Cloud Imaging and Particle Size (CIPS) Instrument Overview

Most recent update: 2 June 2017

The NASA Aeronomy of Ice in the Mesosphere (AIM) mission is designed to study the highest clouds in the earth's atmosphere – clouds at the edge of space (http://aim.hamptonu.edu/). Formally known as noctilucent clouds (NLCs) or polar mesospheric clouds (PMCs), these clouds are located near the polar mesopause in both hemispheres, about 82 km (50 miles) above the surface of the earth. They are composed of water ice crystals, with particle sizes less than about 100 nanometers. They form in the summertime, when ironically the mesopause region is colder than at any other time. Changes observed in PMCs in recent decades are possibly related to anthropogenic effects on the atmosphere. The AIM mission is measuring the clouds and their environment to better understand the underlying chemistry and physics, and the forces that cause them to vary.

The AIM satellite has three instruments that provide information about PMCs and their environment: CIPS, CDE (Cosmic Dust Experiment), and SOFIE (Solar Occultation For Ice Experiment). CDE is an in-situ dust detector that measures cosmic dust input, which is a potential key factor in PMC formation [Poppe et al., 2011]. SOFIE measures vertical profiles of particle extinction as well as temperature, water vapor, ozone, nitric oxide and methane [Gordley et al., 2009]. For more information on CDE and SOFIE, please visit http://aim.hamptonu.edu/instrmt/cde.html and http://sofie.gats-inc.com/sofie/index.php.

The CIPS instrument [McClintock *et al.*, 2009] is a panoramic imager that measures ultraviolet radiation scattered by the clouds and atmosphere. It views in the nadir and off-nadir direction and images the polar atmosphere at a variety of scattering angles in order to determine cloud presence, provide the spatial morphology of the cloud and constrain the microphysical parameters of the cloud particle distribution. The instrument consists of a 2×2 array of cameras operating in a 10 nm passband centered at 265 nm, each with an overlapping field of view (FOV). The total FOV is 80 deg \times 120 deg (960 km \times 1140 km), centered at the sub-satellite point, with the 120 deg axis along the orbit track. Because of slant viewing at the edges of the FOV, the average effective spatial resolution varies from 2 km \times 2.4 km (along-track \times crosstrack) for the nadir cameras and 4.5 km \times 3 km for the forward- and backward-viewing cameras.

CIPS PMC data products include measurements of cloud parameters (albedo, particle radius, ice water content) along each orbit strip, with 25 km² spatial resolution. CIPS images are acquired simultaneously in each camera every 43 seconds in the summer hemisphere between the terminator and a dayside latitude of about 40 degrees. With this cadence, many locations are viewed at as many as seven different scattering angles, providing a direct measurement of the cloud scattering phase function (cloud albedo vs. scattering angle), from which the particle radius is derived. For users interested in a more in-depth and independent look at particle characteristics, these measured phase functions at each location are also available. The near-polar orbit causes the observation swaths to overlap at latitudes higher than about 70 degrees, so that nearly the entire polar cap is mapped with 15-orbit per day coverage. The CIPS retrieval algorithm and error analysis are described in Lumpe *et al.* [2013].

Update on 2 June 2017: In February of 2016, CIPS entered a new mode of operations, referred to as "continuous imaging" (CI) mode. In CI mode, instead of acquiring images every 43 seconds over the summer pole, CIPS acquires images every 3 minutes throughout the orbit. Albedo measurements are thus made globally throughout the sunlit portion of the orbit. This results in less overlap of images in the PMC region, which compromised the measurements of cloud particle phase function and thus particle size. On the other hand, the addition of images in the winter hemisphere enables global investigations of gravity waves (GW), so the CIPS team began producing a "Rayleigh Albedo Anomaly" (RAA) data product, from which GW information near an altitude of 50-55 km can be inferred [Randall et al., 2017]. Users will find information about this data product, as well as browse images and data files, on other pages on this web site.

References:

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Last updated 20170602.