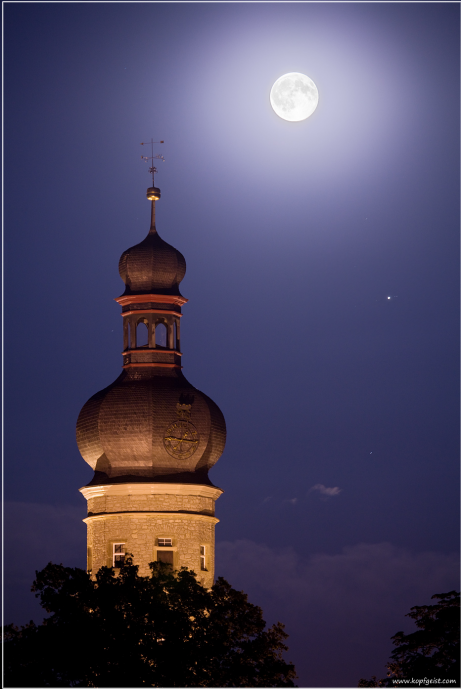


Kepler's laws

Astronomy 101
Syracuse University, Fall 2017
Walter Freeman

September 28, 2017



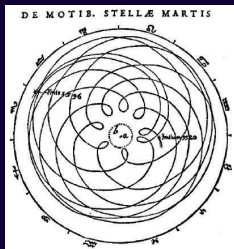
”And yet it moves.”

—Galileo (attributed), on the Earth

- Exam grades posted on Blackboard sometime before Tuesday

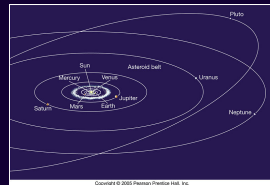
We left our story with two plausible models for the heavens:

The geocentric Ptolemaic model



- The planets (and everything else) revolve around Earth
- Inelegant system of “epicycles” needed to get planets right
- Everything moved in circles (elegant per Greeks)
- Earth and humanity at center (theologically not challenging)
- **Very accurate predictions**

The heliocentric Copernican model



- Earth is one of many planets, all orbiting the Sun
- Apparent motion = motion of Earth + motion of planets
- No (or very small) epicycles
- **Less accurate than Ptolemaic model**
- Matched Galileo's observations:
 - Moons of Jupiter
 - Phases of Venus

The Copernican model had a lot of attractive features, but was still less accurate – less good at actually telling you where the things in the sky would be!

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- What do we do?

What do we do when we don't know what to do?

Maybe our data are wrong...

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Enter Tycho Brahe.

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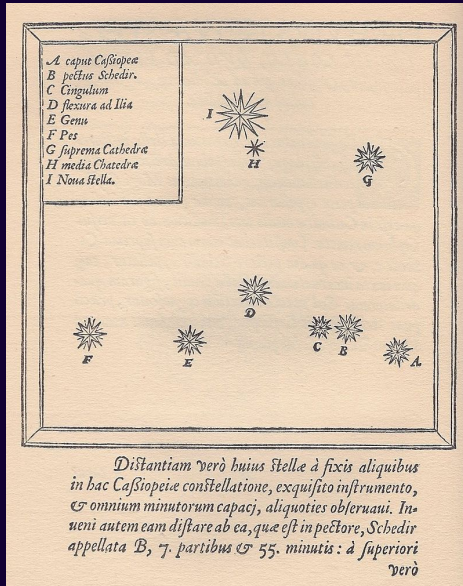
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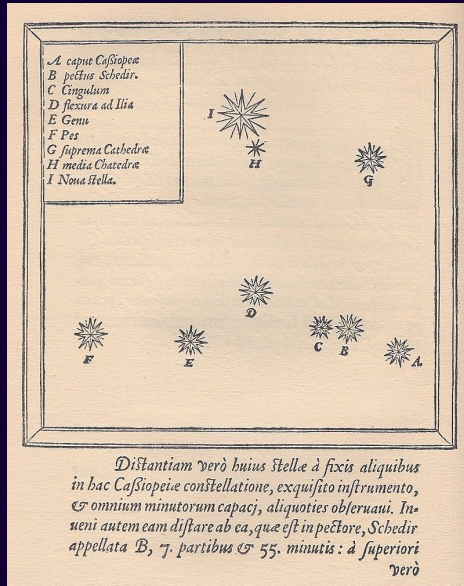
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- Was probably fun at parties (less so after his moose died)

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- Old worldview: world beyond the Moon is eternal
- ... nope: no observed parallax in the supernova → it's very far away!

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- Didn't observe parallax in the distant stars
- Two options:
 - The Earth doesn't move
 - The stars are very far away
- He believed the former
- Proposed another model for the Solar System

Tycho Brahe



- Danish nobleman and astronomer, 1546-1601
- Best known for his precision measurements of the sky from Uraniborg
- Made high precision observations of the motions of the planets and stars
- Even had a crude correction for atmosphere bending light
- Measurements accurate to a few minutes of arc ($1/60$ 'ths of a degree!)
- Made these measurements with his assistant Sophie...
- ... and his later assistant Johannes Kepler, who didn't murder him

You've probably been wondering when we're going to stop this history of false starts and learn how things actually *do* work...

Johannes Kepler

You've probably been wondering when we're going to stop this history of false starts and learn how things actually *do* work...

... here we go. Kepler, Tycho's assistant, finally got it right.



Kepler was a Copernican, and disagreed with his boss.

He tried to improve Copernicus' model, which used circular orbits, and mostly succeeded. But...

- Tycho's data were incredibly precise
- No matter how he rearranged the circles, there was an error of at least $8/60$ of a degree for Mars
- Kepler worked at Uraniborg – he knew how precise the data could be

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Do we:

- A: Reject the belief that Nature must be elegant
- B: Reject the need for our model to match the data precisely
- C: Reject Tycho's data?
- D: Reexamine our ideas about what elegance looks like

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Kepler didn't only want to discover *how the planets moved*; he wanted to know *why*. He didn't figure it out, but he was on the path that led to modern science.

Even if the *answer* doesn't have the perfect elegance of circles, modern science looks for its elegance in *laws*, not in all of their consequences! Kepler discovered the consequences; the laws weren't uncovered yet.

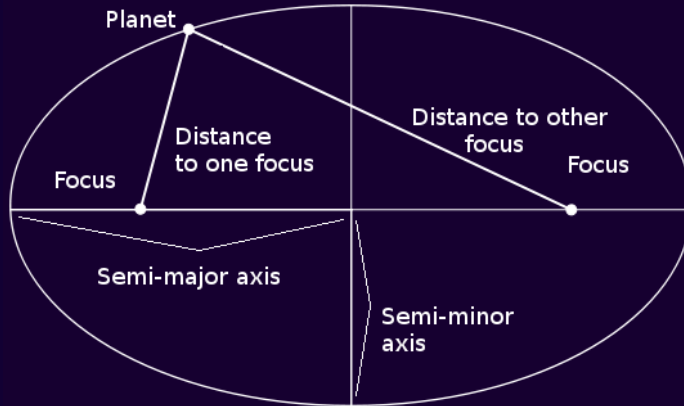
Kepler's laws of planetary motion

- The planets move in *ellipses*, with the Sun at one focus
- The line joining the planet and the Sun sweeps out equal areas in equal times
Alternate formulation: Within its orbit, a planet's speed is inversely proportional to its distance from the Sun
- The square of the orbital period of a planet is directly proportional to the cube of the semi-major axis of the ellipse.

Let's talk about each of these in turn.

Kepler's first law

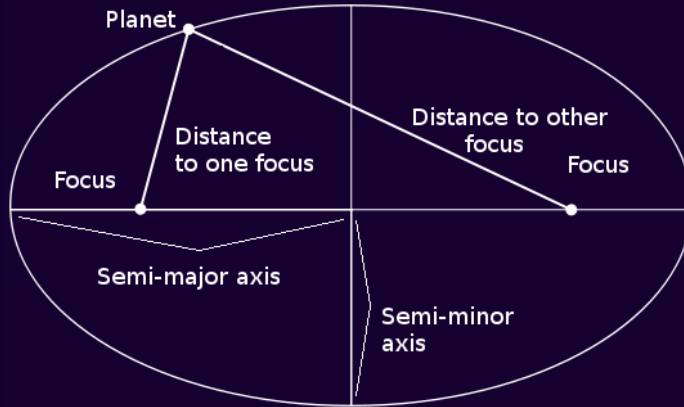
An ellipse is just a stretched circle. Mathematically: it's the curve around two points such that the *sum* of the distances to those points is a constant. A circle is just an ellipse with both foci at the same point.



Some terms:

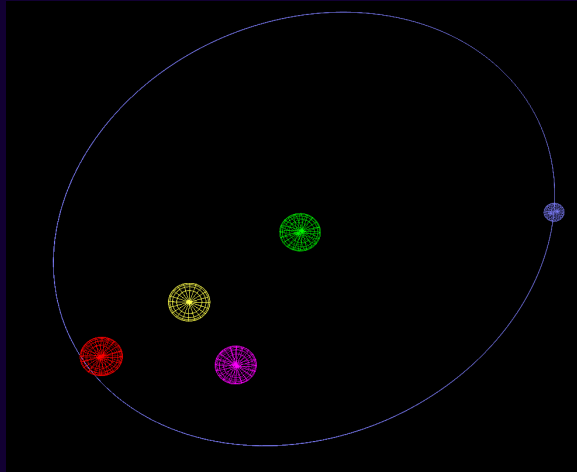
- Focus: One of the two points
- Semimajor axis: the largest distance from the center to the edge
- Eccentricity: how stretched out an ellipse is

Some properties of ellipses



- The two foci always lie along the major axis (“wide axis”)
- The closer together the foci, the less eccentric
- If both foci are exactly at the middle, you get a circle
- Both foci lie inside the ellipse

Here's an orbit. Which is the correct position for the Sun?



A: The red one

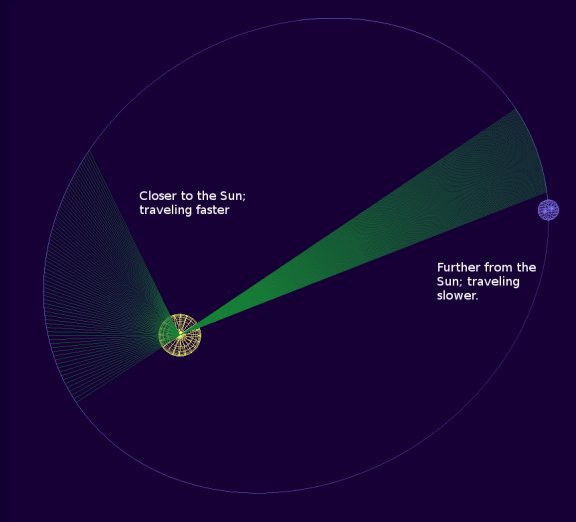
B: The green one

C: The yellow one

D: The purple one

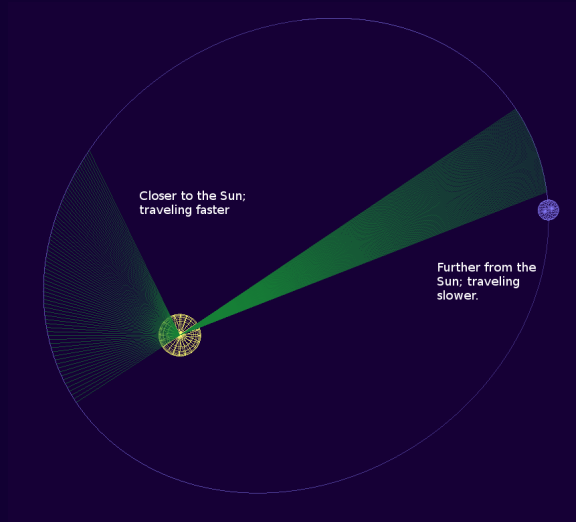
Kepler's second law

In an eccentric orbit, a planet travels fastest when it's nearer the Sun.



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Let's watch this in an animation...

Comets

Comets have highly eccentric orbits. Halley's Comet's furthest point from the Sun – its *aphelion* – is 35 AU away. But its *perihelion* – the nearest point to the Sun – is 0.6 AU away.

Which statement is true?

A: Halley's Comet spends most of its time far from the Sun, and only a little time near the Sun

B: Halley's Comet moves slowly near perihelion, and quickly at aphelion

C: Halley's Comet moves quickly near perihelion, and slowly at aphelion

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(Get creative with your folding...)

Complete *Lecture Tutorials* pp. 21-24.

We will do something else after this.

Kepler's Third Law

Kepler's third law of orbital motion says that the square of a planet's *orbital period* is proportional to the cube of its *semimajor axis*.

Simply put: if a planet is further from the Sun, it takes longer to go around.

If the distance is doubled, the time required *more than doubles*.

Let's watch this...

Complete *Lecture Tutorials* pp. 25-28.