Supplemental material

 ${\it Introduction.}$ — We present below some cumbersome formulations, to be read in conjunction with the text.

The minisuperspace Lagrangian. – Consider the ADM-like extension $ds^2 = u^2 (dt^2 - v^2 d\mathbf{x}^2)$, where the limit $u \mapsto 1$ and $v \mapsto a$ recovers the flat FRW metric. The analogue cleary corresponds to the following gauge

$$b^{a}_{\mu} = u \left(v \left(\delta^{a}_{\mu} - \delta^{a}_{0} \eta_{0\mu} \right) + \delta^{a}_{0} \eta_{0\mu} \right), \quad A^{ab}_{\mu} = u v \delta^{d}_{0} \left(2 \phi \delta^{[b}_{\mu} \delta^{a]}_{d} - \frac{1}{2} \psi \varepsilon_{\mu d}^{ab} \right). \tag{S1}$$

and substitute these into (??) to obtain the minisuperspace Lagrangian of PGT^{q,+}. Once the Gauss-Bonnet term is discarded as a total time derivative, the coupling constants in (??) may be used to express this as

$$\begin{split} \bar{L}_{\mathrm{T}} &= (-3\sigma_{2}u^{2}v\psi^{2} + 24\sigma_{3}u^{2}v\phi^{2} + 3\sigma_{3}u^{2}v\psi^{2} + 3m_{\mathrm{P}}^{2}v_{2}u^{2}v)(\partial_{t}v)^{2} + (-6\sigma_{2}uv^{2}\psi^{2} + 48\sigma_{3}uv^{2}\phi^{2} + 6\sigma_{3}uv^{2}\psi^{2} + 6\sigma_{3}uv^{2}\psi^{2} + 6\sigma_{3}u^{2}v^{2}\psi)\partial_{t}v\partial_{t}\psi + (-3\sigma_{2}v^{3}\psi^{2} + 24\sigma_{3}v^{3}\phi^{2} + 3\sigma_{3}v^{3}\psi^{2} + 3m_{\mathrm{P}}^{2}v_{2}v^{3})(\partial_{t}u)^{2} + 48v^{3}\sigma_{3}u\phi\partial_{t}u\partial_{t}\phi + (-6\sigma_{2}uv^{3}\psi + 6\sigma_{3}uv^{3}\psi)\partial_{t}u\partial_{t}\psi + 24v^{3}\sigma_{3}u^{2}(\partial_{t}\phi)^{2} + (-3\sigma_{2}u^{2}v^{3} + 3\sigma_{3}u^{2}v^{3})(\partial_{t}\psi)^{2} + (-48\sigma_{1}u^{3}v^{3}\phi^{2} + 12\sigma_{1}u^{3}v^{3}\psi^{2} - 3m_{\mathrm{P}}^{2}\alpha_{0}u^{3}v^{3})\partial_{t}\phi + (-48\sigma_{1}u^{3}v^{2}\phi^{3} + 12\sigma_{1}u^{3}v^{2}\phi\psi^{2} - 3m_{\mathrm{P}}^{2}\alpha_{0}u^{3}v^{2}\phi + 6m_{\mathrm{P}}^{2}v_{2}u^{3}v^{2}\phi)\partial_{t}v + (-48\sigma_{1}u^{2}v^{3}\phi^{3} + 12\sigma_{1}u^{2}v^{3}\phi\psi^{2} - 3m_{\mathrm{P}}^{2}\alpha_{0}u^{2}v^{3}\phi + 6m_{\mathrm{P}}^{2}v_{2}u^{2}v^{3}\phi)\partial_{t}u + 3m_{\mathrm{P}}^{2}v_{2}v^{3}u^{4}\phi^{2} + 3v^{3}m_{\mathrm{P}}^{2}\alpha_{0}u^{4}\phi^{2} + 24v^{3}\sigma_{3}u^{4}\phi^{4} + 3v^{3}u^{4}\psi^{2}m_{\mathrm{P}}^{2}v_{1} - 3/4v^{3}m_{\mathrm{P}}^{2}\alpha_{0}u^{4}\psi^{2} + 3/2v^{3}\sigma_{3}u^{4}\psi^{4} - 12\sigma_{2}v^{3}u^{4}\phi^{2}\psi^{2} + \bar{L}_{\mathrm{m}}(\Phi, \Psi; u, v, \phi, \psi). \end{split}$$

The autonomous system. – The variables x and y encode the momentum and position of the canonical inflaton ξ . The Cuscuton ζ constrains the system, and is described by $z^2 = m_{\rm p}^2 W^4 \zeta^2 / 4H^2$. Note that z adopts the rôle of the matter parameter. We further define $\lambda = -m_{\rm p}\partial_{\xi}V/V$ and $\mu = W$. As with x and y, note that λ is a conventional parameter. The $g_{\mu\nu}$ and ξ equations are readily expressed as a coupled first order system in terms of these variables

$$H^{-1}\partial_{t}x = \left(-16\sqrt{2}\mu xy^{2} + 6\sqrt{2}\mu^{3}xy^{2} - 2\sqrt{3}\lambda\mu^{3}x^{2}y^{2} + 16\sqrt{2}\mu xy^{4} - 6\sqrt{2}\mu^{3}xy^{4} - \sqrt{2}\mu^{5}xy^{4} + 32xz^{2} + 16\sqrt{2}\mu xz^{2} - 2\sqrt{2}\mu^{3}xz^{2} - 32xy^{2}z^{2} - 32\sqrt{2}\mu xy^{2}z^{2} + 2\sqrt{2}\mu^{3}xy^{2}z^{2} + 8\mu^{4}xy^{2}z^{2} + 2\sqrt{2}\mu^{5}xy^{2}z^{2} - \sqrt{6}\lambda\mu^{4}x^{2}y^{2}z^{2} + 32xz^{4} + 16\sqrt{2}\mu xz^{4} - 4\mu^{4}xz^{4} - \sqrt{2}\mu^{5}xz^{4}\right)/\left(-2\sqrt{2}\mu^{3} + 2\sqrt{2}\mu^{3}y^{2} + \sqrt{2}\mu^{5}y^{2} - 4\mu^{4}z^{2} - \sqrt{2}\mu^{5}z^{2}\right),$$
(S3a)

$$H^{-1}\partial_{t}y = \left(16\sqrt{2}y - 6\sqrt{2}\mu^{2}y + 2\sqrt{3}\lambda\mu^{2}xy - 32\sqrt{2}y^{3} + 12\sqrt{2}\mu^{2}y^{3} + \sqrt{2}\mu^{4}y^{3} - 2\sqrt{3}\lambda\mu^{2}xy^{3} + 16\sqrt{2}y^{5} - 6\sqrt{2}\mu^{2}y^{5} - \sqrt{2}\mu^{4}y^{5} + 32\sqrt{2}yz^{2} - 2\sqrt{2}\mu^{2}yz^{2} - 12\mu^{3}yz^{2} - 3\sqrt{2}\mu^{4}yz^{2} + 2\sqrt{6}\lambda\mu^{3}xyz^{2} + \sqrt{3}\lambda\mu^{4}xyz^{2} - 32\sqrt{2}y^{3}z^{2} + 2\sqrt{2}\mu^{2}y^{3}z^{2} + 12\mu^{3}y^{3}z^{2} + 2\sqrt{2}\mu^{4}y^{3}z^{2} - \sqrt{6}\lambda\mu^{3}xy^{3}z^{2} + 16\sqrt{2}yz^{4} - 4\mu^{3}yz^{4} - \sqrt{2}\mu^{4}yz^{4}\right)/\left(-2\sqrt{2}\mu^{2} + 2\sqrt{2}\mu^{2}y^{2} + \sqrt{2}\mu^{4}y^{2} - 4\mu^{3}z^{2} - \sqrt{2}\mu^{4}z^{2}\right).$$
(S3b)

To obtain the autonomous system, we must eliminate λ , μ and z. Since W and V depend explicitly on ξ , we easily solve for λ in terms of μ

$$\lambda = -\frac{2\sqrt{2/3}(8\Lambda - 60v_1 + (\Lambda - 12v_1)\mu^2)\sqrt{16 + 10\mu^2 + \mu^4}}{(8 + \mu^2)(8\Lambda - 24v_1 + (\Lambda - 12v_1)\mu^2)},$$
 (S4)

while the ζ equation reduces to a quartic constraint in μ

$$(x^{1} - 1)\mu^{4} + 2\sqrt{2}z\mu^{3} + 2(5x^{2} - z^{2})\mu^{2} + 16x^{2} = 0.$$
 (S5)

Finally z is eliminated for x and y by the $g_{\mu\nu}$ equation, which contains the constraint $x^2 + y^2 - z^2$ and expells the physical portions of the phase space from the unit disc. The overall construction – in particular the root system of (??) – causes the autonomous system to be quite cumbersome: this is an unavoidable feature of Class 2 A* and Class 3 C*.