

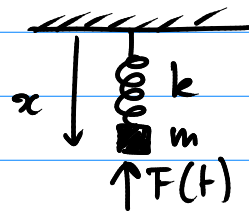
Classical Mechanics (Phys 601) - November 17, 2011

* Final exam:

Wednesday December 7, 2011, between 9 am and 12 pm
in Small Hall 122

* Non-linear systems \rightarrow non-linear equations of motion Mass on a spring:

$$L = \frac{1}{2} m \dot{q}^2 - \frac{1}{2} m \omega^2 q^2$$



$$\hookrightarrow \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = F(t) \Leftrightarrow m \ddot{q} + m \omega^2 q = F(t)$$

Linear system: only terms linear in derivatives of $q(t)$ in the differential equations of motion.

No driving force:

$$m \ddot{q} + m \omega^2 q = 0 \rightarrow \text{homogeneous, linear}$$

If $q_1(t)$ and $q_2(t)$ are solutions (with different initial conditions), then $a q_1(t) + b q_2(t)$ is also a solution.

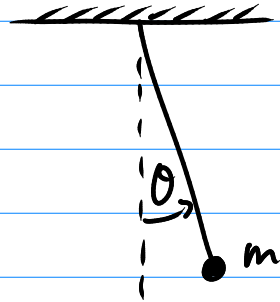
Driving force:

$$m \ddot{q} + m \omega^2 q = F(t) \rightarrow \text{non-homogeneous, still linear}$$

Pendulum:

$$L = \frac{1}{2} m l^2 \dot{\theta}^2 - mgl \cos \theta$$

$$\Rightarrow \ddot{\theta} = -\frac{g}{l} \sin \theta \quad \hookrightarrow \text{non-linear system}$$

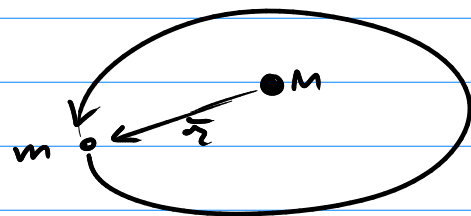


Linearize for θ small: $\sin \theta \approx \theta \Rightarrow \ddot{\theta} = -\frac{g}{l} \theta$

Orbital motion:

$$L = \frac{1}{2} m \vec{v}^2 - GMm \frac{1}{r}$$

$$\Rightarrow m \ddot{\vec{r}} = -GM \frac{\hat{r}}{r^2} \quad \hookrightarrow \text{non-linear system}$$



Linear equations of motion \rightarrow typically solvable

Non-linear equations of motion \rightarrow typically not solvable

\hookrightarrow some exceptions: orbital motion for two-body system

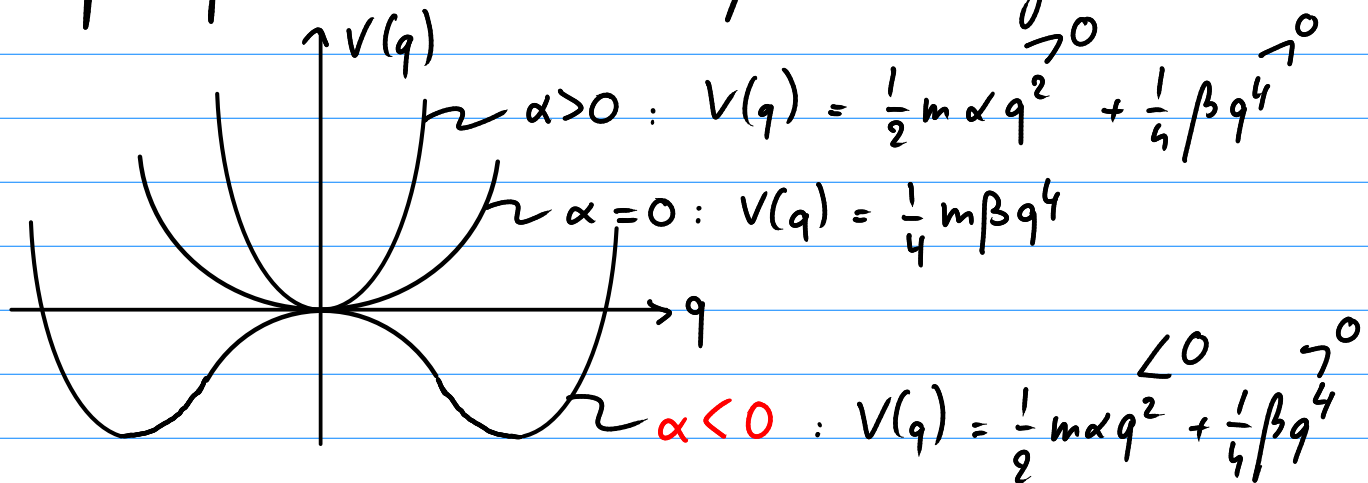
* Duffing oscillator: general form

Consider the following modification to the harmonic oscillator:

$$V(q) = \frac{1}{2} m \alpha q^2 + \frac{1}{4} m \beta q^4 \quad \leftarrow \text{quartic term}$$

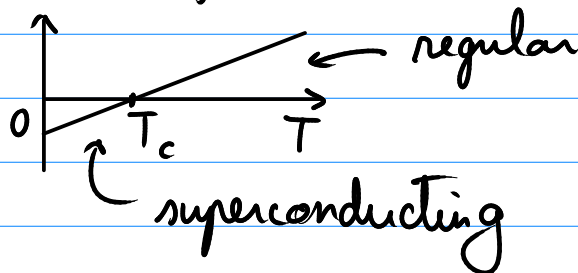
At $|q| \rightarrow \infty$ physical reasons require $V(q)$ to be positive.
 $\Rightarrow \beta$ positive

In principle α can be either positive or negative.



Applications:

- superconductivity phase transitions: $d(T)$



- Higgs mechanism for electroweak symmetry breaking
 - \hookrightarrow minimum energy state is not at $q = 0$
 - \hookrightarrow "hat potential" $V(\varphi)$ where φ is complex

$V(q)$ is stationary when $\frac{\partial V}{\partial q} = 0 \Leftrightarrow m\alpha q + m\beta q^3 = 0$

$$\Leftrightarrow q=0 \quad \text{or} \quad q^2 = -\frac{\alpha}{\beta} \quad \begin{cases} 0, & \text{for } \alpha=0 \\ <0, & \text{for } \alpha>0 \\ >0, & \text{for } \alpha<0 \end{cases}$$

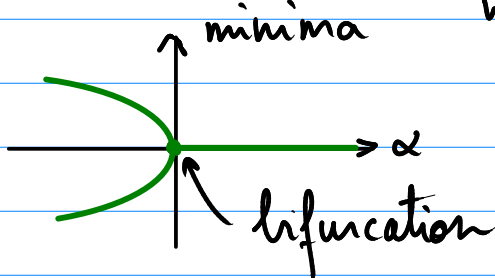
$\alpha > 0$: only one real root as absolute minimum

$\alpha = 0$: $q = 0$ is a triple root as absolute minimum

$$d < 0 : \quad q = 0 \text{ and } q = \pm \sqrt{-\frac{\alpha}{\beta}} = \pm \sqrt{\frac{|d|}{\beta}} \text{ are roots}$$

\uparrow \uparrow
 local maximum absolute minima

⇒ continuous evolution of α leads to **bifurcation** of minima
with energy barrier between them:

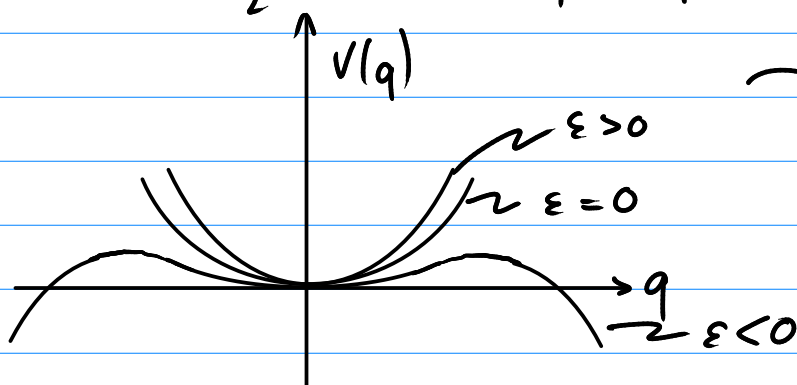


$$V(q^{\pm}) = \frac{1}{2} m \alpha \left(-\frac{\alpha}{\beta} \right) + \frac{1}{4} m \beta \left(-\frac{\alpha}{\beta} \right)^2$$

$$= -\frac{1}{4} m \frac{\alpha^2}{\beta}$$

* Duffing oscillator as perturbation of harmonic oscillator:

$$V(q) = \frac{1}{2} m \omega_0^2 q^2 + \frac{1}{4} \epsilon m q^4 \quad \text{with } \epsilon \text{ small (positive or negative)}$$



→ we will focus on small region around $q = 0$
→ no problems with $q \rightarrow \infty$

$$L = \frac{1}{2} m \dot{q}^2 - \frac{1}{2} m \omega_0^2 q^2 - \frac{1}{4} \varepsilon m q^4$$

$$\Rightarrow \ddot{q} + \omega_0^2 q + \varepsilon q^3 = 0 \rightarrow \text{non-linear!}$$

$$\Rightarrow p = \frac{\partial L}{\partial \dot{q}} = m \dot{q} \Rightarrow H = \frac{p^2}{2m} + \frac{1}{2} m \omega_0^2 q^2 + \frac{1}{4} \varepsilon m q^4$$

Perturbation theory: assume solution of form
 $q(t) = q_0(t) + \varepsilon q_1(t) + \dots$

$$\begin{cases} \text{1st order: } \ddot{q}_0 + \omega_0^2 q_0 = 0 \\ \text{2nd order: } \ddot{q}_1 + \omega_0^2 q_1 + q_0^3 = 0 \\ \vdots \end{cases} \quad \text{with} \quad \begin{cases} q_0(0) = a \\ \dot{q}_0(0) = 0 \\ \vdots \\ q_1(0) = 0 \\ \dot{q}_1(0) = 0 \end{cases}$$

$$\Leftrightarrow \begin{cases} q_0(t) = a \cos \omega_0 t \\ \ddot{q}_1 + \omega_0^2 q_1 = -a^3 \cos^3 \omega_0 t \\ \vdots \end{cases}$$

$$(e^{i\varphi})^3 = \cos 3\varphi + i \sin 3\varphi = \cos^3 \varphi + 3i \cos^2 \varphi \sin \varphi - 3 \cos \varphi \sin^2 \varphi - i \sin^3 \varphi$$

$$\begin{aligned} \text{Re} \quad \Leftrightarrow \cos 3\varphi &= \cos^3 \varphi - 3 \cos \varphi (1 - \cos^2 \varphi) \\ &= -3 \cos \varphi + 4 \cos^3 \varphi \end{aligned}$$

$$\Leftrightarrow \cos^3 \varphi = \frac{1}{4} (\cos 3\varphi + 3 \cos \varphi)$$

$$\Leftrightarrow \begin{cases} q_0(t) = a \cos \omega_0 t \\ \ddot{q}_1 + \omega_0^2 q_1 = \underbrace{-\frac{3}{4} a^3 \cos \omega_0 t}_{\text{driving term}} - \underbrace{\frac{1}{4} a^3 \cos 3\omega_0 t}_{\text{driving term}} \\ \vdots \end{cases}$$

Solution of second order differential equation:

$$a_2 y'' + a_1 y' + a_0 y = e^{ikx}$$

↳ characteristic equation \rightarrow poles k_1 and k_2
 $a_2 k^2 + a_1 k + a_0 = 0$

1) if $k \neq k_1, k_2 \rightarrow y = A e^{ik_1 x} + B e^{ik_2 x}$

2) if $k = k_1 \neq k_2 \rightarrow y = A t e^{ik_1 x} + B e^{ik_2 x} \rightarrow$

3) if $k = k_1 = k_2 \rightarrow y = (A t^2 + B t) e^{ik_1 x} \rightarrow$ *resonance*

↳ substitute to determine A and B of particular solution

$$\Rightarrow q_1(t) = -\frac{a^3}{8\omega_0^2} \left(3\omega_0 t \sin \omega_0 t + \frac{1}{4} (\cos \omega_0 t - \cos 3\omega_0 t) \right)$$

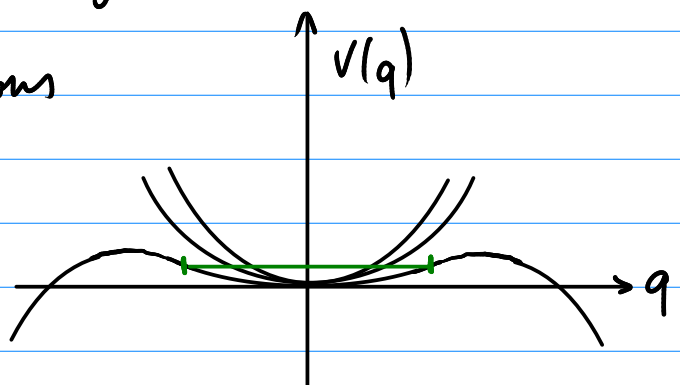
$$\Rightarrow q(t) = a \cos \omega_0 t - \frac{\epsilon a^3}{8\omega_0^2} \left(3\omega_0 t \sin \omega_0 t + \frac{1}{4} (\cos \omega_0 t - \cos 3\omega_0 t) \right)$$

↓

$$+ O(\epsilon^2)$$

This solution is impossible because it indicates that $q(t)$ will increase linearly.

However, energy considerations limit the size of $q(t)$ (if $E < \text{barrier height}$)



Need to be more careful: $\omega(\varepsilon) = \omega_0 + \varepsilon\omega_1 + \dots$

$$\Rightarrow q(t) = a(\varepsilon) \cos(\omega(\varepsilon)t) \quad a(\varepsilon) = a + \varepsilon a_1 + \dots$$

$$= a(\varepsilon) \cos(\omega_0 t + \varepsilon\omega_1 t + \dots) \quad \text{coefficient } \frac{\partial}{\partial \varepsilon} \Big|_{\varepsilon=0}$$

$$= a(\varepsilon) (\cos \omega_0 t - \varepsilon\omega_1 t \sin \omega_0 t + \dots)$$

With this $\omega = \omega_0 + \varepsilon\omega_1 + \dots$ we can write:

$$\ddot{q} + \omega_0^2 q + \varepsilon q^3 = 0 \quad \tau = \omega t$$

$$\Leftrightarrow \omega^2 \frac{d^2 q}{d\tau^2} + \omega_0^2 q + \varepsilon q^3 = 0$$

$$\Leftrightarrow (\omega_0 + \varepsilon\omega_1 + \dots)^2 \frac{d^2 q}{d\tau^2} + \omega_0^2 q + \varepsilon q^3 = 0$$

$$\Leftrightarrow \left(\omega_0^2 \frac{d^2 q}{d\tau^2} + \omega_0^2 q \right) + \left(2\varepsilon\omega_0\omega_1 \frac{d^2 q}{d\tau^2} + \varepsilon q^3 \right) + \dots = 0$$

$$\Leftrightarrow \left(\omega_0^2 \frac{d^2 q_0}{d\tau^2} + \omega_0^2 q_0 \right) + \left(\omega_0^2 \varepsilon \frac{d^2 q_1}{d\tau^2} + \omega_0^2 \varepsilon q_1 \right) + \left(2\varepsilon\omega_0\omega_1 \frac{d^2 q_0}{d\tau^2} + \varepsilon q_0^3 \right) + \dots = 0$$

$$\Leftrightarrow \begin{cases} \omega_0^2 \left(\frac{d^2 q_0}{d\tau^2} + q_0 \right) = 0 \\ \omega_0^2 \left(\frac{d^2 q_1}{d\tau^2} + q_1 \right) = -2\omega_0\omega_1 \frac{d^2 q_0}{d\tau^2} - q_0^3 \\ \vdots \end{cases}$$

first order equation \rightarrow harmonic oscillator
 $\Rightarrow q_0(\tau) = a \cos \tau$

second order equation after substituting $q_0(\tau)$

$$\omega_0^2 \left(\frac{d^2 q_1}{d\tau^2} + q_1 \right) = \underbrace{\left(2\omega_0 \omega_1 - \frac{3}{4} a^2 \right)}_{\text{additional term}} a \cos \tau - \frac{1}{4} a^3 \cos 3\tau$$

resonant driving term

Remove resonant driving term by choosing $\omega_1 = \frac{3}{8} \frac{a^2}{\omega_0}$

$$\hookrightarrow \omega_0^2 \left(\frac{d^2 q_1}{d\tau^2} + q_1 \right) = -\frac{1}{4} a^3 \cos 3\tau$$

$$\Rightarrow q_1(\tau) = -\frac{a^3}{32\omega_0^2} (\cos \tau - \cos 3\tau) \quad \tau \simeq \omega_0 t$$

$$\text{or } q_1(t) \simeq -\frac{a^3}{32\omega_0^2} (\cos \omega_0 t - \cos 3\omega_0 t) + O(\varepsilon)$$

Full solution:

$$q(t) = a \cos \left(\left(\omega_0 + \varepsilon \frac{3a^2}{8\omega_0} \right) t \right) - \frac{\varepsilon a^3}{32\omega_0^2} (\cos \omega_0 t - \cos 3\omega_0 t)$$

\rightarrow no term blows up anymore, but expand this in ε :

$$q(t) = a \cos \omega_0 t - \varepsilon \frac{3a^3}{8\omega_0} t \sin(\omega_0 t) - \frac{\varepsilon a^3}{32\omega_0^2} (\cos \omega_0 t - \cos 3\omega_0 t) + O(\varepsilon^2)$$

⇒ by changing $\omega_0 \rightarrow \omega_0 + \varepsilon\omega$, for a specific ω ,
we could avoid infinities in our perturbation theory
calculation

↳ concept behind parts of quantum field theory
and renormalization !
and running couplings •

+ Perturbations of periodic Hamiltonian systems

If we have a Hamiltonian of the form

$$H(p, q, t) = H_0(p, q, t) + \varepsilon V(p, q, t)$$

Can we use the methods we developed for linear systems on non-linear perturbations?

Review:

Canonical transformation:

$$\begin{cases} H_0 \\ p, q \end{cases} \xrightarrow{S_0} \begin{cases} \tilde{H}_0 \\ P, Q \end{cases} \quad \text{with } S_0(q, P, t)$$

$$\begin{cases} \tilde{H}_0(P, Q, t) = H_0(p, q, t) + \frac{\partial S_0}{\partial t}(q, P, t) \\ p = \frac{\partial S_0}{\partial q}(q, P, t) \\ Q = \frac{\partial S_0}{\partial P}(q, P, t) \end{cases}$$

Canonical transformation preserves Hamilton's equations:

$$\begin{cases} \dot{Q} = \frac{\partial \tilde{H}_0}{\partial P}(P, Q, t) \\ \dot{P} = -\frac{\partial \tilde{H}_0}{\partial Q}(P, Q, t) \end{cases}$$

Hamilton-Jacobi equation:

Now transform to P, Q such that $\tilde{H}_0(P, Q, t) \equiv 0$
 $\Rightarrow P$ and Q are constants of motion

$$\text{Condition on } S_0: H_0(p, q, t) + \frac{\partial S_0}{\partial t}(q, P, t) = 0$$

with $p = \frac{\partial S_0}{\partial q}$

$$\Rightarrow H_0\left(\frac{\partial S_0}{\partial q}, q, t\right) + \frac{\partial S_0}{\partial t}(q, P, t) = 0$$

$$\text{If } \frac{\partial H}{\partial t} = 0 \rightarrow S_0(q, P, t) = W_0(q, P) - Et$$

(note: we called $P_i = \alpha_i$ earlier, now we will just use P)

For one-dimensional problems: $S_0(q, E, t) = W(q, E) - Et$

$$\Rightarrow H_0\left(\frac{\partial S_0}{\partial q}, q, t\right) = E$$

Action-angle variables for periodic systems:

Define action variable (now with 2π)

$$J = \frac{1}{2\pi} \int_{\text{cycle}} p dq = \frac{1}{2\pi} \int_{\text{cycle}} \frac{\partial W_0}{\partial q}(q, E) dq = J(E)$$

\uparrow
 $p = \frac{\partial S_0}{\partial q} = \frac{\partial W_0}{\partial q}$

If we assume that $J(E)$ can be inverted, then

$$\bar{S}_0(q, J, t) = S_0(q, E(J), t)$$

describes a canonical transformation to p and

$$\bar{Q} = \frac{\partial \bar{S}_0}{\partial J}(q, J, t) = \underbrace{\frac{\partial \bar{W}_0}{\partial J}(q, J)}_{\text{angle variable } \varphi} - \underbrace{\frac{\partial E_0(J)}{\partial J}}_{\text{frequency } \omega_0} t \equiv \bar{\beta}$$

$$\Rightarrow \varphi = \omega_0 t + \bar{\beta}$$

\Rightarrow action variable is constant
angle variable increases linearly with time

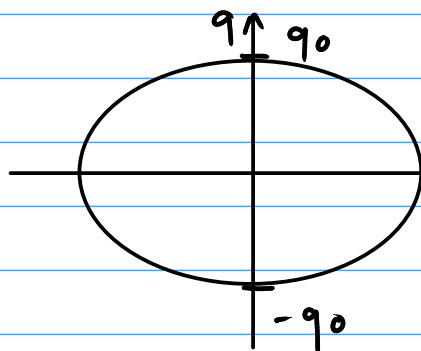
For perturbative system this will not be true anymore!

Harmonic oscillator:

$$H_0 = E_0 = \frac{p^2}{2m} + \frac{1}{2} m \omega_0^2 q^2 \Rightarrow p = \pm \sqrt{2m} \sqrt{E_0 - \frac{1}{2} m \omega_0^2 q^2}$$

$$\Rightarrow J = \frac{1}{2\pi} \int_{\text{cycle}} p dq = \frac{4\sqrt{2m}}{2\pi} \int_{\text{quadrant}} \sqrt{\frac{1}{2} m \omega_0^2 (q_0^2 - q^2)}$$

$$= \frac{2q_0 \sqrt{2mE_0}}{\pi} \int_0^1 dx \sqrt{1-x^2}$$



$$= \frac{E_0}{\omega_0} \Rightarrow E_0(J) = \omega_0 J$$

$$\Rightarrow \omega = \frac{\partial E_0}{\partial J} = \omega_0$$

$$\begin{aligned}
H_0\left(\frac{\partial W}{\partial q}, q, t\right) &= E_0 \\
\Rightarrow W(q, E) &= \sqrt{2m} \int dq \sqrt{E_0 - \frac{1}{2} m \omega_0^2 q^2} + W_0(E_0) \\
\Rightarrow \bar{W}(q, E_0(J)) &= \sqrt{2m} \int dq \sqrt{\omega_0 J - \frac{1}{2} m \omega_0^2 q^2} + \bar{W}_0(J) \\
\Rightarrow \varphi = \frac{\partial \bar{W}}{\partial J}(q, J) &= \frac{\omega_0}{2} \sqrt{2m} \int dq \frac{1}{\sqrt{\omega_0 J - \frac{1}{2} m \omega_0^2 q^2}} + \varphi_0(J) \\
&= \sin^{-1} \frac{q}{q_0}
\end{aligned}$$

$$\Rightarrow q = \sqrt{\frac{2J}{m\omega_0}} \sin \varphi \quad \text{and} \quad p = \sqrt{2m\omega_0 J} \cos \varphi, \quad \varphi = \omega_0 t$$

Perturbation to H_0

$$\begin{aligned}
H(J, \varphi) &= H_0(J, \varphi) + \varepsilon H_1(J, \varphi) \\
&= E_0(J, \varphi) + \varepsilon H_1(J, \varphi) \\
&= \omega_0 J + \varepsilon \frac{1}{4} m \left(\frac{2J}{m\omega_0}\right)^2 \sin^4 \varphi
\end{aligned}$$

\Rightarrow Hamilton's equations:

$$\begin{cases} \dot{J} = -\frac{\partial H}{\partial \varphi} = -\varepsilon m \left(\frac{2J}{m\omega_0}\right)^2 \sin^3 \varphi \cos \varphi \\ \dot{\varphi} = \frac{\partial H}{\partial J} = \omega_0 + \varepsilon \frac{2J}{m\omega_0^2} \sin^4 \varphi \end{cases}$$