

Quantum Field Theory I - Prof. Gustavo Burdman

Homework 1

Due 20/03/20

1. Relativity Gymnastics

- (a) Show that the invariance of the interval between two events $(t^2 - |\vec{x}|^2)$ under Lorentz transformations is satisfied if we define the position four-vector $x^\mu \equiv (t; \vec{x})$ such that

$$x'^\mu = \Lambda^\mu_\nu x^\nu$$

where we defined the Lorentz transformations as Λ^μ_ν . The inverse Lorentz transformation is defined by $x^\nu = \Lambda^\nu_\mu x'^\mu$, and it satisfies

$$\Lambda^\mu_\nu \Lambda^\nu_\rho = \delta^\mu_\rho$$

- (b) The Minkowski space metric is given by

$$g^{\mu\nu} = g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix},$$

- i. Show that the interval is given by the position four-vector squared defined by $x.x \equiv x^\mu g_{\mu\nu} x^\nu$.
 - ii. A contra-variant four-vector is denoted as $v^\mu = (t; \vec{v})$. Show that if we define a covariant four-vector as $v_\mu \equiv (t; -\vec{v})$, we can write $v.v = v^\mu v_\mu$.
 - iii. Verify that $v_\mu = g_{\mu\nu} v^\nu$, and $v^\mu = g^{\mu\nu} v_\nu$.
- (c) Show that the relativistic dispersion relation $E^2 - |\vec{p}|^2 = m^2$, where m is the particle's mass, is consistent with defining the momentum four-vector as $P^\mu = (E; \vec{p})$.

(d) Define the differential operators

$$\frac{\partial}{\partial x^\mu} \equiv \partial_\mu \equiv \left(\frac{\partial}{\partial t}; \vec{\nabla} \right) \quad \frac{\partial}{\partial x_\mu} \equiv \partial^\mu \equiv \left(\frac{\partial}{\partial t}; -\vec{\nabla} \right)$$

i. Construct the Klein-Gordon operator using the identifications

$$E \rightarrow i \frac{\partial}{\partial t} \quad \vec{p} \rightarrow -i \vec{\nabla}$$

ii. Show that if we define the four-vector current $J^\mu \equiv (\rho; \vec{j})$, where ρ and \vec{j} the charge density and current respectively, the continuity equation in electrodynamics can be written as the conservation of the four-current as

$$\partial_\mu J^\mu = 0$$

2. About those Delta Functions

Using the defining property of the Delta function

$$\int_{-\infty}^{+\infty} f(x) \delta(x) dx = f(0)$$

for some function $f(x)$, and that $\delta(-x) = \delta(x)$

(a) Show that

$$\int_{-\infty}^{+\infty} f(x) \delta(ax) dx = \frac{1}{|a|} f(0)$$

for an arbitrary constant a .

(b) Using all these show that for a function $g(x)$

$$\int_{-\infty}^{+\infty} f(x) \delta(g(x)) dx = \sum_a \frac{1}{|g'(a)|} \int_{-\infty}^{+\infty} f(x) \delta(x-a) dx ,$$

where a are the zeroes of $g(x)$ (i.e. $g(a) = 0, \forall a$) while $g'(a) \neq 0$.

(c) Now verify that for $\omega_p = \sqrt{(\vec{p})^2 + m^2} > 0$ we have

$$\int d^4p \delta(P^2 - m^2) f(P) = \int d^3p \int dp_0 \delta(p_0^2 - (\vec{p})^2 - m^2) f(P) = \frac{1}{2\omega_p} \int d^3p f(\omega_p, \vec{p}) ,$$

where you need to use the fact that since the four-momentum is always time-like, the sign of p_0 is Lorentz invariant. Thus we need only integrate p_0 from 0 to ∞ , which in turn means we only need the positive root.

3. Electromagnetism:

The covariant form of the Lagrangian density for the electromagnetic field in the absence of sources is

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} ,$$

where $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ is the field strength tensor defined in terms of the potential 4-vector A_μ .

- (a) Derive the equations of motion for $A_\mu(x)$ assuming it is the dynamical field, using Euler-Lagrange. Show that these are equivalent to Maxwell's equations in the vacuum. (It is useful to remember that $E^i = -F^{0i}$ and $\epsilon^{ijk} B^k = -F^{ij}$.)
- (b) If we now consider sources in the form of a current j_μ , what is the form of the interaction term between A_μ and j_μ we must add to \mathcal{L} in order to obtain Maxwell's equations in the presence of the source $j_\mu = (\rho, \vec{j})$?
- (c) The Lagrangian in the absence of sources is invariant under the gauge transformation

$$A_\mu(x) \rightarrow A_\mu(x) + \partial_\mu \alpha(x) ,$$

with $\alpha(x)$ an arbitrary function. What is the condition that j_μ must fulfill for the Lagrangian to remain gauge invariant in the presence of sources ?

QFT I - Homework 1

Víctor Muñoz

19

P11 (a) the invariance of the interval is proved by showing that $S'^2 = X'^2 = X^2$.

$$\begin{aligned} X'^2 &= X'^{\mu} X'_{\mu} = \Lambda^{\mu}_{\nu} X^{\nu} g_{\mu\alpha} \Lambda^{\alpha}_{\beta} X^{\beta} \\ &= \Lambda^{\mu}_{\nu} g_{\mu\alpha} \Lambda^{\alpha}_{\beta} X^{\nu} X^{\beta} \end{aligned}$$

3

$$\text{but } \Lambda^{\mu}_{\nu} g_{\mu\alpha} \Lambda^{\alpha}_{\beta} = g_{\nu\beta} \quad (*)$$

$$\Rightarrow X'^2 = g_{\nu\beta} X^{\nu} X^{\beta} = X^2 \quad \text{so that } S'^2 = S^2.$$

let us verify $(*)$ in the simplest case: $\nu = \beta = 0$, then it should be true that

$$\Lambda^{\mu}_0 g_{\mu\alpha} \Lambda^{\alpha}_0 = 1$$

As the only non null component of $g_{\mu\nu}$ are $g_{\lambda\lambda}$ with $\lambda = 0, 1, 2, 3$ we have:

$$\Lambda^0_0 g_{00} \Lambda^0_0 + \dots + \Lambda^3_0 g_{33} \Lambda^3_0 = (\Lambda^0_0)^2 - (\Lambda^1_0)^2 - (\Lambda^2_0)^2 - (\Lambda^3_0)^2$$

$$\text{in the case of a boost in the } X \text{ direction: } \Lambda = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\text{with } \gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \quad \text{and } \beta = \frac{v}{c}$$

So we have: $(\Lambda^0_0)^2 - (\Lambda^1_0)^2 = \gamma^2 - \beta^2 \gamma^2$

$$= \gamma^2 (1 - \beta^2)$$

$$= \frac{(1 - \beta^2)}{(1 - \beta^2)}$$

$$= 1. \quad \text{As expected.}$$

The other terms are calculated in similar way.

(b) i. $x \cdot x = x^\mu g_{\mu\nu} x^\nu$

$$= x^0 g_{00} x^0 + x^1 g_{10} x^0 + x^2 g_{20} x^0 + x^3 g_{30} x^0$$

$$= x^0 g_{00} x^0 + x^1 g_{11} x^1 + x^2 g_{22} x^2 + x^3 g_{33} x^3$$

$$= x^0^2 - x^1^2 - x^2^2 - x^3^2$$

The only non null components of $g_{\mu\nu}$ are $g_{\lambda\lambda}$, $\lambda=0,1,2,3$

which is precisely s^2 ✓

ii. $v^\mu v_\mu = v^0 v_0 + v^1 v_1 + v^2 v_2 + v^3 v_3$

$$= t^2 + (v_x)(-v_x) + (v_y)(-v_y) + (v_z)(-v_z)$$

$$= t^2 - \vec{v}^2$$

$$= v \cdot v$$

iii. $v_\mu = g_{\mu\nu} v^\nu$, so

$$v_0 = g_{0\nu} v^\nu = t$$

$$v_1 = g_{1\nu} v^\nu = -v_x$$

$$v_2 = g_{2\nu} v^\nu = -v_y$$

$$v_3 = g_{3\nu} v^\nu = -v_z$$

similarly $v^\mu = g^{\mu\nu} v_\nu$, so

$$v^0 = g^{0\nu} v_\nu = t$$

$$v^1 = g^{1\nu} v_\nu = v_x$$

$$v^2 = g^{2\nu} v_\nu = -v_y$$

$$v^3 = g^{3\nu} v_\nu = -v_z$$

(c) $P^\mu = (E; \vec{p})$, so $P^\mu P_\mu = P \cdot P = E^2 - \vec{p}^2$

on the other hand the 4-momentum is defined by $P^\mu = m U^\mu$

where U^μ is the 4-velocity such that $U^2 = 1$.

Then $P^2 = m^2$ and matching these two results we have :

$E^2 - \vec{p}^2 = m^2$ which gives the correct dispersion relation.

(d) from the relativistic dispersion relation : $E^2 - \vec{p}^2 - m^2 = 0$

And using the identifications : $E \rightarrow i \partial_t$ and $\vec{p} \rightarrow -i \vec{\nabla}$

we have: $(i \partial_t)(i \partial_t) - (-i \vec{\nabla})(-i \vec{\nabla}) - m^2 = -\partial_t^2 + \vec{\nabla}^2 - m^2$

which is the Klein - Gordon operator.

P2 (a) consider $\int_{-\infty}^{+\infty} f(x) \delta(ax) dx$, making the change of variable:

$x = ay$, with a some constant .

2

$dx = a dy$

we have: $\int_{-\infty}^{+\infty} f(x) \delta(ax) dx = \int_{-\infty}^{+\infty} f(y/a) \delta(y) dy \cdot \frac{1}{|a|}$

Note: the absolute value over "a" stands because if "a" is minor than zero, then the limits of integration are reversed providing a global minus sign in the equation.

$= \frac{1}{|a|} f(0)$

$= \frac{1}{|a|} \int_{-\infty}^{+\infty} f(y) \delta(y) dy$

(b) From the previous answer we have:

$\delta(ax) = \frac{1}{|a|} \delta(x)$. We will conclude that this is just a special case from $\delta(g(x)) = \frac{\delta(x-x_0)}{|g'(x_0)|}$, with x_0 satisfying $g(x_0)=0$.
For example? $g(x) \approx g(a) + (x-a)g'(a) + \dots$

For $\int_{-\infty}^{+\infty} f(x) \delta(g(x)) dx$, let $y = g(x)$, $dy = g'(x) dx$
 $x = g^{-1}(y)$, $a = f(x \rightarrow -\infty)$, $b = f(x \rightarrow +\infty)$

$$\Rightarrow \int_{-\infty}^{+\infty} f(x) \delta(g(x)) dx = \int_a^b f(g^{-1}(y)) \delta(y) \frac{dy}{|g'(g^{-1}(y))|}$$

$$= \frac{f(g^{-1}(0))}{|g'(g^{-1}(0))|} = \frac{f(x_0)}{|g'(x_0)|}$$

We are assuming that the condition $y=0$ is satisfied with $x=x_0$ for which $g(x_0)=0$.

The last result can be obtained from:

$$\int_{-\infty}^{+\infty} f(x) \frac{\delta(x-x_0)}{|g'(x)|} dx = \frac{f(x_0)}{|g'(x_0)|}$$

So that by comparison we have: $\delta(g(x)) = \frac{1}{|g'(x_0)|} \delta(x-x_0)$, with $g(x_0)=0$

And the obvious generalization when $g(x)$ possess multiple (simple) zeros is :

$$\delta(g(x)) = \sum_{x_i} \frac{1}{|g'(x_i)|} \delta(x-x_i) \quad , \quad g(x_i) = 0$$

so that

$$\int_{-\infty}^{+\infty} f(x) \delta(g(x)) dx = \sum_{x_i} \frac{1}{|g'(x_i)|} \int_{-\infty}^{+\infty} f(x) \delta(x-x_i) dx$$

$$(c) \quad \int d^4P \delta(P^2 - m^2) F(P) = \int d^3P \int dP_0 \delta(P_0^2 - \vec{P}^2 - m^2) F(P)$$

here $g(P_0) = P_0^2 - \vec{P}^2 - m^2 \approx P_0^2 - \omega_P^2$ here $g(P_0)$ has two zeros

At $P_0 = \pm \omega_P$, And $g'(P_0) = 2P_0$. We are interested only in the case where $\omega_P > 0$, so we have:

$$\begin{aligned} \int d^4P \delta(P^2 - m^2) F(P) &= \int d^3P \int dP_0 \frac{\delta(P_0 - \omega_P)}{2\omega_P} F(P_0, \vec{P}) \\ &= \int d^3P \frac{F(\omega_P, \vec{P})}{2\omega_P} \end{aligned}$$

P3] (a) First notice that :

(4)

$$\begin{aligned} F^{\mu\nu} F_{\mu\nu} &= (\partial_\mu A_\nu - \partial_\nu A_\mu) (\partial^\mu A^\nu - \partial^\nu A^\mu) \\ &= \partial_\mu A_\nu (\partial^\mu A^\nu - \partial^\nu A^\mu) - \partial_\nu A_\mu (\partial^\mu A^\nu - \partial^\nu A^\mu) \\ &= \partial_\mu A_\nu (\partial^\mu A^\nu - \partial^\nu A^\mu) - \partial_\mu A_\nu (\partial^\nu A^\mu - \partial^\mu A^\nu) \\ &= 2 \partial_\mu A_\nu (\partial^\mu A^\nu - \partial^\nu A^\mu) \end{aligned}$$

Therefore $\mathcal{L} = -\frac{1}{2} \partial_\mu A_\nu (\partial^\mu A^\nu - \partial^\nu A^\mu)$

The E-L equations are $\frac{\partial \mathcal{L}}{\partial A_\nu} = \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} \right)$

with $\frac{\partial \mathcal{L}}{\partial A_\nu} = 0$ and $\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} = -\frac{1}{2} (\partial^\mu A^\nu - \partial^\nu A^\mu) - \frac{1}{2} \partial_\mu A_\nu \times \frac{\partial}{\partial (\partial_\mu A_\nu)} (\partial^\mu A^\nu - \partial^\nu A^\mu)$

using that $\partial^\mu A^\nu = \partial_\mu A_\nu \eta^{\mu\alpha} \eta^{\nu\beta}$ and $\partial^\nu A^\mu = \partial_\mu A_\nu \eta^{\mu\alpha} \eta^{\nu\beta}$

we get

$$\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} = F^{\mu\nu}, \text{ so that}$$

$$\partial_\mu F^{\mu\nu} = 0$$

hence

$$\partial_0 F^{0\nu} + \partial_1 F^{1\nu} + \partial_2 F^{2\nu} + \partial_3 F^{3\nu} = 0$$

As $F^{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & B_z & -B_y \\ E_y & -B_z & 0 & B_x \\ E_z & B_y & -B_x & 0 \end{pmatrix}$ (we are using $c=1$)

We have for $\nu=0$: $\boxed{\vec{\nabla} \cdot \vec{E} = 0}$ ✓

For $\nu=1$: $\partial_t E_x = \partial_z B_y - \partial_y B_z$

For $\nu=2$: $\partial_t E_y = \partial_x B_z - \partial_z B_x$

For $\nu=3$: $\partial_t E_z = \partial_x B_y - \partial_y B_x$

Adding the last three equations, we get : $\boxed{\vec{\nabla} \times \vec{B} = \partial_t \vec{E}}$ ✓

* Note: The other Maxwell's equations are obtained from the dual tensor: $\tilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta}$ whose divergence is will: $\partial_\mu \tilde{F}^{\mu\nu} = 0 \Rightarrow \vec{\nabla} \cdot \vec{B} = 0$ and $\vec{\nabla} \times \vec{E} = -\partial_t \vec{B}$.

(b) It's easy to see that if we add the term

$A_\nu J^\nu$ to the previous Lagrangian, then we will have

$\frac{\partial \mathcal{L}}{\partial A_\nu} = J^\nu$ and then we will have $\partial_\mu F^{\mu\nu} = J^\nu$ ✓

which are the Maxwell's equations in the presence of a source.

(c) The action: $S = \int d^4x \left(-\frac{1}{4} F^{\mu\nu} F_{\mu\nu} - A_\nu J^\nu \right)$ is gauge

invariant if and only if J^ν is a conserved current: $\partial_\nu J^\nu = 0$.

Because for $A_\nu \rightarrow A_\nu(x) + \partial_\nu \alpha(x)$ we have:

$$S' = S - \int d^4x \partial_\nu \alpha(x) J^\nu(x) \quad \text{if } \alpha(x) \rightarrow 0 \text{ when } x \rightarrow \pm\infty$$

then by the Stokes' theorem we have : $\int d^4x \partial_\mu (\alpha J^\mu) = 0$

$$\Rightarrow \int d^4x \partial_\nu \alpha(x) J^\nu = - \int d^4x \partial_\nu J^\nu \alpha(x) \quad \text{So that}$$

$$S' = S + \int d^4x \alpha(x) \partial_\nu J^\nu \quad \text{and then} \quad S' = S \Leftrightarrow \partial_\nu J^\nu = 0$$