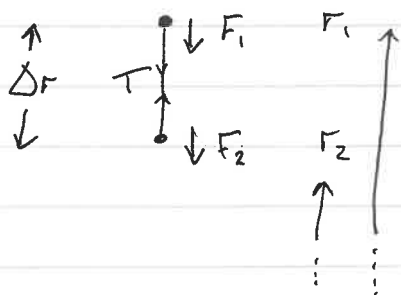


6)



$$F_1 = \frac{GMm}{r_1^2}$$

$$F_2 = \frac{GMm}{r_2^2}$$

$$r_1 = r_2 + \Delta r$$

$$F_1 + T = F_2 - T$$

$$2T = F_2 - F_1 = GMm \left(\frac{1}{r_2^2} - \frac{1}{r_1^2} \right)$$

$$T = \frac{GMm}{2} \left(\frac{r_1^2 - r_2^2}{r_1^2 r_2^2} \right)$$

$$= \frac{GMm}{2} \left(\frac{r_2^2 + 2r_2 \Delta r + \Delta r^2 - r_2^2}{r_1^2 r_2^2} \right)$$

$$= \frac{GMm \Delta r}{2} \frac{(2r_2 + \Delta r)}{r_1^2 r_2^2} = \frac{GMm \Delta r}{2} \frac{(r_1 + r_2)}{r_1^2 r_2^2}$$

At the horizon, assuming Schwarzschild:

$$r_s = \frac{2GM}{c^2}$$

Thus $T = \frac{GMm l}{2} \frac{(2r_s)}{r_s^3}$ assuming $l \ll r_s$

$$= \frac{GMm l}{(2GM/c^2)^3} = \frac{mc^6 l}{8GM^2}$$

For stellar mass, this gives $F \sim 5 \times 10^{11} \text{ N}$
supermassive $F \sim 5 \times 10^{-5} \text{ N}$

↑
Note that this scales as M^{-2}
so could be several
orders of mag. different.