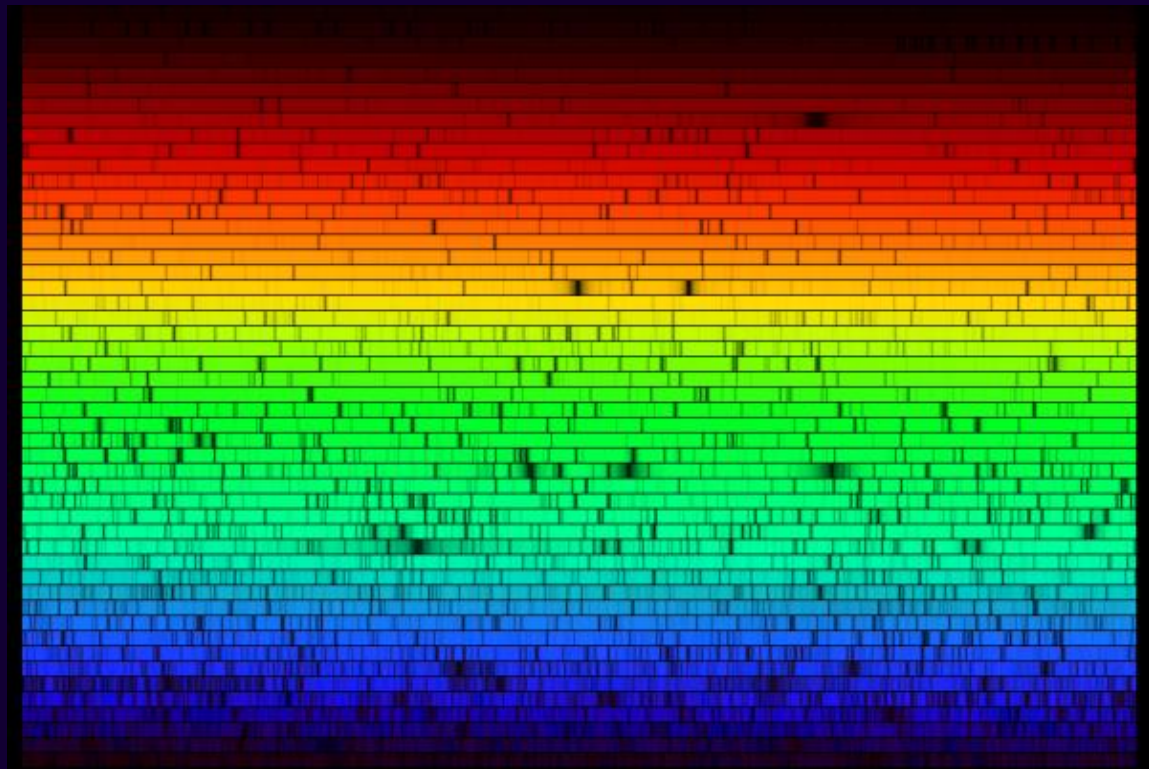


Light and matter

Astronomy 101
Syracuse University, Fall 2016
Walter Freeman

October 27, 2016



Announcements

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Chemistry done quick[†]

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But a *confined* wave can have only very particular frequencies – the resonant frequencies of the box you put it in.

Quantum mechanics says that electrons have wave properties, and that frequency corresponds to energy!

- In an atom, the electric force confines the electron-waves to a region near the nucleus (I can't show you this)

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Quantum mechanics says that electrons have wave properties, and that frequency corresponds to energy!

- In an atom, the electric force confines the electron-waves to a region near the nucleus (I can't show you this)
- In a drum, the frame of the drum confines the drumhead-waves to a circular cavity. I **can** show you this!
- The shape is different, but there is a close analogy here:
- The different wave-shapes in a drum correspond to the different atomic energy levels in chemistry!

Chemistry done quick[†]

A remarkable consequence of quantum mechanics:

Electrons in an atom can only have **very particular** amounts of energy!

These energy levels correspond to the resonant frequencies of the “box” they’re trapped in.

We measure this energy in “electron volts” (eV).

- Usually all the electrons live in the lowest available levels
- There’s a limit to how many electrons can be in each level
- Atoms “fill up” the levels starting from the bottom
- This process leads to the periodic table

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[†] Does not replace your introductory chemistry class on your transcript

Chemistry done quick[†]

- As in the drum, the energy levels in atoms are controlled by two numbers:
 - “Inside-to-outside wiggleness”: **principal quantum number n**
 - “Around-and-around wiggleness”: **angular quantum number l**

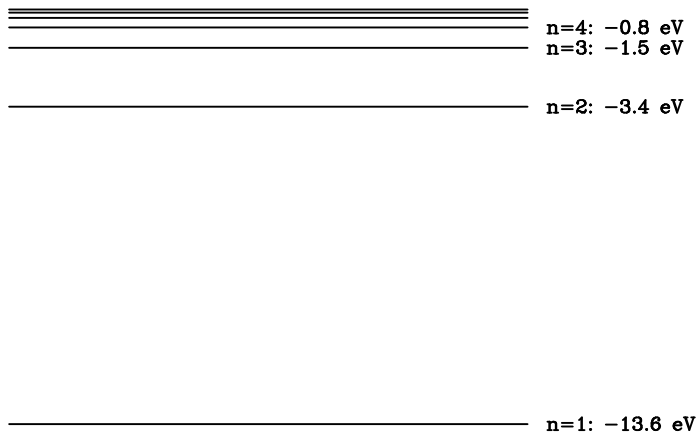
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 - “Inside-to-outside wiggleness”: **principal quantum number n**
 - “Around-and-around wiggleness”: **angular quantum number l**
- Chemists call $l = 0$ the S-orbitals, $l = 1$ the P-orbitals, $l = 2$ the D-orbitals, and so on.
- In hydrogen, the value of l doesn't affect the energy much at all.
- In other elements, it does a little bit.

Hydrogen

Here's a diagram of the allowed energies for hydrogen.

$E > 0$: electron escapes from atom



Chemistry: all I want you to know

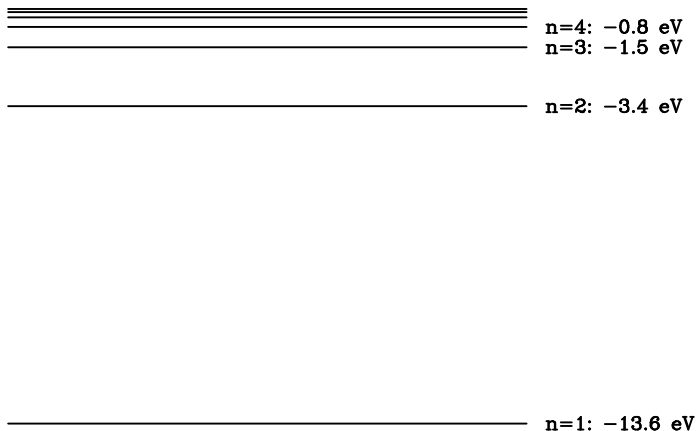
- Electrons occupy certain **energy levels**
- The particular energies that these levels have is **unique** to particular elements: hydrogen has different allowed energies than mercury or neon or sodium etc.

... that's it. :)

How do electrons get to different energy levels?

- They can *absorb* a photon and move to a higher energy level
- They can *emit* a photon and move to a lower energy level

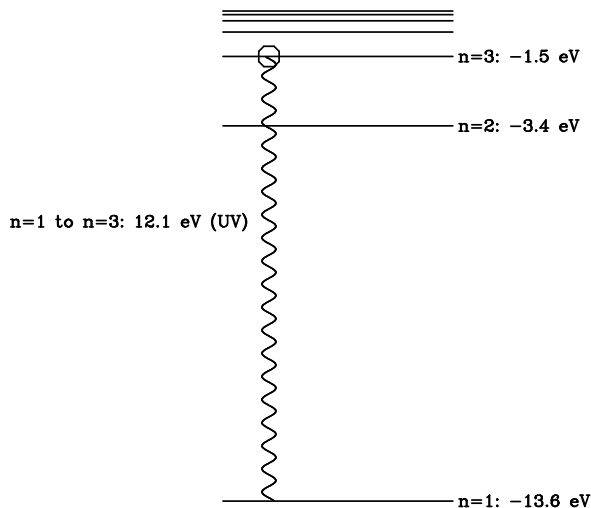
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Chemistry and light

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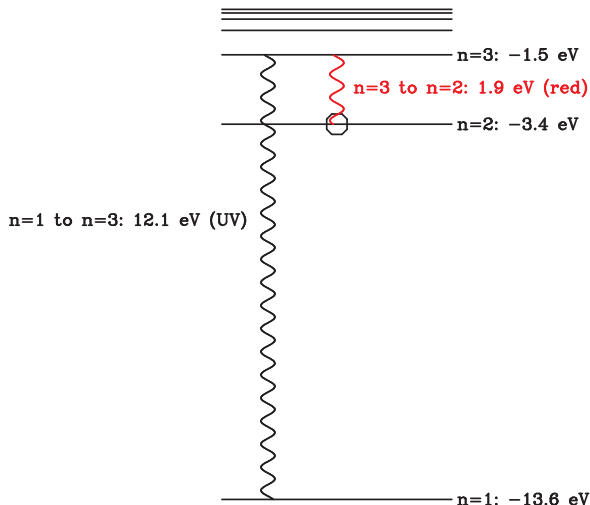
Absorb a 12.1 eV photon and move from $n = 1$ to $n = 3$; this photon is UV.



Chemistry and light

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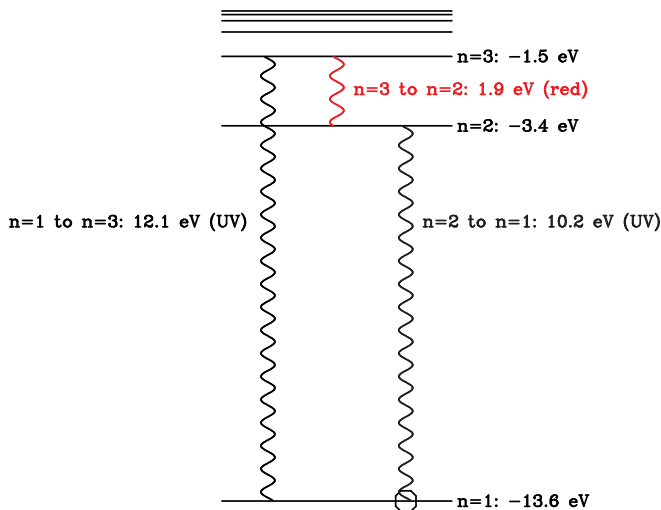
Emit a 1.9 eV photon and move from $n = 3$ to $n = 2$. This is red!



Chemistry and light

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- They can *emit* a photon and move to a lower energy level

Emit a 10.2 eV photon and move from $n = 2$ to $n = 1$. (UV, again)



If I take hydrogen and tear the electrons off of the atoms with an electric current, they'll "fall" back down, going through the energy levels down to $n = 1$.

Sometimes they'll skip energy levels; sometimes they'll go in sequence.

If I do this to hydrogen, what color will we see?

A: UV: we won't see it, since the transitions down to $n = 1$ are in the UV

B: Several shades of red: we'll see the transitions down to $n = 2$, which are red

C: Infrared: the transitions at the top are very low energy, corresponding to infrared light which we can't see

D: UV, IR, and red, all at once: all the transitions happen, but we only see the red photons because of the limits of our eyes

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E: Orange, because this is Syracuse, darnit!

Complete *Lecture Tutorials* pp.65-69.

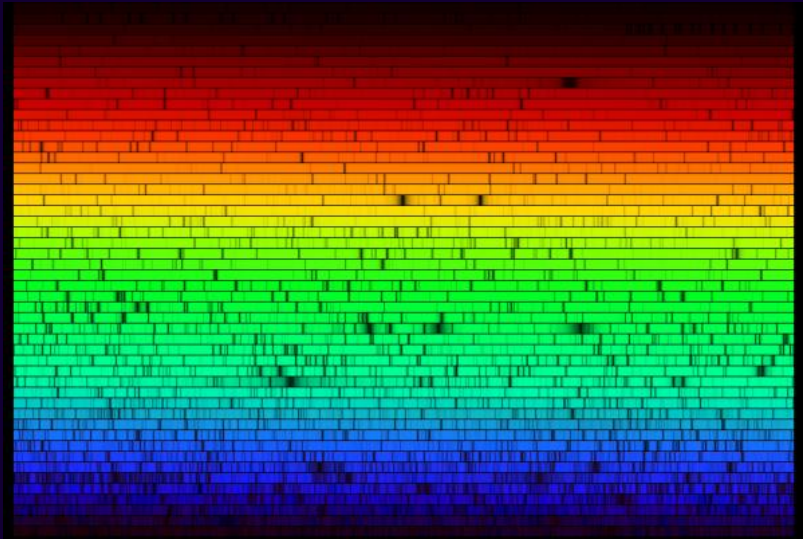
After this, we'll talk about another application of this idea.

Emission spectra

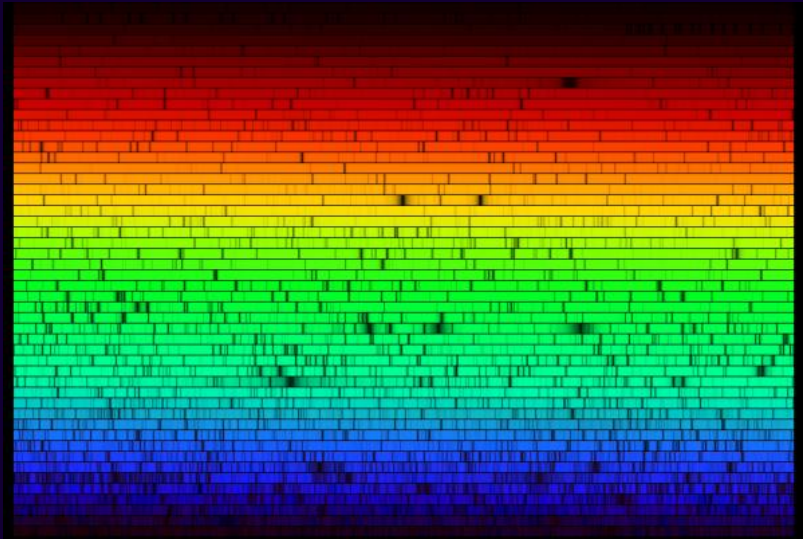
Every chemical element has a unique *spectrum*: the colors of light that it can emit and absorb.

Other colors simply pass through.

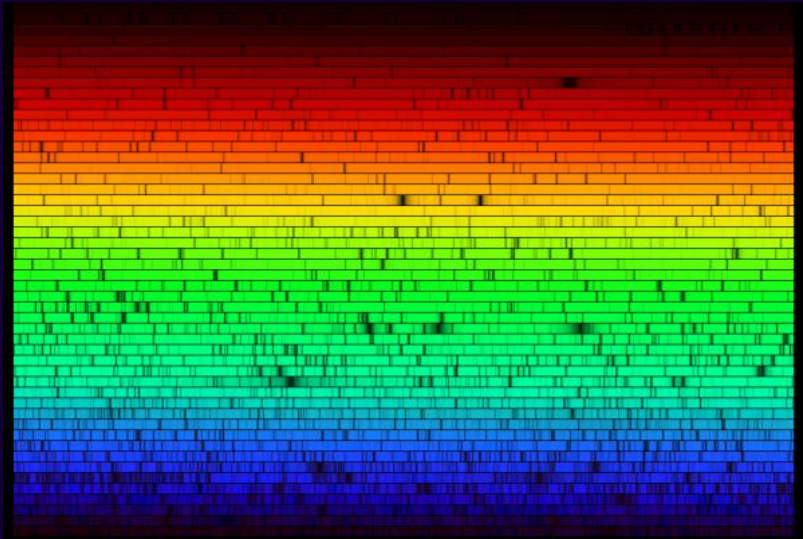
(Molecules have these spectra too: their electron energy levels are more complicated.)



- The hot core of the Sun emits light of all wavelengths (you'll learn why next class)



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- The gases in the cooler atmosphere absorb light of their particular wavelengths



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This picture tells us what's in the Sun!

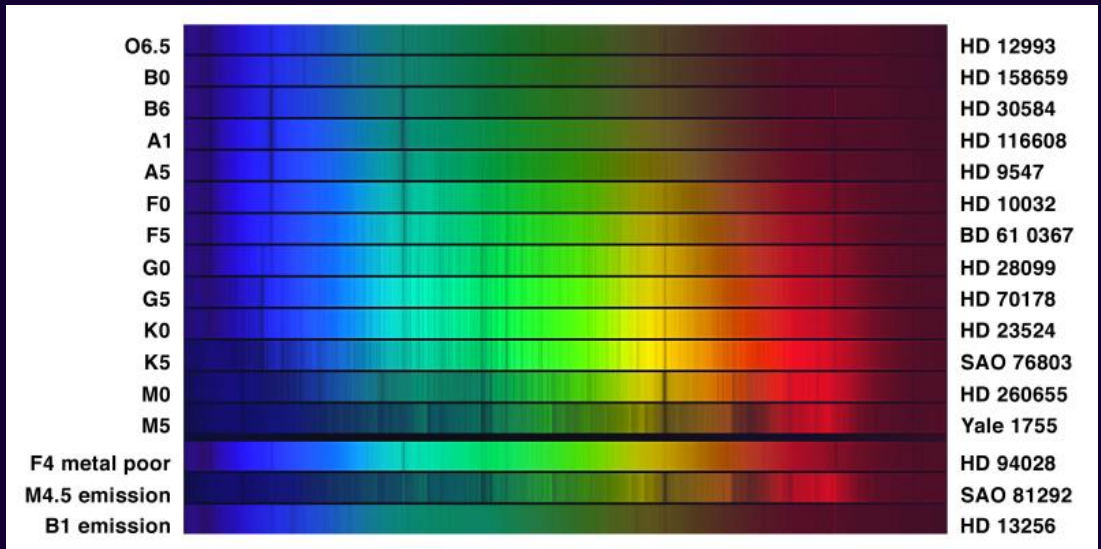
You discover lines in the solar spectrum that don't correspond to any known element. What do you conclude?

A: Something about quantum mechanics is different in the Sun

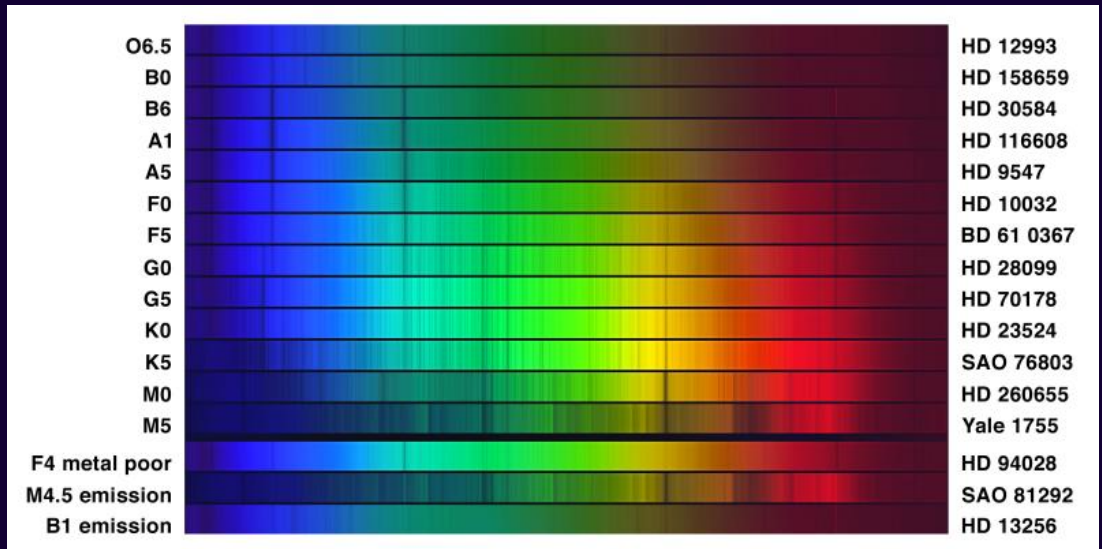
B: Something about light is different in the Sun

C: There's an element in the Sun that's not on Earth – call it **sunium**

D: The extreme temperature of the Sun causes new lines to appear in its gas



All the stars are made of the same stuff – the same stuff as we are.



All the stars are made of the same stuff – the same stuff as we are.

“The cosmos is also within us. We are made of star-stuff. We are a way for the universe to know itself.”

–Carl Sagan, *Cosmos*

What a lucky accident!

We're very lucky that atomic transitions happen to lie in our visual range!

There are others that are very interesting to astronomers:

- Molecular vibrations: infrared

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- Molecular vibrations: infrared
- Molecular *rotations*: microwave
- “Hyperfine structure” energy levels in hydrogen: 21 cm radio waves

This last is particularly interesting: it is a very particular frequency, echoing out from all corners of the Universe, that says: hydrogen is here. (Hydrogen is 75% of the universe.)

