

$$8) \quad \frac{1}{2} M \dot{r}^2 + \frac{1}{2} M (r \dot{\phi})^2 \left(1 - \frac{2GM}{c^2 r} \right) - \frac{GM M}{r} = \frac{1}{2} M c^2 (k^2 - 1) \quad \dot{\phi} = 0$$

and $\dot{r}(\infty) = 0 \Rightarrow k = \pm 1$

$$\Rightarrow \frac{1}{2} M \dot{r}^2 = \frac{GM M}{r}$$

$$\dot{r} = \pm \left(\frac{2GM}{r} \right)^{1/2} \quad (\text{must be zero at } \infty, \text{ negative}).$$

$$\frac{1}{(2GM)^{1/2}} \int_{r_s}^{\infty} dr \, r^{1/2} = - \int_0^{\tau} d\tau'$$

$$\frac{1}{(2GM)^{1/2}} \left[\frac{2}{3} r^{3/2} \right]_{r_s}^{\infty} = -\tau$$

$$\frac{1}{(c^2 r_s)^{1/2}} \cdot \frac{-2}{3} r_s^{3/2} = -\tau$$

$$c\tau = \frac{2}{3} r_s$$

For solar mass, $\tau \sim 7 \times 10^{-6} \text{ s}.$