## PGF5107 - Quantum Field Theory I

## Homework 4

Due 29/04/2016

## 1. Correlation Functions in Perturbation Theory:

Defining the field in the Heisenberg picture as

$$\tilde{\phi}(t, \vec{x}) = e^{iHt}\phi(0, \vec{x})e^{-iHt}$$

and the field in the interaction picture as

$$\phi(t, \vec{x}) = e^{iH_0t}\phi(0, \vec{x})e^{-iH_0t}$$

- (a) Show that  $\phi(t, \vec{x})$  is a free field. For instance, if this is a real scalar field, show that  $\phi(t, \vec{x})$  satisfies the homogenous Klein-Gordon equation.
- (b) Show that

$$\tilde{\phi}(x) = \Omega(t) \, \phi(x) \, \Omega^{\dagger}(t)$$

where we defined

$$\Omega(t) = e^{iHt} e^{-iH_0t}$$

- (c) Show that for any time order
  - $U(t_3, t_1) = U(t_3, t_2)U(t_2, t_1)$ ,
  - $U^{\dagger}(t_1, t_2) = U(t_2, t_1)$ .

with

$$U(t_1, t_2) = \Omega^{\dagger}(t_1) \Omega(t_2)$$

(d) Using the properties of the U operators and the transformation of the  $\tilde{\phi}$ 's, show that

$$\tilde{\phi}(x_1)\dots\tilde{\phi}(x_n) = U^{\dagger}(\infty,0)U(\infty,t_1)\phi(x_1)U(t_1,t_2)\phi(x_2)\dots U(t_{n-1},t_n)\phi(x_n)U(t_n,-\infty)U(-\infty,0)$$

(e) Replacing  $H_0$  by  $H_0(1-i\epsilon)$ , with  $\epsilon > 0$ , real and infinitesimal, show that

$$U(-\infty,0)|\tilde{0}\rangle = \langle 0|\tilde{0}\rangle|0\rangle$$
,

and

$$\langle \tilde{0}|U^{\dagger}(\infty,0) = \langle 0|\langle \tilde{0}|0\rangle$$
,

where  $|\tilde{0}\rangle$  and  $|0\rangle$  are the perturbed and unperturbed vacuum respectively, and in the Heisenberg picture  $e^{-iHt}|\tilde{0}\rangle = |\tilde{0}\rangle$ . To prove this you want to use the fact that we can expand the perturbed vacuum in terms of eigenstates of  $H_0$ ,  $|n\rangle$ , such that  $|\tilde{0}\rangle = \sum_n |n\rangle\langle n|\tilde{0}\rangle$ . As long as  $\epsilon > 0$  then the ground state will dominate.

(f) Show that

$$\langle \tilde{0} | T \tilde{\phi}(x_1) \dots \tilde{\phi}(x_n) | \tilde{0} \rangle = \langle 0 | U(\infty, t_1) \phi(x_1) U(t_1, t_2) \phi(x_2) \dots U(t_{n-1}, t_n) \phi(x_n) U(t_n, -\infty) | 0 \rangle \times |\langle \tilde{0} | 0 \rangle|^2$$

(g) Finally,

i. Show

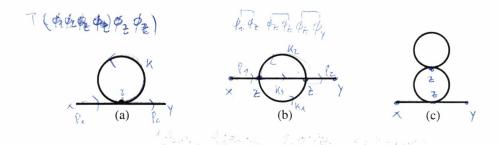
$$\langle \tilde{0} | T \tilde{\phi}(x_1) \dots \tilde{\phi}(x_n) | \tilde{0} \rangle = \langle 0 | T \phi(x_1) \dots \phi(x_n) e^{-i \int_{-\infty}^{+\infty} dt V(t)} | 0 \rangle \times |\langle \tilde{0} | 0 \rangle|^2$$
where you use that  $U(\infty, -\infty) = e^{-i \int_{-\infty}^{+\infty} dt V(t)}$ ,

ii. Also show that

$$|\langle \tilde{0}|0\rangle|^2 = \frac{1}{\langle 0|Te^{-i\int_{-\infty}^{+\infty} dtV(t)}|0\rangle}$$

In this way the final answer is

$$\langle \tilde{0}|T\tilde{\phi}(x_1)\dots\tilde{\phi}(x_n)|\tilde{0}\rangle = \frac{\langle 0|T\phi(x_1)\dots\phi(x_n)e^{-i\int_{-\infty}^{+\infty}dtV(t)}|0\rangle}{\langle 0|Te^{-i\int_{-\infty}^{+\infty}dtV(t)}|0\rangle}$$



## 2. Wick's Theorem and Feynman Diagrams:

Consider the 3 Feynman diagrams in the figure, for a theory with

$$\mathcal{H}_I = \frac{\lambda}{4!} \, \phi^4 \ .$$

- (a) Using Wick's Theorem compute the symmetry factors for each, i.e. compute all the possible contractions, etc.
- (b) Using the Feynman rules in momentum space, compute the contributions of Figure (a) to the two-point correlation function. Consider the ultra-violet (UV) limit  $E \gg m$ : how does the answer behave with E? Is it finite for  $E \to \infty$ ?
- (c) Do the same for Figure (b).

(b) 
$$\frac{(-i\lambda)^2}{6} \int \frac{d^4k_2}{(2\pi)^4} \int \frac{d^4k_3}{(2\pi)^4} \frac{1}{k_2^2 m^2 + i\epsilon} \frac{1}{k_3^2 - m^2 + i\epsilon} \frac{1}{(\kappa_2 + k_3)^2 + m^2 + i\epsilon}$$

$$\int d^{D}k \frac{1}{(h^{2}+2p-k-m^{2}+i\epsilon)^{n}} = \frac{i(-1)^{n} \pi^{D}z}{T(n)(m^{2}+p^{2})^{n}-b_{2}} T(n-D_{2})$$

$$+(a) = \frac{2}{2} (-i \pi^2 m^2) \tau (-1)$$

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11 Correlation Luctions in Perturbation Theory:

9.75

(a) Let us apply K-G operator to  $\phi(u,\vec{x}) = e^{iHot} \phi(o,\vec{x}) e^{-iHot}$  $(\partial_t z - \vec{\nabla}^z + w^z) \phi(t,\vec{x}) = -H_0^z e^{iHot} \phi(o,\vec{x}) + H_0 e^{iHot} \phi(o,\vec{x}) H_0 e^{-iHot}$   $+ H_0 e^{iHot} \phi(o,\vec{x}) H_0 e^{-iHot} - e^{iHot} \phi(o,\vec{x}) H_0^z e^{-iHot}$ 

+ e iHot { - \$\vec{7}^2 + m^2} \phi(0,\vec{x}) e - iHot

Now notice that [Ho, etiHot] =0, and that

From it  $\frac{d}{dt} \phi(x) = [\phi(x), H_0]$ , we have that  $[\phi(x), H_0] = 0$ .

And obviously for t=0, is satisfied  $(-\vec{\nabla}^2 + m^2) \phi(0,\vec{x}) = 0$ .

Thus

 $(\partial^2 + m^2) \phi(t, \vec{x}) = 0$ , and  $\phi(t, \vec{x})$  is a free field.

(b)  $\tilde{\phi}(x) = \mathcal{L}(t) \phi(x) \mathcal{L}^{\dagger}(t)$ 

= eint e-inot p(x) { eint e-intot} +

=  $e^{iHt}e^{-iHot}\phi(x)e^{iHot}e^{-iHt}$ , As  $\phi(x)=e^{iHot}e^{-iHot}$ 

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= e int p(o,x) e - int

(c) As 
$$V(ts, tr) = \Omega^+(ts)\Omega(tr)$$
 and  $V(tz, ta) = \Omega^+(tz)\Omega(ta)$   
then
$$V(ts, tr)V(tz, ta) = \Omega^+(ts)\Omega(ta)\Omega^+(tz)\Omega(ta)$$

$$= \Omega^+(ts)\Omega(ta)$$

$$= U(ts, ta).$$

Ou the other hand we have;

(d) 
$$\widetilde{\phi}(x_1) - \widetilde{\phi}(x_n) = \mathcal{N}(t_n) \varphi(x_n) \mathcal{N}^{\dagger}(t_n) \mathcal{N}(t_n) \varphi(x_n) \mathcal{N}^{\dagger}(t_n)$$

Notice: 
$$U^{\dagger}(\omega,0) U(\omega,t_1) = U(0,\infty) U(\omega,t_1)$$
  

$$= U(0,t_1)$$

$$= \mathcal{I}^{\dagger}(0) \mathcal{I}(t_1)$$

$$= \mathcal{I}(t_1).$$

Besides 
$$V(t_n, -\infty) V(-\infty, 0) = V(t_n, 0)$$
  
=  $\mathcal{N}^+(t_n) \mathcal{N}(0)$   
=  $\mathcal{N}^+(t_n)$ .

Using this, we can write:

 $\widetilde{\phi}(x_1) - \widetilde{\phi}(x_1) = U^{\dagger}(\omega, 0) U(\omega, t_1) \varphi(x_1) U(t_1, t_2) \varphi(x_2) - \cdots U(t_{m-1}, t_m) \varphi(x_m) U(t_m, \omega) U(-\omega, 0)$ 

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$$(e)$$
  $((-\infty,0)(\tilde{0}) = \mathcal{N}^+(-\infty)\mathcal{N}(0)(\tilde{0})$ 

As Holh) = Enly), and taking Ho -> Ho (1-ie), with E>O;

$$U(-\infty,0)|\tilde{O}\rangle = \lim_{\gamma \to -\infty} e^{i\tilde{E}_0\gamma} e^{\xi\tilde{E}_1\gamma} |0\rangle \langle 0|\tilde{O}\rangle + \sum_{n \neq 0} e^{i\tilde{E}_n\gamma} e^{\xi\tilde{E}_n\varepsilon\gamma} |n\rangle \langle n|\tilde{O}\rangle.$$

We can drop the terms with  $N \neq 0$  because they die faster than the ground state. Also notice that we can take  $E_0=0$ , so that  $e^{E_0 \Upsilon(i+\epsilon)} \rightarrow 0$ .

Thus

Proceeding in a similar way, we get

Where we have used the same argument as before.

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(f) using the result that we found in (d), we can write:
\langle \tilde{o} \mid T \tilde{\varphi}(x_1) - \tilde{\varphi}(x_n) \mid \tilde{o} \rangle = \langle \tilde{o} \mid U^{\dagger}(\omega_{i} \circ) \cup (\omega_{i} \circ) \varphi(x_i) \cup (t_n t_n) \varphi(x_n) - \tilde{\varphi}(x_n) \rangle
                                                        U (tn, -00) U (-00,0) ( 0)
   but \langle \tilde{\sigma} | U^{\dagger}(\omega_{1}0) = \langle 01 \langle \tilde{\sigma}_{1}0 \rangle, \langle U(-\omega_{1}0) | \tilde{\sigma} \rangle = \langle 01 \tilde{\sigma} \rangle | 0 \rangle,
                                                                                                    = <610)* (0)
thus
2\tilde{o}[T\tilde{\phi}(x_n)...\tilde{\phi}(x_n)|\tilde{o}\rangle = 20|U(\omega_0)\phi(x_n)...\phi(x_n)U(t_{n_1}-\omega)|0\rangle|(\tilde{o}|0\rangle)^2
(g) (i) Notice that U(\infty,0) \phi(x_1) - U(t_{n-1},t_n) \phi(x_n) U(t_{n-1}-\infty)
  is already temporary ordered (recall (d)).
  So we can write:
 U(\omega,0) \not \varphi(x_n) U(t_n,t_2) \varphi(x_2) - \cdots \varphi(x_n) U(t_n,-\infty) = T \varphi(x_n) \varphi(x_2) - \cdots \varphi(x_n) U(\omega,t_1)
                                                                                     U(trite) --- U(tri-00)
  how we can use that
  U(t_a,t_b) = U(t_a,t_p)U(t_p,t_b) (2) rsb)
  repeatedly, so we found:
 U(\infty,0) \phi(x_n) - \cdots \phi(x_n) U(t_n,-\infty) = T \phi(x_n) - \cdots \phi(x_n) U(t_\infty,-\infty)
  As V(\omega_1 - \omega) = T e^{-i \int_{-\omega}^{\infty} dt V(t)}
  Then
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 $\angle \tilde{o} | T \tilde{\varphi}(x_n) - \tilde{\varphi}(x_n) | \tilde{o} \rangle = \angle o | T \varphi(x_n) - \tilde{\varphi}(x_n) e^{-i \int dt \, V(t)} | o \rangle | \langle \tilde{o} | o \rangle |^2$ 

by the left by  $U(0,-\infty)$ :

 $U(0,-\infty)$   $U(-\infty,0)$   $|\tilde{0}\rangle = \langle 0|\tilde{0}\rangle$   $U(0,-\infty)$   $|0\rangle$ 

how Consider  $Z\tilde{o}IU^{\dagger}(\omega_{1}o)=ZoIZ\tilde{o}Io)$  and unitiply this by the right by  $U(+\omega_{1}o)$ , to get

 $\langle \tilde{O} | = \langle \tilde{O} | 0 \rangle \langle 0 | U(\omega_{10}), \quad \text{(we used that } U^{\dagger}(\omega_{10}) = U(0,\infty) \text{)}$ 

Thus

 $\langle \tilde{o} | \tilde{o} \rangle = |\langle \tilde{o} | o \rangle|^2 \langle o | U(+\infty, o) U(0, -\infty) | o \rangle$ 

 $= |26/05|^2 20|0(+\infty,-\infty)|0$ 

= 12510512 201 Te-is de V(t) (0)

Using that 10% is normalized, we get:

(0) T e-ij dt V(t) (0)

21 Wick's theorem and Feynman Diagrams: (4.75)

(a) From Wick's theorem and the definition of Normal order, we know that the only terms that will contribute in the VA CUUM expectation value of time-ordered product of Fields, will be those that are totally contracted.

Thus we know that

is related with the contraction:

Ø(x) Ø(z) Ø(z) Ø(z) Ø(x)

but there is no reason to belive that Ø(Z) are distinguishable so it's also true that the contraction:

Ø(x) Ø(z) Ø(z) Ø(z) Ø(z) Ø(y) will work

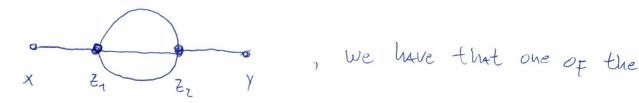
So Ø(x) may be contracted with any of the Four Ø(z), and Q(x) with any of the other three Q(Z) remaining. We will call P the wunber of possible contractions, so that the symmetry factor (S.F.) will be calculated as:

(S.F.) = P (4!)"

where n is the number of vertex OF the diagram.

So for we have 
$$IP = 4.3$$

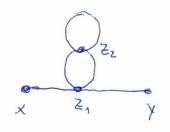
thus 
$$(5.F.)^{-1} = \frac{4.3}{4!} = \frac{1}{2}$$



possible contractions is:  $\phi_{x}$   $\phi_{z_1}$   $\phi_{z_1}$   $\phi_{z_1}$   $\phi_{z_2}$   $\phi_{z_2}$   $\phi_{z_2}$   $\phi_{z_2}$   $\phi_{z_3}$ 

it follows that IP = 4.4.3.2,

thus 
$$(5.F.)^{-1} = \frac{4.4.3.2}{(4!)^2} = \frac{1}{6}$$



We have that one of the possible contractions is:

Øx Øz, Øz, Øz, Øz, Øz, Øz, Øz, Øy

then P = 4.3.12

$$(5.E)^{-1} = \frac{4 \cdot 3 \cdot 12}{(4!)^2} = \frac{1}{4}$$

$$(6) \longrightarrow \frac{1}{2} = -\frac{1}{2} \int \frac{d^4K}{(2\pi)^4} \cdot \frac{1}{\kappa^2 - m^2 + i\epsilon} = I$$

Making a continuation to evolidery momentum:

$$T = -\frac{i\lambda}{2} \int \frac{d^4k_E}{(2\pi)^4} \cdot \frac{1}{K_E^2 + m^2}$$
, Using four dimensional spherical coordinates;

$$= -\frac{i\lambda}{2} \int dx_{4} \int \frac{dk_{E}}{(2\pi)^{4}} \frac{k_{E}^{3}}{k_{E}^{3} + M^{2}}$$

, impose momentum space cutoff 1:

$$= -\frac{c2}{2} \cdot \frac{2\pi^{2}}{(2\pi)^{2}} \cdot \frac{1}{2} \left\{ K_{\epsilon}^{2} = m^{2} \ln (m^{2} + K_{\epsilon}^{2}) \right\}^{2}$$

$$= \frac{-i\lambda}{32\pi^2} \left\{ \Lambda^2 - M^2 \ln \left( \frac{\Lambda^2 + M^2}{m^2} \right) \right\}$$

for A large compared to un:

$$T = -\frac{c\lambda}{32\pi^2} \left\{ \Lambda^2 - m^2 \ln \left( \frac{\Lambda^2}{m^2} \right) \right\}$$

We see that this expression diverges for 1 -> 00

(3)

 $=\frac{(-i\lambda)^{2}}{6}\int \frac{d^{4}K_{2}}{(2\pi)^{4}} \frac{d^{4}K_{3}}{(2\pi)^{4}} \frac{i}{K_{2}^{2}-m^{2}+i\epsilon} \frac{i}{K_{3}^{2}-m^{2}+i\epsilon} \frac{i}{(\kappa_{2}+\rho-\kappa_{3})^{2}+m^{2}+i\epsilon}$ 

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