Physics 251-8th Class - Thursday Feb. 1, 2024

A. More on Approximation - Approximate SHM and Molecular Radiation

1. Change of Power Series Expansion "reference point" [K-text,]

Consider, e.g., the function $f(x) = x^2$. As stated, it is a one-term "power series" around x = 0. But, suppose that we want to expand that function around, say, x = 3. Then, the power series expansion is

(1) $\chi^2 = 9 + 6(\chi - 3) + \frac{2}{2!}(\chi - 3)^2$

For that expansion, multiplying out the right-hand side, we arrive at

 $\chi^2 = 9 + 6\chi - 18 + \chi^2 - 6\chi + 9 = \chi^2$

Thus, eqn. (1) provides a three term power series around the point x=3 that is actually equal to x^2 !

On, consider the function $f(x) = \frac{a}{\chi 9} - \frac{b}{\chi}$ where a and b are both positive constants and where we only consider value of $\chi \ge 0$. As written, $f(\chi) = \frac{a}{\chi 9} - \frac{b}{\chi}$ is a two-term

Alries in inverse powers of x expanded around x=0. As you can show by setting $\frac{df}{dx}$ equal to zero, the function $\frac{a}{x^9} - \frac{b}{x}$ has a minimum at x = X where $X^8 = 9\frac{a}{b}$ $\Rightarrow \chi_{min}$ is at $\sqrt[8]{9\frac{a}{b}}$. As we will soon see, in applications to molecular vibrations of diatomic molecules, expanding $f(x) = \frac{a}{x^9} - \frac{b}{x}$ around the point χ_{min} (i.e., "ne-referencing the series to $\chi = \chi$) can be very helpful.

2. Molecular Vibrational Oscillation (K-text, sect. 3-7)

Context: Molecules were discovered in interstellar Space. How?

Each molecular species has its own emitted radiation spectrum.

The details depend on quantum mechanics, but we can use classical mechanics as a rough guide for what frequency range the radio telescope should look in.

We'll restrict today's discussion to ionically bonded molecules and a case in which one atom is much heavier than the other (HCI). The classical picture is then that the heavy ion can be considered essentially stationary and the lighter



atom oscillates against it. As the lighter ion is charged

and accelerating, it radiates EM waves at the oscillation frequency. We wish to estimate the expected frequency. To do this, it is certainly of interest to know if the motion is (at least approximately) simple harmonic. If it is, then you know how to pick off the frequency from the equation of motion.

In order to determine things like this, we need to know something about the force between ions in a molecule. The details of this would come from advanced quantum mechanics, but we can content ourselves here with just asking about some general features.

Actually, it is more convenient for us to discuss the

potential energy U(r) [r is the distance between the ion centers].

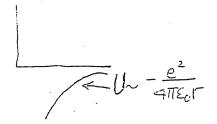
Recall that force and potential

energy are related by
$$F(r) = -\frac{dU(r)}{dr}$$

What are some general features of U(r)?

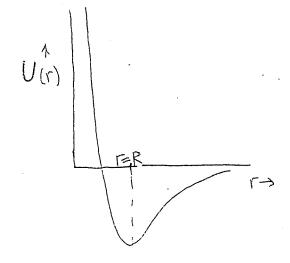
1. When the ions are far apart, they attract according to Coulombis law.

So part of U is
$$U_{far}(r) = \frac{-e^2}{4\pi \epsilon_0 r}$$



2. When close there is a strong repulsive force of quantum nechanical origin. (Otherwise, the molecule would collapse in rechanical origin. (response to the coulomb attraction of the oppositely changed ions).

So U(r) must look qualitatively like



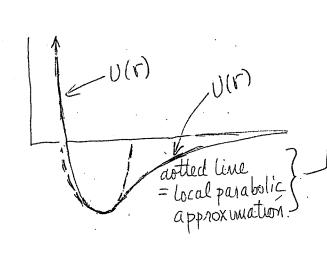
To match the two parts of U(r), there must be a minimum. Since the force is zero at the minimum (why?), that value of ionic separation (call it V=R) is equilibrium.

Then, expanding U(R) around the minimum at F=R, we obtain

$$U(r) = U(R) + \frac{1}{2} \left(\frac{d^2 U(r)}{dr^2} \right)_{r=R} + \frac{1}{6} \left(\frac{d^3 U(r)}{dr^3} \right)_{r=R} + \frac{1}{4} + \frac{1}{6} \left(\frac{d^3 U(r)}{dr^3} \right)_{r=R} + \frac{1}{4} + \dots$$
where $\psi = r - R$. [Since $\frac{d}{dr} |_{r=R} = 0$].

Several Lessons from this:

1. If U(r) is a continuous function with a local minimum at r=R, if $\frac{d^2U(r)}{dr^2}\Big|_{r=R} > 0$ then U(r) is arbitrarily well represented by a parabolic function [VIZ. $U(r)=U(R)+\frac{1}{2}\left(\frac{d^2U(r)}{dr^2}\right)_{r=R}(r-R)^2$] as $r\to R$ (as long as $\frac{d^3U(r)}{dr^3}\Big|_{r=R}$ is finite);



Since the curve of U(r)

Is much shallower to the

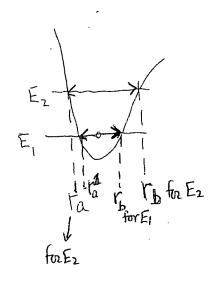
right of r=R than to the left

(Why?) the averaging parabolic approx. must be steeper in

slope than U(r) to the right

of r=R and shallower to the left.]

As you will recall from the Phys. 151 topic "Reading a Potential Evergy Plot", if, in the situation under discussion here, if the evergy of the oscillation of the lighter ion is somewhat greater than the value of U(r) at the minimum (r=R), then the lighter ion can only oscillate back and for the between the two points ("ra" and "rs" in figure below) for which $E = U(r) \implies K.E. = 0$ at r = ra, r = rb.



(next page >

3. If E is only slightly greater than Umin, this oscillatory motion is simple harmonic.

This is because simple harmonic motion is motion in a parabolic potential

- recall for, say, a mass on a spring

$$4 \text{ m}$$
 $U(x) = \frac{1}{2} k x^2 = \text{parabolic}$

[note
$$U(x) = \frac{1}{2}k\chi^2 \Rightarrow F(x) = \frac{-dV(x)}{dx} = -k\chi$$
]

4. Comparing this to our series expansion (top of last page), we see that for "small oscillations" (E only slightly greater than Vmin, so to not much different than Ta), the "chemical bond" can be conceptually replaced by a simple spring (!) with "effective spring constant"

$$k = \left(\frac{d^2 V c}{dr^2}\right)$$
eff $\left(\frac{d^2 V c}{dr^2}\right)$

And the expected angular hequency of radiation will be

$$W = \sqrt{\frac{k_{eff}}{k}}$$
 denom= μ (reduced mass of the pair)

Now, if we knew U as a function of r, then we could simple calculate $\frac{d^2U(r)}{dr^2}\Big|_{r=R}$. But, the

detailed functional form of U(r) can only come from quantum mechanics.

While we aren't in a position to derive U(r) from quantum mechanics in this course, we can at least get some practice with the spirit of how to fit simple approximate forms for U(r) to extract info:

Example: A relatively simple functional form that is sometimes used in this context as a rough approximation is

$$U(r) = \frac{-e^2}{4\pi\epsilon_0 r} + \frac{B}{r^9}$$
Coulombic repulsive and attraction very strong at small r .

where B is a constant. This form agrees with our expectations, since it has both a Coulombic attraction (between the ions) piece and also a repulsive piece (+ sign in potential energy of interaction indicates regulsion) that is weak at large r but very strong at very small r, since it is $\propto \frac{1}{19}$. Again, this form is only a simple approximation to the "real U(r)"; as well, since we wish only an (approximate) frequency prediction, were really only interested in fitting a parabola for small 1r-R1.

Let's take this form for U(r) and run with it.

A plot of it (for a particular value of B) is shown below:

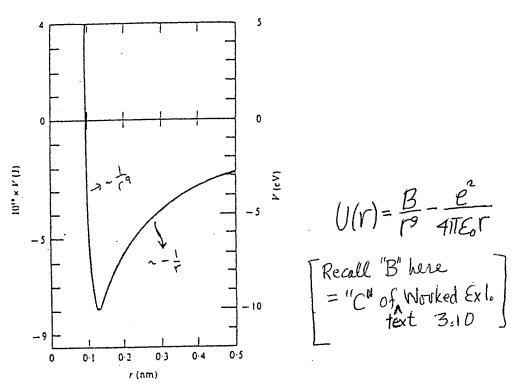


Fig. 2.6 The potential energy curve for a pair of ions: formula (2.15), with B given by (2.16) and R chosen as 0.13 nm (the measured value for HCI). The zero of potential energy is the value when the ions are infinitely far apart. (It is customary to shift the curve upwards by an amount which allows for the energy required to form the ions from neutral atoms, but this makes no difference to the vibrational behaviour discussed in this book.)
Figure from Vibrations and Waves in Physics, 2nd Ed. by Jain Main (Cambridge).

1. Let's see if we can figure out what B should be. Letting r = R be the equilibrium separation of the ions, we know that 11113

hat
$$\frac{dV_{i0}}{dr}\Big|_{r=R} = 0$$
, which is $\frac{P^2}{4\pi\epsilon_0 R^2} - \frac{9B}{R^{10}} = 0$

$$\Rightarrow$$
 (2) $B = \frac{e^2R^8}{36\pi\epsilon_0}$ [Same as lext eqn. (3.27) on page 3-53 in Worked Exl. 3.10]

2. Now, it would appear that all we've done is trade off one unknown ("B") for another ("R", the equilibrium separation of the ions in the molecule). Let's continue anyway: To obtain keff for this molecule, we use

(3)
$$R_{eff} = \frac{d^2 V(r)}{dr^2} \Big|_{r=R} = -\frac{2e^2}{4 \pi \epsilon_0 R^3} - \frac{9(10) B}{R^{11}} = \frac{("B" \text{kere})}{= "C" \text{ in text ext.}}$$

(same as Worked Ext 3.10, eqn. (3.28), page 3-53). Using eqn. (2) above for B in eqn. (3), we find

(4)
$$R_{eff} = \frac{8e^2}{4\pi\epsilon_0 R^3} \approx 1.8 \times 10^{-27} N_0 M^2$$

3. Now we need to input [into (4)] a value for R. Rather than input "the accepted value for R" [as in Worked exl. 3.10], here we use "the method of guessing":

Radius of an atom or ion $\sim 0.5 \, \text{Å} = 0.5 \times 10^{-10}$ guess center-to center separation $\sim 3 \times \text{nadius of ion}$ \Rightarrow guess (nough) $R \sim 1.5 \, \text{Å}$ Potting that cuto (4) above $\Rightarrow \text{keff} \approx \frac{1.8 \times 10^{-27}}{3.4 \times 10^{-30}} \, \frac{N}{M}$,

 $R_{eff} \sim 530 \frac{N}{m}$

Pause to think:

 $k_{eff} \sim 530 \frac{N}{m}$: How much is that ? Is that "alot"?

To get an idea, consider a spring with spring constant 530 m on which we have an "everyday" amound of mass

- Say, 1 kg. The stretch caused by this is given from

$$\Rightarrow \Delta \chi = \frac{mg}{R} \approx \frac{1(10)}{510}$$
 meters

The "effective spring constant" that models
the quantum mechanical chemical bond in HCI
in the small-oscillations regime is comparable
to that of a spring used in everyday application!

Anyway - back to our estimate of expected radiation frequency for HC1:

$$W = \sqrt{\frac{k_{eff}}{\mu}}$$

in HCI, CI um is about 35 times as heavy as H ton

$$\Rightarrow$$
 CT is essentially stationary, $M \approx M_{H^+} = mass of proten (Mp)$.

 $M_p \approx 1.67 \times 10^{-27} \text{kg}$

$$\Rightarrow W^{\text{pred.}} \sqrt{\frac{530 \text{ N/m}}{1.67 \times 10^{-27} \text{ Kg}}} \sim 5.6 \times 10^{14} \text{ nad/sec}$$
(for our "rough gress")

$$\Rightarrow f''^{\text{predicted}"} = \frac{1}{2\pi} W^{\text{pred}} \sim \frac{9.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times 10^{13} \text{ Hz}} \approx \frac{3.0 \times 10^{13} \text{ Hz}}{900 \times$$

whereas f experimental $\sim 8.9 \times 10^{13} HZ$

The closeness of our guess to the actual experimental value is mostly fortuitous - we are lucky to have even gotten a result within a factor of 5 of the expt. value, but the method (an order of magnitude estimate method) has pedagogical value.

B. Superposition Theorem; General Solutions

Linear Differential Equations - Homogeneous Case - Superposition Clearly, differential equations have been unportant un our work. Now we need more: A differential equation is "linear" if it is linear in the dependent Variable. We consider linear d.e.'s involving functions of (t). For simplicity we consider second order equations, but our result applies to equations of any order. [Why are we so interested in second order equations?]. The most general znd order linear differential equation is.

(alt) if (t) + b(t) if (t) + c(t) if (t) = d(t)

Here alt, blt), c(t), d(t) are any functions of t (e.g., 6t2).

Thus, by "linear" is weant linear in the dependent variable (4).

If the right side is zero (i.e, dlt) = 0), the equation is said to be "homogeneous." Our first theorem here is:

a. Superposition Principle: Consider the general second order liveur, howagen eous d.e.

(1) $a(t)\dot{\psi}(t) + b(t)\dot{\psi}(t) + c(t)\dot{\psi}(t) = 0$

If 4,(t) and $4_2(t)$ are any two solutions of (1), then, for any constants C_1 and C_2 ,

(2) $\psi(t) = c_1 \psi_1(t) + c_2 \psi_2(t)$ is also a solution of (1).

Proof: Let
$$\psi(t) = c_1 \psi_1(t) + c_2 \psi_2(t)$$
. Then

$$a(t) \psi(t) = a(t) \left[c_1 \psi_1(t) + c_2 \psi_2 \right]' = c_1 a(t) \psi_1' + c_2 a(t) \psi_2'$$

$$b(t) \psi(t) = b(t) \left[c_1 \psi_1 + c_2 \psi_2 \right]' = c_1 b(t) \psi_1' + c_2 b(t) \psi_2'$$

$$Also, c(t) \psi = c(t) \left[c_1 \psi_1 + c_2 \psi_2 \right] = c_1 c(t) \psi_1 + c_2 c(t) \psi_2'$$

$$alt) \psi' + b(t) \psi(t) + c(t) \psi$$

$$= c_1 alt) \psi'_1 + c_1 b(t) \psi'_1 + c_1 \psi_1$$

$$+ c_2 a(t) \psi'_2 + c_2 b(t) \psi'_2 + c_2 \psi_2$$

$$= c_1 \left[a(t) \psi'_1 + b(t) \psi'_1 + c(t) \psi_1 \right]$$

$$+ c_2 \left[a(t) \psi'_2 + b(t) \psi'_2 + c(t) \psi_2 \right]$$

$$= 0$$

q.e.d.

Comment. Note that the surple harmonic oscillator equation is of this form.

Question: What are a(t), b(t), c(t) for the "paradigm" simple harmonic oscillator?

Here's another very important theorem concerning I.h.d.e.'s:

b. Theorem: Suppose that 4(t) and 4(t) are two livearly independent solutions of (1) on an interval [a,b]. Then, on [a,b], the general solution to the d.e. (1) can be expressed in the form (2) 4(t) = 64(t) + 624(t)

for arbitrary constants C_1 and C_2 -1.e, on [a,b] every soln. of eqn. (1) can be expressed in the form (2) for some choice of (C_1,C_2) .

We leave the proof of this theorem to your d.e. course; here we only note that it is plausible (as we already discussed earlier)

- 1. Second order eqn ⇒ loosely must integrate given

 if (t) twice to obtain 4(t) ⇒ two constants of integration

 set by initial conditions; (2) has two different constants

 multiplying two linearly independent functions.
- 2. Exl: s.h.o., 4(t) = sin wt, 42(t) = coswt,
 general, 4(t) = Csin wt + Dcoswt

2. Now consider a nontwear homogeneous différential equi

 $\dot{\psi}(t) = -\frac{9}{4} \sin \psi$

$$\varphi(t) = -c \psi + \alpha \psi^2 + b \psi^3$$

For any of these (udeed, for any nonlinear differential equation), the superposition theorem is not obeyed - the sum of two solutions is not a solution.

We will have you show that for homework, however, here we offer a plausibility argument:

Linear, homogeneous eqn is expressible as \$[4]=0

where I is a linear operator [e.g., L= 3t \frac{d}{dt^2} + 4t \frac{d}{dt}]

As a summary of our proof of the Superposition theorem:
Since
$$f$$
 is linear, $f(4, +4_2) = f(4,) + f(4_2) = 0 + 0 = 0$

Now consider nonlinear homogeneous eqn.

$$\mathcal{N}(4) = 0$$
 where \mathcal{N} is a nonlinear operator [e.g., $\mathcal{N}(4) = e^{4}$]

 $N(4+42) \neq N(4) + N(42) = 0$

$$N(4) = \dot{\psi}^2$$
; $(4_1 + 4_2)^2 = \dot{4}_1^2 + \dot{4}_2^2 + 2\dot{4}_1\dot{4}_2 \neq \dot{4}_1^2 + \dot{4}_2^2$ $(N(4) = \dot{4}^2)$

Superposition of Solutions for Inhomogeous I.d.e.s?

Consider the second order linear inhomogeneous equ

(4)
$$a(t) \dot{\psi}(t) + b(t) \dot{\psi}(t) + c(t) \psi(t) = a(t)$$

Then, if $4_1(t)$ and $4_2(t)$ are linearly independent solus, then $4(t) = 4_1(t) + 4_2(t)$ is not a solu of (4).

We leave the proof of that to you.

D. Inhomogeneous Linear Differential Equations, "Superposition of Driving Forces"
Consider the pair of unhomogeneous linear d.e.'s

(5)
$$a(t) \dot{\psi}(t) + b(t) \dot{\psi}(t) + c(t) \psi(t) = r(t)$$

(6) alt
$$\mathring{\psi}(t) + b(t) \mathring{\psi}(t) + c(t) \psi(t) = s(t)$$
,

For the same functions alt), b(t), c(t), but different r(t), s(t),

Then, if 4(t) and 4(t) are two indep. solus of (5) and (6), respectively, then 4(t) = 4(t) + 4(t) is a solu. of

E. Introduction to Normal Modes

We come now to something very generally important in physics - the concept of a "normal mode". To begin looking at this, we ask you to consider the following system:

("general" configuration).

We assume the two masses are equal - Ma=Mb=M.

Exercize: Show that the differential equations for this system are

$$m\frac{d^2t_{att}}{dt^2} = -kt_{a}(t) + k[t_{g}(t) - t_{a}(t)]$$
 governing d.e.'s.

$$m \frac{d^2 + (t)}{dt^2} = -k [+_b(t) - +_a(t)] - k +_b(t)$$

These equations are somewhat complicated because the "fa equation" Contains 45 and the "45 equation" contains 4a - we say

that the equations are coupled. This makes sense because the motion of the affects that of b, and vice versa, due to the coupling spring.

This is actually not the first time we've seen coupling between masses - in the "diatomic molecule" example, it is certainly true that the motion of one of the masses affects (via the coupling spring) the motion of the other!

also, the general motion of the system (the two masses) is nother complicated - and, in fact, many different notions are possible. Each "notion of the system" in characterized by a particular pair of functions {4a(t), 4b(t)}. As you might imagine, the particular functions 4a(t) and 4b(t) are determined by the initial conditions - 4a(0), 4b(0), 4a(0), 4b(0) - 100, 4b(0) the two initial positions and the two initial velocities.

^{*} At least from our present naive point of view!

Now, we could try to attack the coupled equations by a purely mathematical method - a method that might proceed by Trying to find appropriate linear combinations of 4a and 4b to use as new coordinates obeying uncoupled equations ("smart coordinates")

We will try that wethod later. First, however, we will try thinking physically about the system.

We begin this physical approach by asking-"of the many possible bairs of functions discribing motions of the system - are there any that are "simple" (*)?" We ask, in particular, about simple harmonic notion - is there a possible motion in which both masses undergo simple harmonic nuction simultaneously? For such a motion of the masses we might have, say,

 $\psi_{a}(t) = A \cos \omega t$ $\psi_{a}(t) = B \cos \omega t$

A motion like this, in which both masses undergo smultaneous simple harmonic motion at the same heguency and with fixed relative phase, is called a <u>Normal Mode</u>.

^{*} The word "sumple" is a technical term in physica- it means "of that which we already kmow"

Let us see if a Normal Mode is possible for this system. Suppose we begin by displacing both masses an equal amount to the right from their equilibrium positions.

So $\psi_a(t=0) = \psi_b(t=0)$.

Now let us ask about the forces on the masses the instant we let go. We have already learned

that only forces in excess of those used to maintain equilibrium contribute to motion - the forces on each mass used to maintain equilibrium cancel out (that's why there's no motion in equilibrium).

Notice that the extension of the central spring has remained constant at the equilibrium value a. Therefore, at this instant, it contributes no excess face on either mass. Thus, at this instant the only excess force on the left hand mass is that due to the left most spring -

excess force on Ma at t=0 = - fe Ha(t=0). Likewise, excess force on MB at t=0 = - & \$\psi_b(t=0)\$ from right most spring. Since Ya(t=0) = 46(t=0), these are equal. Thus the masses will start to move with identical accelerations to the left. After a short time At, both masses more an equal distance and the center spring remains at extension a. Therefore the masses still have identical forces acting on them, and will move further identically. By continuing this kind of reasoning it becomes clear that the motions of the two masses are identical for all t>0 and that (ignoring friction) the motion is stable.

^{*} As previously, these equations are true regardless of the signs of ta and to-take a moment and convince yourself of that.

Suce the middle spring always remains at constant length, it is easy to see that at any t>0, for motion in this mode

$$m \frac{d^2 + dt}{dt^2} = -k + a(t)$$

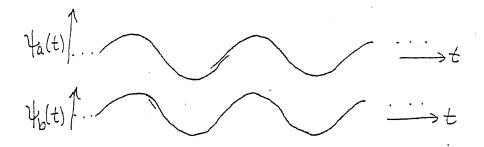
$$m \frac{d^2 \psi_b(t)}{dt^2} = -k \psi_b(t)$$

(not the governing differential equs. since only valid for present set of physical initial conditions

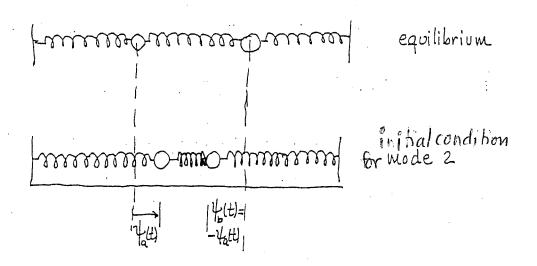
Thus in this mode, the differential equations have become much simpler - they are uncoupled. (This makes sense, Since the coupling device, the middle spring, is not contributing to the motion). The solution to the pair of equations that obeys the prescribed initial conditions is

which verifies that this motion is a possible, stable Normal Mode.

In this mode the two masses oscillate in lock-step at the common frequency w=1 k/m.



Now let us ask if there are any other normal modes for this system. It turns out that this system does have one other normal mode. Let us try to guess this second mode.



From the symmetry, we guess that if a and b move equally and oppositely we might have a mode. In particular, if at any matant $t_b = -t_a$, the masses are subjected to equal magnitude (oppositely directed) excess forces.

Consider the left-hand mass a. It is pulled to the left by the left hand spring with an excess force F=-Rya. But it is also pushed to the left by the middle spring with an excess force $F = -2 k \gamma_a$. (The factor of two comes in because the central spring. is compressed by a total amount = 2 /a). Thus the net force on Wa is - 3R Va. A similar analysis (Hink it through) shows that the net force on b is - 3R4b. Bout since, by hypothesis 4=-4a, at any instant the forces on the two masses are the same in magnitude but oppositely duected. We have, at any t>0, for this motion

$$M \frac{d^2 \psi_a}{dt^2} = -3k \psi_a$$

$$M \frac{d^2 \psi_b}{dt^2} = -3k \psi_b$$

 $\Rightarrow \quad \forall_a(t) = A \cos \omega t$ $\forall_b(t) = -A \cos \omega t$

pount: Still true even

if signs of both tha

the are simultaneously

neversed (different
IC'S)

$$W = \sqrt{\frac{3k}{m}}$$

(The relative minus sign comes from the nutial conditions).

Thus we have verified that this system has a second normal mode. In this higher frequency normal mode

the two masses oscillate 180° out of phase with each other ("against each other").

