

# **Explicit Electromagnetic Particle-In-Cell Simulation KEMPO1 & 2 :**

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# Software for Space Plasma Simulations

<http://space.rish.kyoto-u.ac.jp/software/>

# **References: TERRAPUB e-Library**

## **Advanced Methods for Space Simulations (ISSS-7)**

“One-dimensional Electromagnetic Particle Code: KEMPO1”  
Edited by H. Usui and Y. Omura (2007)

## **Computer Space Plasma Physics: Simulation Technique and Software (ISSS-4)**

“KEMPO1: Technical Guide to one-dimensional  
electromagnetic particle code”  
Edited by H. Matsumoto and Y. Omura (1993)

## **Computer Simulation of Space Plasmas (ISSS-1)**

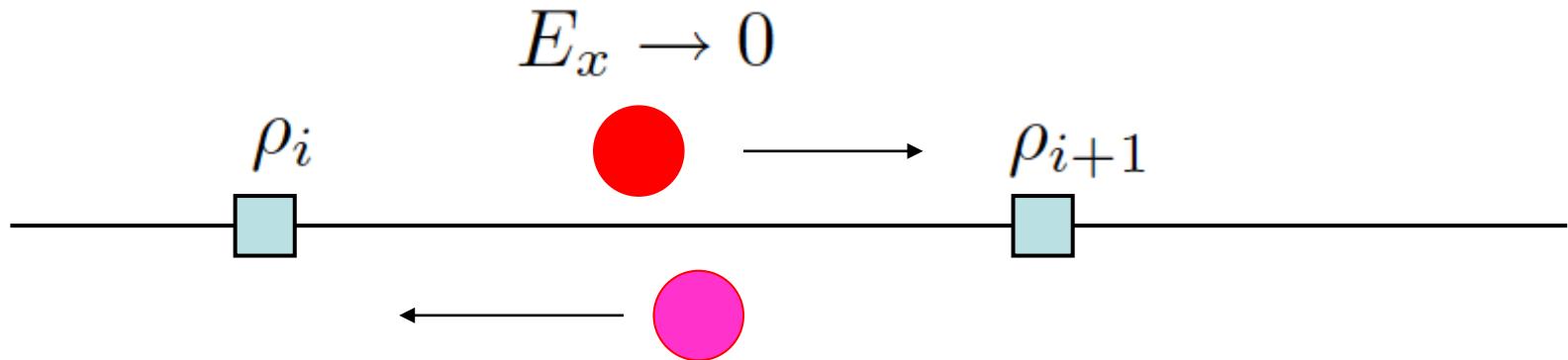
“Particle simulation of electromagnetic waves and its  
application to space plasmas”

Edited by H. Matsumoto and T. Sato (1985)

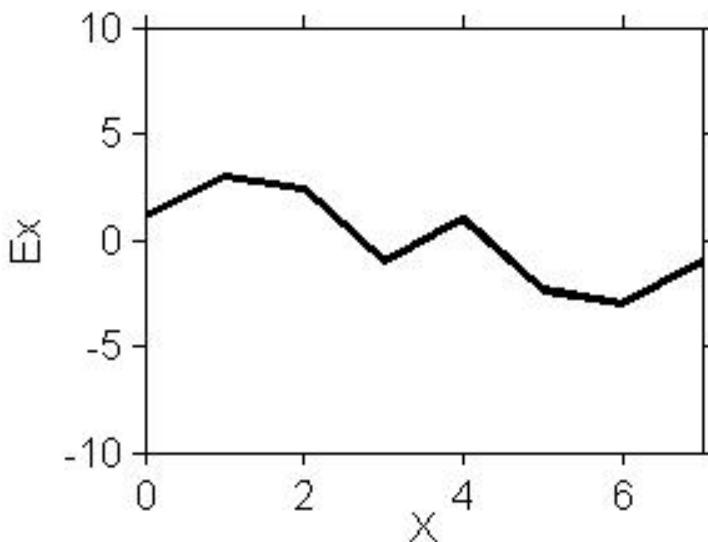
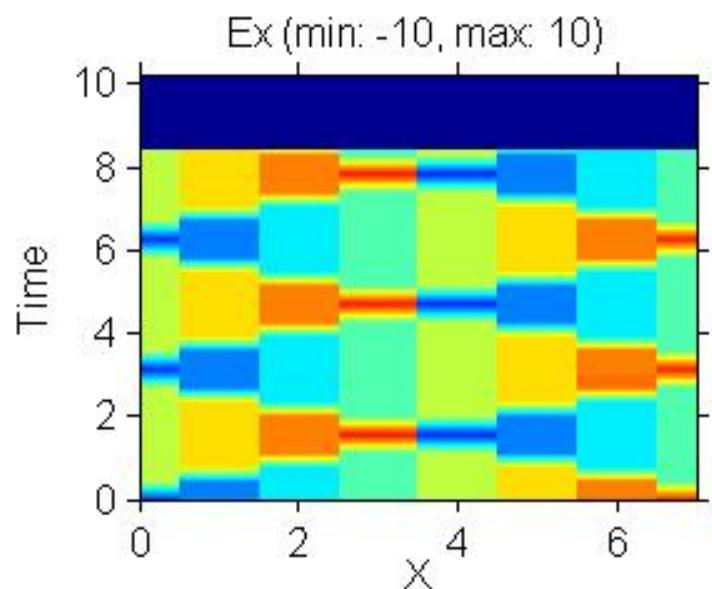
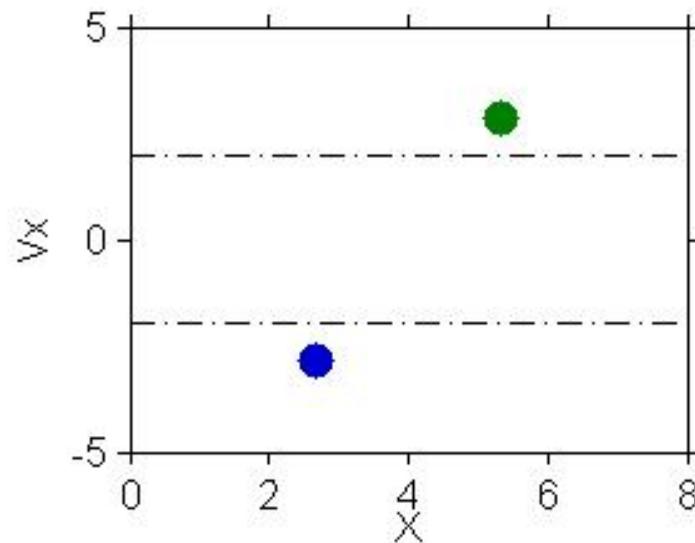
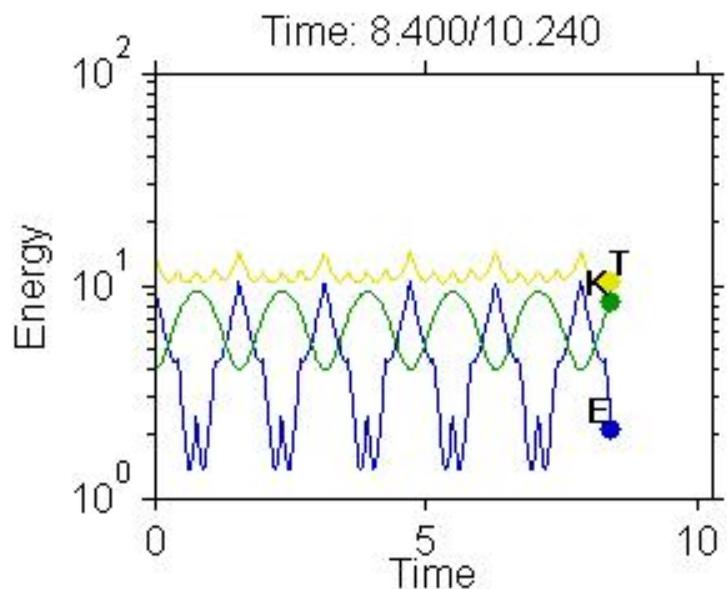
# PIC code for Space Plasmas

- Space Plasmas: **Collisionless**
- Particle-In-Cell Code
- Particles:  $x(t)$ ,  $v(t)$
- Fields:  $E(t, X)$ ,  $B(t, X)$

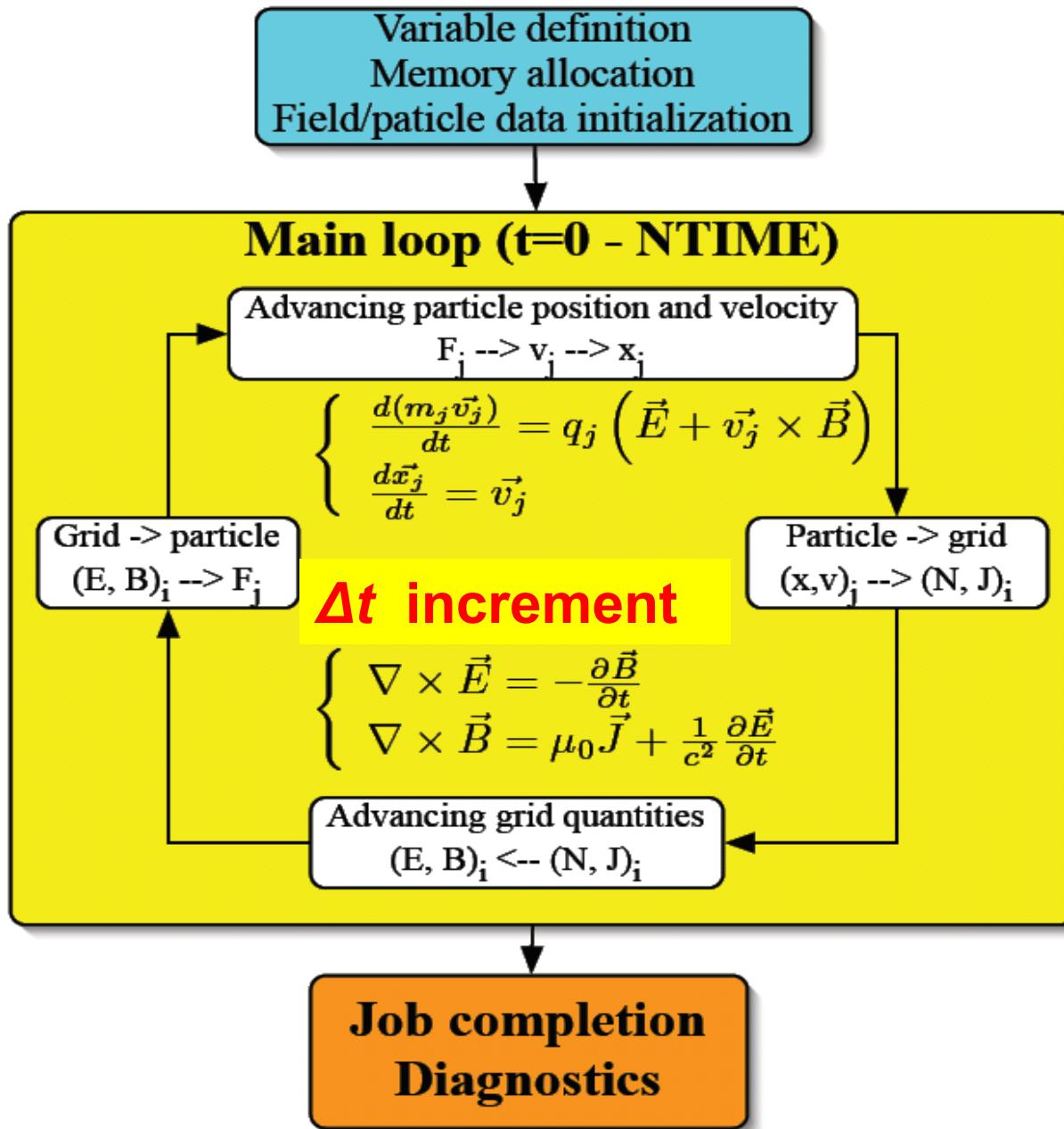
$E$  and  $B$  are defined on grid points, and calculated from  $\rho$  and  $J$ . The electrostatic force between two particles in the same cell disappears.



# KEMPO1



# Basic procedure of PIC simulation code



# Maxwell's Equations

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t}$$

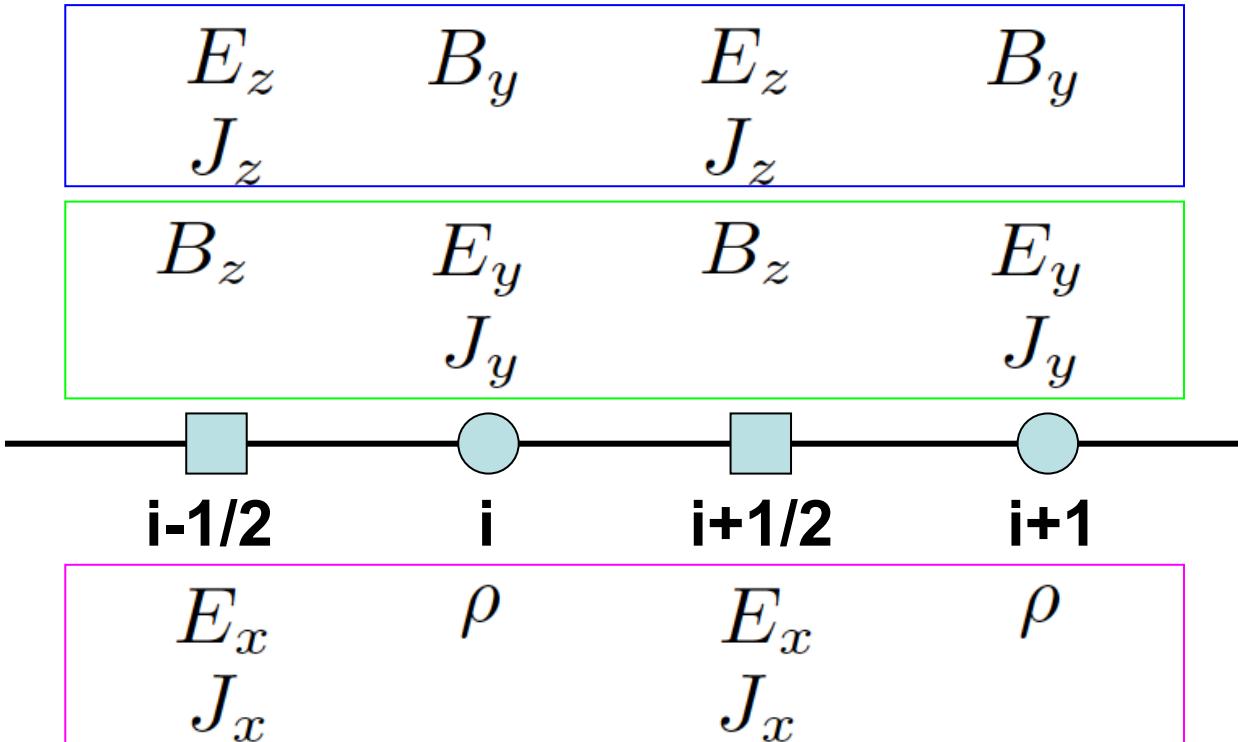
$$\nabla \times \mathbf{E} = - \frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\varepsilon_0}$$

$$\nabla \cdot \mathbf{B} = 0$$

where  $\varepsilon_0 \mu_0 = \frac{1}{c^2}$

# Grid Assignment



$$\frac{\partial E_z}{\partial t} = c^2 \frac{\partial B_y}{\partial x} - J_z$$

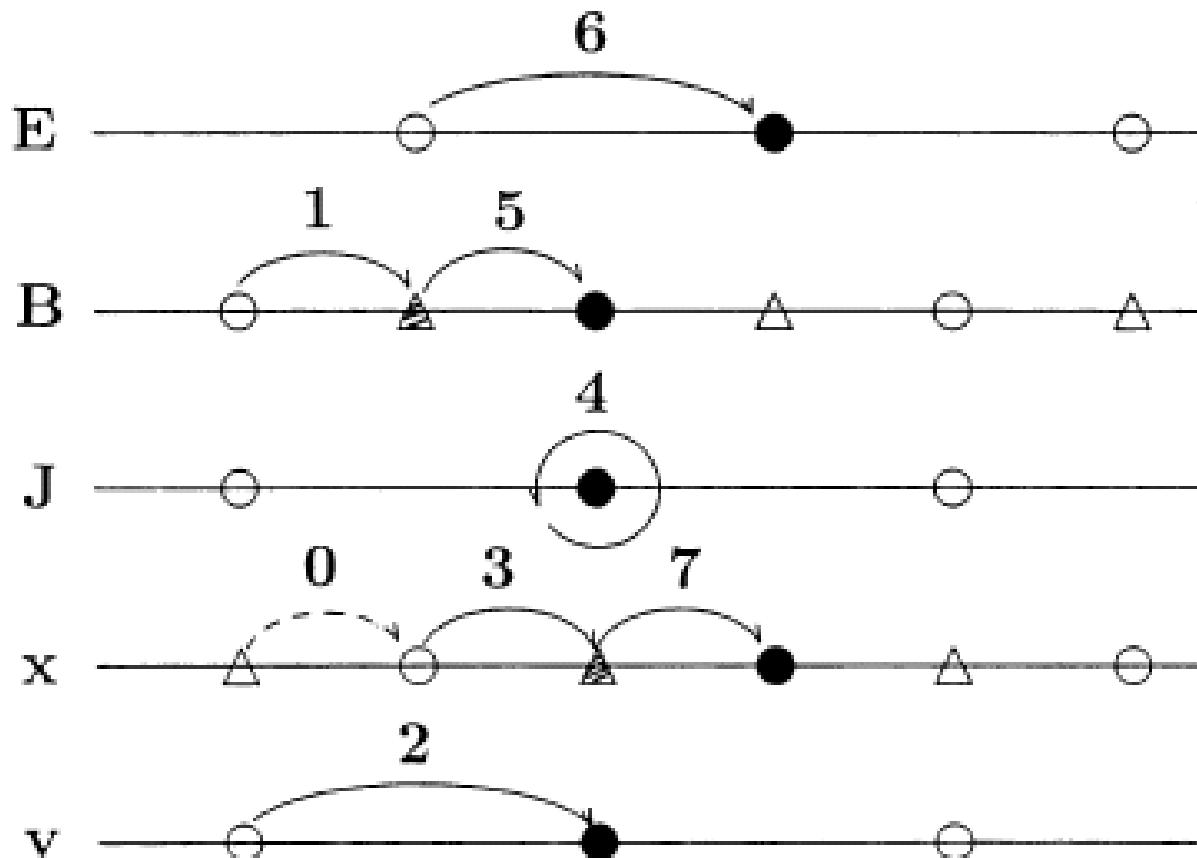
$$\frac{\partial B_y}{\partial t} = \frac{\partial E_z}{\partial x}$$

$$\frac{\partial E_y}{\partial t} = -c^2 \frac{\partial B_z}{\partial x} - J_y$$

$$\frac{\partial B_z}{\partial t} = -\frac{\partial E_y}{\partial x}$$

$$\frac{\partial E_x}{\partial t} = -J_x$$

# Time Step Chart

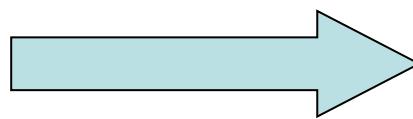


# Centered Difference Scheme

$$E(X_i, t) = E_o \exp(ikX_i - i\omega t)$$

$$\begin{aligned}\frac{\partial E(X_i, t)}{\partial x} &= \frac{E(X_i + \Delta x/2, t) - E(X_i - \Delta x/2, t)}{\Delta x} \\ &= \frac{1}{\Delta x} [\exp(ik\Delta x/2) - \exp(-ik\Delta x/2)] E(X_i, t) \\ &= i \frac{\sin(k\Delta x/2)}{\Delta x/2} E(X_i, t) = iK E(X_i, t)\end{aligned}$$

$k$



$$K = \frac{\sin(k\Delta x/2)}{\Delta x/2}$$

$\omega$



$$\Omega = \frac{\sin(\omega\Delta t/2)}{\Delta t/2}$$

# Courant Condition

Electromagnetic modes in vacuum

$$\omega^2 = c^2 k^2$$

Centered Difference Scheme in space and time

$$\Omega^2 = c^2 K^2 \quad K = \frac{\sin(k\Delta x/2)}{\Delta x/2}$$

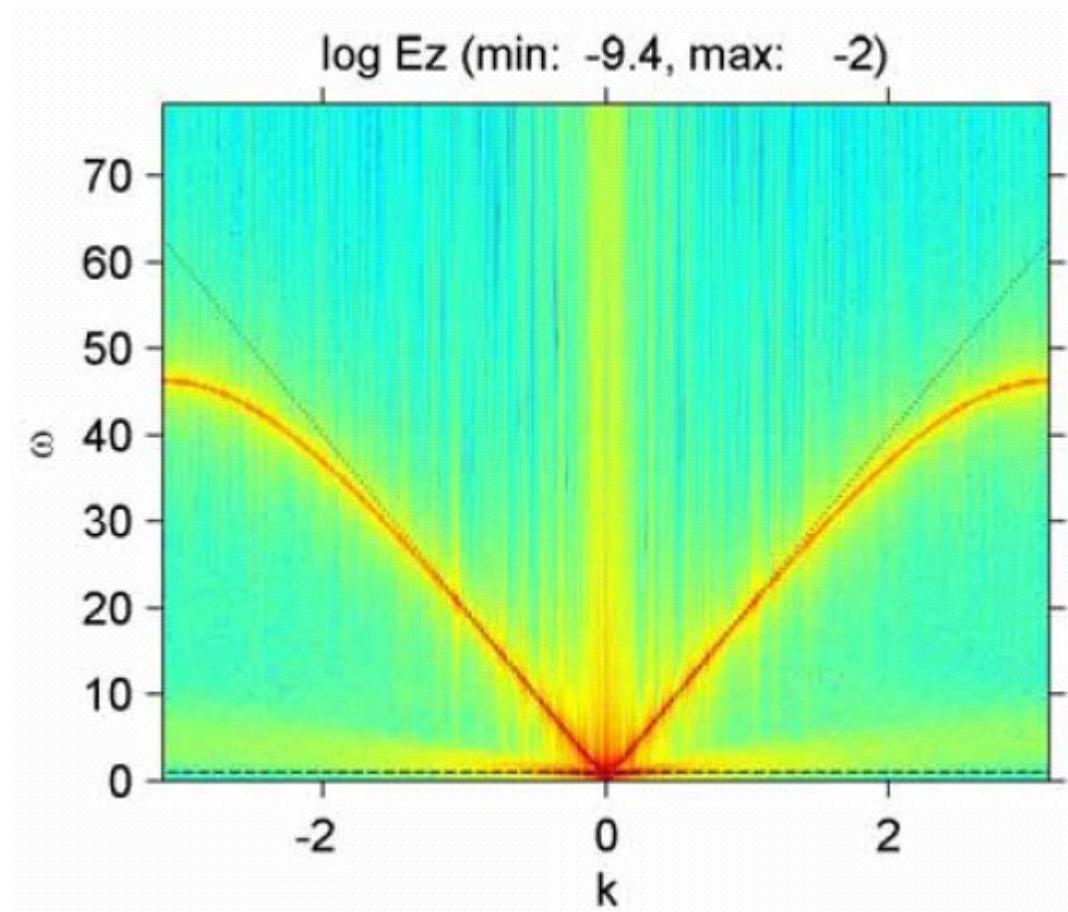
For  $k = \frac{\pi}{\Delta x}$  we have  $\sin(\frac{\omega\Delta t}{2}) = \frac{\Delta t}{\Delta x}c < 1$

Courant Condition

$$c\Delta t < \Delta x$$

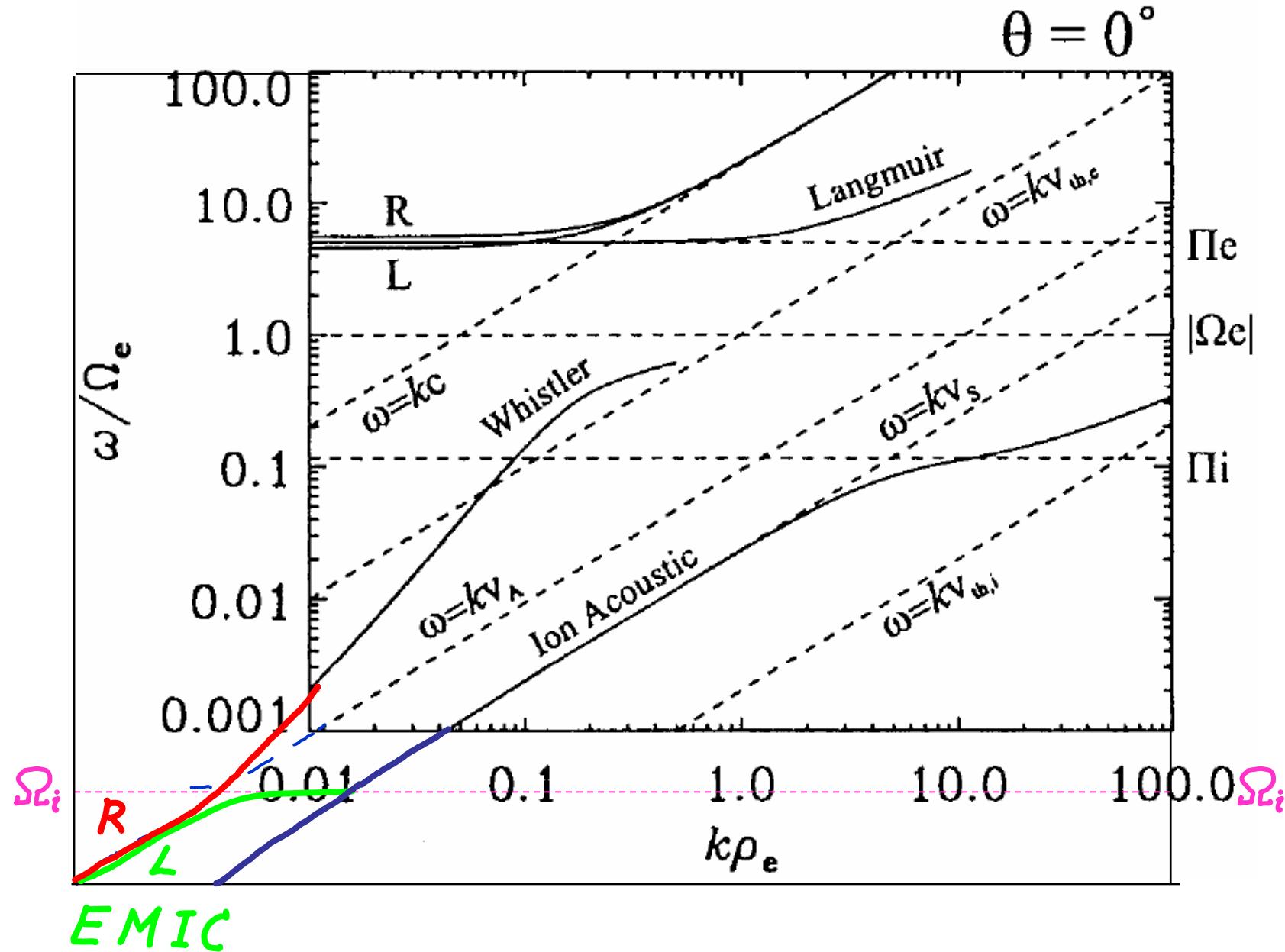
# Dispersion Relation of Light Mode

$$\Omega^2 = c^2 K^2$$



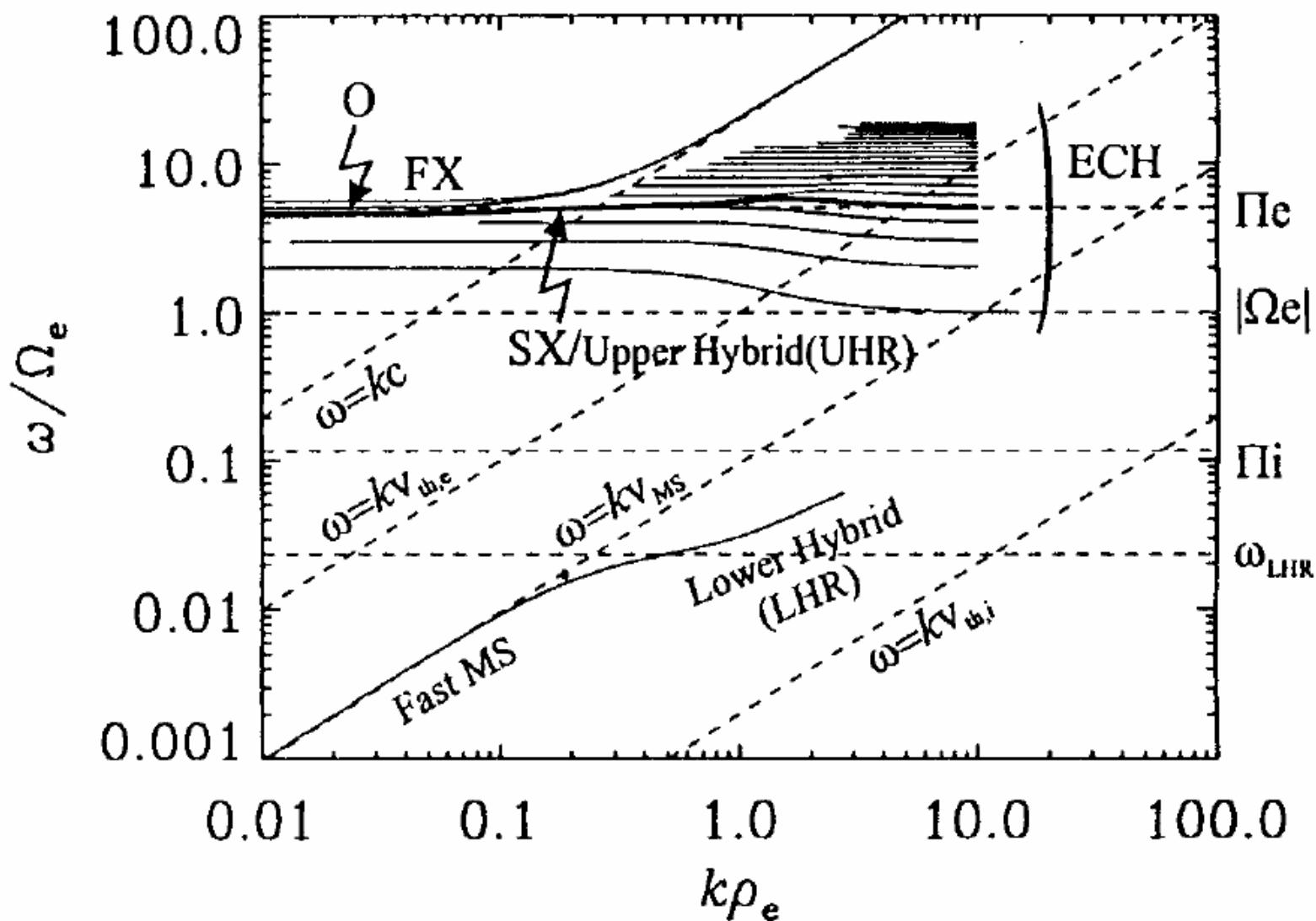
$$\Omega = \frac{\sin(\omega \Delta t / 2)}{\Delta t / 2}, \quad K = \frac{\sin(k \Delta x / 2)}{\Delta x / 2}$$

# Waves in Plasmas (Parallel Propagation)



# Waves in Plasmas (Perpendicular Propagation)

$$\theta = 90^\circ$$



# Charge Density

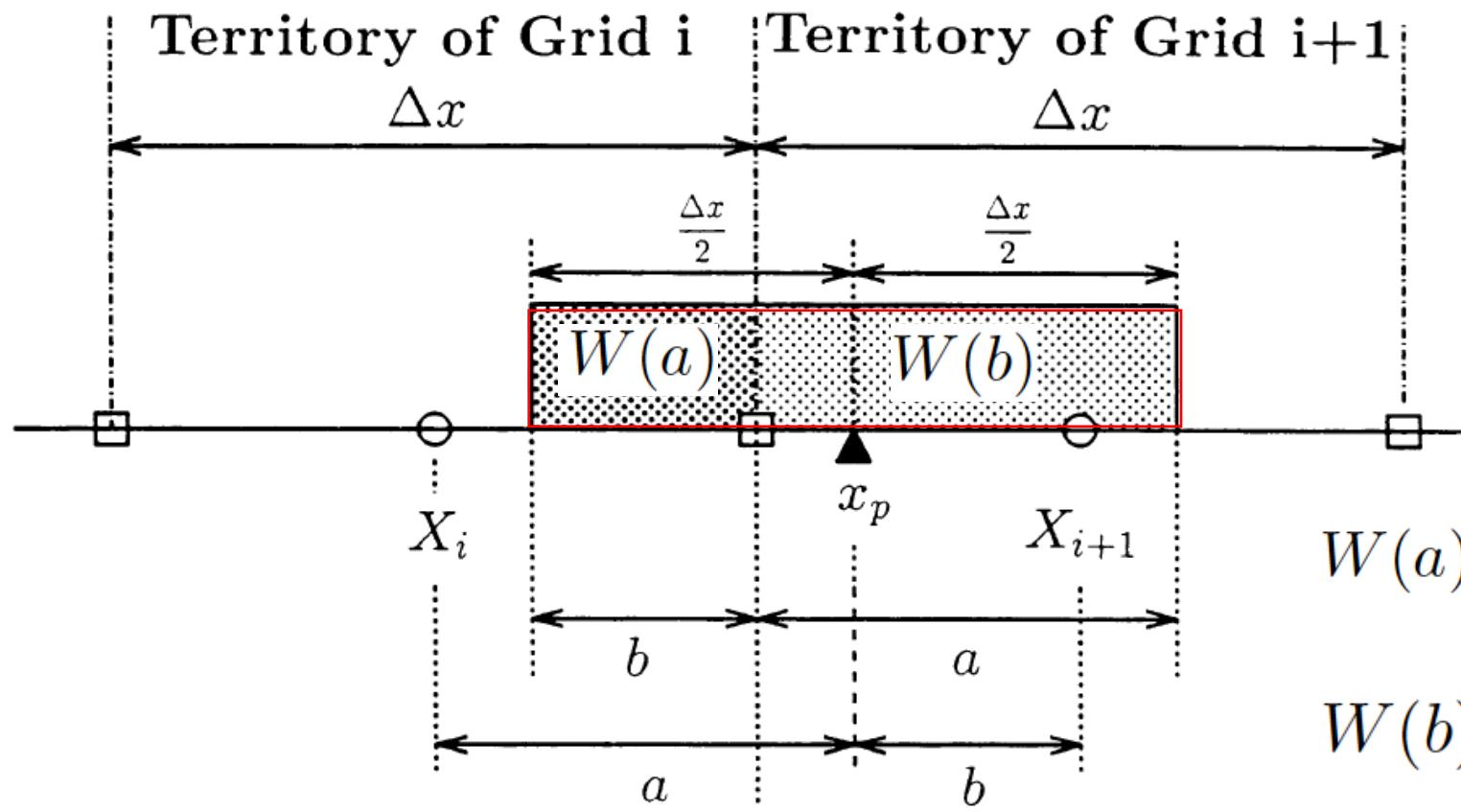
$$\rho_i = \frac{1}{\Delta x} \sum_j^{N_p} q_j W(x_j - X_i)$$

## Shape Function

$$W(x) = 1 - \frac{|x|}{\Delta x}, |x| \leq \Delta x$$
$$= 0, \quad |x| > \Delta x$$

Np: Number of Particles

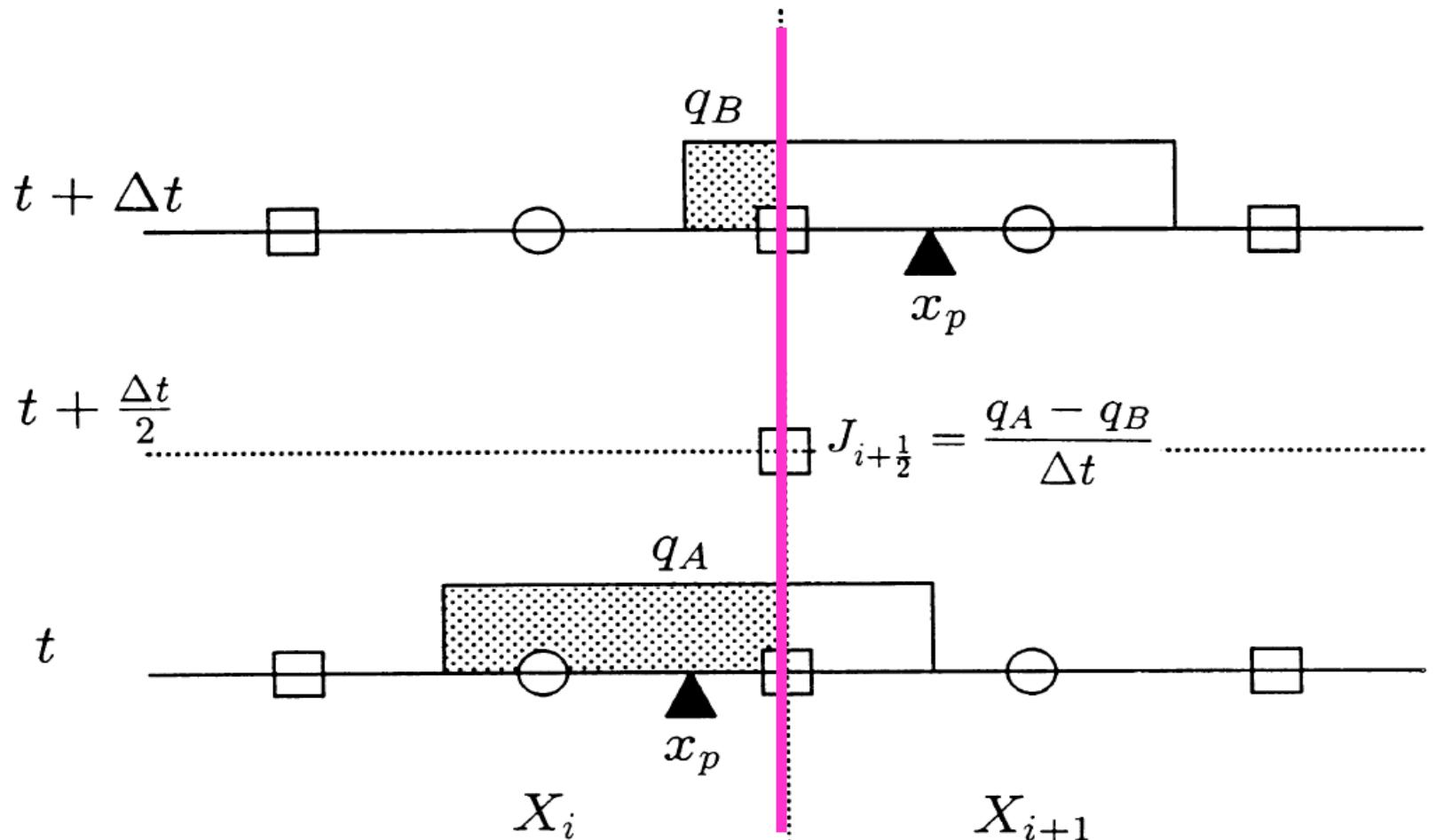
**“fat particle”**



$$W(a) = \frac{b}{\Delta x}$$

$$W(b) = \frac{a}{\Delta x}$$

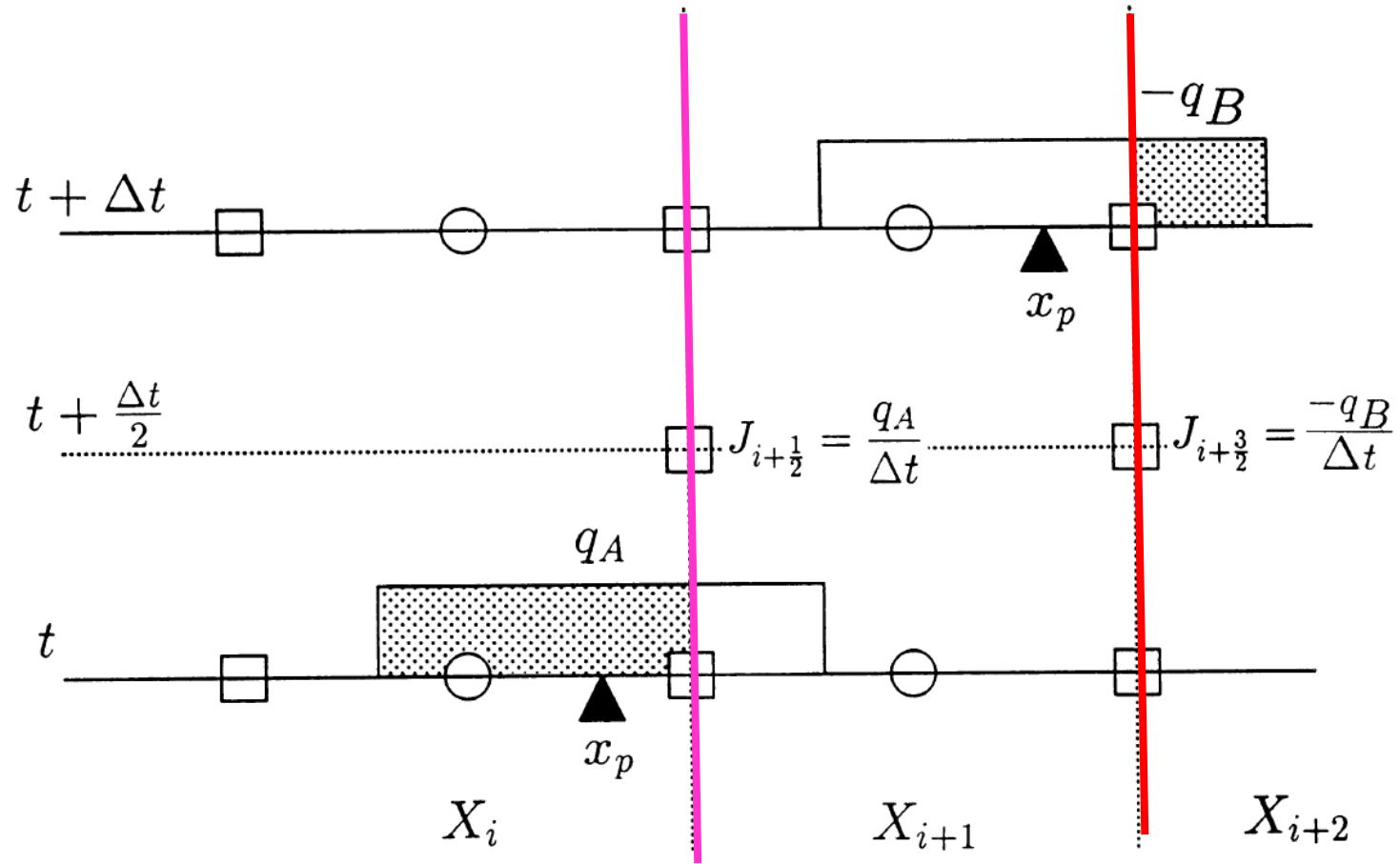
# Current Density $J_x$ : Case 1



$$q_A = q \frac{X_{i+1} - x_p(t)}{\Delta x}$$

$$q_B = q \frac{X_{i+1} - x_p(t + \Delta t)}{\Delta x}$$

# Current Density $J_x$ : Case 2



$$q_A = q \frac{X_{i+1} - x_p(t)}{\Delta x}$$

$$q_B = q \frac{X_{i+1} - x_p(t + \Delta t)}{\Delta x}$$

# Program for Current Density Computation

$$J_{x,i+1/2}^{t+\Delta t/2} - J_{x,i-1/2}^{t+\Delta t/2} = -\frac{\Delta x}{\Delta t}(\rho_i^{t+\Delta t} - \rho_i^t)$$

```
for m = (n1+1):n2
%-- charge conversion method --
qhs = qh * sign(vx(m));
avx = abs(vx(m));
x1 = x(m) + 2.0 -avx;
x2 = x(m) + 2.0 +avx;
i1 = floor(x1);
i2 = floor(x2);
ajx(i1) = ajx(i1) + (i2 - x1)*qhs;
ajx(i2) = ajx(i2) + (x2 - i2)*qhs;
end
```

Only valid for particles satisfying  $|v\Delta t| < \Delta x$

which is automatically satisfied in the relativistic code.

# Cancellation of Uniform Current

$$\frac{\partial \mathbf{J}_u}{\partial t} = \frac{n_e e^2}{m_e} \mathbf{E}_u$$

$$\frac{\partial \mathbf{E}_u}{\partial t} = -\mathbf{J}_u$$

$$\mathbf{J}_u = \mathbf{J}_o \exp(i\omega_{pe}t) \quad \mathbf{E}_u = \frac{i}{\omega_{pe}} \mathbf{J}_o \exp(i\omega_{pe}t)$$

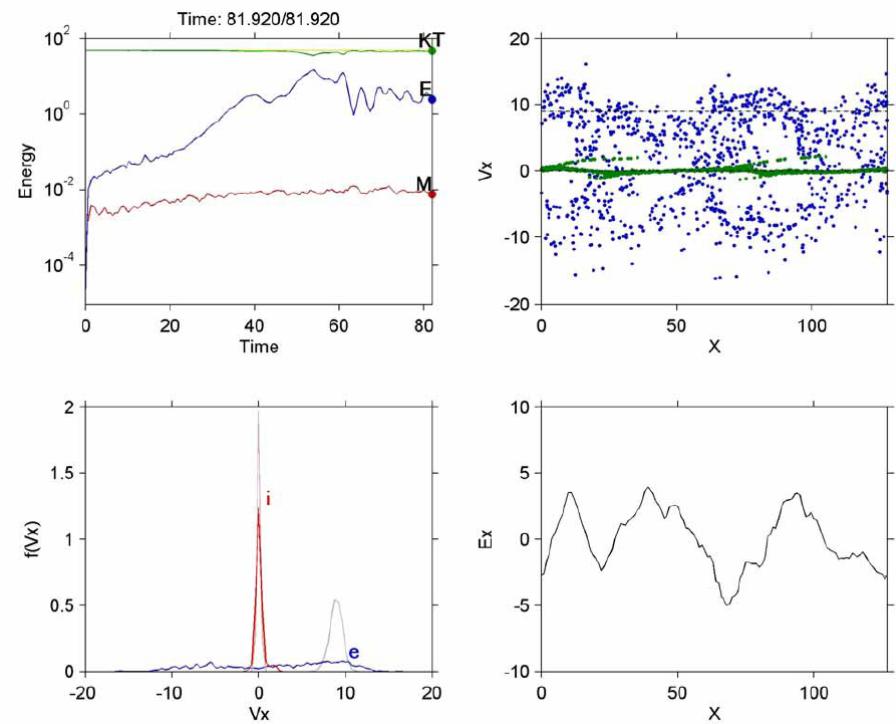
**The uniform component must be subtracted.**

$$\mathbf{J}_u = \frac{1}{N_x} \sum_{i=1}^{N_x} \mathbf{J}_i \quad \mathbf{J}_{i,sub} = \mathbf{J}_i - \mathbf{J}_u$$

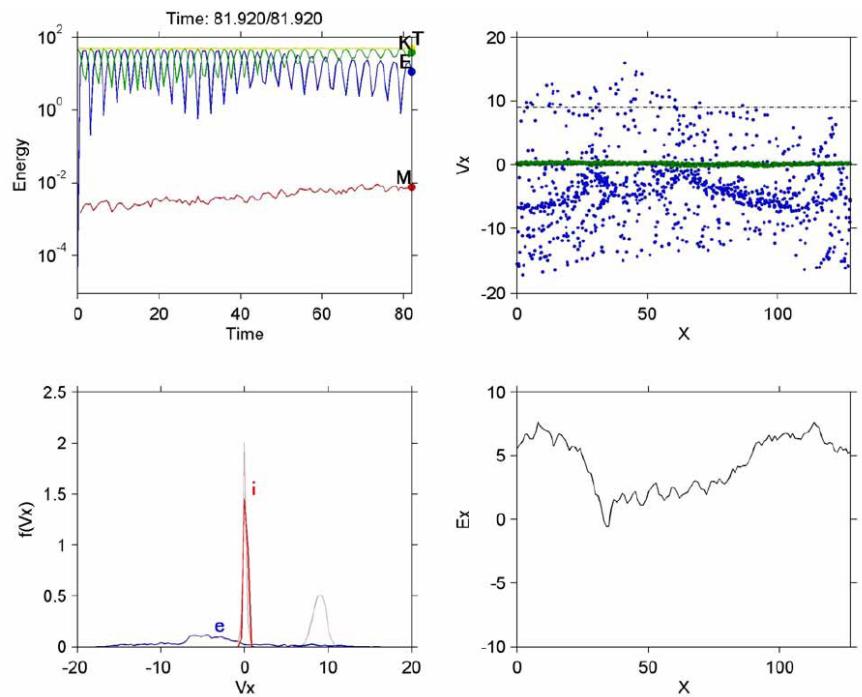
“Periodic systems are current-neutral.” (V. Decyk)

# KEMPO1

## Without Uniform Current



## With Uniform Current

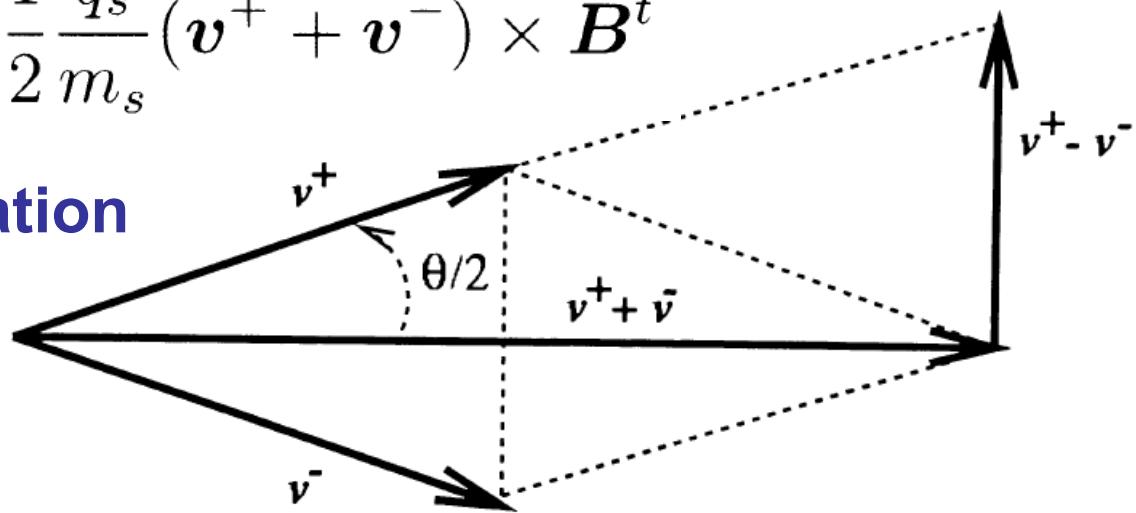


# Buneman-Boris Method

$$\frac{\mathbf{v}^{t+\Delta t/2} - \mathbf{v}^{t-\Delta t/2}}{\Delta t} = \frac{q_s}{m_s} (\mathbf{E}^t + \frac{\mathbf{v}^{t+\Delta t/2} + \mathbf{v}^{t-\Delta t/2}}{2} \times \mathbf{B}^t)$$

$$\mathbf{v}^- = \mathbf{v}^{t-\Delta t/2} + \frac{q_s}{m_s} \mathbf{E}^t \frac{\Delta t}{2} \quad \mathbf{v}^+ = \mathbf{v}^{t+\Delta t/2} - \frac{q_s}{m_s} \mathbf{E}^t \frac{\Delta t}{2}$$

$$\frac{\mathbf{v}^+ - \mathbf{v}^-}{\Delta t} = \frac{1}{2} \frac{q_s}{m_s} (\mathbf{v}^+ + \mathbf{v}^-) \times \mathbf{B}^t$$



**Kinetic Energy Conservation**

$$(\mathbf{v}^+)^2 = (\mathbf{v}^-)^2$$

**Small Phase Delay**

$$\Omega_c = \frac{\tan^{-1} \omega_c \Delta t / 2}{\Delta t / 2}$$

$$\Omega_c / \omega_c = 0.9967 \text{ with } \omega_c \Delta t = 0.2$$

# Relativistic Equation of Motion

$$\frac{d}{dt}(m\mathbf{v}) = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

$$m = \gamma m_0$$
$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

$$\mathbf{u} = \frac{c}{\sqrt{c^2 - |\mathbf{v}|^2}} \mathbf{v}$$

$$\mathbf{B}_u = \frac{c}{\sqrt{c^2 + |\mathbf{u}|^2}} \mathbf{B}$$

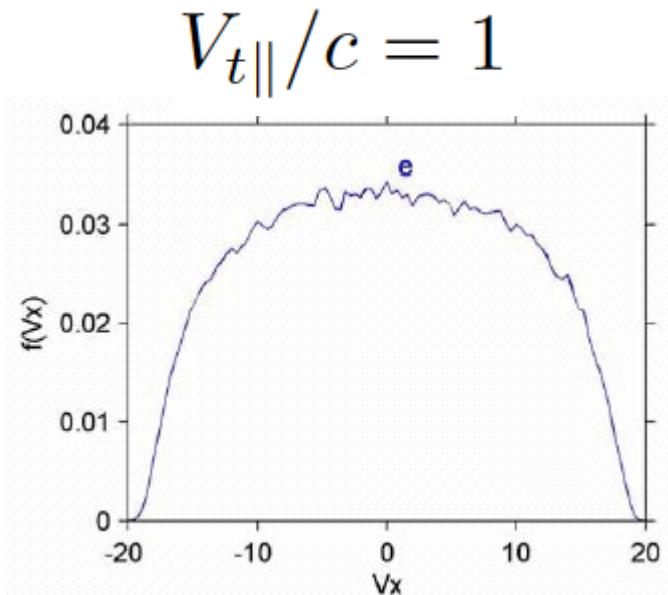
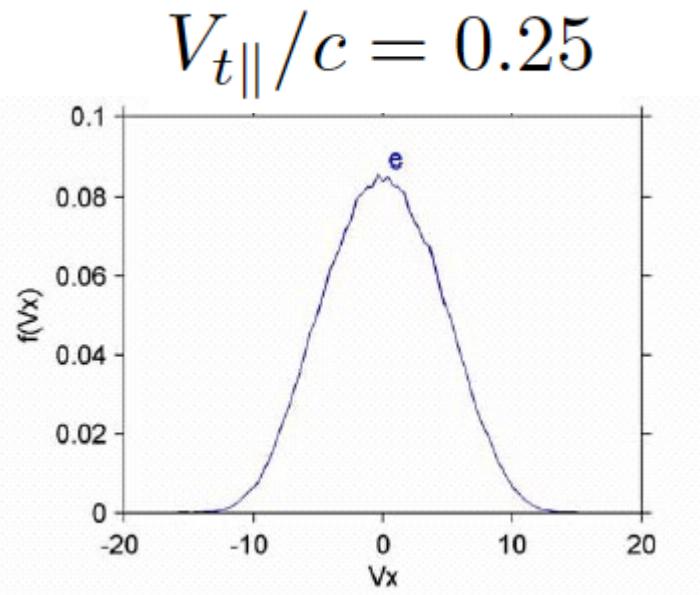
$$\frac{d\mathbf{u}}{dt} = \frac{q}{m_0} (\mathbf{E} + \mathbf{u} \times \mathbf{B}_u)$$

$$\mathbf{v} = \frac{c}{\sqrt{c^2 + |\mathbf{u}|^2}} \mathbf{u}$$

# Initial Velocity Distribution Function

$$f(u_{\parallel}, u_{\perp}) \propto \exp\left(-\frac{(u_{\parallel} - V_{d\parallel})^2}{2V_{t\parallel}^2} - \frac{(u_{\perp} - V_{d\perp})^2}{2V_{t\perp}^2}\right)$$

$$\boldsymbol{v} = \boldsymbol{u}/\gamma = \frac{\boldsymbol{c}}{\sqrt{c^2 + u_x^2 + u_y^2 + u_z^2}} \boldsymbol{u}$$
$$V_{d\parallel} = V_{d\perp} = 0$$

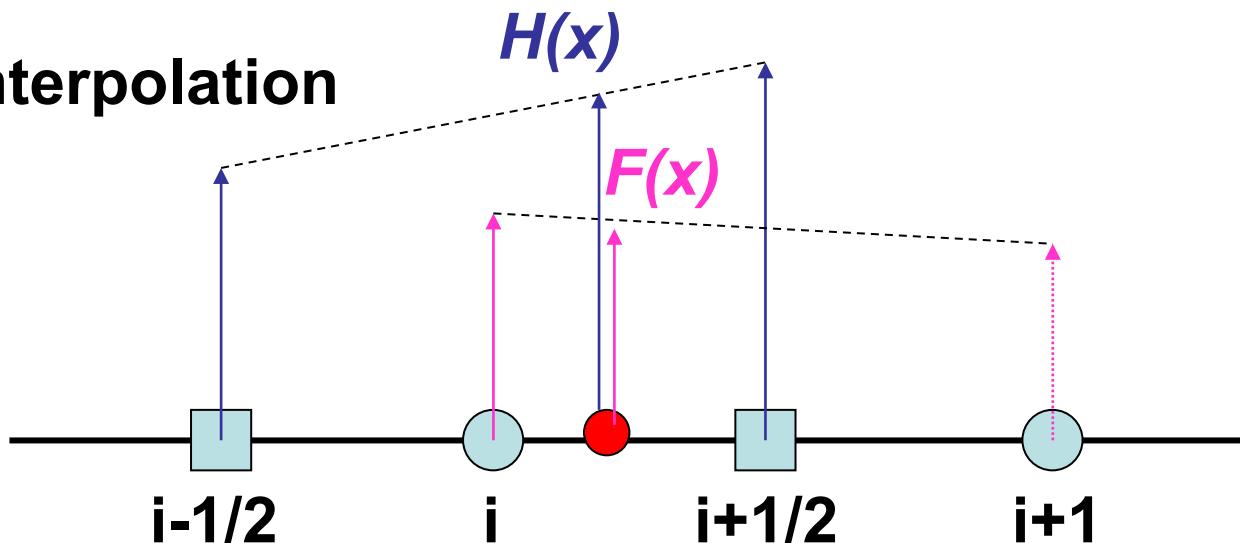


# Field Interpolation to Particle Position

$$F(x) = \sum_{i=1}^{N_x} F_i W(x - X_i)$$

$$H(x) = \sum_{i=1}^{N_x} H_{i+1/2} W(x - X_{i+1/2})$$

Linear Interpolation



# Electrostatic Self-force Cancellation

$$\rho_i = \frac{1}{\Delta x} \sum_j^{N_p} q_j \underline{W(x_j - X_{\boxed{i}})}$$

$$\frac{E_{x,i+1/2} - E_{x,i-1/2}}{\Delta x} = \frac{\rho_i}{\varepsilon_0}$$

$$E_x(x) = \sum_{i=1}^{N_x} E_{x,i+1/2} \underline{W(x - X_{\boxed{i+1/2}})}$$



**Self-force**

**Relocation**

$$E_{x,i} = \frac{E_{x,i-1/2} + E_{x,i+1/2}}{2}$$

$$E_x(x) = \sum_{i=1}^{N_x} E_{x,i} \underline{W(x - X_{\boxed{i}})}$$



**No Self-force**

# Magnetostatic Self-force Cancellation

Magnetostatic equation (Ampere's Law)  $\nabla \times \mathbf{B} = \mu_0 \mathbf{J}$

$$\mathbf{J}_{yz,i+1/2} = \frac{1}{\Delta x} \sum_j^{N_p} q_j \mathbf{v}_{yz} W(x_j - X_{i+1/2})$$

## Source Relocation

$$J_{y,i} = \frac{J_{y,i-1/2} + J_{y,i+1/2}}{2}$$

$$\frac{B_{z,i+1/2} - B_{z,i-1/2}}{\Delta x} = -\mu_o J_{y,i}$$

## Field Relocation

$$\frac{B_{y,i+1} - B_{y,i}}{\Delta x} = \mu_o J_{z,i+1/2}$$

$$B_{y,i+1/2} = \frac{B_{y,i} + B_{y,i+1}}{2}$$

# Relative Unit System

We assume in KEMPO1

$$\varepsilon_0 = 1, \quad \mu_0 = \frac{1}{c^2}$$

## Reference Values for Space and Time

Speed of light:  $c$

Characteristic frequency:  $\omega \max(\omega_p, \omega_c)$

We choose  $\Delta t$  so that  $\omega \Delta t \ll 2\pi$

We choose  $\Delta x$  so that  $c \Delta t < \Delta x$

# Charge Neutrality Condition

Number Density  $n_i = \frac{N_p}{N_x \Delta x}$  Nx: Number of Grid Points

Particle Charge  $q_i = \frac{\rho_i}{n_i} = \frac{\omega_{pi}^2}{(q_i/m_i)n_i}$

$q_i/m_i$  is an arbitrary value.

For simplicity, -1 for electrons.

0.0625, 0.01 or 1/1836 for protons.

$$\sum_i \rho_i = \sum_i q_i n_i = \boxed{\sum_i \frac{\omega_{pi}^2}{q_i/m_i} = 0}$$

# Enhanced Thermal Fluctuation

## Electrostatic Field Energy Density

$$F_E = \frac{T}{2} \int_{-\infty}^{\infty} \frac{1}{1 + k^2 \lambda_D^2} \frac{dk}{2\pi} = \frac{T}{4\lambda_D}$$

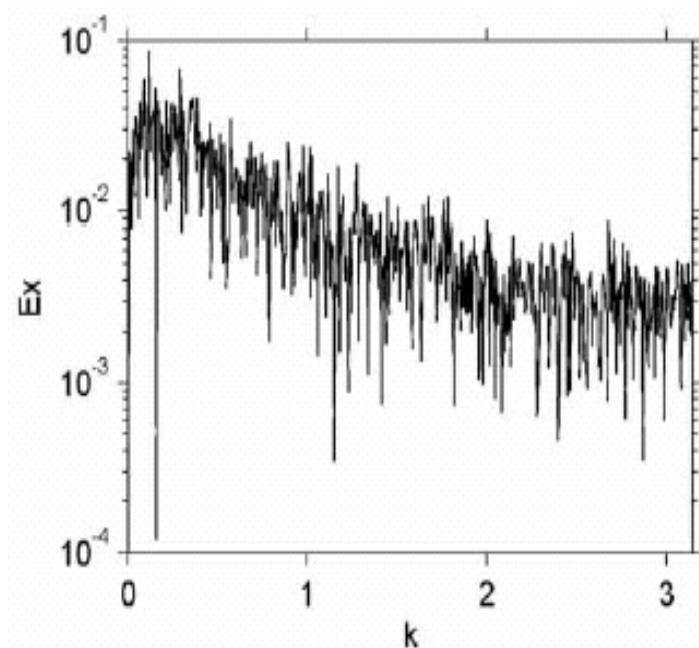
where  $T (\equiv m V_t^2)$

## Thermal Energy Density

$$T_E = \frac{1}{2} n T$$

## Normalized Thermal Fluctuation

$$\frac{F_E}{T_E} = \frac{1}{2} \frac{1}{n \lambda_D} = \frac{1}{2} \frac{N_x}{N_p} \frac{\Delta x}{\lambda_D}$$



# Debye Length / Grid Spacing

**Density Perturbation with a small potential  $\delta\phi$**

$$n \sim n_0 \exp\left(-\frac{q\delta\phi}{T}\right)$$

**Inserting into Poisson's Equation,**

$$\frac{\partial^2 \delta\phi}{\partial x^2} \sim \frac{q^2 n}{\varepsilon_o T} \delta\phi = \frac{\omega_p^2}{V_t^2} \delta\phi = \frac{1}{\lambda_D^2} \delta\phi$$

**Fourier transformation**

$$(k^2 - \frac{1}{\lambda_D^2}) \delta\phi_k \sim 0$$

$$|k| \sim \frac{1}{\lambda_D}$$

**Replacing  $k$  with  $K$**

$$|\sin(k\Delta x/2))| \sim \frac{\Delta x}{2\lambda_D} < 1$$

$$\boxed{\Delta x < 2\lambda_D}$$

# GUI for Input Parameters

KEMPO 1

File(F) Help(H)

**KEMPO 1**

**Input parameters**

DX	1	NX	256
DT	0.04	NTIME	16384
CV	20		
WC	-1	ANGLE	0
NS	2	Species 1	
QM	-1	WP	2
VPE	1	VPA	1
VD	0	PCH	0
NP	4096	Preview	
AJAMP	0	WJ	0

**IEX**  0: Ex=0  
 1: EM  
 2: ES

**Diagnostics**

Panel 1	Energy
Panel 2	VyzEByz-X
Panel 3	By(w,k)
Panel 4	Bz(t,X)
NPLOT	512
Vmax	20
Emax	10
Bmax	0.2
NV	100

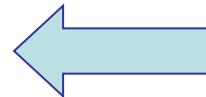
P. Color     Param

LOAD    SAVE    START    EXIT

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# Renormalization

Simulation Unit System



Relative Unit System

distance

$$x_S = (1/\Delta x)x_R$$

time

$$t_S = (2/\Delta t)t_R$$

velocity

$$v_S = (\Delta t/2)(1/\Delta x)v_R$$

electric field

$$E_S = (\Delta t/2)^2(1/\Delta x)E_R$$

magnetic field

$$B_S = (\Delta t/2)B_R$$

charge density

$$\rho_S = (\Delta t/2)^2\rho_R$$

current density

$$J_S = (\Delta t/2)^3(1/\Delta x)J_R$$

energy density

$$\sigma_S = (\Delta t/2)^4(1/\Delta x)^2\sigma_R$$

number density

$$n_S = \Delta x n_R$$

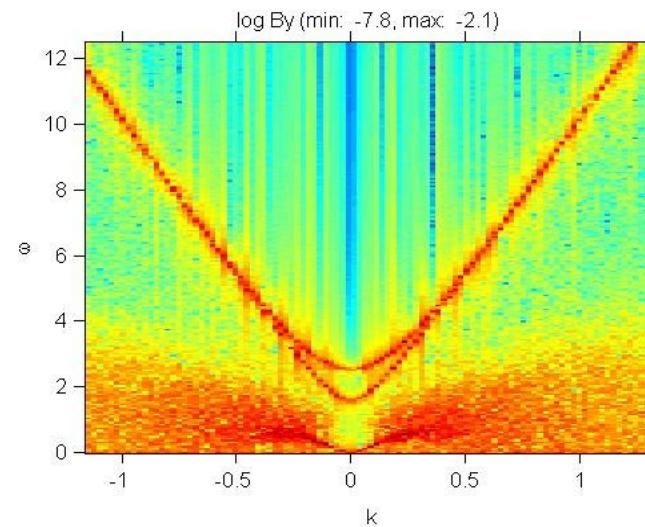
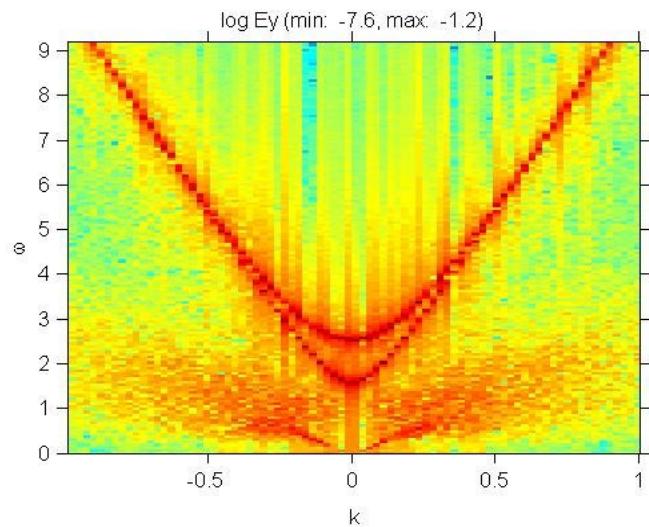
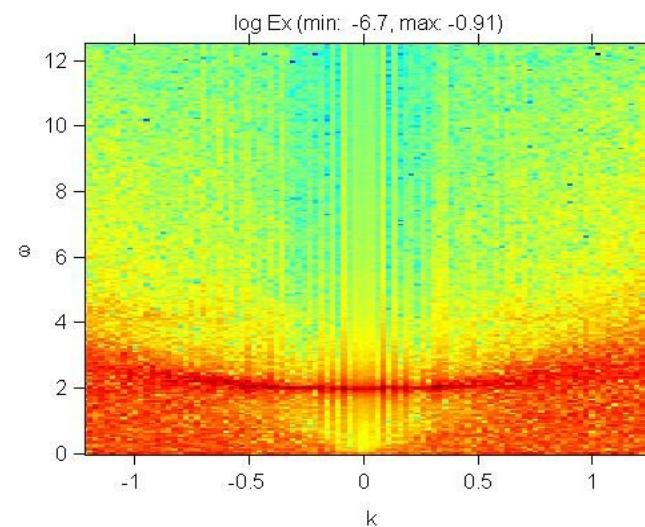
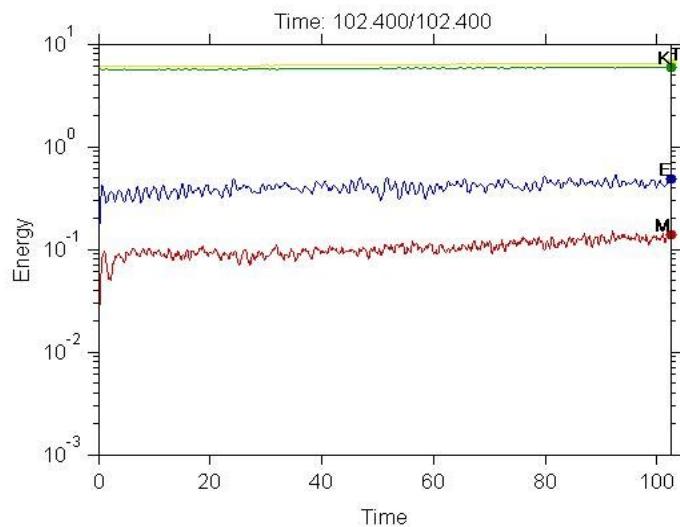
charge

$$q_S = (\Delta t/2)^2(1/\Delta x)q_R$$

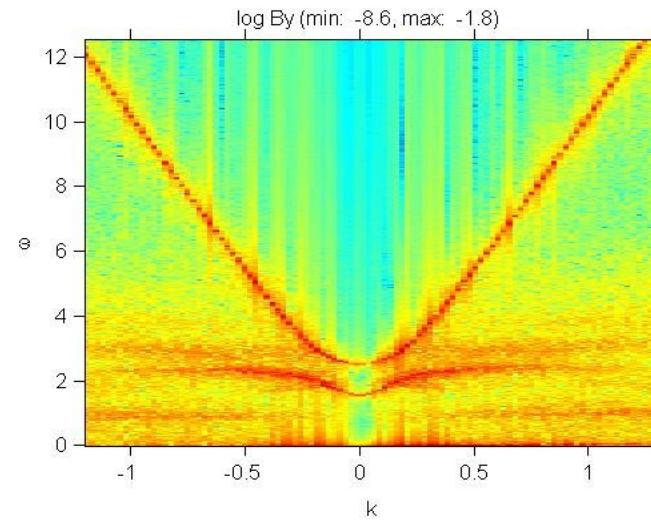
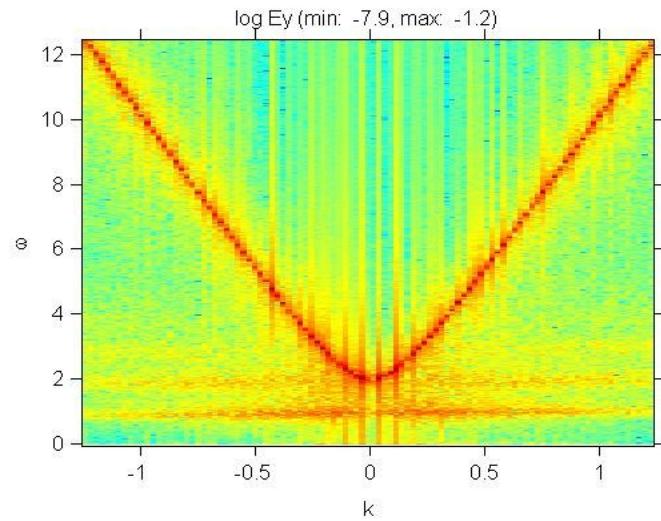
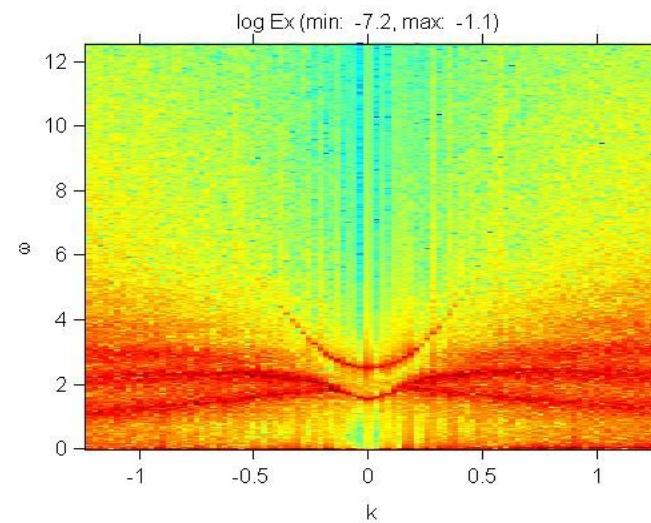
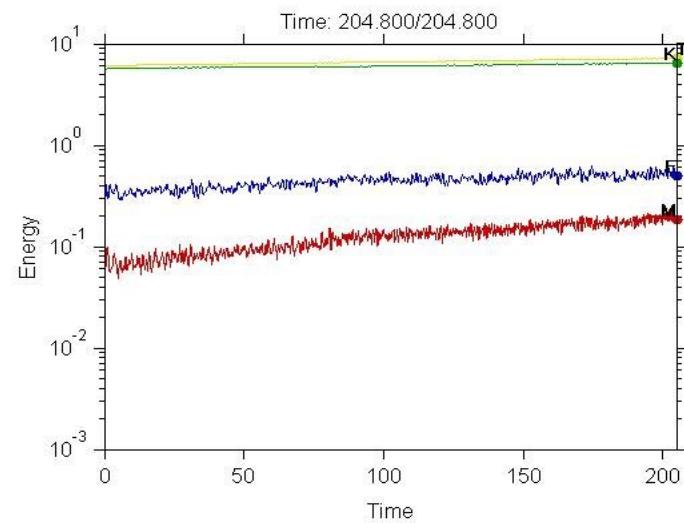
mass

$$m_S = (\Delta t/2)^2(1/\Delta x)m_R$$

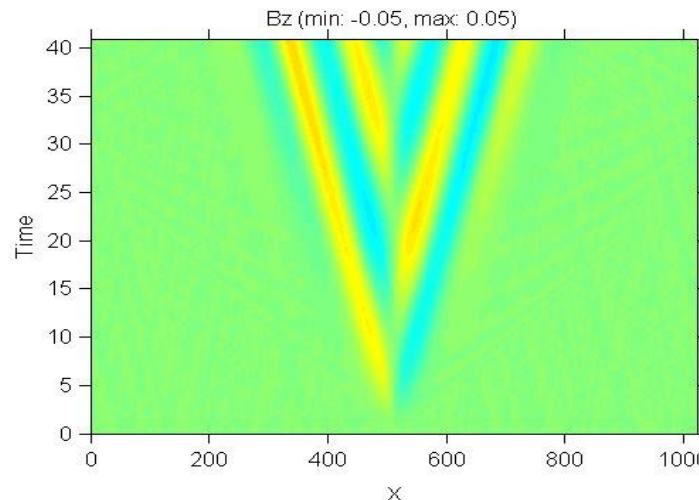
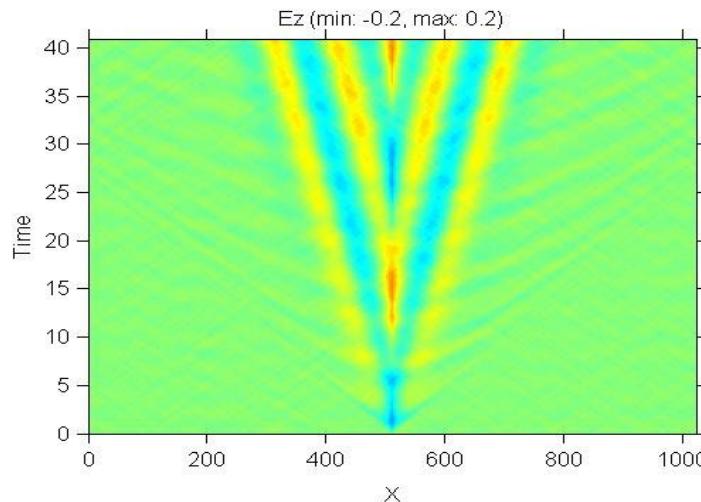
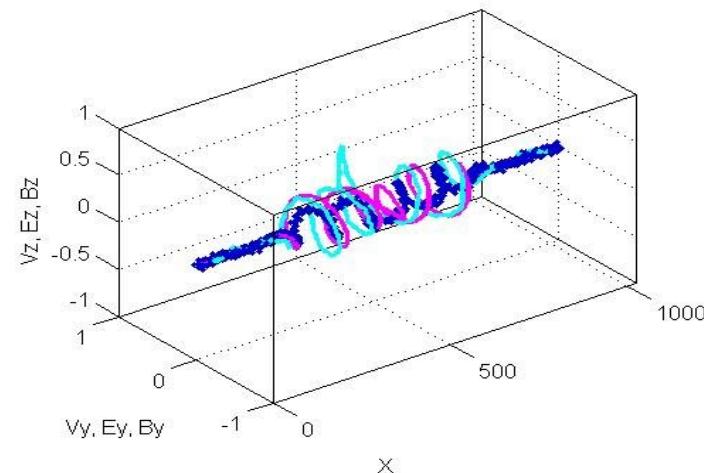
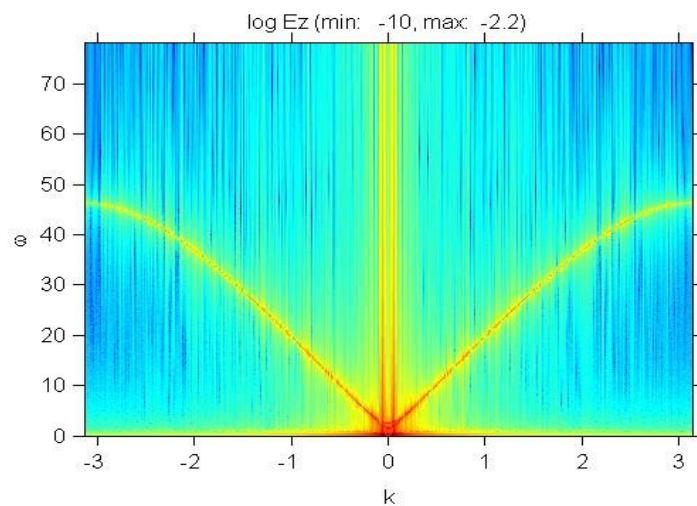
# Normal Modes in the Parallel Direction



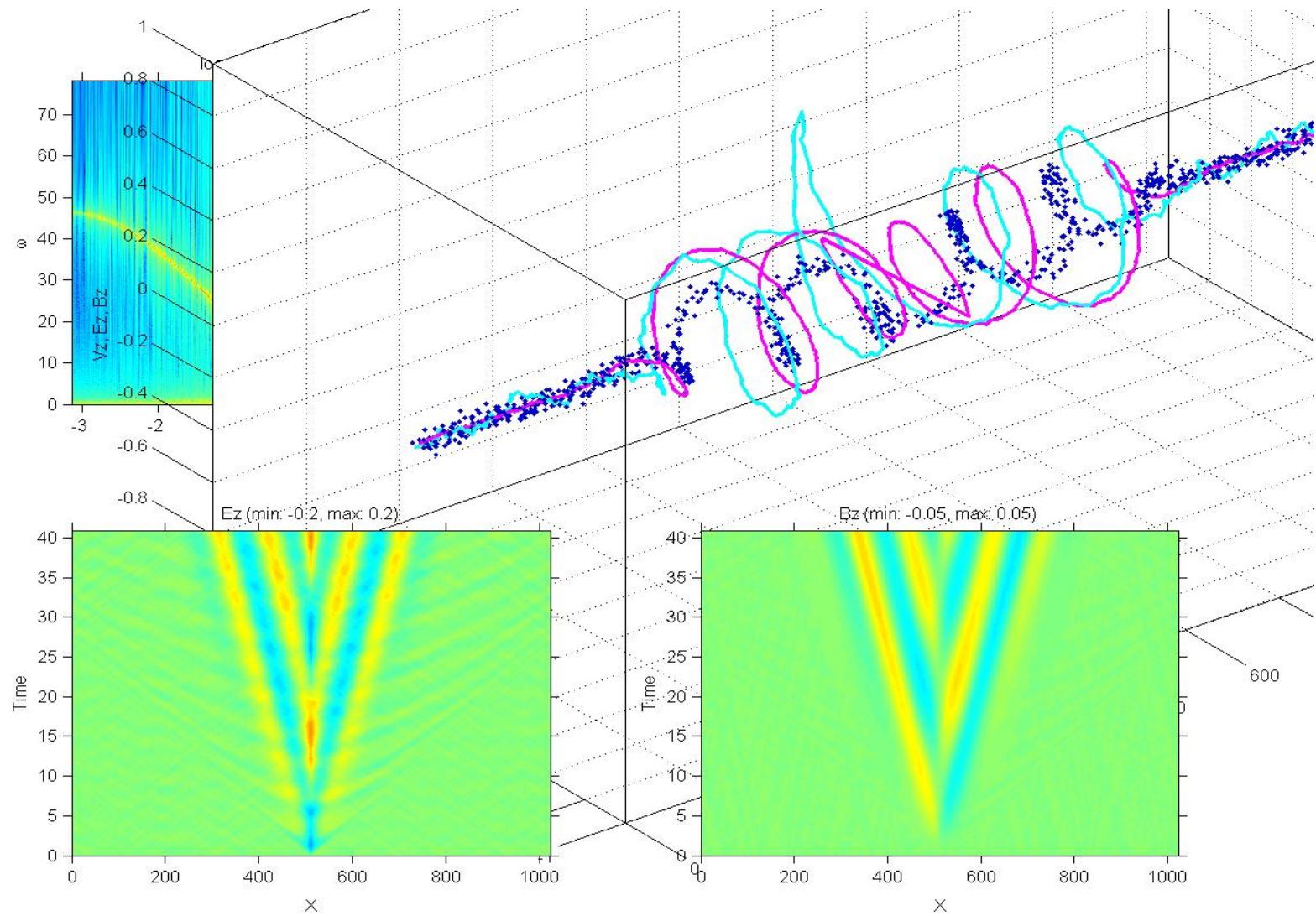
# Normal Modes in the Perpendicular Direction



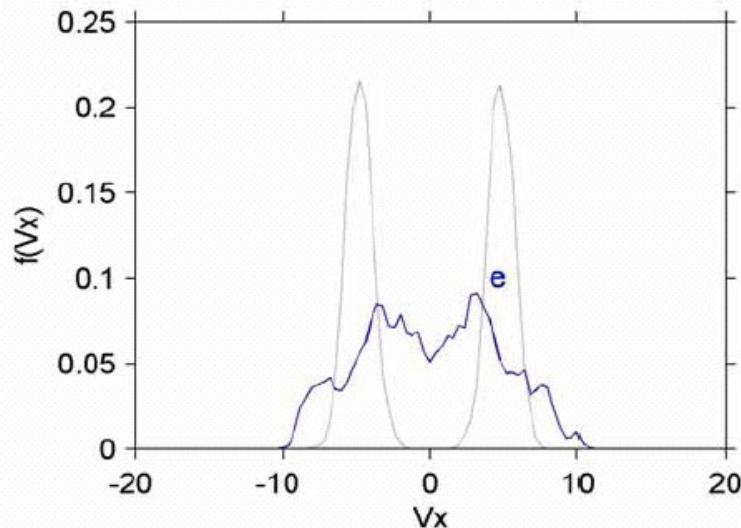
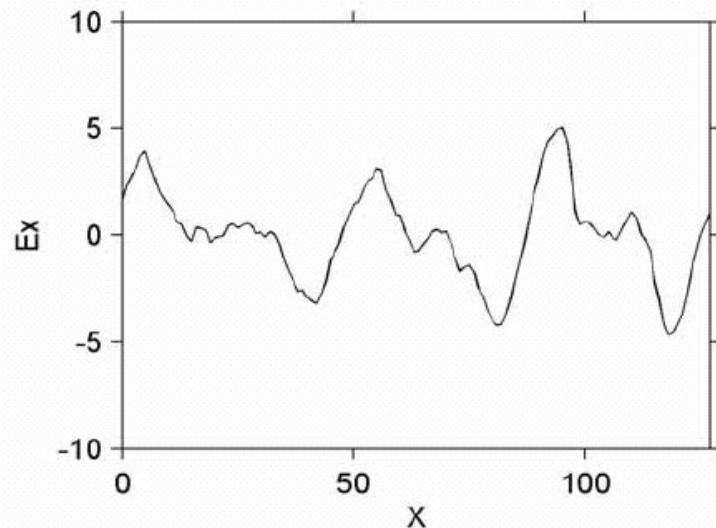
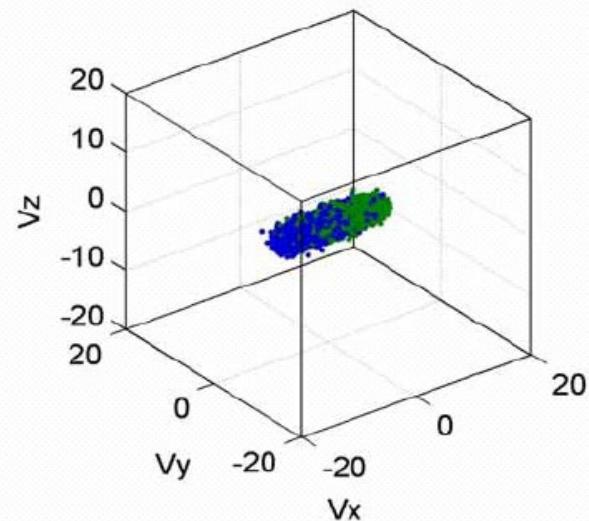
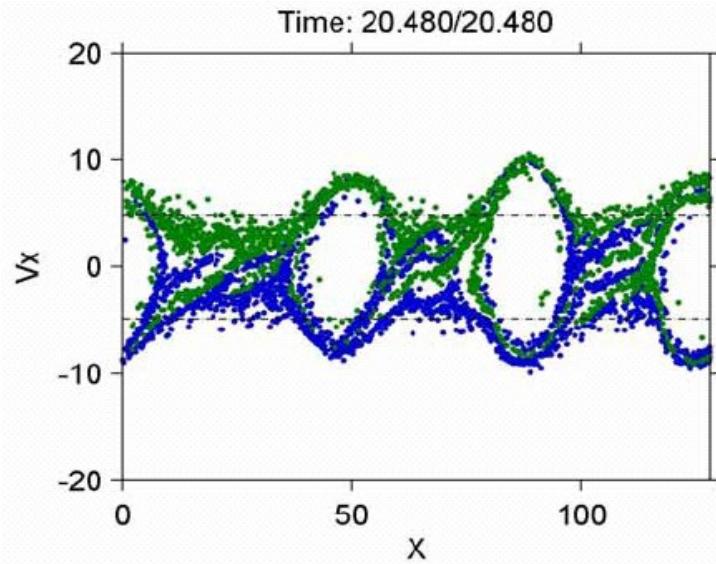
# Whistler Mode Wave Radiation



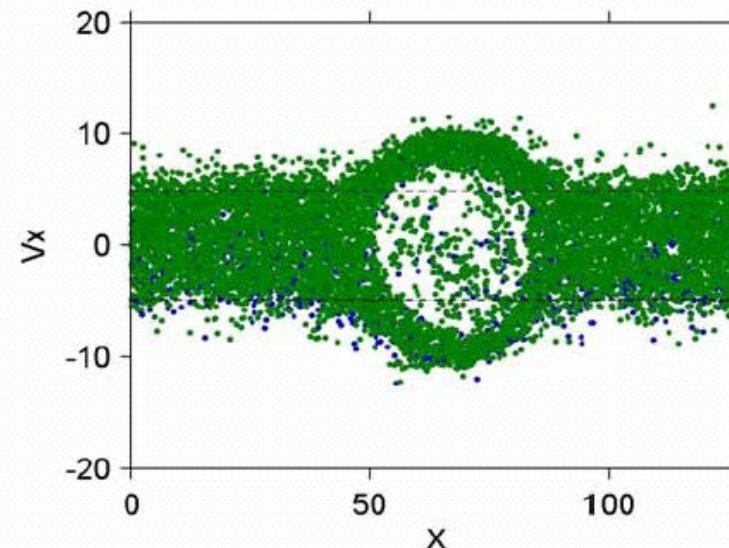
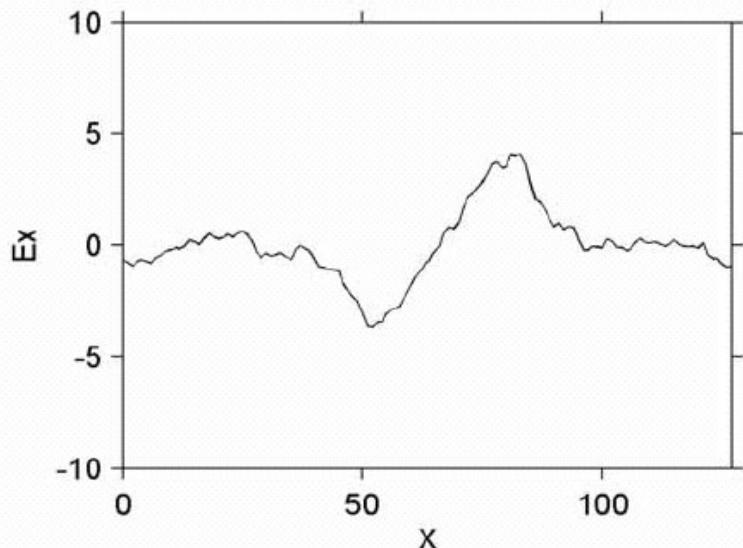
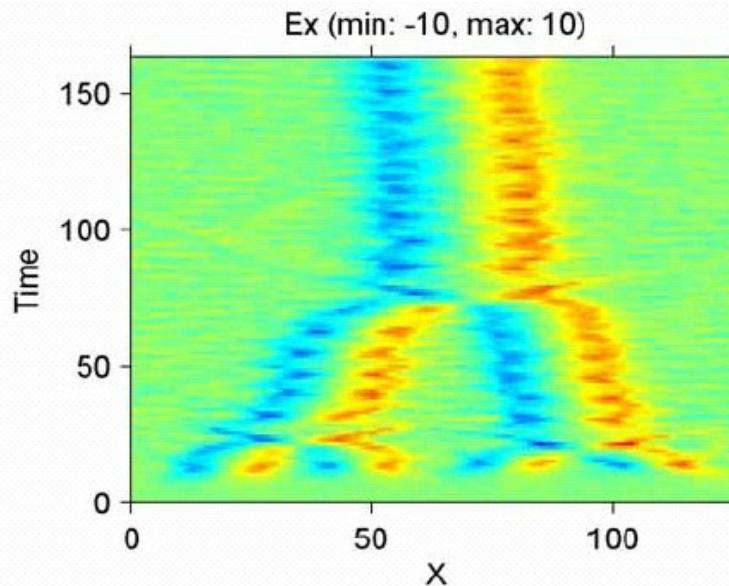
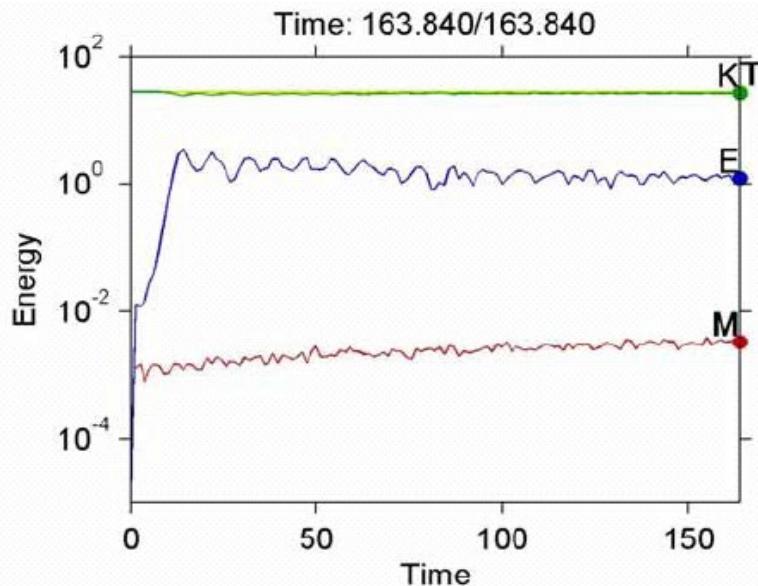
# Circularly Polarized Wave



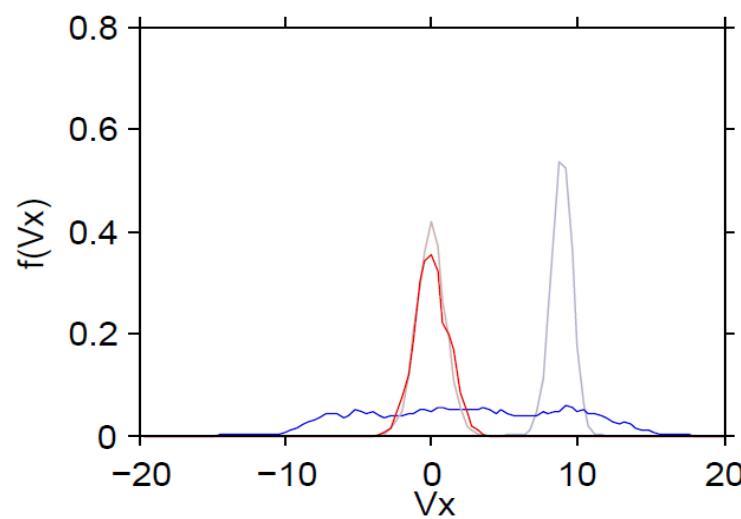
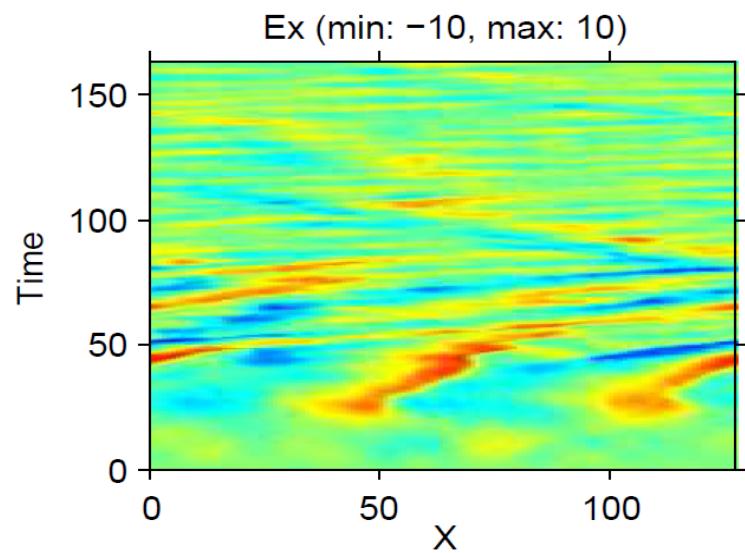
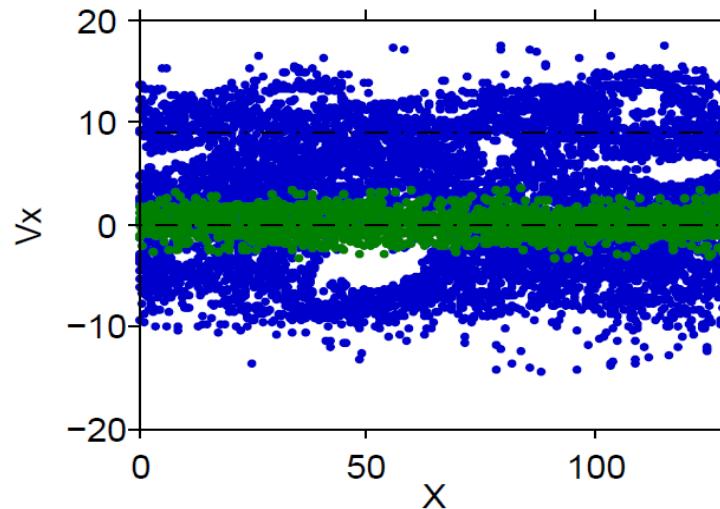
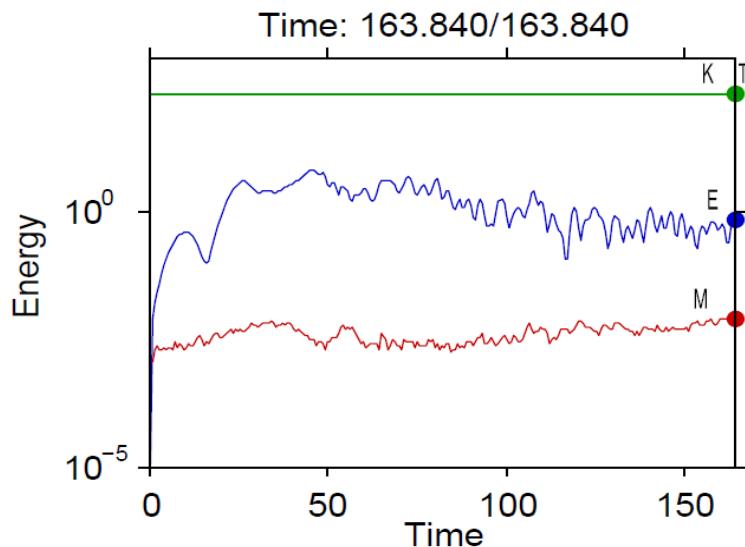
# Electron Two-stream Instability



# Electrostatic Solitary Waves



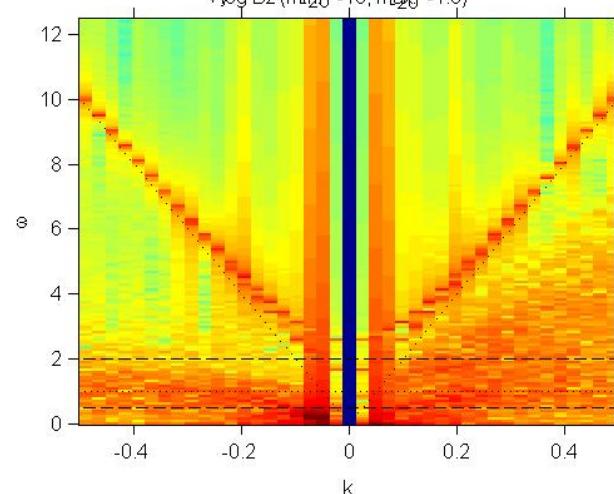
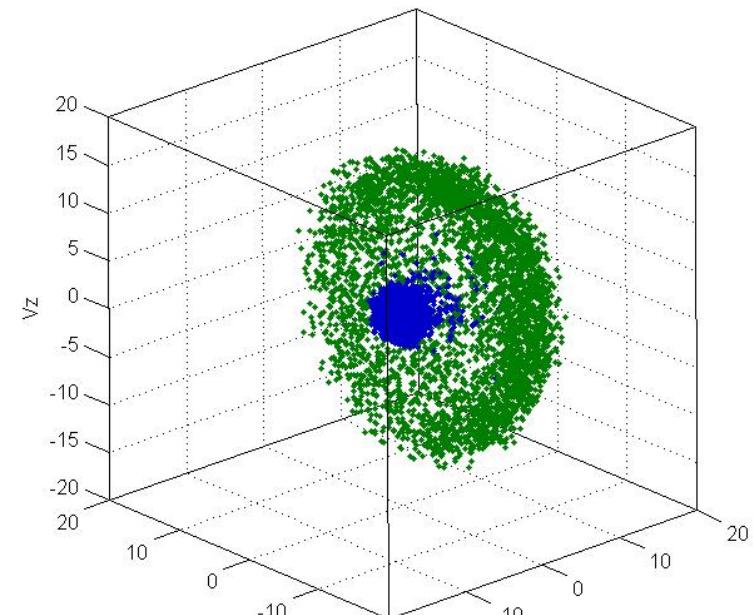
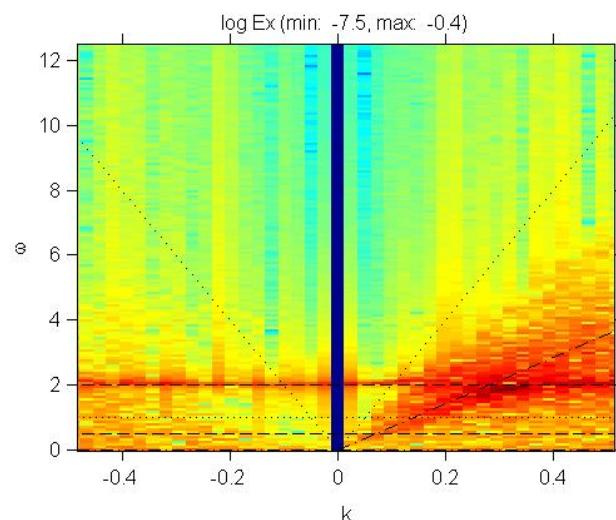
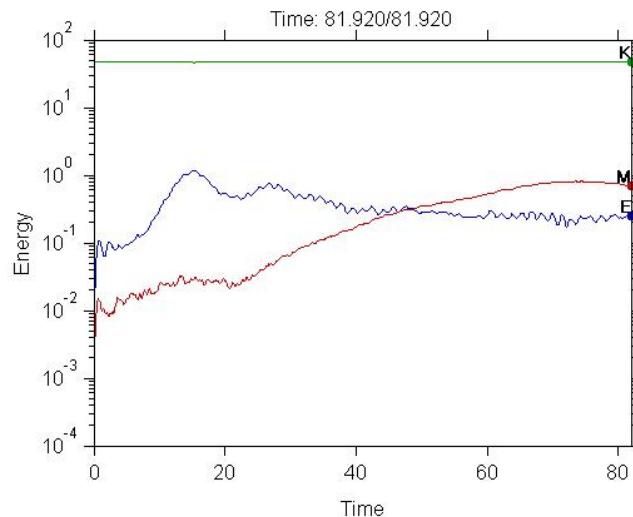
# Buneman Instability



[e.g., Omura et al., JGR, 2003]

# Electrostatic and Whistler Instabilities

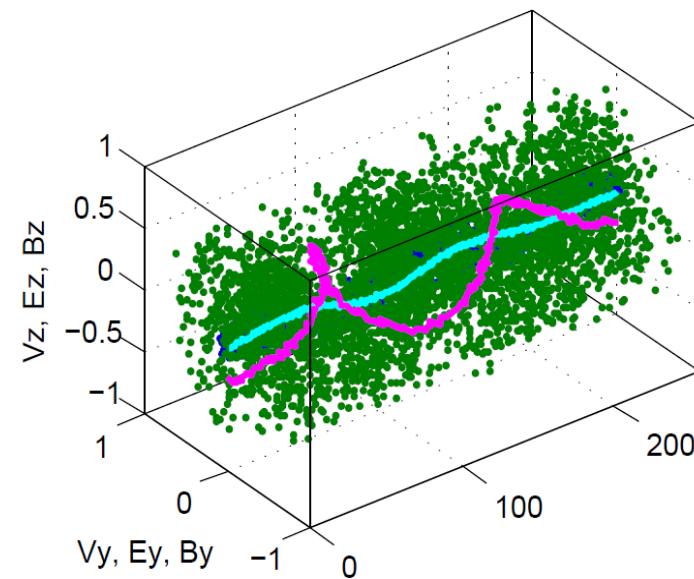
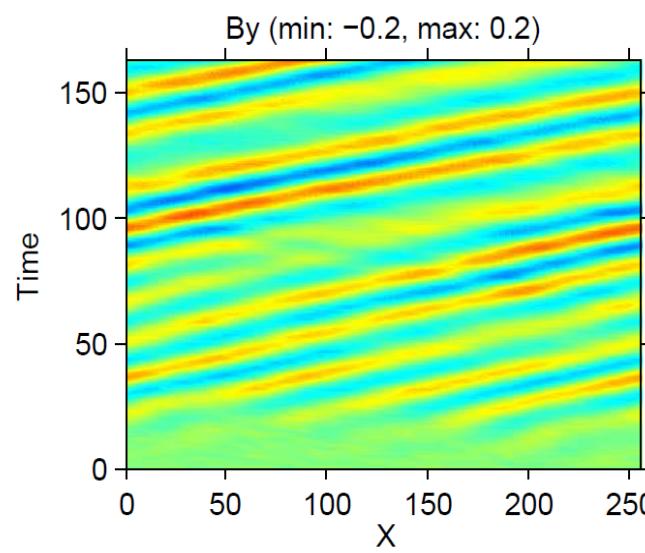
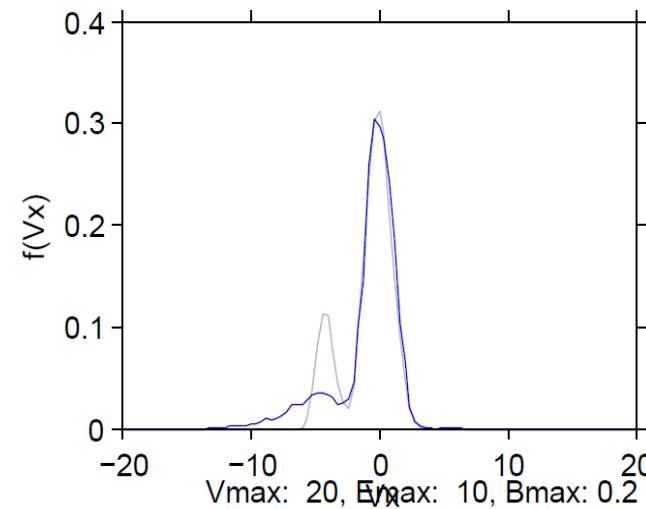
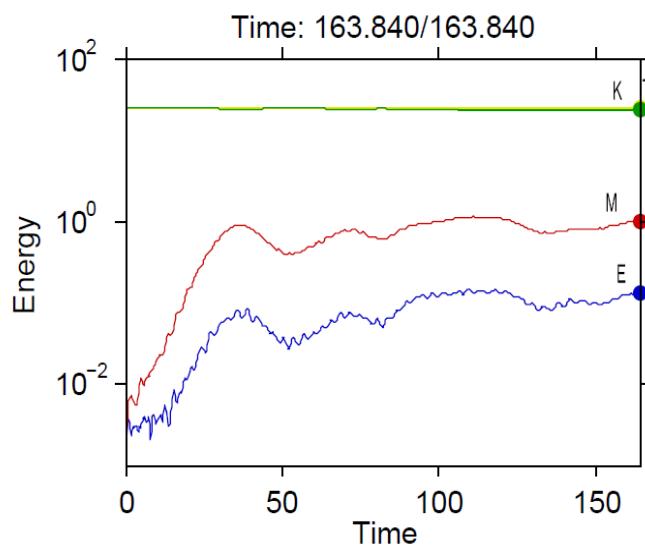
$\theta = 0 \text{ deg}$



[Omura and Matsumoto, JGR, 1987]

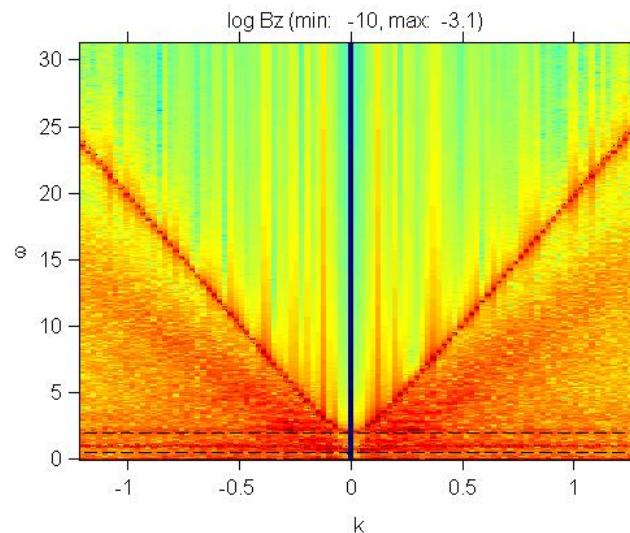
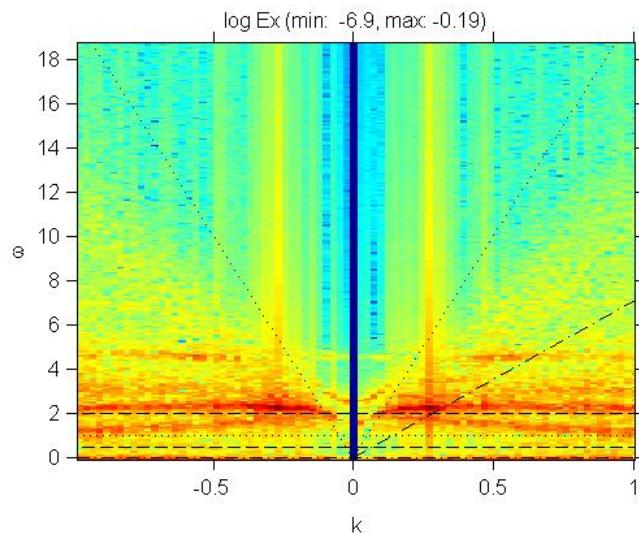
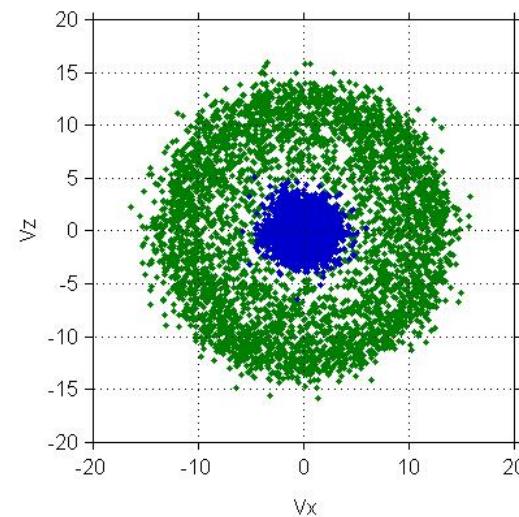
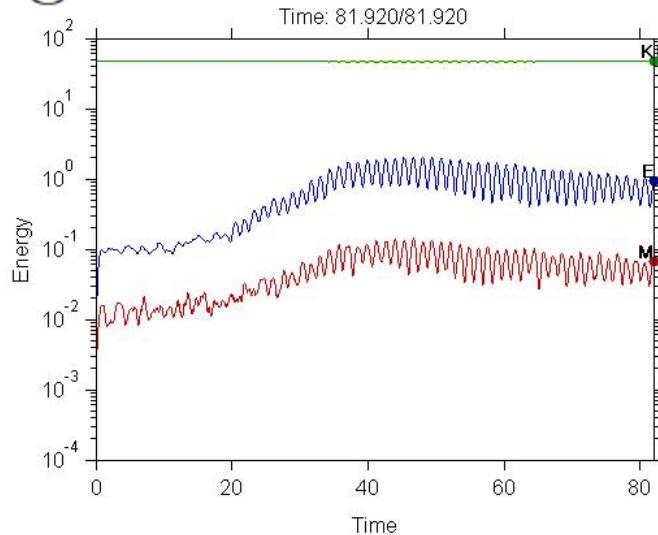
# Electron Beam Whistler Instability

$E_x = 0$

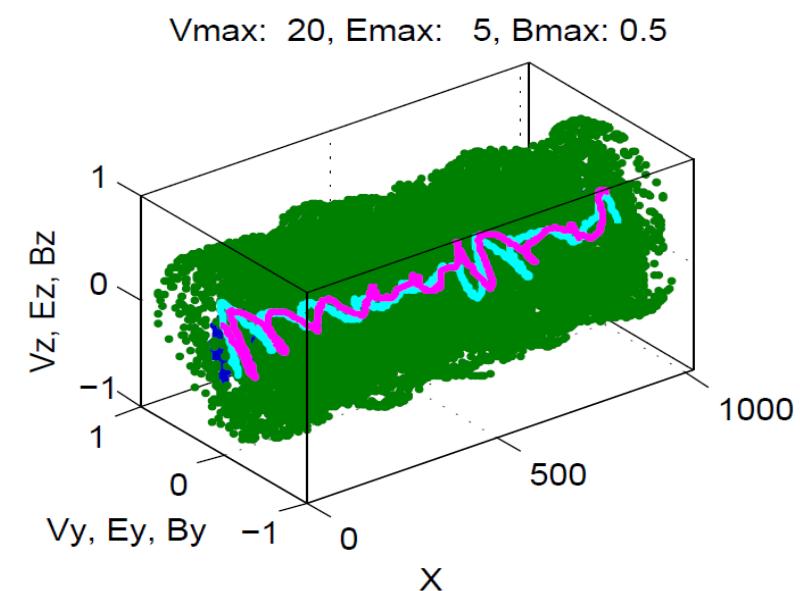
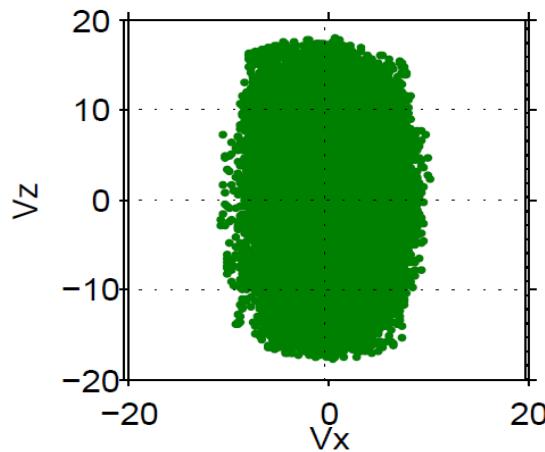
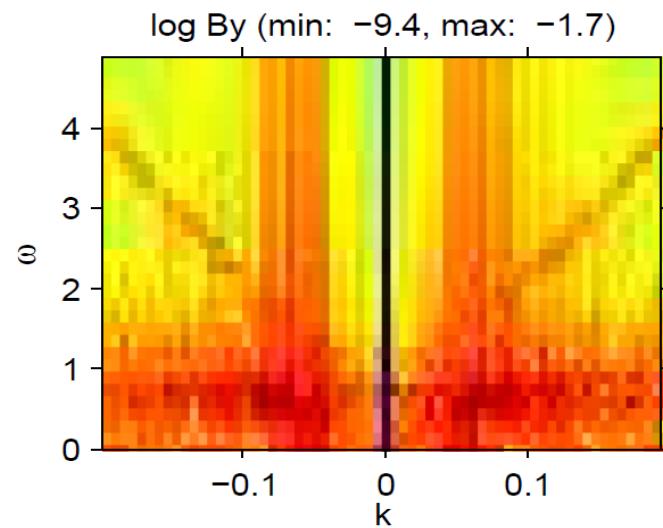
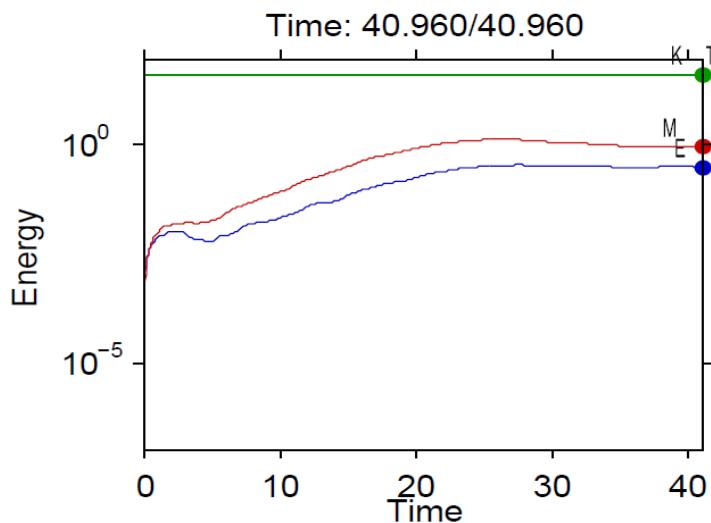


# Electrostatic Cyclotron Wave Instability

$\theta = 90 \text{ deg}$



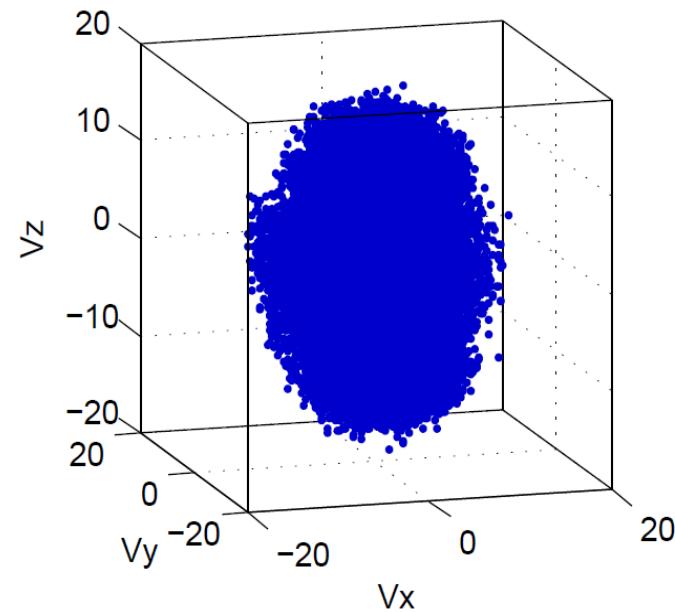
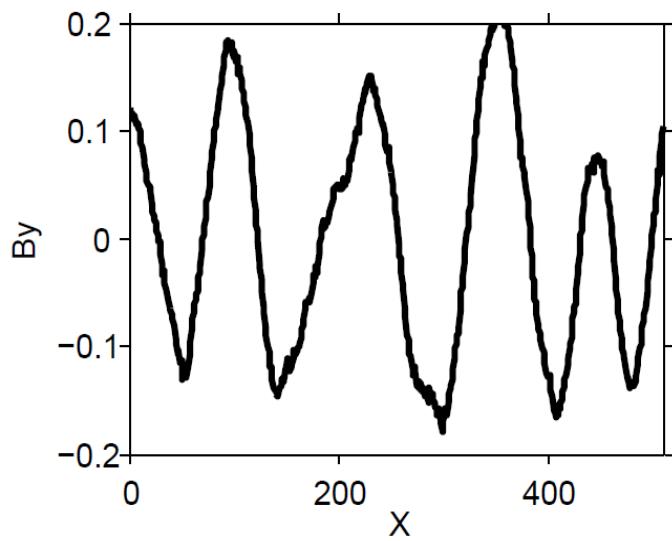
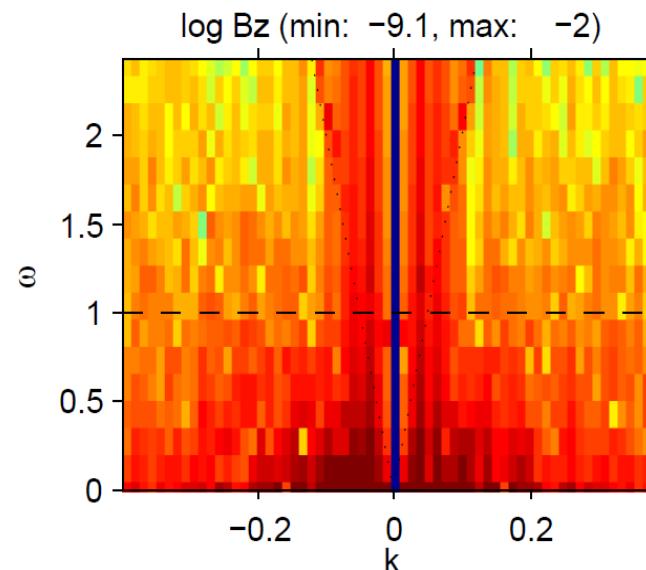
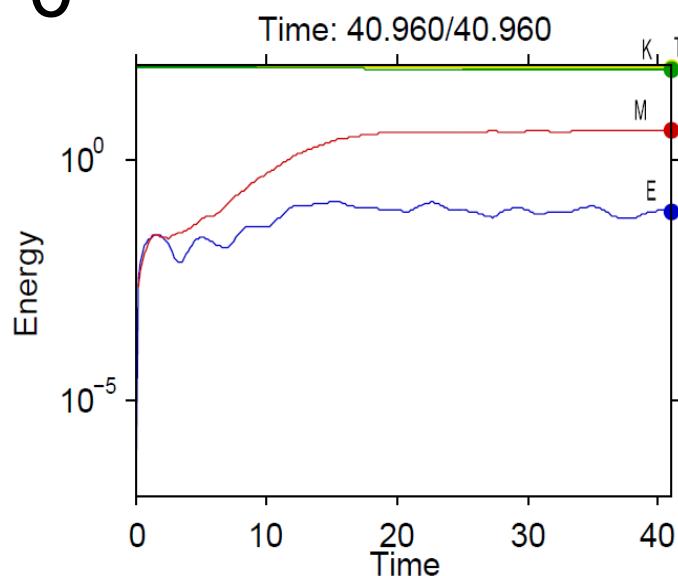
# Relativistic Ring Beam Instability (Cyclotron Maser Instability)



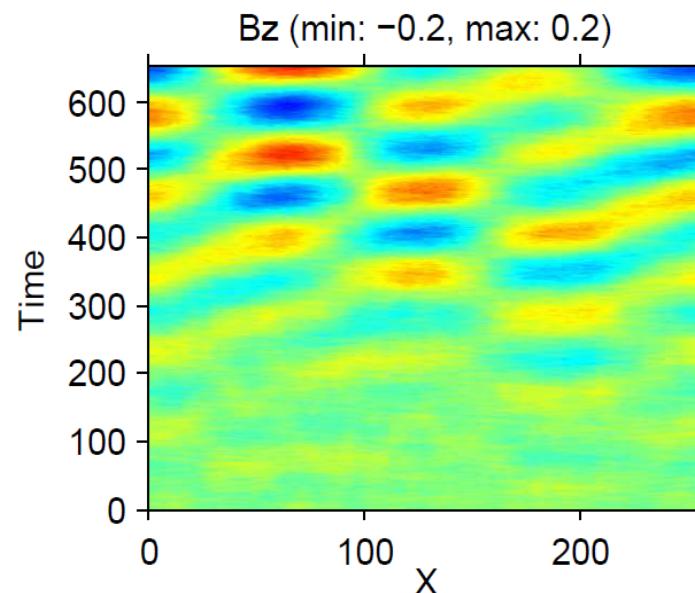
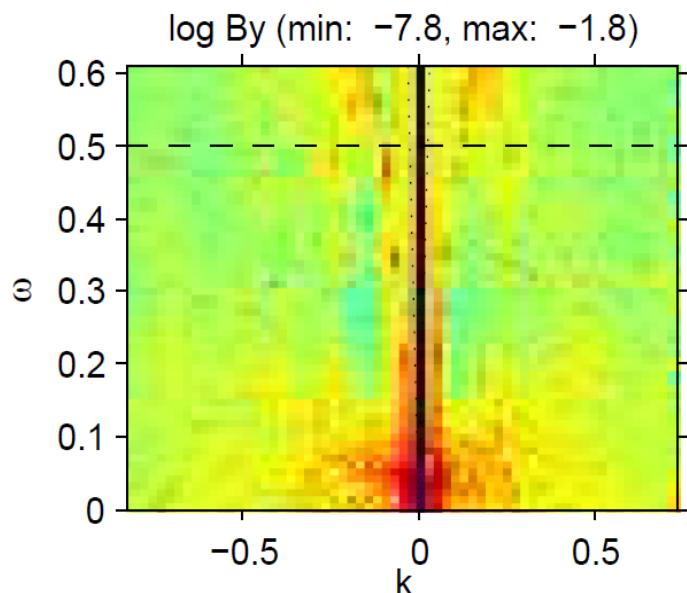
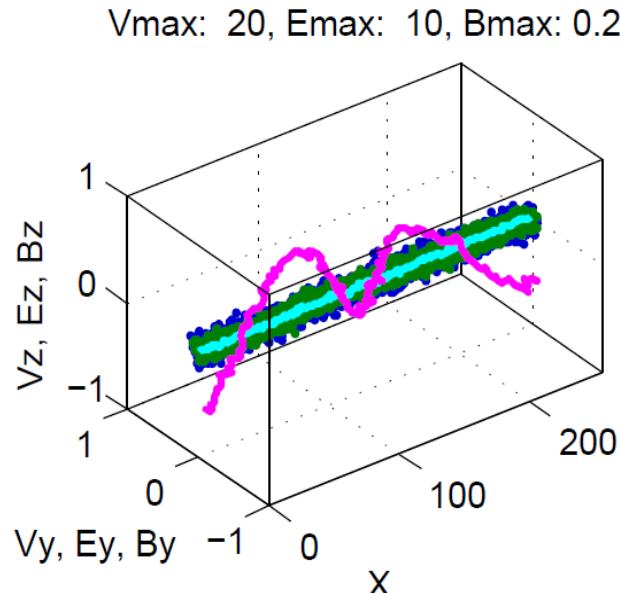
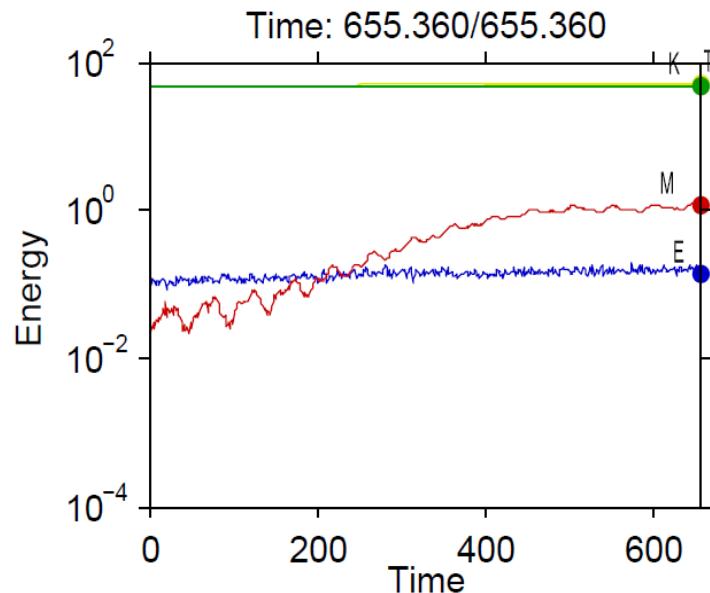
[cf. Lee et al., Phys. Plasmas, 2012]

# Weibel Instability

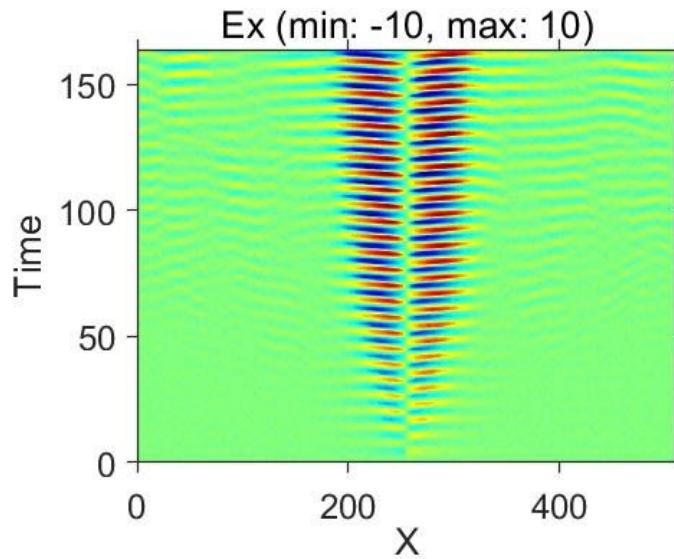
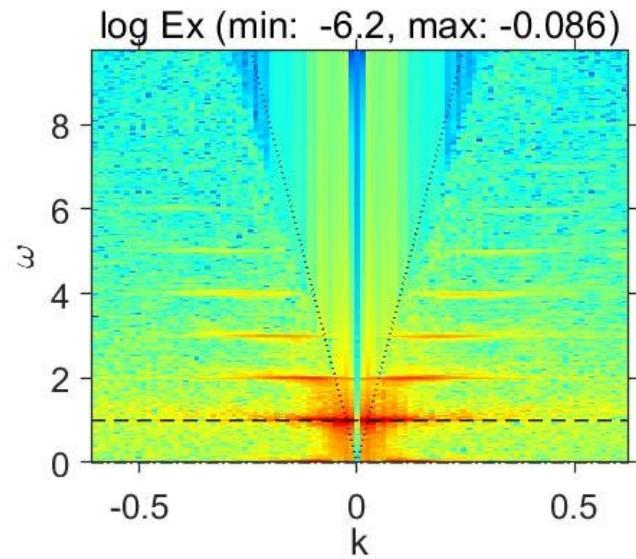
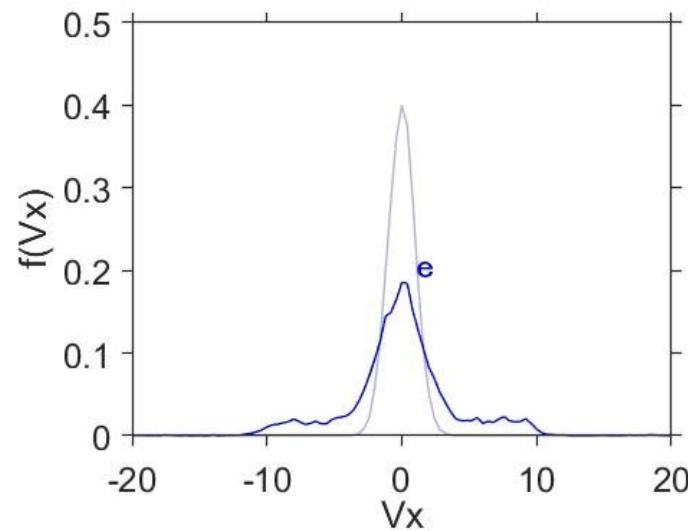
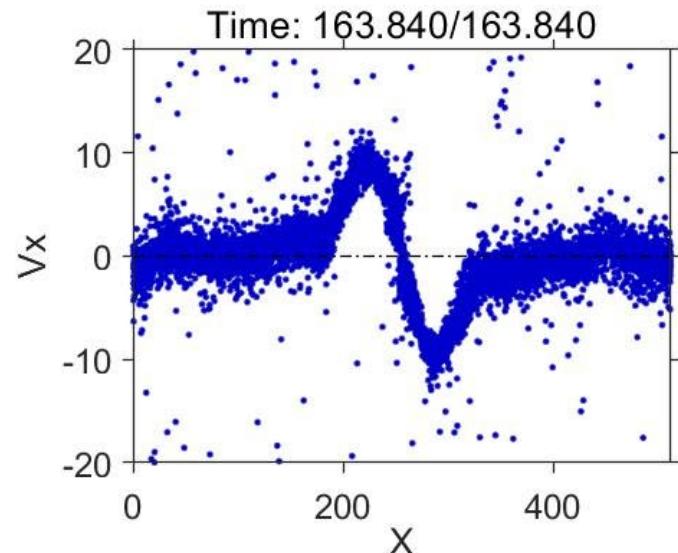
$B_0 = 0$



# EMIC Wave Instability



# Harmonic Langmuir Waves



[cf. Yoon et al., Phys. Plasmas, 2003]

# KEMPO2

## References

H. Matsumoto and Y. Omura, Particle Simulations of Electromagnetic Waves and its Applications to Space Plasmas, Computer Simulations of Space Plasmas, edited by H. Matsumoto and T. Sato, Terra Pub. and Reidel Co., pages 43-102, 1984.

T. Umeda, Y. Omura, T. Tominaga, and H. Matsumoto, A new charge conservation method in electromagnetic particle-in-cell simulations, Computer Physics Communications, Vol.156, No.1, pp.73-85, 2003.

# Two-Dimensional Simulation System (x, y)

$$\frac{\partial}{\partial z} = 0 \quad \nabla = e_x \frac{\partial}{\partial x} + e_y \frac{\partial}{\partial y}$$

$$\mathbf{E} = \mathbf{E}_{xy} + \mathbf{E}_z$$

$$\mathbf{B} = \mathbf{B}_{xy} + \mathbf{B}_z$$

$$\mathbf{J} = \mathbf{J}_{xy} + \mathbf{J}_z$$

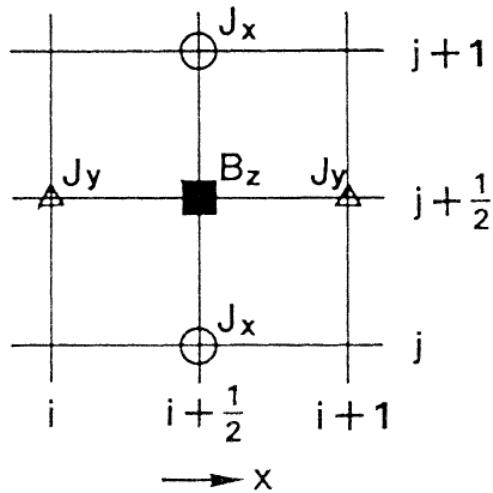
## Courant Condition

$$\Omega = \frac{\sin(\omega \Delta t / 2)}{\Delta t / 2} \quad K_x = \frac{\sin(k_x \Delta x / 2)}{\Delta x / 2} \quad K_y = \frac{\sin(k_y \Delta y / 2)}{\Delta y / 2}$$

$$\Omega^2 = c^2(K_x^2 + K_y^2) \quad \sqrt{2}c\Delta t < \Delta x$$

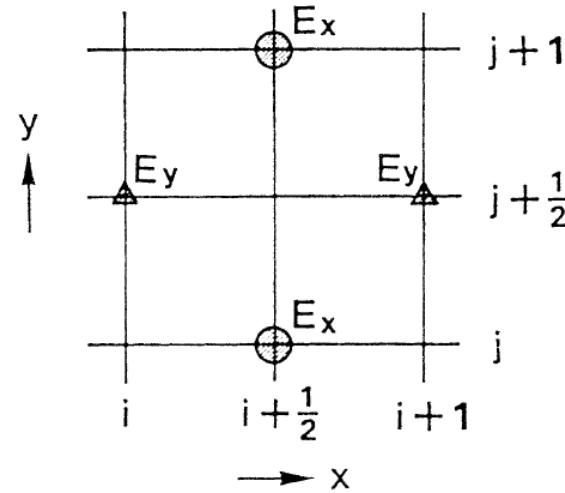
$$\frac{\partial \mathbf{B}_z}{\partial t} = -\nabla \times \mathbf{E}_{xy}$$

$t=(n-1/2)\Delta t$



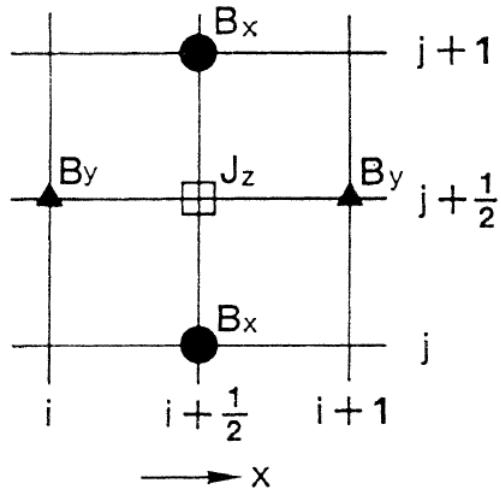
$$\frac{\partial \mathbf{E}_{xy}}{\partial t} = c^2 \nabla \times \mathbf{B}_z - \mathbf{J}_{xy}$$

$t=n\Delta t$



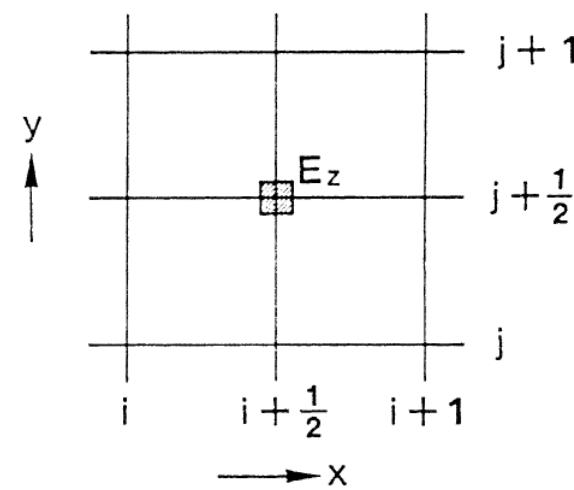
$$\frac{\partial \mathbf{B}_{xy}}{\partial t} = -\nabla \times \mathbf{E}_z$$

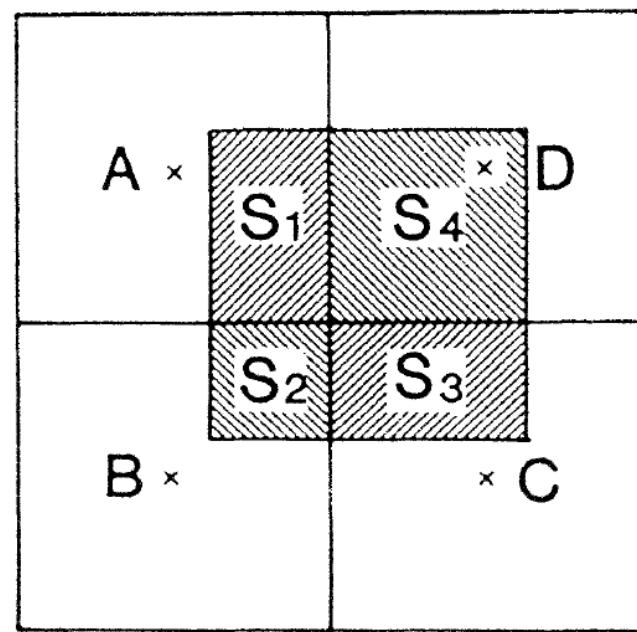
$t=(n-1/2)\Delta t$



$$\frac{\partial \mathbf{E}_z}{\partial t} = c^2 \nabla \times \mathbf{B}_{xy} - \mathbf{J}_z$$

$t=n\Delta t$





charge      current

- A:  $\frac{S_1}{S} q_{s,i}^{(s)}, \frac{S_1}{S} q_{s,i}^{(s)} \forall s,i$
- B:  $\frac{S_2}{S} q_{s,i}^{(s)}, \frac{S_2}{S} q_{s,i}^{(s)} \forall s,i$
- C:  $\frac{S_3}{S} q_{s,i}^{(s)}, \frac{S_3}{S} q_{s,i}^{(s)} \forall s,i$
- D:  $\frac{S_4}{S} q_{s,i}^{(s)}, \frac{S_4}{S} q_{s,i}^{(s)} \forall s,i$

Where

$$S = S_1 + S_2 + S_3 + S_4$$

# Relocations

Source Relocation

$$Jx_{i+1/2,j} = \frac{1}{2}(Jx_{i,j} + Jx_{i+1,j})$$

$$Jy_{i,j+1/2} = \frac{1}{2}(Jy_{i,j} + Jy_{i,j+1})$$

$$Jz_{i+1/2,j+1/2} = \frac{1}{4}(Jz_{i,j} + Jz_{i+1,j} + Jz_{i,j+1} + Jz_{i+1,j+1})$$

Field Relocation (E & B)

$$PEx_{i,j} = \frac{1}{2}(Ex_{i-1/2,j} + Ex_{i+1/2,j})$$

$$PEy_{i,j} = \frac{1}{2}(Ey_{i,j-1/2} + Ey_{i,j+1/2})$$

$$PEz_{i,j} = \frac{1}{4}(Ez_{i-1/2,j-1/2} + Ez_{i+1/2,j-1/2} + Ez_{i-1/2,j+1/2} + Ez_{i+1/2,j+1/2})$$

# Solution of Electrostatic Field (Correction of Electric Field)

$$\frac{\partial \rho}{\partial t} + \boxed{\nabla \cdot \mathbf{J}} = 0$$

$$\nabla \cdot (\mathbf{E} + \mathbf{E}_c) = \rho \quad \rho_c = \rho - \nabla \cdot \mathbf{E}$$

$$\mathbf{E}_c = -\nabla \psi \quad \nabla \cdot \mathbf{E}_c = \rho_c$$

$$\frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} = -\rho_c$$

$$\frac{\psi_{i+1,j} - 2\psi_{i,j} + \psi_{i-1,j}}{(\Delta x)^2} + \frac{\psi_{i,j+1} - 2\psi_{i,j} + \psi_{i,j-1}}{(\Delta y)^2} = -\rho_{c,i,j}$$

$$\tilde{\rho}_{kx,ky} = \text{FFT} \{ \rho_c(x, y) \} \quad K_x = \frac{\sin(k_x \Delta x / 2)}{\Delta x / 2} \quad K_y = \frac{\sin(k_y \Delta y / 2)}{\Delta y / 2}$$

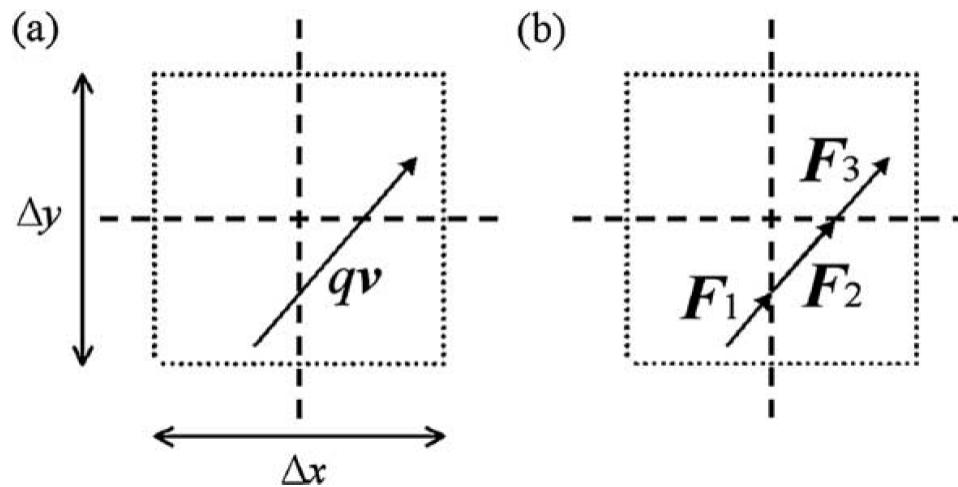
$$(K_x^2 + K_y^2) \tilde{\psi}_{kx,ky} = \tilde{\rho}_{kx,ky} \quad \tilde{\psi}_{kx,ky} = \frac{\tilde{\rho}_{kx,ky}}{(K_x^2 + K_y^2)}$$

$$\psi(x, y) = \text{IFFT} \left\{ \tilde{\psi}_{kx,ky} \right\}$$

$$\mathbf{E} = \mathbf{E} - \nabla \psi$$

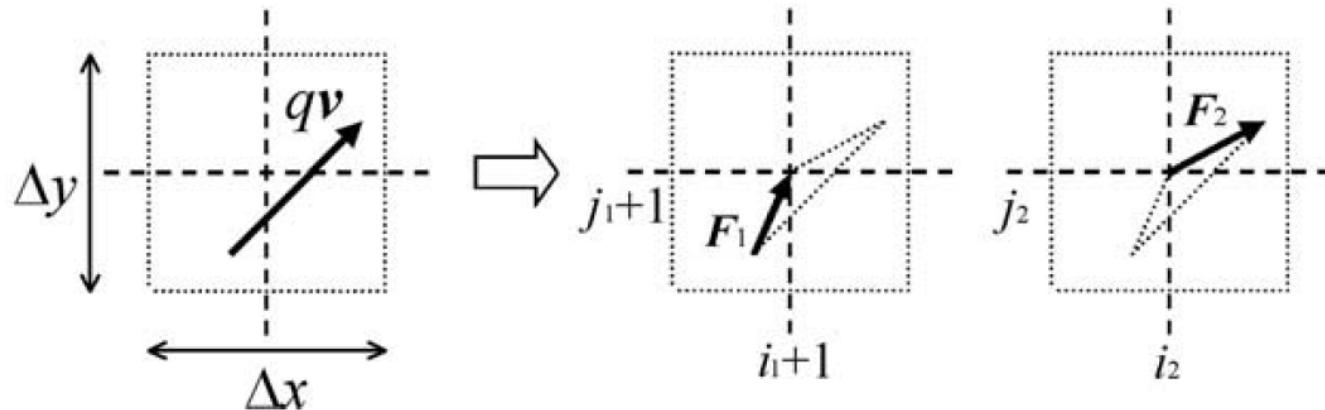
$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0$$

$$\begin{aligned} & \frac{J_x^{t+\Delta t/2}(i + \frac{1}{2}, j) - J_x^{t+\Delta t/2}(i - \frac{1}{2}, j)}{\Delta x} + \frac{J_y^{t+\frac{\Delta t}{2}}(i, j + \frac{1}{2}) - J_y^{t+\Delta t/2}(i, j - \frac{1}{2})}{\Delta y} \\ &= \frac{\rho^t(i, j) - \rho^{t+\Delta t}(i, j)}{\Delta t}. \end{aligned}$$

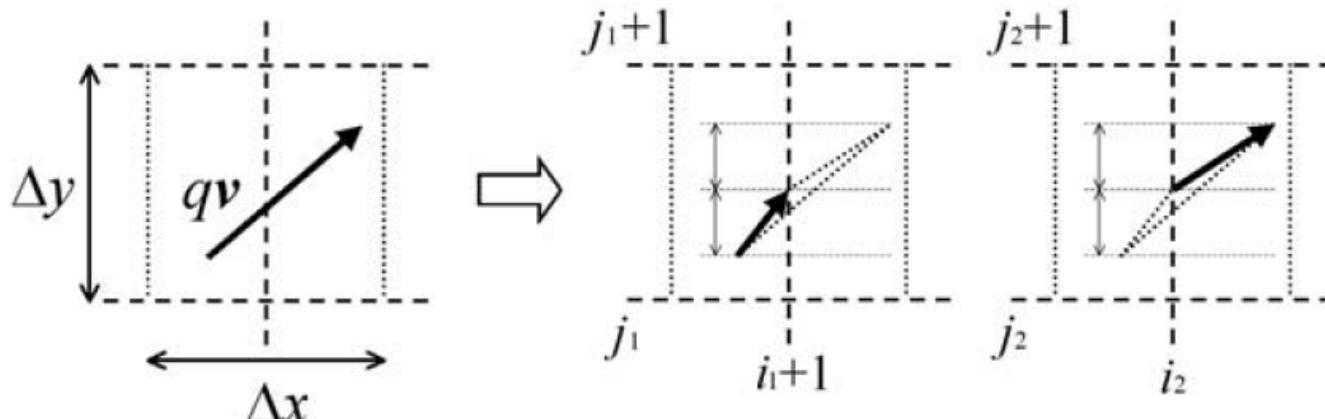


# Zigzag Method (current2)

(a)  $i_1 \neq i_2$  and  $j_1 \neq j_2$



(b)  $i_1 \neq i_2$  and  $j_1 = j_2$  or (c)  $i_1 = i_2$  and  $j_1 \neq j_2$



# KEMPO2: Main Program

## Method 1: Electrostatic Field Correction Method

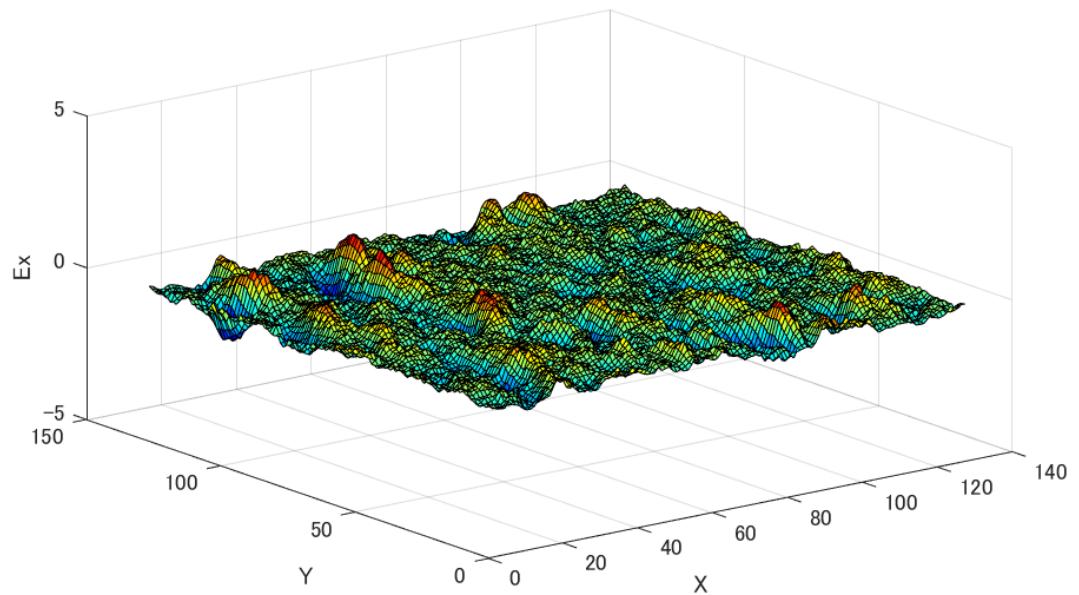
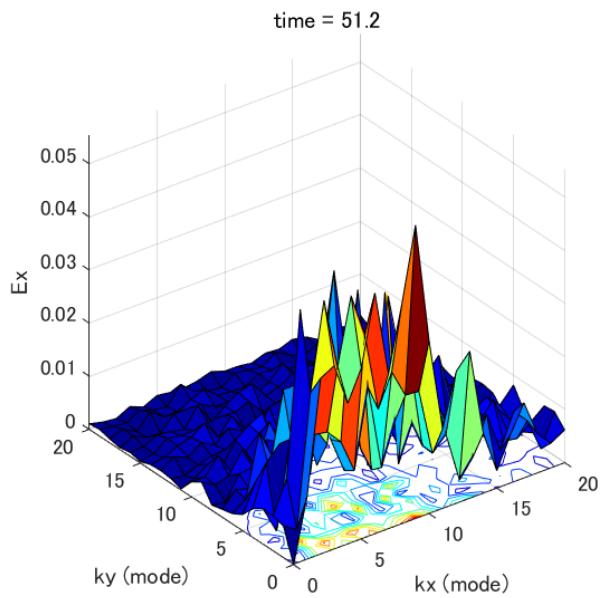
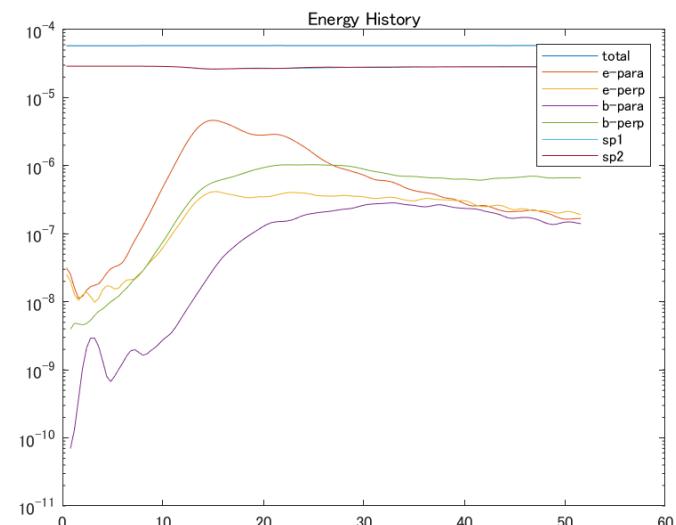
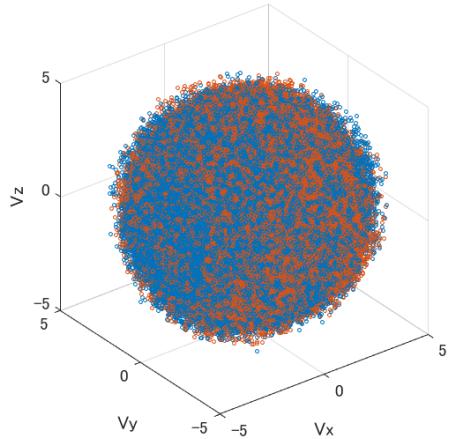
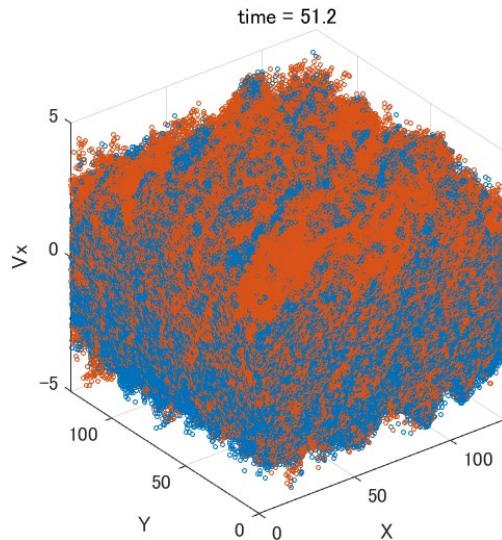
```
charge;  
potential;  
for itime = 1:ntime  
    jtime = jtime +1;  
    bfield;  
    rvelocity;  
    position;  
    current;  
    position;  
    bfield;  
    efield;  
    charge;  
    potential;  
    diagnostics;  
end
```

## Method 2: Charge Conservation (zigzag)

```
charge;  
potential;  
for itime = 1:ntime  
    jtime = jtime +1;  
    bfield;  
    rvelocity;  
    position;  
    current2;  
    position;  
    bfield;  
    efield;  
    diagnostics;  
end
```

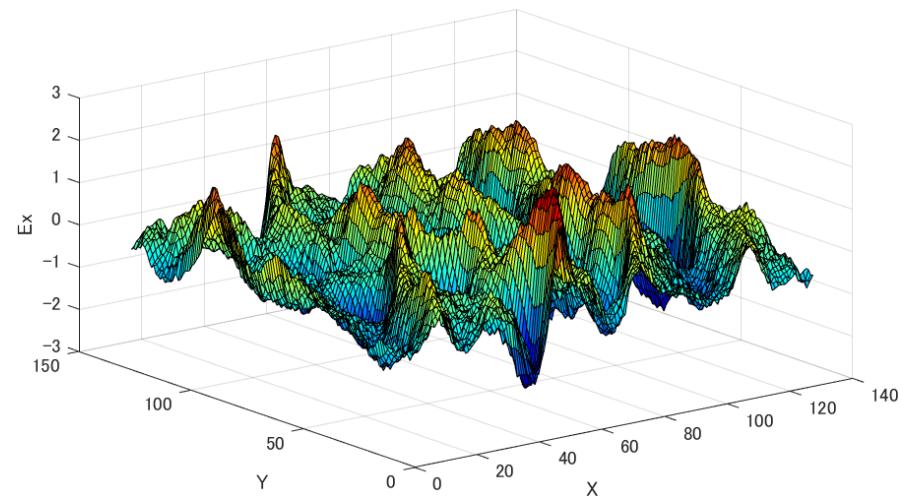
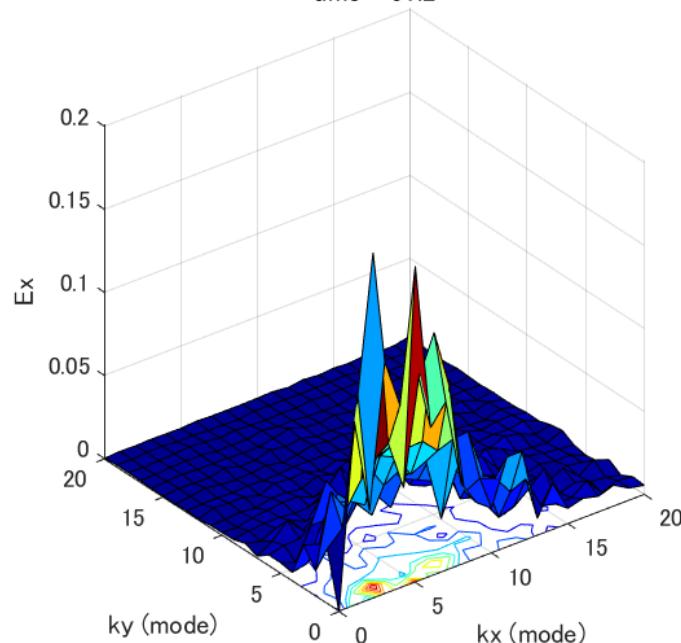
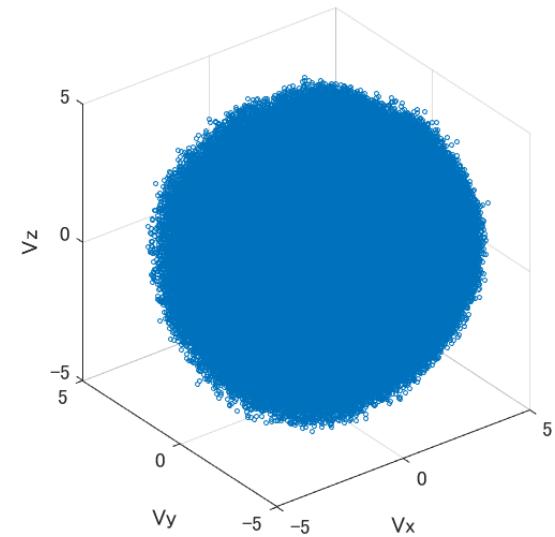
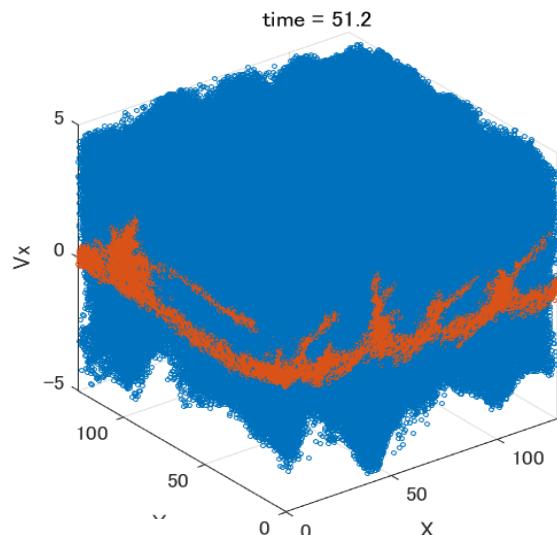
% Computation times for Methods 1 and 2 are nearly the same.

# Electron Two-stream Instability

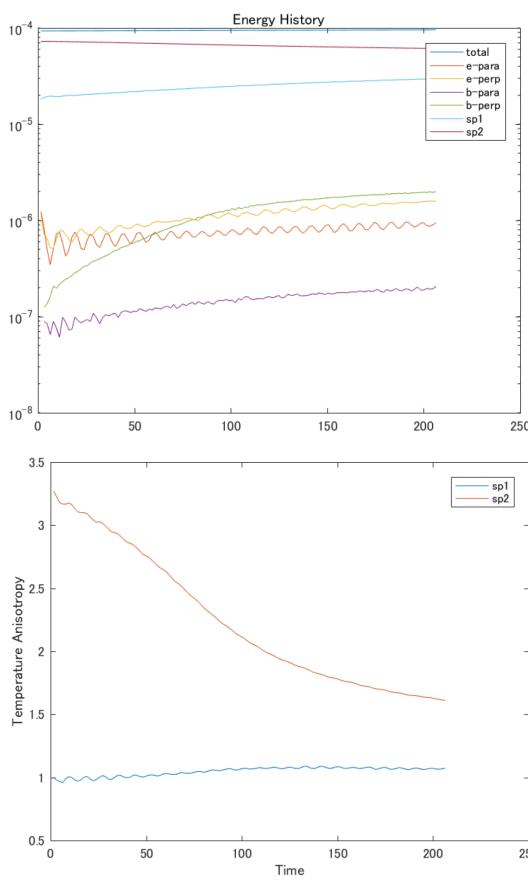
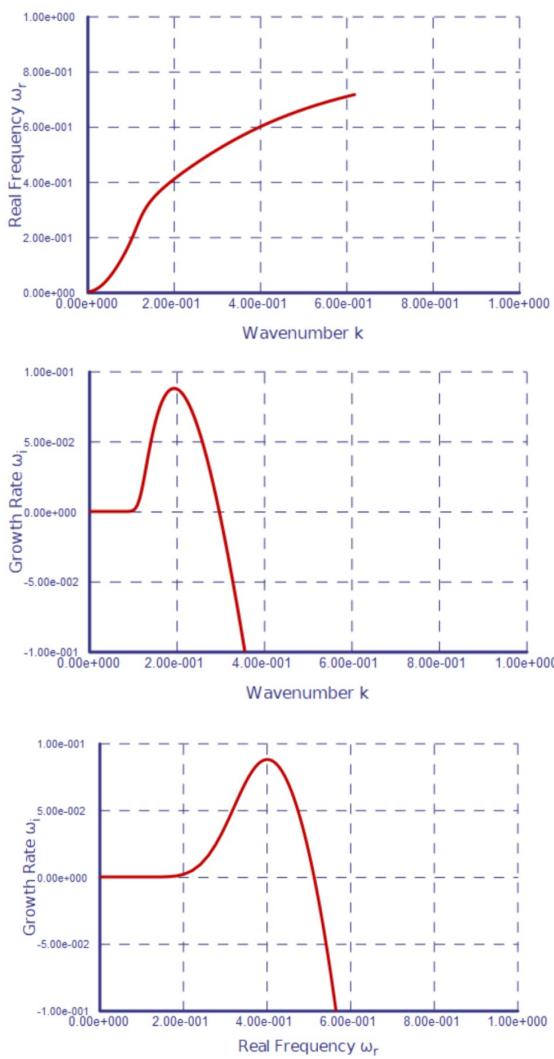


# Buneman Instability

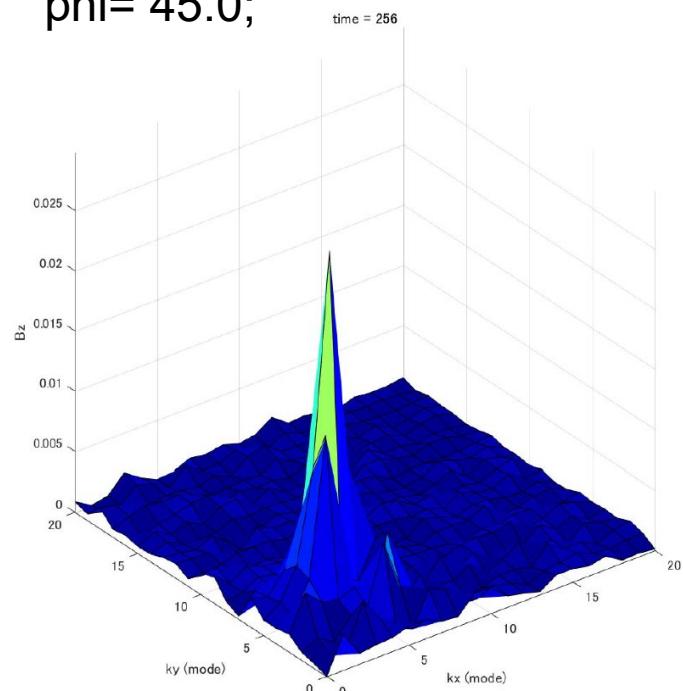
$m_i / m_e = 25$



# Whistler Mode Instability Driven by Temperature Anisotropy



% direction of the static magnetic field  
 $\theta = 90.0$ ;  
 $\phi = 45.0$ ;

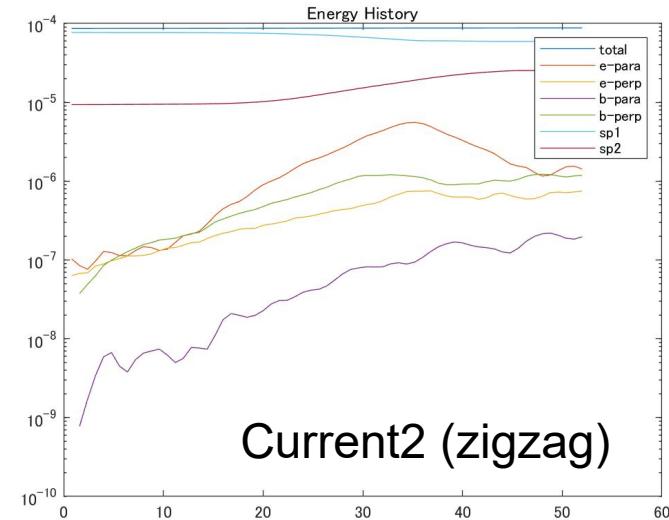
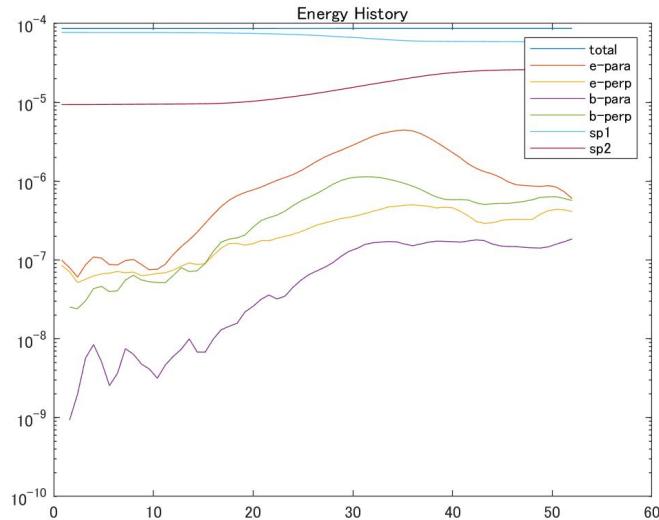


```

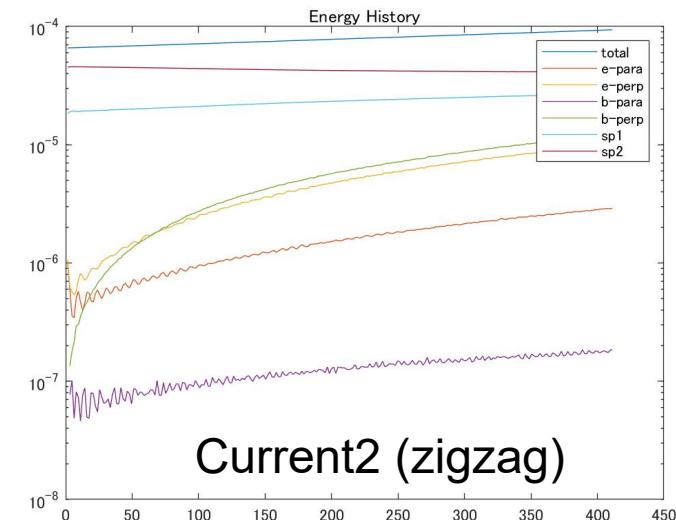
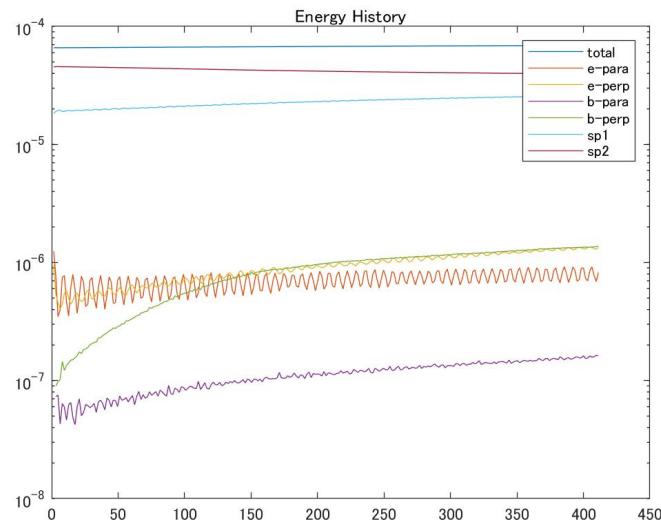
vs = sum(vx(m).^2 + vy(m).^2 + vz(m).^2);
vspara = sum(((vx(m)*bx0+vy(m)*by0+vz(m)*bz0)/b0).^2);
%Temperature Anisotropy T_perp/T_para
At(it,k) = 0.5*(vs/vspara - 1);

```

# Buneman Instability



# Whistler Instability



# Comment on Zigzag Method

The Zigzag method for computation of the current density is slightly faster than the electrostatic correction method with charge density computation. It can enhance the electromagnetic fluctuations because of the non-physical trajectories of particles assumed to simplify the computation of the current.