# Dynamics of Large-Scale Atmospheric Flows

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# **Basics**

#### Notation

$$\frac{D_h}{Dt} = \frac{\partial}{\partial t} + u \frac{\partial}{\partial x} + v \frac{\partial}{\partial y}$$
$$\frac{D}{Dt} = \frac{D_h}{Dt} + w \frac{\partial}{\partial z}$$
$$i = (1,0,0)$$
$$j = (0,1,0)$$
$$k = (0,0,1)$$

#### Coordinate system

#### Spherical coordinate system

$$x = r \sin \phi \cos \lambda$$
$$y = r \sin \phi \sin \lambda$$
$$z = r \cos \phi$$

#### Unit vectors on the sphere

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} : i = \begin{pmatrix} -\sin \lambda \\ \cos \lambda \\ 0 \end{pmatrix}$$
$$j = \begin{pmatrix} -\sin \theta \cos \lambda \\ -\sin \theta \sin \lambda \\ \cos \theta \\ \cos \theta \cos \lambda \\ \cos \theta \sin \lambda \\ \sin \theta \end{pmatrix}$$

#### Wind vectors on the sphere

$$\frac{Di}{Dt} := u = r \cos \phi \frac{D\lambda}{Dt}$$

$$\frac{Dj}{Dt} := v = r \frac{D\phi}{Dt}$$

$$\frac{Dk}{Dt} := w = \frac{Dr}{Dt}$$

#### **Basic** equations

6 equations for 6 unknowns  $u, v, w, \rho, p, T$ 

#### Equations of motion (Navier-Stokes)

x,y,z are spherical unit coordinates (eastward, northward, vertical). Equations are simplified through  $\beta$ -plane.

#### $\beta$ -plane approximation

$$f = f_0 + \beta y$$
,  $\beta = \frac{2\Omega\cos\phi_0}{a}$   
a = Earth radius

$$\frac{Du}{Dt} - fv = -\frac{1}{\rho} \frac{\partial p}{\partial x}$$
$$\frac{Dv}{Dt} + fu = -\frac{1}{\rho} \frac{\partial p}{\partial y}$$
$$\frac{Dw}{Dt} = -\frac{1}{\rho} \frac{\partial p}{\partial z} - g$$

#### Continuity equation

$$\frac{D\rho}{Dt} + \rho(\nabla \cdot \boldsymbol{u}) = 0$$

#### Equation of state

$$pV = nRT$$
$$\Leftrightarrow p = \rho RT$$

#### Thermodynamic equation

$$\frac{Dln\theta}{Dt} = \frac{1}{\theta} \frac{D\theta}{Dt} = \frac{1}{c_p T} \frac{DH}{\underbrace{Dt}}_{\substack{\sum \text{diabatic} \\ \text{processes}}}$$

# Large-scale approximations (Synoptic-Scale Motions)

Large-scale approximations hold for the sea breeze and cumulus scale.

#### Vertical component decomposition

Due to the strong vertical stratification of the atmosphere, it is useful to decompose a field variable  $\chi$  as follows.

$$\chi = \chi_0(z) + \chi^*(x, y, z, t)$$

At a certain height,

$$\chi_0 \gg |\chi^*|$$

This holds for  $\chi = p, \rho, \theta, T$ 

Equation of state approx

$$\frac{\rho^*}{\rho_0} = \frac{p^*}{p_0} - \frac{T^*}{T_0}$$

Potential temperature approx

$$\frac{\rho^*}{\rho_0} \approx -\frac{\theta^*}{\theta_0}$$

Vertical momentum equation approx

$$\frac{Dw}{Dt} \approx -\left(\frac{\partial}{\partial z} - \frac{N^2}{g}\right) \frac{p^*}{p_0} + g \frac{\theta^*}{\theta_0}$$

#### Hydrostatic approximation

Using: Vertical momentum equation

$$\frac{Dw}{Dt} \approx 0 \Rightarrow \frac{\partial p}{\partial z} = -\rho g$$

#### Geostrophic wind

Using: Horizontal momentum equation

$$-fv \approx -\frac{1}{\rho} \frac{\partial p}{\partial x}$$
$$fu \approx -\frac{1}{\rho} \frac{\partial p}{\partial y}$$

Equivalently,

$$v_G = \begin{pmatrix} u \\ v \\ 0 \end{pmatrix} = k \times \frac{1}{f\rho} \nabla_h p$$

 Geostrophic wind field is nondivergent

$$\nabla_h(\rho_0 v_G) = -\left(\frac{1}{f} \frac{\partial f}{\partial y'}\right) \rho_0 v_G < 0$$

 Geostrophic approximation not valid for large Rossby numbers (Ro >> 1)

#### Thermal wind

Using: Hydrostatic & geostrophic approximations & equation of state

$$\frac{\partial v_G}{\underline{\partial z}} = \frac{g}{fT} k \times \nabla_{\mathbf{h}} T$$

- Left turning of geostrophic wind => Cold air advection
- Right turning of geostrophic wind => Warm air advection

Aka thermal wind relationship is the following

$$\frac{g}{\theta_0} \frac{\partial \theta^*}{\partial x} = f_0 \frac{\partial v_G}{\partial z}$$

# Vorticity

 $\zeta$ : Vertical component of vorticity

$$\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} = \mathbf{k} \times (\nabla \times \mathbf{u})$$

$$\zeta = \begin{cases} \text{cyclonic circulation,} & \zeta > 0\\ \text{anticyclonic circulation,} & \zeta < 0 \end{cases}$$

Vorticity equation

$$\frac{D\zeta}{Dt} + \beta v = \underbrace{-\frac{(\zeta + f)(\nabla_h v)}{\text{divergence effect}}}_{\text{divergence effect}} - \underbrace{-\frac{\partial w}{\partial x} \frac{\partial v}{\partial z} - \frac{\partial w}{\partial y} \frac{\partial u}{\partial z}}_{\text{twisting/tilting}} + \underbrace{\frac{1}{\rho^2} \left(\frac{\partial \rho}{\partial x} \frac{\partial p}{\partial y} - \frac{\partial p}{\partial y} \frac{\partial p}{\partial x}\right)}_{\text{solenoidal effect}}$$

Synoptic scale approximation

$$\frac{D_h}{Dt}\zeta + \beta v = -f_0(\nabla_h v)$$

 $\beta v$ : Meridional excursion

 $-f_0(\nabla_h \mathbf{v})$ : Flow convergence / divergence

Assumptions

- Invalid in frontal regions with  $L \approx 10^5$ ,  $\zeta \approx f$ ,  $w \approx 10^{-1} \text{ms}^{-1}$
- Caution:  $f \le 10^{-5} \text{s}^{-1}$  in tropical and equatorial regions
- If  $\nabla_h v \approx 0$ , then the vorticity equation is reduced to the **barotropic vorticity** equation.

# Quasi-geostrophic (QG) approximation

**Simplifications** 

$$f \approx f_0$$

$$\zeta \approx \zeta_G$$

$$\frac{D_h}{Dt} \approx \frac{D_G}{Dt}$$

$$\Psi = \frac{p^*}{f_0 \rho_0}$$

Subscript G denotes the usage of the geostrophic wind, i.e.  $f_G(\mathbf{u}) = f(\mathbf{u}_G)$ 

QG Equations

Geostrophic approximation 
$$u_G = -\frac{\partial \Psi}{\partial y}; v_G = \frac{\partial \Psi}{\partial x}$$

$$\begin{array}{ll} \text{Hydrostatic} & g \, \frac{\theta^*}{\theta_0} = f_0 \, \frac{\partial \Psi}{\partial z} \\ \text{Vorticity} & \frac{D_G}{Dt} \, \zeta_G + \beta \, \frac{\partial \Psi}{\partial x} = -f_0(\nabla_h v) \\ \text{Mass} & \nabla_h v + \frac{1}{\rho_0} \frac{\partial}{\partial z} (\rho_0 w) = 0 \\ \text{Thermodynam} & \frac{D_G}{Dt} \Big( f_0 \, \frac{\partial \Psi}{\partial z} \Big) + N^2 w = 0 \end{array}$$

# Implications on vorticity / westward tilt

- $\zeta_G = \nabla_h^2 \Psi \propto -p^*$ Positive / negative values of  $\zeta_G$  are associated with low / high pressure
- $\frac{\partial}{\partial z} \zeta_G = \nabla_h^2 \left( \frac{\partial}{\partial z} \Psi \right) = \frac{g}{f_0 \theta_0} \nabla_h^2 \theta^* \propto -\theta^*$ Positive / negative values of  $\frac{\partial}{\partial z} \zeta_G$  are associated with low / high temperature
- Both effects together lead to the westward slope of cyclones and anticyclones.

#### Ageostrophic wind

Split velocity in basic state and perturbation. The geostrophic wind  $oldsymbol{u}_G$  is the basic state while the ageostrophic wind  $oldsymbol{u}_a$  is the perturbation.

$$\mathbf{u} = \underbrace{(u_G, v_G, 0)}_{\mathbf{u}_G} + \underbrace{(u_a, v_a, w)}_{\mathbf{u}_a}$$

#### Solution of QG inconsistency

Geostrophic approximation  $\Rightarrow \nabla_h v_G = 0$ , but Mass conservation  $\Rightarrow \nabla_h v \neq 0$ 

Solution: 
$$\mathbf{v} = (u_G, v_G) + (u_a, v_a)$$
 where  $\nabla_h(u_a, v_a) \neq 0$ , so  $\nabla_h v = \underbrace{\nabla_h \mathbf{v}_G}_{=0} + \nabla_h \mathbf{v}_a \neq 0$ 

#### Some properties

- $div v_G = 0$
- $div v_a \neq 0$
- $|u_a| \ll |v_c|$

# Linkage of ageo- and geostrophic wind

Assumptions:  $N^2 = const.$ ,  $\beta = 0$ ,  $\rho_0 = const.$ 

From thermodynamic equation and vorticity equation using hydrostatic approximation.

$$N^2 \frac{\partial w}{\partial x} - f_0^2 \frac{\partial u_a}{\partial z} = 2Q_1$$

$$N^{2} \frac{\partial w}{\partial y} - f_{0}^{2} \frac{\partial v_{a}}{\partial z} = 2Q_{2}$$

$$Q_{1} = f_{0} \left( \frac{\partial u_{G}}{\partial z} \frac{\partial v_{G}}{\partial x} + \frac{\partial v_{G}}{\partial z} \frac{\partial v_{G}}{\partial y} \right)$$

$$= -\frac{g}{\theta_{0}} \left( \frac{\partial u_{G}}{\partial x} \frac{\partial \theta^{*}}{\partial x} + \frac{\partial v_{G}}{\partial x} \frac{\partial \theta^{*}}{\partial y} \right)$$

$$Q_{2} = -f_{0} \left( \frac{\partial u_{G}}{\partial z} \frac{\partial u_{G}}{\partial x} + \frac{\partial v_{G}}{\partial z} \frac{\partial u_{G}}{\partial y} \right)$$

$$= -\frac{g}{\theta_{0}} \left( \frac{\partial u_{G}}{\partial y} \frac{\partial \theta^{*}}{\partial x} + \frac{\partial v_{G}}{\partial y} \frac{\partial \theta^{*}}{\partial y} \right)$$

# Diagnostic w equation (vertical wind)

$$2 \nabla \cdot \boldsymbol{Q} = N^2 (\nabla_h^2 w) + f_0^2 \left( \frac{\partial^2 w}{\partial z^2} \right)$$

With 
$$2\nabla \cdot \boldsymbol{Q} = 2\left(\frac{\partial}{\partial x}Q_1 + \frac{\partial}{\partial y}Q_2\right)$$

# Ageostrophic wind is completely determined by geostrophic wind!

#### 4-step golden rule

- 1.  $\nabla \cdot O \approx \nabla^2 w \approx -w$
- 2.  $\frac{D_G}{Dt} \zeta_G \propto \frac{\partial}{\partial z} (\rho_0 w)$ 3.  $\frac{D_G}{Dt} \zeta_G \begin{cases} > 0 & \text{Cyclogenesis} \\ < 0 & \text{Anticyclogenesis} \end{cases}$
- Combination of vorticity equation and mass conservation
- 3. Approximating  $\zeta$  with  $\frac{D_G}{D_A} \zeta$ ?

#### Finding $\nabla \cdot Q$

- 1. Find largest Q-vector. Arrowhead is zone of convergence ( $\nabla \cdot Q < 0$ ), tail is zone of divergence
- 2.  $Q = -\frac{g}{\theta_0} |\nabla_h \theta^*| \left( \mathbf{k} \times \frac{\partial}{\partial \xi} \mathbf{v}_{\mathbf{G}} \right)$
- 3. Find largest temperature gradient and a strong wind change along the isentrope

#### Thermal steering effect

Neighboring low and high pressure cells tend to move perpendicular to the isentropes. This coincides with the direction of thermal wind.

#### Development on the left exit of a jet

From a jetstream's point of view, on the left side of its exit, there is a zone of divergence and associated with it, there is upward motion.

#### QG Potential Vorticity

- Combination of vorticity equation and mass conservation.
- 2. Combine this with thermodynamic equation
- 3. Interchange  $\frac{\partial}{\partial z}$  and  $\frac{D_G}{Dt}$

$$\frac{D_G}{Dt} \left[ \underbrace{\zeta_G + \frac{f_0^2}{\rho_0} \frac{\partial}{\partial z} \left( \frac{\rho_0}{N^2} \frac{\partial \Psi}{\partial z} \right)}_{=:a = OGPV} + \beta y \right] = 0$$

- $q \approx \nabla^2 \Psi$
- q = relative vorticity + static stability
  - $\circ \quad \text{Static stability: From hydrostatic} \\ \text{equation: } \frac{\partial}{\partial z}\theta^* \propto \frac{\partial^2}{\partial z^2}\Psi$
- Only valid for adiabatic flow.
- Boundary condition:  $\frac{D_G}{Dt} \left( \underbrace{f_0 \frac{\partial \Psi}{\partial z}}_{g \theta^* / \theta_0} \right) = 0$

#### QG Prognostic system

- 1.  $q(t_0)$  known with B.C.  $\theta^*(t_0, z = 0)$
- 2. Main diagnostic step
  - a. Solve  $q = \nabla^2 \Psi$
  - b.  $\Psi$  yields  $u_G$ ,  $v_G$  and  $\theta^*$
  - c. These yield the Q-vectors
  - d. The Q-vectors yield w
- 3. Do prognostic step using passive advection of  $\frac{D_G}{Dt}(q + \beta y) = 0$  and go to 2.

#### QGPV key properties

- QG flow determined by interior distribution of QGPV and surface distribution of  $\theta^*$
- Conservation:  $\frac{D_G}{Dt}(q + \beta y) = 0$
- Invertibility:  $q = \nabla^2 \Psi$
- Partition and attribution:  $q = \sum_i q_i$

#### Prototype vortex for QG flow

Since QGPV can be decomposed into "atomic" vortices, we can analyse prototype vortices.

Given:

•  $q = \begin{cases} p, & r < a \\ 0, & r > a \end{cases}$ 

Then:

•  $v = \begin{cases} \Omega R, \ r < a \\ \frac{k}{r^2}, \ r \ge a \end{cases}$ 

 Here, r is the spherical symmetrical coordinate, and R is the "x" coordinate

#### Ertel PV

Idea: Analogue derivation to QGPV, but starting from full NSE.

Applicable only for inviscid and adiabatic flow.

$$\frac{D}{Dt} \left[ \underbrace{\frac{1}{\rho} \boldsymbol{\eta} \cdot \boldsymbol{\nabla} \theta}_{\text{Extel PV}} \right] = 0$$

•  $\eta = rot \ u + 2\Omega$ , absolute vorticity

#### Isentropic PV

$$\frac{D}{Dt} \left[ \underbrace{\frac{1}{\rho} (\zeta + f) \cdot \frac{\partial}{\partial z} \theta}_{IPV} \right] = 0$$

- IPV = z-component of PV
- $\bullet \quad \frac{1}{\zeta} \propto \frac{\partial}{\partial z} \theta$

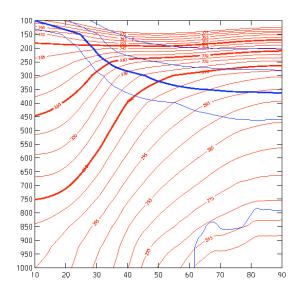
#### Preliminaries for PV chart analysis

- Unit for PV: 1 pvu. 2 pvu are defined as the tropopause.
- $\frac{\partial \theta}{\partial z} > 0 \& \theta$  increases towards equator.
- Max(PV) at equator due to Coriolis parameter.
- $\frac{\partial PV}{\partial z} > 0$ , because  $\frac{\partial \theta}{\partial z} > 0 \& \frac{\partial}{\partial z} \left(\frac{1}{\rho}\right) > 0$
- For adiabatic and inviscid flow

$$\frac{D\theta}{Dt} = 0 & \frac{DPV}{Dt} = 0$$

#### PV anomaly / cyclogenesis

- Adiabatic movement of air parcel towards south => stays on isentrope.
- As it must retain PV, it "drags" its PV value along the isentrope, especially when intersecting the 2 pvu tropopause. A PV anomaly occurs.
- Because of far-field effect of a strong PV gradient (=> strong wind), a cyclone forms at ground.



#### Idealised PV situations

Effect: Strong PV gradient w.r.t latitude => Strong wind.

Reason: High PV => CCW flow, low PV => CW flow. Strongest wind at strongest gradient. C.f. Jetstream and PV

Effect: Positive / negative upper-level PV anomaly ⇒ CCW / CW circulation and cold / warm below, warm / cold above Reason: Positive upper-level PV anomaly => higher  $\zeta$  (=> CCW flow ) and lower  $\frac{\partial \theta}{\partial z}$ 

Effect: Positive / negative surface temp. anomaly ⇒ CCW / CW circulation and warm / cold above

Reason: Negative sfc temp: Isentropes squeezed, higher  $\frac{\partial \theta}{\partial z}$  => lower  $\zeta$  => CW flow

Effect: Diabatically produce low-level PV anomalies

Reason: Microphysical processes

Effect: PV tower. If all anomalies are aligned on top of each other, they intensify and form a devastating storm.

#### Diabatic PV

$$\frac{D}{Dt}PV \approx -g(f+\zeta)\frac{\partial\dot{\theta}}{\partial p}$$

**Important**:  $\frac{D}{Dt}PV$  is only proportional to  $\frac{\partial \dot{\theta}}{\partial p'}$ , not  $\dot{\theta}$ 

Insert: Image with diabatic PV through a cloud

# QG Wave theory

#### **Examples**

- Large scale flows in mid and uppertroposphere
- Planetary scale quasi-stationary and quasi-steady wave features in stratosphere

## Waves on a uniform zonal flow

- Set:  $\Psi(y) = -\mathcal{U}y$
- Split quantities in basic state and perturbations

$$\psi = \Psi(y) + \psi'(x,y,z,t), |\psi'| \ll |\Psi|$$

From the QG potential vorticity equation, inserting  $\psi$  and linearizing (i.e. discarding terms like a'b') yields the perturbation equation:

$$\left(\frac{\partial}{\partial t} + \mathcal{U}\frac{\partial}{\partial x}\right)q' + \frac{\partial \bar{q}}{\partial y}v' = 0$$

- $\Psi(y) = -\mathcal{U}y$ ,  $\nabla_h^2 \Psi = 0$  and  $\frac{\partial}{\partial z} \Psi = 0$ , thus  $\bar{q} = 0 + \beta y = \beta y$
- $q' = \nabla_h^2 \psi' + \frac{f_0^2}{\rho_0} \frac{\partial}{\partial z} \left( \frac{\rho_0}{N^2} \frac{\partial \psi'}{\partial z} \right)$ QGPV definition

#### 2d wave solution (aka Rossby wave)

Solution on a mid-latitude band (all longitudes, latitudes 30-90°N)

$$\psi' = A\sin(kx - \omega t)\sin(l(y+d))$$

#### Dispersion relationship

$$\omega = \mathcal{U}k - \frac{\beta k}{k^2 + l^2}$$

 $\bullet \quad k = \frac{2\pi}{\mathcal{L}} = \frac{2m}{a}$ 

East-West wavelength:  $m\mathcal{L} = \pi a$ 

•  $l = \frac{2\pi}{\mathcal{M}} = \frac{3m'}{a}$ • North-South half-wavelength:  $m'\left(\frac{\mathcal{M}}{2}\right) =$ 

#### Phase velocities

• 
$$u_p = \frac{\omega}{k} = \mathcal{U} - \frac{\beta}{k^2 + l^2}$$
  
•  $v_p = \frac{\omega}{l}$ 

• 
$$v_p = \frac{\alpha}{l}$$

**Effects** 

m'=1;m	1	2	3	4
β	-35.5	-18.5	-6.0	-1.7
$-\frac{1}{k^2+l^2}$				

- Short waves (large m) typically move with u
- Long waves (small m) move against  $oldsymbol{u}$

#### Group velocities

• 
$$u_g = \frac{\partial \omega}{\partial k} = \mathcal{U} + \frac{\beta(k^2 - l^2)}{(k^2 + l^2)^2}$$

• 
$$v_g = \frac{\partial \omega}{\partial l} = 2\beta \frac{kl}{(k^2 + l^2)^2}$$

#### **Effects**

• 
$$k^2 < l^2 \Rightarrow (u_a - \mathcal{U}) < 0$$

• 
$$k^2 > l^2 \Rightarrow (u_q - \mathcal{U}) > 0$$

- Longitudinal waves propagate eastwards
- Latitudinal waves propagate westwards

#### More effects

$$\frac{v_g}{u_g} = \tan \alpha = \frac{2kl}{k^2 - l^2}$$

Westward translation of a sinusoidal displacement of a fluid

#### 3d wave solution

$$\psi'(x,y,z,t)=\chi'(x,y,z,t)e^{\frac{z}{2H_0}}$$

Yields a slightly different perturbation equation with solutions of the form:

$$\chi' = \psi(z)\sin(kx - \omega t)\sin l(y + d)$$

With the following constraints

$$\bullet \quad \psi_{zz} + n^2 \psi = 0$$

• 
$$n^2 = \left(\frac{N}{af_0}\right)^2 \left[\frac{\beta a^2}{u-c} - (4m^2 + 9m'^2 + \alpha^2 a^2)\right]$$

• 
$$(\omega - \mathcal{U}k)\left(\psi_z + \frac{1}{2H_0}\psi\right) = 0$$

#### Effect: Vertical propagability

• 
$$\psi \propto e^{\pm inz}$$
 for  $n^2 > 0$ 

$$\begin{array}{ll} \bullet & \psi \propto e^{\pm inz} & \quad \text{for } n^2 > 0 \\ \bullet & \psi \propto e^{\mp nz} & \quad \text{for } n^2 < 0 \end{array}$$

Discard exponential solution. Vertically propagating waves only occur for  $n^2 > 0$ 

$$n^2 > 0 \Leftarrow$$

- $(\mathcal{U}-c)>0$ : Wave velocity c must be easterly relative to zonal flow  ${\mathcal U}$
- $(\mathcal{U}-c)<rac{\beta}{k^2+l^2+lpha^2}$ : For typical values
- The above criteria are only satisfied for low zonal wavenumbers m, since

Zonal	β
wavenumber	$\overline{k^2 + l^2 + \alpha^2}$
m=1	$55  ms^{-1}$
m=2	$38  ms^{-1}$
m=3	$24  ms^{-1}$

Real effect: From the troposphere, only m=1 or m=2 waves propagate to the stratosphere.

#### Orographic forcing

Setting: Stationary, incompressible wave  $(\frac{D}{Dt} =$ 0,  $ho_0=const.$  ), subject to sinusoidal terrain

#### PV equation

$$\mathcal{U}\frac{\partial}{\partial x}\left[\psi_{xx}'+\psi_{yy}'+\frac{f_0^2}{N^2}\psi_{zz}'\right]+\beta\frac{\partial\psi'}{\partial x}=0$$

#### Terrain $(\eta)$ and Thermodynamic B.C.

$$\eta = \eta_0 \cos(kx) \sin\left[\frac{\pi}{2d}(y+d)\right]''$$
$$\frac{\partial \eta}{\partial x} = \frac{\omega'}{1}$$

#### Solution

$$\psi' \propto -\sin(kx + nz) \sin\left[\frac{\pi}{2d}(y+d)\right], n^2 > 0$$

$$\psi' \propto \cos(kx)e^{-\mu z} \sin\left[\frac{\pi}{2d}(y+d)\right], n^2 < 0$$

$$m = 1, 2, 3; m' = 1, 3;$$

There are large Rossby waves induced by (large) mountains such as the Rocky mountains, Himalaya or Greenland.

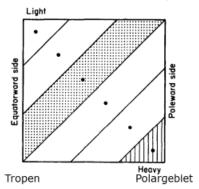
They propagate vertically for low wavenumbers (m <= 3) and decay vertically for high wavenumbers (m > 3)

#### Diabatic forcing

Large planetary scale diabatic heating distributions generate vertically propagating waves at high altitudes. At the surface, the low and high pressure centres are displaced

about ¼ wavelength eastward from centres of the diabatically heated and cooled regions.

# Baroclinic instability



Baroclinic atmosphere

#### Eady problem

Set:  $\mathcal{U} = \frac{\Lambda z}{d}$ , where d = height of atmosphere

#### **PV** equation

$$\left(\frac{\partial}{\partial t} + \mathcal{U}\frac{\partial}{\partial x}\right)\left(\frac{\partial^2 \psi'}{\partial x^2} + \frac{f_0^2}{N^2}\frac{\partial^2 \psi'}{\partial z^2}\right) = 0$$

#### Thermodynamic boundary condition

$$\left(\frac{\partial}{\partial t} + \mathcal{U}\frac{\partial}{\partial x}\right) f_0 \frac{\partial \psi'}{\partial z} - \left(f_0 \frac{\Lambda}{d}\right) \frac{\partial \psi'}{\partial x} = 0$$

#### **Ansatz**

$$\psi' = \Psi(z)e^{i(kx-\omega t)}$$

#### Solution

$$\Psi(z) = A \sinh(\mu z) + B \cosh(\mu z)$$

Where 
$$\mu^2 = \frac{k^2 N^2}{f_0^2}$$

$$\omega = k\sqrt{\gamma}$$

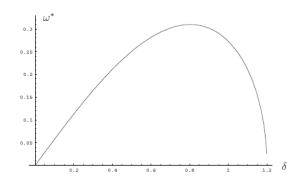
Where 
$$\gamma = -\left(\frac{\Lambda}{ud}\right)^2 (\delta - \tanh \delta)(\coth \delta - \delta)$$

Where  $\delta = \frac{1}{2}\mu d$ 

#### Exponential growth solution

If 
$$\sqrt{\gamma} \in \mathbb{C} \ (\Leftrightarrow \gamma < 0)$$
, then  $\psi' \propto e^{\omega t}$ 

$$\gamma < 0 \text{ for } 0 < \delta < 1.1997$$



Eady growth rate

# Nature of solution

- Wavelength  $(2\pi/k)$  for  $\delta=1.1997$   $\mathcal{L}\approx 2500~km$
- Wavelength for  $\delta_{max}$   $\mathcal{L} \approx 4000 \ km$
- Growth rate for  $\delta_{max}$

$$T_e = \frac{1}{\omega} \approx 1.1 \ days$$

 All perturbations of wavelength greater than 2500 km are unstable

#### **Notions**

- Cyclonic = Counter clockwise = Left turn
- Anticyclonic = Clockwise = Right turn
- Zonal wavenumber = m
- North-south petal count = m'
- Large waves ⇔ small m
- Short waves ⇔ large m
- $\mathbf{u} = (u, v, w)$
- $\bullet \quad \boldsymbol{v} = (u, v) = \boldsymbol{u}(0,1)$

#### Numerical values

Earth radius	а	6.37 · 10 <sup>6</sup> m
Gas constant of air	R	2.87
		$\cdot 10^{2} \text{JK}^{-1} \text{kg}^{-1}$
Specific heat at	$c_V$	7.17
constant volume		$\cdot 10^{2} \text{JK}^{-1} \text{kg}^{-1}$
Specific heat at	$c_p$	10.04
constant pressure	•	$\cdot 10^{2} \text{JK}^{-1} \text{kg}^{-1}$
Coriolis parameter	f(45°)	≈ 10 <sup>-4</sup>
on mid-latitude		

#### Quantities

Coriolis parameter

Full form

 $f = 2\Omega \sin \varphi$ 

 $\Omega$ : Rotation rate of Earth

 $\varphi$ : Latitude

 $\beta$ -plane approximation

$$f = f_0 + \beta y$$

$$f_0 = 2\Omega \sin \phi_0$$

$$\beta = \frac{2\Omega}{a}\cos\phi_0$$

a = radius of earth

 $f \approx 10^{-4}$  for mid-latitudes

f > 0 for the Northern hemisphere

f < 0 for the Southern hemisphere

#### Brunt-Vaisala frequency

$$N^2 = \frac{g}{\theta_0} \frac{d\theta_0}{dz}$$

#### Rossby number

Ratio of horizontal advection to Coriolis term

$$\frac{u\frac{\partial u}{\partial x}}{fv} \approx \frac{U}{fL} \equiv Ro$$

# Geopotential height

#### Geopotential

$$\Phi(h) = \int_0^h g(\phi, z) dz$$

 $\phi$ :latitude ,z: geometric height

The geopotential could also be expressed as a function of pressure

#### **Geopotential height**

As a function of geometric height

$$Z_g^h(h) = \frac{\Phi(h)}{g_0}$$

 $g_0$ : Standard gravity at mean sea level

As a function of pressure

$$Z_g^p(p) = \frac{\Phi(p)}{q_0}$$

The geostrophic wind  $v_G$  is parallel to the  $Z_a^p$  contours, and its magnitude is proportional to the distance between the  $Z_q^p$  contours.

#### Potential temperature

Temperature which an air parcel would acquire if adiabatically brought from level 1 to level 0.

$$T_0 = T_1 \left(\frac{p_0}{p_1}\right)^{\kappa} \equiv \theta(T_1, p_1)$$

$$\kappa \equiv R/c_p = 0.286$$
,  $p_0 = 1000 \text{ hPa}$ 

- $\frac{\partial \theta}{\partial z} > 0 \Rightarrow$  Stable atmosphere  $\frac{\partial \theta}{\partial z} < 0 \Rightarrow$  Unstable atmosphere
- $\theta(300K, p_0) = 300$  $\theta(220K, 200hPa) = 348$

# Circulation and vorticity

#### Vorticity

$$\zeta = k \cdot (\nabla \times u)$$

#### Mean vorticity

$$\bar{\zeta} = \frac{1}{S} \iint_{S} \zeta dS$$

## Circulation

$$C = \iint_{S} \zeta dS = \bar{\zeta}S$$

$$C = \oint v_q dr$$

# Mathematical tricks

Interchange of 
$$\frac{\partial}{\partial z}$$
 and  $\frac{D}{Dt}$ 

$$\frac{\partial}{\partial z}\frac{D}{Dt} = \frac{D}{Dt}\frac{\partial}{\partial z} + \frac{\partial u}{\partial z}\frac{\partial}{\partial x} + \frac{\partial v}{\partial z}\frac{\partial}{\partial y} + \frac{\partial w}{\partial z}\frac{\partial}{\partial z}$$

#### Product rule reversed

The product rule is normally used to expand a derivative of a product to a product of derivatives. This can be done the other way around, too.

# Neglecting small terms

$$\Psi = \overline{\Psi} + \Psi^*$$

where  $\Psi^* \ll \overline{\Psi}$  and  $\Psi$  is an arbitrary quantity.  $\overline{\Psi}$  is the basic state and  $\Psi^*$  is the perturbation.

$$\Psi \cdot \Phi = (\overline{\Psi} + \Psi^*)(\overline{\Phi} + \Phi^*)$$
$$\approx \overline{\Psi}\overline{\Phi} + \overline{\Psi}\Phi^* + \overline{\Phi}\Psi^*$$

#### Approximate functions