. **113**, 117203 (2014).
- <sup>50</sup> M. Greiner, O. Mandel, T. Hänsch, and I. Bloch, Nature (London) **419**, 51 (2002).
- <sup>51</sup> T. Kinoshita, T. Wenger, and D. S. Weiss, Nature (London) **440**, 900 (2008).
- <sup>52</sup> S. Hofferberth, I. Lesanovsky, B. Fischer, T. Schumm, and J. Schmiedmayer, Nature (London) **449**, 324 (2007).
- <sup>53</sup> I. Bloch, J. Dalibard, and W. Zwerger, Rev. Mod. Phys. **80**, 885 (2008).
- <sup>54</sup> S. Trotzky, Y.-A. Chen, A. Flesch, I. P. McCulloch, U. Schollwöck, J. Eisert, and I. Bloch, Nature Phys. **8**, 325 (2012).
- <sup>55</sup> M. Gring, M. Kuhnert, T. Langen, T. Kitagawa, B. Rauer, M. Schreitl, I. Mazets, D. A. Smith, E. Demler, and J. Schmiedmayer, Science **337**, 6100 (2012).
- <sup>56</sup> M. Cheneau, P. Barmettler, D. Poletti, M. Endres, P. Schaua, T. Fukuhara, C. Gross, I. Bloch, C. Kollath, and S. Kuhr, Nature (London) **481**, 484 (2012).
- <sup>57</sup> U. Schneider, L. Hackeruller, J. P. Ronzheimer, S. Will, S. Braun, T. Best, I. Bloch, E. Demler, S. Mandt, D. Rasch, and A. Rosch, Nature Phys. **8**, 213 (2012).
- <sup>58</sup> M. Kuhnert, R. Geiger, T. Langen, M. Gring, B. Rauer, T. Kitagawa, E. Demler, D. Adu Smith, and J. Schmiedmayer, Phys. Rev. Lett. **110**, 090405 (2013).
- <sup>59</sup> T. Langen, R. Geiger, M. Kuhnert, B. Rauer, and J. Schmiedmayer, Nature Phys. **9**, 640 (2013).
- <sup>60</sup> F. Meinert, M. J. Mark, E. Kirilov, K. Lauber, P. Weinmann, A. J. Daley, and H.-C. Nagerl, Phys. Rev. Lett. **111**, 053003 (2013).
- <sup>61</sup> T. Fukuhara, A. Kantian, M. Endres, M. Cheneau, P. Schaua, S. Hild, C. Gross, U. Schollwöck, T. Giamarchi, I. Bloch, and S. Kuhr, Nature Phys. **9**, 235 (2013).
- <sup>62</sup> J. P. Ronzheimer, M. Schreiber, S. Braun, S. S. Hodgman, S. Langer, I. P. McCulloch, F. Heidrich-Meisner, I. Bloch, and U. Schneider, Phys. Rev. Lett. **110**, 205301 (2013).
- <sup>63</sup> S. Braun, M. Friesdorf, S. Hodgman, M. Schreiber, J. Ronzheimer, A. Riera, M. del Rey, I. Bloch, J. Eisert, and U. Schneider, arXiv:1403.7199.
- <sup>64</sup> J. M. Deutsch, Phys. Rev. A **43**, 2046 (1991).
- <sup>65</sup> M. Srednicki, Phys. Rev. E **50**, 888 (1994).
- <sup>66</sup> M. Srednicki, J. Phys. A **29**, L75 (1996).
- <sup>67</sup> M. Srednicki, J. Phys. A **32**, 1163 (1999).
- <sup>68</sup> S. Goldstein, J. L. Lebowitz, R. Tumulka, and N. Zanghi, Phys. Rev. Lett. **96**, 050403 (2006).
- <sup>69</sup> S. Goldstein, J. L. Lebowitz, C. Mastrodonato, R. Tumulka, and N. Zanghi, Proc. R. Soc. A **466**, 3203 (2010).
- <sup>70</sup> S. Goldstein, J. L. Lebowitz, R. Tumulka, and N. Zanghi, Eur. Phys. J. H **35**, 173 (2010).
- <sup>71</sup> T. N. Ikeda, Y. Watanabe, and M. Ueda, Phys. Rev. E **84**, 021130 (2011).
- <sup>72</sup> T. N. Ikeda, Y. Watanabe, and M. Ueda, Phys. Rev. E **87**, 012125 (2013).
- <sup>73</sup> R. Steinigeweg, J. Herbrych, and P. Prelovšek, Phys. Rev. E **87**, 012118 (2013).
- <sup>74</sup> W. Beugeling, R. Moessner, and M. Haque, Phys. Rev. E **89**, 042112 (2014).
- <sup>75</sup> R. Steinigeweg, A. Khodja, H. Niemeyer, C. Gogolin, and J. Gemmer, Phys. Rev. Lett. **112**, 130403 (2014).
- <sup>76</sup> S. Sorg, L. Vidmar, L. Pollet, and F. Heidrich-Meisner, arXiv:1405.5404v2.
- <sup>77</sup> W. Beugeling, R. Moessner, and M. Haque, arXiv:1407.2043.
- <sup>78</sup> V. Khemani, A. Chandran, H. Kim, and S. L. Sondhi, arXiv:1406.4863.
- <sup>79</sup> H. Kim, T. N. Ikeda, and D. Huse, arXiv:1408.0535.
- <sup>80</sup> L. Bonnes, F. H. L. Essler, and A. M. Läuchli, arXiv:1404.4062 (2014).
- <sup>81</sup> J.-S. Caux and J. Mossel, J. Stat. Mech. (2011) P02023.
- <sup>82</sup> V. Alba, M. Fagotti, and P. Calabrese, J. Stat. Mech. (2009) P10020.
- <sup>83</sup> N. Kitanine, J. M. Maillet, and V. Terras, Nucl. Phys. B **554**, 647 (1999).
- <sup>84</sup> N. Kitanine, J. M. Maillet, and V. Terras, Nucl. Phys. B **567**, 554 (2000).
- <sup>85</sup> L. Amico, R. Fazio, A. Osterloh, and V. Vedral, Rev. Mod. Phys. **80**, 517 (2008).
- <sup>86</sup> M. Takahashi, *Thermodynamics of one-dimensional solvable models*, Cambridge University Press 1999.
- <sup>87</sup> C. N. Yang and C. P. Yang, J. Math. Phys. **10**, 1115 (1969).
- <sup>88</sup> M. Takahashi, Prog. Theor. Phys. **46**, 401 (1971).
- <sup>89</sup> M. P. Grabowski and P. Mathieu, Ann. Phys. N.Y. **243**, 299 (1995).
- <sup>90</sup> J. Eisert, M. Cramer, and M. B. Plenio, Rev. Mod. Phys. **82**, 277 (2009).
- <sup>91</sup> P. Calabrese, J. Cardy, and B. Doyon Eds., Special issue: Entanglement entropy in extended systems, J. Phys. A **42**, 50 (2009).
- <sup>92</sup> P. Calabrese and J. Cardy, J. Phys. A **42** 504005 (2009).
- <sup>93</sup> V. E. Korepin, N. M. Bogoliubov, and A. G. Izergin, *Quantum Inverse Scattering Methods and Correlation Functions*, Cambridge University Press 1997.
- <sup>94</sup> X. Zotos and P. Prelovšek, Phys. Rev. B **53**, 983 (1996).
- <sup>95</sup> H. Castella and X. Zotos, Phys. Rev. B **54**, 4375 (1996).
- <sup>96</sup> X. Zotos, F. Naef, and P. Prelovšek, Phys. Rev. B **55**, 11029 (1997).
- <sup>97</sup> F. C. Alcaraz, M. I. Berganza, and G. Sierra, Phys. Rev. Lett. **106**, 201601 (2011).
- <sup>98</sup> I. Pizorn, arXiv:1202.3336.
- <sup>99</sup> M. I. Berganza, F. C. Alcaraz, and G. Sierra, J. Stat. Mech. (2012) P01016.
- <sup>100</sup> G. Wong, I. Klich, L. A. P. Zayas, and D. Vaman, JHEP **12** (2013) 020.
- <sup>101</sup> M. Storms, and R. R. P. Singh, Phys. Rev. E **89**, 012125 (2014).

- <sup>102</sup> R. Berkovits, Phys. Rev. B **87**, 075141 (2013).
- <sup>103</sup> F. H. L. Essler, A. M. Läuchli, and P. Calabrese, Phys. Rev. Lett. **110**, 115701 (2013).
- <sup>104</sup> M. Nozaki, T. Numasawa, T. Takayanagi, Phys. Rev. Lett. **112**, 111602 (2014).
- <sup>105</sup> G. Ramirez, J. Rodriguez-Laguna, and G. Sierra, arXiv:1402.5015.
- <sup>106</sup> F. Ares, J. G. Esteve, F. Falceto, and E. Sánchez-Burillo, arXiv:1401.5922.
- <sup>107</sup> Y. Huang, and J. Moore, arXiv:1405.1817.
- <sup>108</sup> T. Pálmai, arXiv:1406.3182.
- <sup>109</sup> J. Mölter, T. Barthel, U. Schollwöck, and V. Alba, arXiv:1407.0066.
- <sup>110</sup> H.-H. Lai and K. Yang, arXiv:1409.1224
- <sup>111</sup> J. Sato, B. Aufgebauer, H. Boos, F. Göhmann, A. Klümper, M. Takahashi, and C. Trippe, Phys. Rev. Lett. **106**, 257201 (2011).
- <sup>112</sup> M. Fagotti and P. Calabrese, Phys. Rev. A **78**, 010306 (2008).
- <sup>113</sup> V. Gurarie, J. Stat. Mech. (2014) P02014.
- <sup>114</sup> M. Collura, M. Kormos, and P. Calabrese, J. Stat. Mech. (2014) P01009.
- <sup>115</sup> M. Kormos, L. Bucciantini, and P. Calabrese, EPL **107**, 40002 (2014).
- <sup>116</sup> J.-S. Caux and J.-M. Maillet, Phys. Rev. Lett. **95**, 077201 (2005).
- <sup>117</sup> J.-S. Caux, R. Hagemans and J.-M. Maillet, J. Stat. Mech. P09003 (2005).
- <sup>118</sup> J.-S. Caux, J. Math. Phys. **50**, 095214 (2009).