

SUPPLEMENTARY MATERIAL

Turbulence in compact to giant H II regions

J. García Vázquez^{1★}, William J. Henney² and H. O. Castañeda^{1†}

¹*Escuela Superior de Física y Matemáticas, Instituto Politécnico Nacional, U.P. Adolfo López Mateos, Zacatenco, Ciudad de México, México C.P. 07738*

²*Instituto de Radioastronomía y Astrofísica, Universidad Nacional Autónoma de México, Antigua Carretera a Pátzcuaro #8701, Ex-Hda. San José de la Huerta, Morelia, Michoacán, México C.P. 58089*

Accepted XXX. Received YYY; in original form ZZZ

APPENDIX D: ADDITIONAL COVARIANCE CORNER PLOTS FOR MODEL FITS

Supplementary online-only material.

This paper has been typeset from a T_EX/L^AT_EX file prepared by the author.

Table D1. Fitting parameters.

H II Region	No. points merged	Relatively uncertainty	Weight / No. points Small scales	Weight / No. points Large scales	Max. separation	Reduced chi-square	Autocorrelation time	Acceptance fraction
Orion	3	0.02	–	–	0.5L	0.82	60	0.55
EON	–	0.03	2.0/3	3.0/3	0.7L	0.62	40	0.64
Lagoon	2	0.08	2.5/16	2.5/8	0.5L	0.94	72	0.50
Carina	2	0.055	3.0/14	3.0/4	0.5L	0.94	72	0.50
30 Dor	5	0.07	–	–	0.9L	0.98	84	0.54
NGC 346	3	0.02	–	–	0.5L	0.83	64	0.54
Hubble V	–	0.05	2.0/3	–	0.6L	0.86	340	0.44
Hubble X	–	0.032	2.0/4	2.0/6	0.5L	0.73	79	0.49
NGC 595	–	0.055	4.5/3	3.0/5	0.5L	0.83	67	0.51
NGC 604	–	0.08	2.0/3	1.5/2	0.5L	0.92	112	0.45

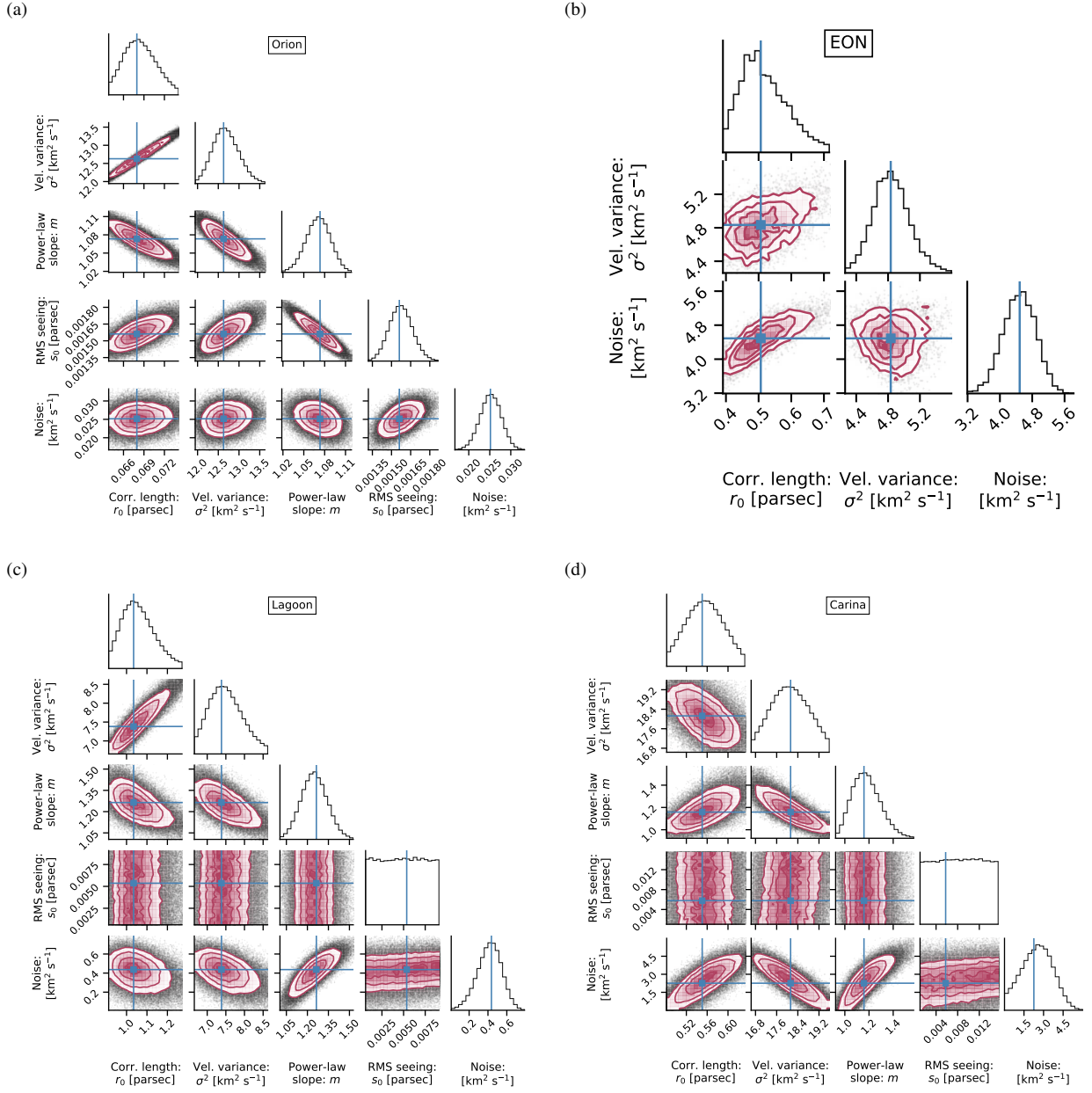


Figure D1. Corner plots of covariances between fitted model parameters of H α structure function for Galactic H II regions.

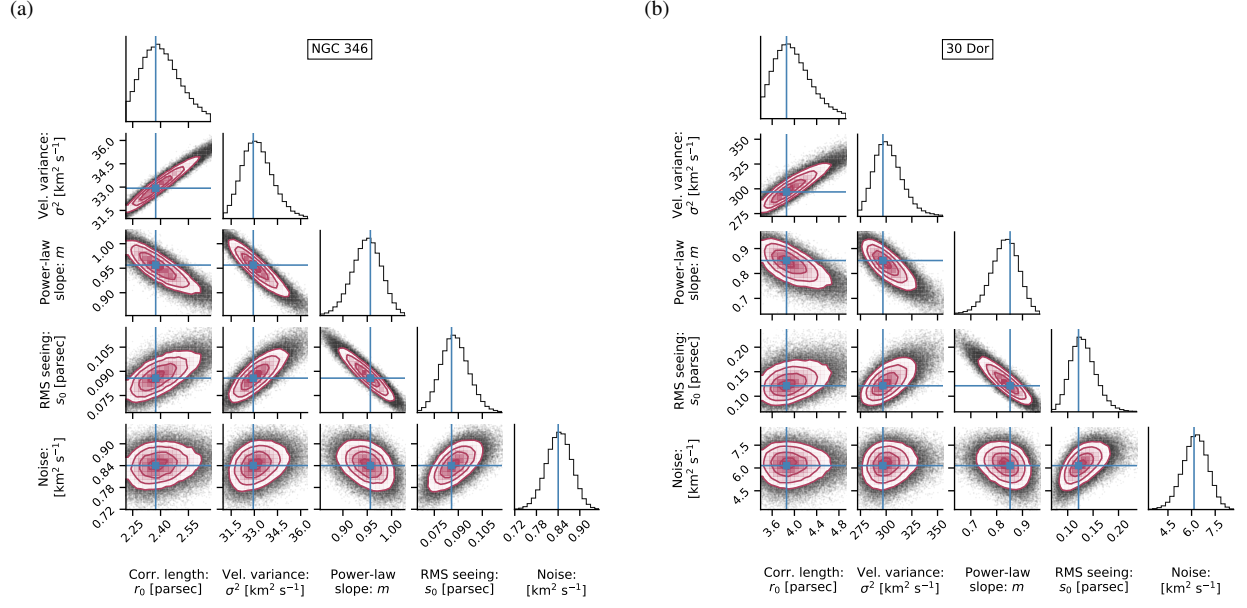


Figure D2. Same as figure D1 except for Magellanic Cloud H II regions.

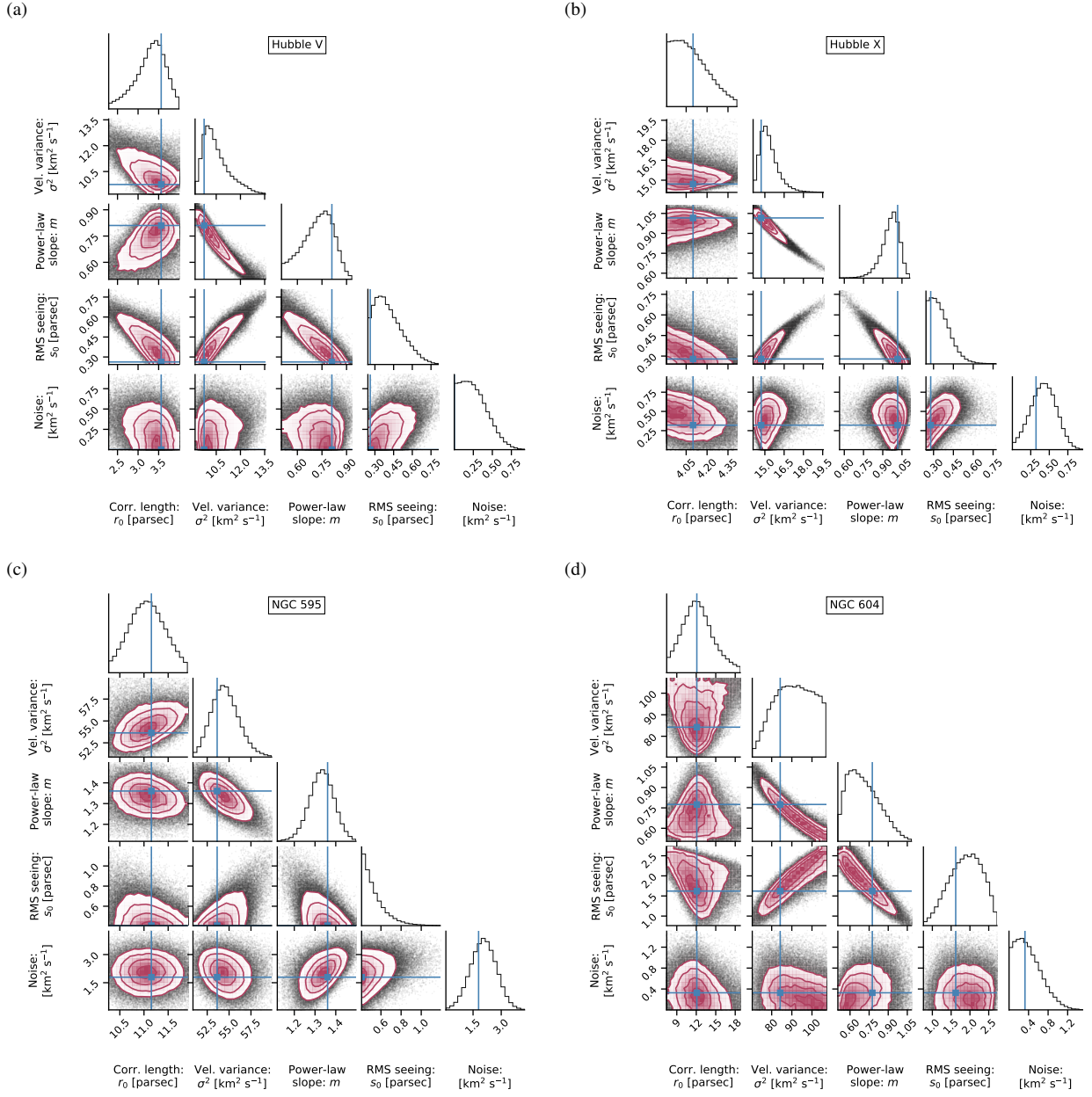


Figure D3. Same as figure D1 except for H II regions in more distant Local Group Galaxies.