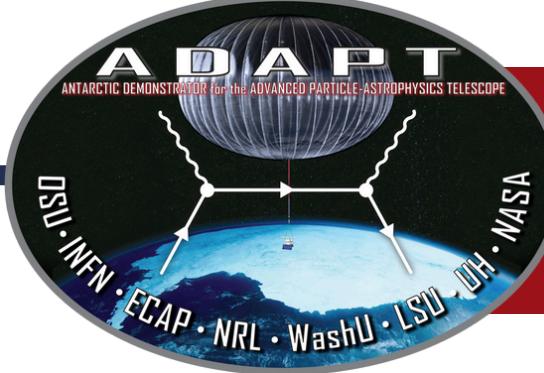


# AIRIS: HIGH-PRECISION OPTICAL FOLLOW-UP TELESCOPE FOR GAMMA-RAY BURST OBSERVATION WITH ADAPT

WashU Satellite Team, James Buckley, Ph.D.



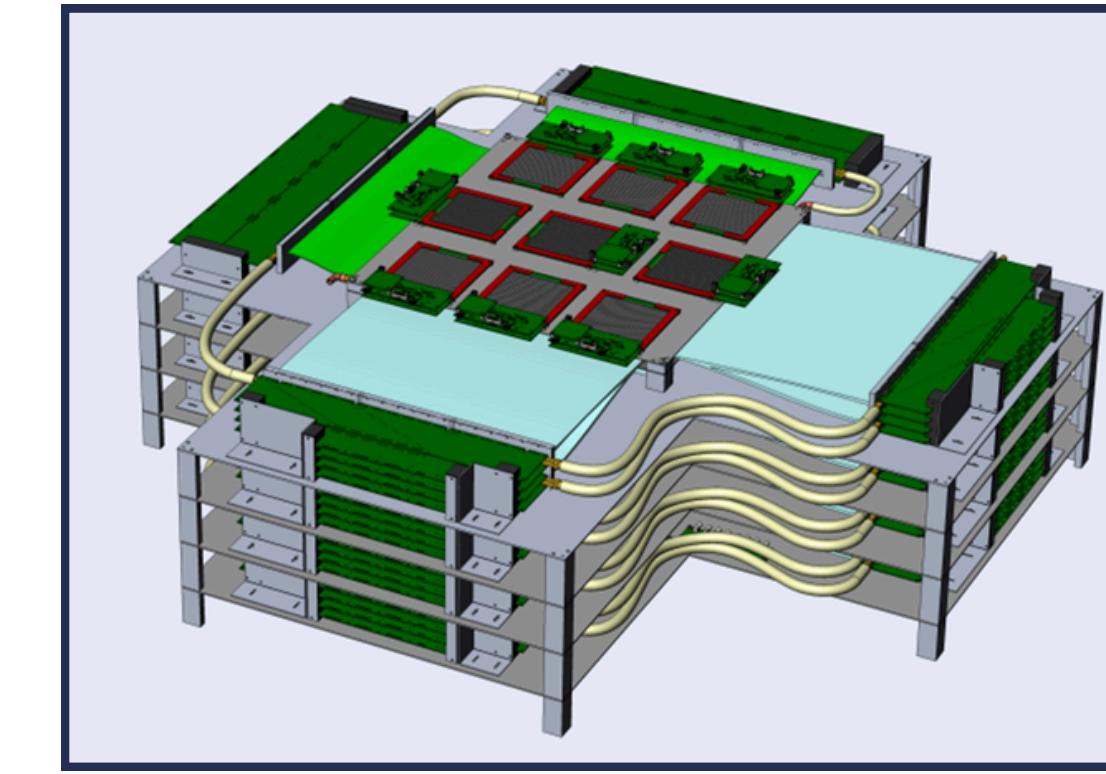
**AIRIS** delivers  $\leq 30$  s **optical follow-up** of **ADAPT** detected **GRBs**, improving localizations from **degrees to sub-arcseconds** and capturing **early afterglow evolution**.



## ADAPT: Full Sky Detection of GRBs

**ADAPT (Antarctic Demonstrator for the Advanced Particle-astrophysics Telescope)** is a NASA ballooning mission led by PI James Buckley, of WashU in St. Louis. Demonstrates a **Compton scattering detector** for future space-based **gamma-ray/cosmic-ray observer**.

- Aims to deliver degree-scale localizations and polarization constraints on fairly short timescales (sub-second to seconds).
- **Not designed for fine localization or on-board optical imaging/follow-up. Optical afterglow detection and evolution is outside its scope.**
- $\sim 1 \text{ MeV/cm}^2$  fluence with  $\sim 7.5^\circ$  accuracy (68% containment) and  $\sim 15^\circ$  accuracy (95% containment)



The ADAPT Detector Stack

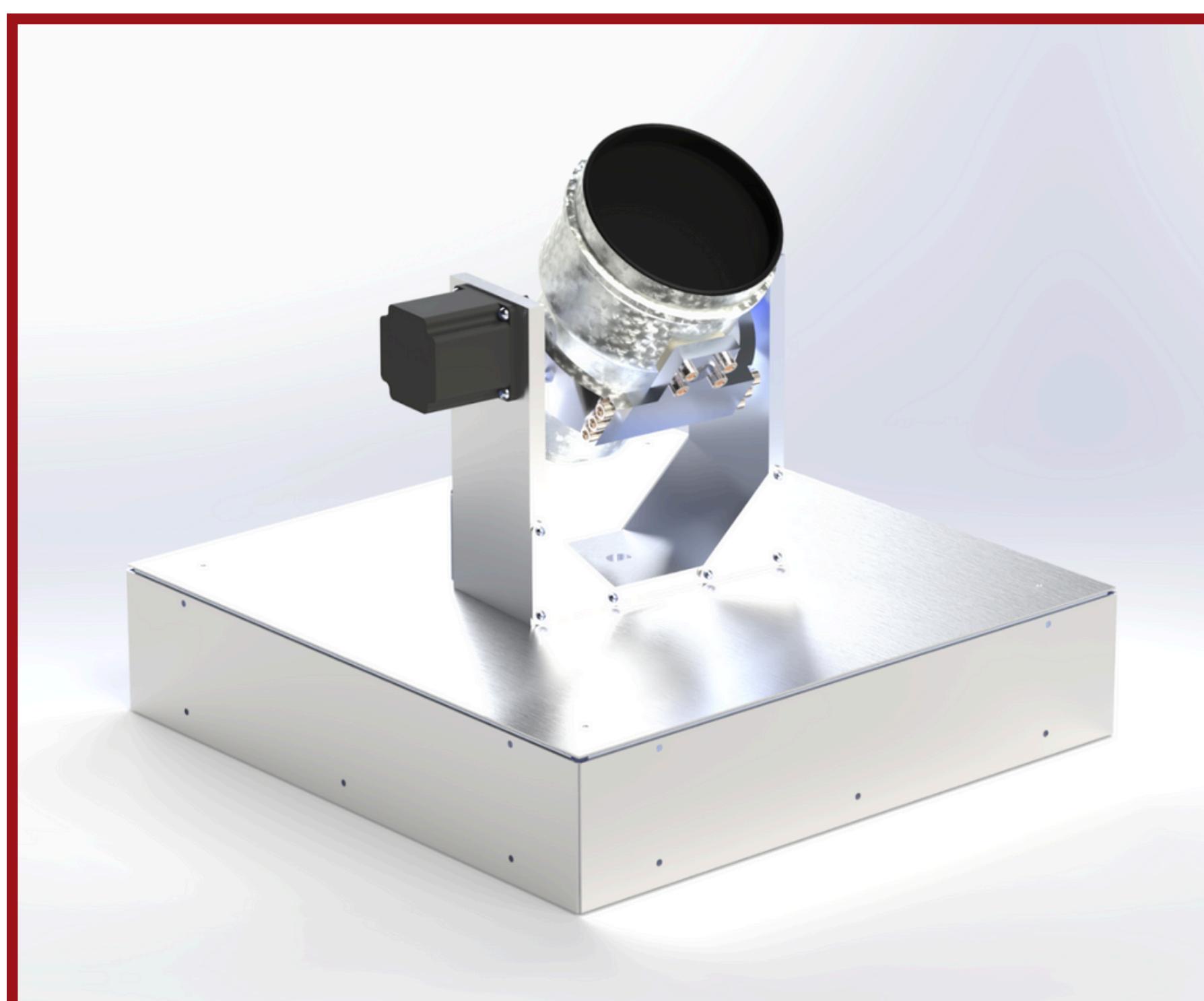


## AIRIS: Telescope for GRB Optical Afterglows

**AIRIS (ADAPT Incidence Resolution & Imaging Subsystem)** is a fast-slewing optical telescope mounted on the same platform as ADAPT.

- **Goal:** Capture burst optical data within 30 seconds of an ADAPT trigger. AIRIS is designed to reach any point in the sky within  $\sim 10$  seconds of slewing.
- **Response & localization:** On ADAPT triggers, AIRIS begins an imaging pipeline with stacked exposures to actively search the error region for an optical afterglow.
- **Data products:** Burst localization + time-evolution tracking (early light curves, positions) suitable for rapid follow-up.
- **Community alerts:** Results are pushed to NASA's General Coordinates Network (GCN) to coordinate global, multi-messenger observations.

Lens	Canon 200mm f/1.8
Bandpass Filter	<b>645–675 nm</b> Suppresses bright blue/green sky; avoids defocus in NIR
Sensor	<b>Sony IMX455 mono</b> High QE, single-electron read noise (for stacking)



AIRIS CAD Model

## Current Work

- **Target regime:** Early optical afterglows at  $m \approx 12\text{--}16$ .
- **Limiting-mag sims:** Physically accurate Python pipeline (PSF + Poisson shot noise + read/readout + thermal noise) with options for 3x3 binning, moving exposures, cosmic-ray hits, and Rayleigh-sky background; to be validated empirically with the flight sensor + lens.
- **Current performance (sim):**  $\text{SNR} \geq 3$  for  $m \lesssim 14$  in  $\gtrsim 10$  s effective exposure (non-stacking).
- **Onboard processing:** CMOS  $\rightarrow$  carrier  $\rightarrow$  NVIDIA Jetson GPU for real-time deblur & astrometric registration vs. a stored starmap; short individual exposures enable rapid tiling.
- **Map fusion & detection:** Frames accumulated in HEALPix; probability map compared at each step to the ADAPT localization PDF to flag transient candidates.
- **Follow-up & alerting:** On high-probability detection, AIRIS locks, continues photometry, and issues a GCN alert for community follow-up.

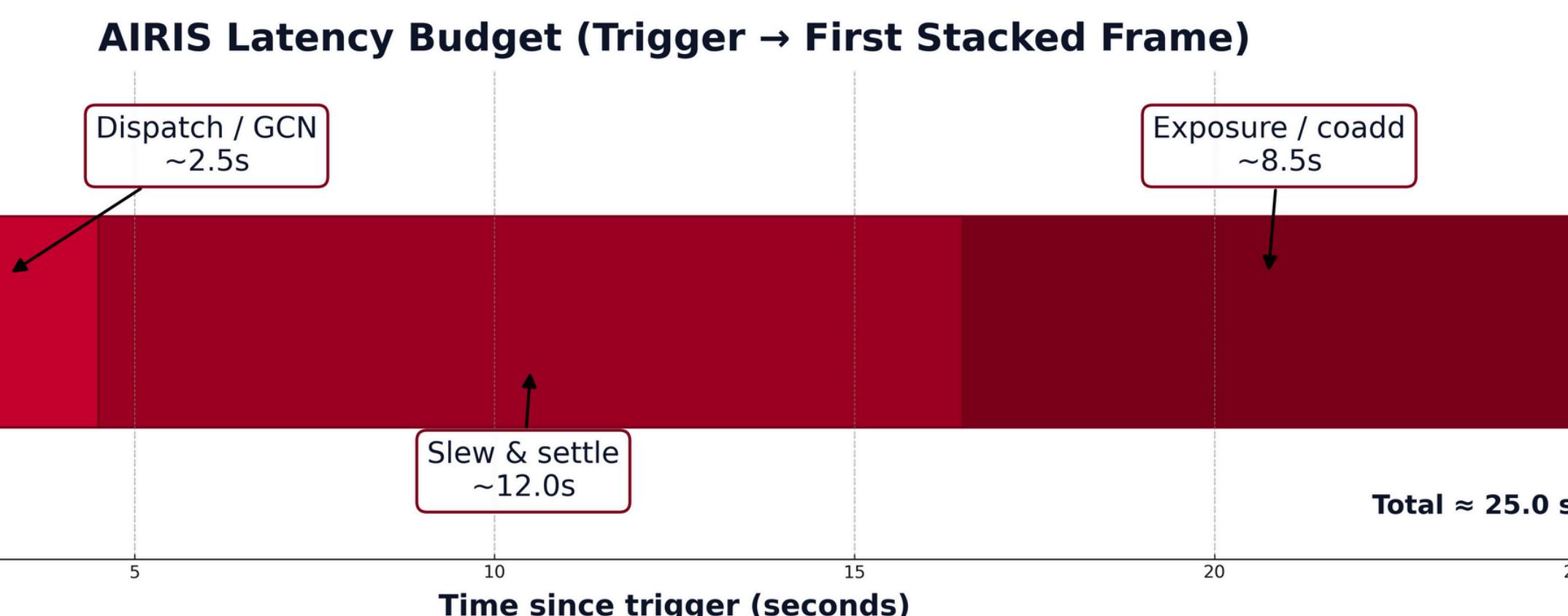
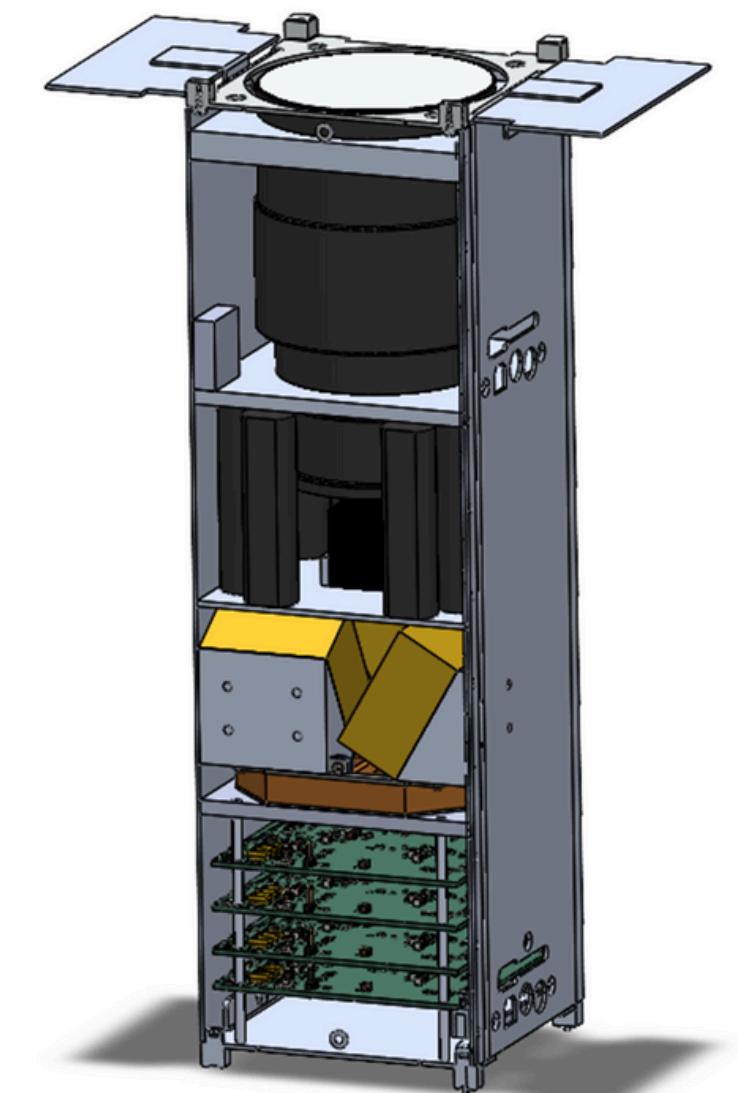


Fig. AIRIS latency budget showing  $\sim 25$  s from trigger to first stacked frame.

## Future Work

- **Algorithm development:** Advance tiling, detection, and ranking for multi-messenger triggers; optimize PSF-matched, variability-aware search.
- **Control system:**  $\leq 30^\circ/\text{s}$  slews,  $< 2$  s settle,  $\leq 1''$  pointing knowledge; vibration isolation so AIRIS does not disturb ADAPT/gondola.
- **Search algorithm:** Implement and validate an afterglow localization pipeline (HEALPix tiling + candidate scoring + false-positive controls).
- **Software pipeline:** On-Jetson deblur  $\rightarrow$  register  $\rightarrow$  detect  $\rightarrow$  alert; evaluate motion-blur removal and search strategies against injections.
- **Technology demonstration:** De-risk hardware/software to pave the way for WashU's VECTOR CubeSat proposal.



VECTOR CAD Model

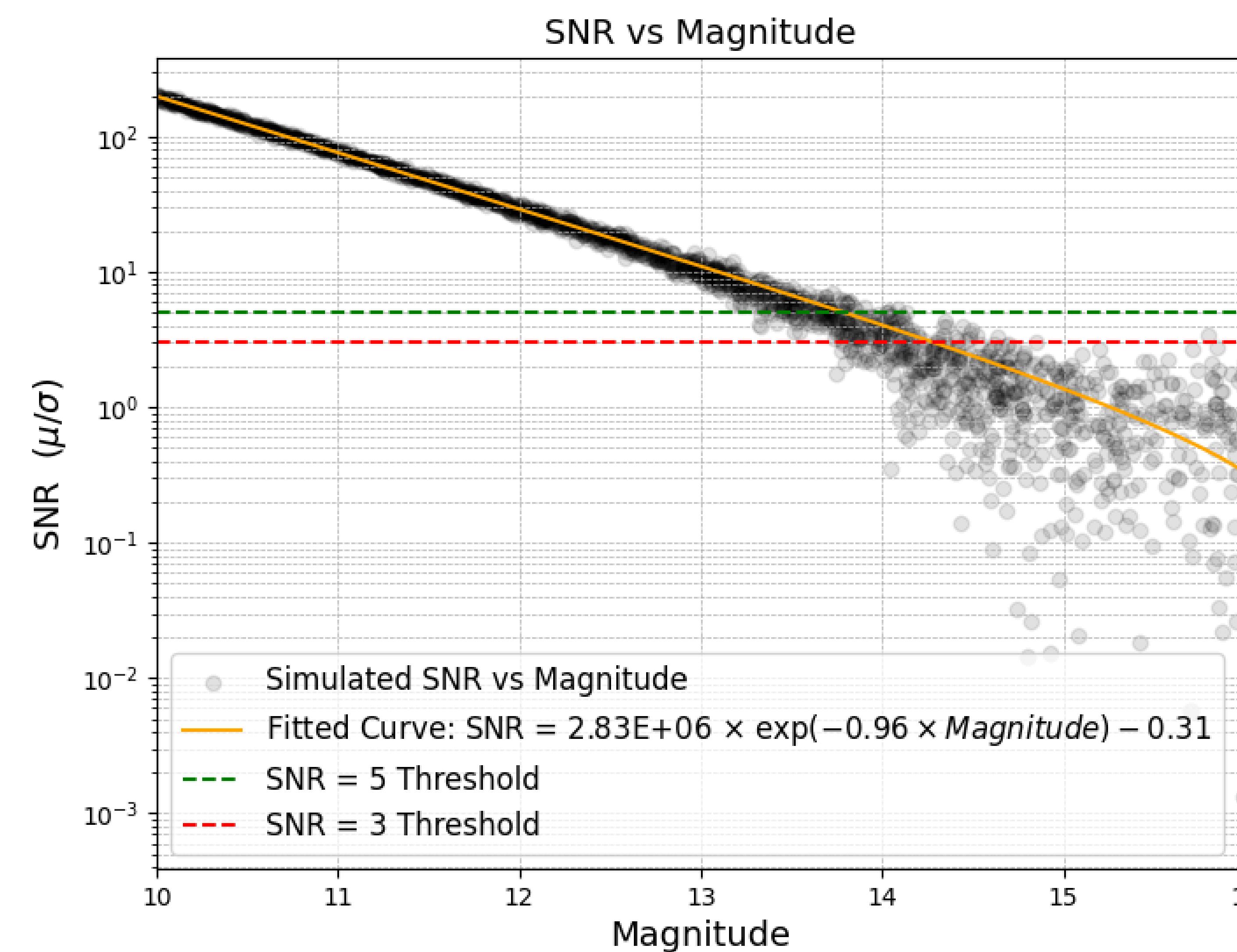
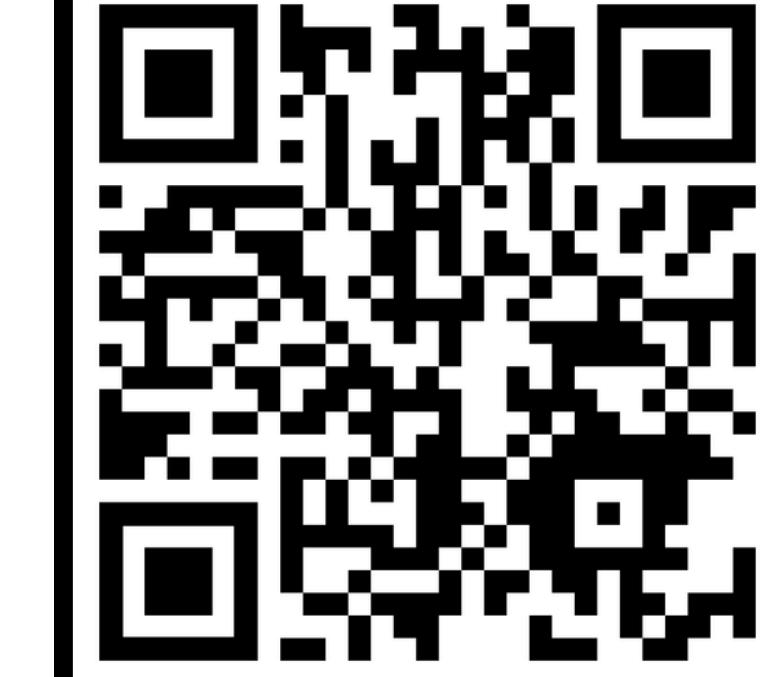


Fig. Simulated SNR vs. magnitude with best-fit curve and detection thresholds at  $\text{SNR} = 3$  and 5.

## Collaborators

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## Get in Touch



## References

