

Electrodynamics Notes

Phillips Academy Andover

PRO601 Physics IP: Advanced Electrodynamics with Jeffrey Shi
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Introduction

This document contains my notes for my Spring Term Physics IP (Independent Project) entitled *Advanced Electrodynamics*, taken during my last term at Phillips Academy Andover. Jeffrey Shi was my partner for the group IP and our mentor was Ms. Fei Yao. Our main reference text is *Introduction to Electrodynamics*, 4th edition by Griffiths.

I have also included some of our solutions to selected problems from the text. The numbering of these problems follows the book, rather than this document. We tried to tackle every problem marked with a bullet • or an exclamation point !. We also tackled any other problems that we found interesting.

1 Vector Analysis

§1.1 Vector Algebra

§1.1.1 Vector Operations

§1.1.2 Vector Algebra: Component Form

These first two subsections reviews the basics of vectors; for the purpose of these notes, I will assume the reader already has a strong basic understanding of vectors, so they will be left empty.

§1.1.3 Triple Products

The output of a cross product $\vec{b} \times \vec{c}$ is a vector, so it can be dotted or crossed with another vector \vec{a} to form a triple product.

The **scalar triple product** $\vec{a} \cdot (\vec{b} \times \vec{c})$ gives the signed volume of the parallelepiped generated by $\vec{a}, \vec{b}, \vec{c}$. Therefore,

$$\text{Claim 1.1.1} \text{ — } \vec{a} \cdot (\vec{b} \times \vec{c}) = \vec{b} \cdot (\vec{c} \times \vec{a}) = \vec{c} \cdot (\vec{a} \times \vec{b}).$$

Notice that we preserved the cyclic order of the vectors in this case; if we flip the order (e.g. $\vec{a} \cdot (\vec{c} \times \vec{b})$), then we obtain the same value with opposite sign. In component form, we have

$$\vec{a} \cdot (\vec{b} \times \vec{c}) = \det \begin{bmatrix} a_x & a_y & a_z \\ b_x & b_y & b_z \\ c_x & c_y & c_z \end{bmatrix},$$

The **vector triple product** $\vec{a} \times (\vec{b} \times \vec{c})$ can be simplified by the so-called “*bac – cab* rule”:

$$\text{Claim 1.1.2} \text{ — } \vec{a} \times (\vec{b} \times \vec{c}) = \vec{b}(\vec{a} \cdot \vec{c}) - \vec{c}(\vec{a} \cdot \vec{b}).$$

Note that cross-products are not associative, so the order of parentheses matters. By repeated application of this fact, it is never necessary for an expression to contain more than one cross product in any term. For example,

$$(\vec{a} \times \vec{b}) \cdot (\vec{c} \times \vec{d}) = (\vec{b} \times (\vec{c} \times \vec{d})) \cdot \vec{a} = (\vec{c}(\vec{b} \cdot \vec{d}) - \vec{d}(\vec{b} \cdot \vec{c})) \cdot \vec{a} = (\vec{a} \cdot \vec{c})(\vec{b} \cdot \vec{d}) - (\vec{a} \cdot \vec{d})(\vec{b} \cdot \vec{c}),$$

where in the first equality we utilized the scalar triple product identity $\vec{a} \cdot (\vec{b} \times \vec{c}) = \vec{b} \cdot (\vec{c} \times \vec{a})$.

§1.1.4 Position, Displacement, and Separation Vectors

Definition 1.1.3. Given a coordinate system with origin \mathcal{O} , the vector from \mathcal{O} to some point is the **position vector**

$$\vec{r} := x\hat{x} + y\hat{y} + z\hat{z},$$

where (x, y, z) are the Cartesian coordinates of the point.

Note that the distance from the point to the origin is

$$r = |\vec{r}| = \sqrt{x^2 + y^2 + z^2}$$

and the unit vector in the direction of \vec{r} is

$$\hat{r} = \frac{\vec{r}}{r} = \frac{x\hat{x} + y\hat{y} + z\hat{z}}{\sqrt{x^2 + y^2 + z^2}}.$$

Definition 1.1.4. The **infinitesimal displacement vector** from (x, y, z) to $(x + dx, y + dy, z + dz)$ is denoted by

$$d\vec{\ell} = dx\hat{x} + dy\hat{y} + dz\hat{z}.$$

Oftentimes in electrodynamics, we encounter problems involving two points:

- A **source point** \vec{r}' , where an electric charge is located,
- A **field point** \vec{r} , where we are calculating the electric or magnetic field.

In the situation when $\vec{r}' \neq \vec{0}$, it is useful to adopt a shorthand notation for the **separation vector** from the source point to the field point; we will use the script letter

$$\vec{r} := \vec{r} - \vec{r}'.$$

§1.1.5 How Vectors Transform

What exactly is a vector? You might have learned it as “a quantity with a magnitude and direction,” but this isn’t exactly satisfactory...

Perhaps a 3-dimensional vector is simply anything with three real components that combine properly under addition. However, what if we considered some silly example like a barrel with N_x pear, N_y apples, and N_z bananas? Is $\vec{N} = N_x\hat{x} + N_y\hat{y} + N_z\hat{z}$ suddenly a vector? Surely not—what is the direction? What is wrong with it?

The answer is that \vec{N} *doesn’t transform properly when you change coordinates*.

In three dimensions, if a x, y, z system is rotated to some $\bar{x}, \bar{y}, \bar{z}$ system, then the coordinates of some vector \vec{a} undergo a matrix transformation

$$\begin{bmatrix} \bar{a}_x \\ \bar{a}_y \\ \bar{a}_z \end{bmatrix} = \begin{bmatrix} R_{xx} & R_{xy} & R_{xz} \\ R_{yx} & R_{yy} & R_{yz} \\ R_{zx} & R_{zy} & R_{zz} \end{bmatrix} \cdot \begin{bmatrix} a_x \\ a_y \\ a_z \end{bmatrix} \quad \text{or} \quad \vec{\bar{a}} = R\vec{a},$$

where R_{ij} are some products of trigonometric functions (or it could simply be any arbitrary linear transformation of \mathbb{R}^3). More compactly, we can write it as

$$\bar{a}_i = \sum_{j=1}^3 R_{ij} a_j,$$

where index 1 stands for x , etc.

The components of \vec{N} clearly don’t transform in this way.

Definition 1.1.5. Formally, a **vector** is any set of three components that transforms in the same manner as a displacement when you change coordinates.

Therefore, displacement is the *model* for the behavior of all vectors.

Definition 1.1.6. A **second-order tensor** is a quantity with *nine* components: $T_{xx}, T_{xy}, T_{xz}, T_{yx}, \dots, T_{zz}$. These transform with *two* factors of R :

$$\bar{T}_{ij} = \sum_{k=1}^3 \sum_{\ell=1}^3 R_{ik} R_{j\ell} T_{k\ell}.$$

You can also rewrite this in matrix form as $\bar{T} = RTR^T$.

In general, an n th-order tensor has n indices and 3^n components (in \mathbb{R}^3), and transforms with n factors of R . Therefore, a scalar is a tensor of order zero (indeed, note that scalars never change under coordinate changes) and a vector is a tensor of order one.

§1.2 Differential Calculus

§1.2.1 “Ordinary” Derivatives

Suppose we have a single-variable function $f(x)$. What does the derivative $\frac{df}{dx}$ do?

The derivative of a function df/dx tells us how rapidly the function f varies when we change x by a tiny amount dx .

In other words, if x is incremented by an infinitesimal amount dx and f changes by an amount df , then the derivative is the proportionality factor:

$$df = \left(\frac{df}{dx} \right) dx.$$

From a geometric perspective, the derivative $\frac{df}{dx}$ gives the *slope* at any point on the graph of f against x .

§1.2.2 Gradient

That’s all well and good for one variable. What about functions of three variables, say temperature $T(x, y, z)$ in \mathbb{R}^3 ?

How do we generalize the notion of the derivative to functions like T ?

We want the derivative to tell us how T changes when we move a tiny distance. However, T may change at a different rate depending on the direction we move in! In a room, nudging z will probably have a much higher affect on the temperature than nudging x or y the same amount.

Fortunately, the situation is not so bad: it turns out that when we zoom in enough, a differentiable function T will always look like a plane, linear in all variables:

$$dT = \left(\frac{\partial T}{\partial x} \right) dx + \left(\frac{\partial T}{\partial y} \right) dy + \left(\frac{\partial T}{\partial z} \right) dz.$$

The proportionality constants are *partial derivatives*, which measure the change in T when you hold all variables constant except for the one in the denominator: therefore, $\frac{\partial T}{\partial x}$ measures the rate of change of T when you move only x by an infinitesimal amount.

This is a rather nice result: we don’t need to compute the rate of change of T among all possible directions; three directions x, y, z suffice.

We can also use it to define another concept. The equation is reminiscent of a dot product:

$$dT = \left(\frac{\partial T}{\partial x} \hat{x} + \frac{\partial T}{\partial y} \hat{y} + \frac{\partial T}{\partial z} \hat{z} \right) \cdot (dx \hat{x} + dy \hat{y} + dz \hat{z}).$$

We give the first vector in this dot product a name:

Definition 1.2.1. The **gradient** of a scalar function T is the vector quantity

$$\nabla T := \frac{\partial T}{\partial x} \hat{x} + \frac{\partial T}{\partial y} \hat{y} + \frac{\partial T}{\partial z} \hat{z}.$$

Then, $dT = \nabla T \cdot d\vec{\ell}$. Wait... what just happened? This means that if we move by a small vector $d\vec{\ell}$, then the change in T is given by taking the dot product of this small shift vector by the gradient of T at that point.

Let's dissect this a little more: we can expand the dot product to get

$$dT = \nabla T \cdot d\vec{\ell} = |\nabla T| |d\vec{\ell}| \cos \theta,$$

where θ is the angle between this new gradient vector we just defined and the small shift $d\vec{\ell}$ that we take. Suppose we *fix* the magnitude of $d\vec{\ell}$, so we're always stepping the same tiny amount. Then, if we search various directions by varying θ , notice the *maximum* change in T occurs when $\theta = 0$.

Focusing on this case, we extract the following intuition on what exactly the gradient is:

The gradient ∇T points in the direction of maximum increase of the function T . Moreover, its magnitude $|\nabla T|$ gives the rate of increase of T along this maximal direction.

If we point $d\vec{\ell}$ in the opposite direction of the gradient ($\theta = 180^\circ$), then the rate of change is precisely the negative, which makes sense. Also note if we point perpendicular to the gradient ($\theta = 90^\circ$), T doesn't change along this direction at all.

Here's a nice picture to keep in your mind: suppose you're on a large, bumpy mountain. We can characterize the surface of this mountain by the *two*-variable function $h(x, y)$ which gives the height of the mountain above any point (x, y) in the plane. Standing at any point (x, y) , you can look around and find the direction of steepest ascent—that is the *direction* of the gradient. If you calculate the slope of the mountain in that direction (rise over run)—that is the *magnitude* of the gradient.

As a person standing on the mountain, if you zoom in enough, you can disregard any large scale bumps in the mountain: it simply looks like a plane; this is what allows the dot product formula $dT = \nabla T \cdot d\vec{\ell}$ to work (with this picture in your mind, reason through the $\theta = 180^\circ$ and $\theta = 90^\circ$ cases again. This is analogous to the one-variable case, where taking the derivative assumes that when you zoom far in enough on f , you get something that looks like a straight line.

Definition 1.2.2. We call points where $\nabla T = \vec{0}$ **stationary points**.

These can be maximums (summits), minimums (valleys), saddle points (a pass), or “shoulders.”

§1.2.3 The Del Operator

Definition 1.2.3. We define the **del operator** as the vector operator

$$\nabla := \hat{x} \frac{\partial}{\partial x} + \hat{y} \frac{\partial}{\partial y} + \hat{z} \frac{\partial}{\partial z}.$$

The del operator ∇ is not really a vector; it's more of an instruction. When given a scalar function T to act upon, it produces the gradient ∇T .

Nonetheless, ∇ can mimic the behavior of an ordinary vector in virtually every way, if we translate “multiply” as “act upon.” It is great notational simplification, even if it's a bit of an abuse. Indeed, we can mimic all three types of vector products:

1. On a scalar function T , ∇ can act to give the gradient ∇T ;
2. On a vector function \vec{v} , via the dot product, to give the **divergence** $\nabla \cdot \vec{v}$;
3. On a vector function \vec{v} , via the cross product, to give the **curl** $\nabla \times \vec{v}$.

The following sections will analyze the divergence and curl.

§1.2.4 The Divergence

Definition 1.2.4. The **divergence** of a vector function \vec{v} is the scalar quantity

$$\nabla \cdot \vec{v} = \frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} + \frac{\partial v_z}{\partial z}.$$

The name *divergence* is well-chosen: $\nabla \cdot \vec{v}$ is a measure of how much the vector field \vec{v} spreads out (i.e. diverges) from a given point.

Imagined a lake with streams and eddies where \vec{v} represented the velocity of the water at any given point. If you drop some particles into the lake at any given point, if they spread out, that's a point of positive divergence, while if they collect together, that's a point of negative divergence. The points of positive divergence are sources, while the points of negative divergence are sinks. (This analogy is not perfect, since water is generally incompressible, i.e. has zero divergence; however, you can imagine there are literal underwater faucets or drains adding or removing water from certain places.)

The divergence of a vector field \vec{v} measures how much the vectors “spread out” or “collect in” at any given point.

§1.2.5 The Curl

Definition 1.2.5. The **curl** of a vector function \vec{v} is the vector quantity

$$\begin{aligned} \nabla \times \vec{v} &= \det \begin{bmatrix} \hat{x} & \hat{y} & \hat{z} \\ \partial/\partial x & \partial/\partial y & \partial/\partial z \\ v_x & v_y & v_z \end{bmatrix} \\ &= \left(\frac{\partial v_z}{\partial y} - \frac{\partial v_y}{\partial z} \right) \hat{x} + \left(\frac{\partial v_x}{\partial z} - \frac{\partial v_z}{\partial x} \right) \hat{y} + \left(\frac{\partial v_y}{\partial x} - \frac{\partial v_x}{\partial y} \right) \hat{z}. \end{aligned}$$

The name *curl* is also well-chosen: $\nabla \times \vec{v}$ measures how much the vector field \vec{v} rotates (i.e. curls) around a given point. The direction of the vector $\nabla \times \vec{v}$ is given by the right hand rule: if you curl your fingers in the same direction as the vectors are curling, your thumb points in the direction of $\nabla \times \vec{v}$.

In the lake analogy, if you drop a toothpick, if it starts to rotate, then there is nonzero curl at that point. If it rotates counterclockwise, the curl points upward (positive z), while if it rotates clockwise, the curl points downward (negative z).

The curl of a vector field \vec{v} measure how much the vectors “swirl” around any given point, and also give the direction of this rotation by the right hand rule.

§1.2.6 Product Rules

The vector derivatives of sums and scalar products are rather simple; they're just what you would expect.

Proposition 1.2.6 (Vector derivative sum rules)

We have that

- $\nabla(f + g) = \nabla f + \nabla g,$
- $\nabla \cdot (\vec{a} + \vec{b}) = (\nabla \cdot \vec{a}) + (\nabla \cdot \vec{b}),$
- $\nabla \times (\vec{a} + \vec{b}) = (\nabla \times \vec{a}) + (\nabla \times \vec{b}).$

Proposition 1.2.7 (Vector derivative scalar multiplication rules)

We have that

- $\nabla(kf) = k\nabla f,$
- $\nabla \cdot (k\vec{a}) = k(\nabla \cdot \vec{a}),$
- $\nabla \times (k\vec{a}) = k(\nabla \times \vec{a}).$

The product rules, however, are not so simple. There are two ways to construct a scalar function as a product of two functions: either by multiplying two scalar functions to get fg or by taking the dot product of two vector functions to get $\vec{a} \cdot \vec{b}$. In addition, there are two ways to construct a vector function as a product of two functions: either by multiplying a scalar function by a vector function to get $f\vec{a}$, or by taking the cross product of two vector functions to get $\vec{a} \times \vec{b}$.

You can take the gradient of the two products that output a scalar function, as well as the divergence or curl of the two products that output a vector function, resulting in *six* product rules.

In general, product rules involving a scalar function are easy, and exactly what you expect (in the divergence/curl formulas, the ∇ on the scalar function turns into a gradient). The product rules involving dot and cross products of two vector functions are more complicated, however.

Proposition 1.2.8 (Gradient product rules)

We have that

- $\nabla(fg) = f\nabla g + g\nabla f,$
- $\nabla(\vec{a} \cdot \vec{b}) = \vec{a} \times (\nabla \times \vec{b}) + \vec{b} \times (\nabla \times \vec{a}) + (\vec{a} \cdot \nabla)\vec{b} + (\vec{b} \cdot \nabla)\vec{a}.$

Proposition 1.2.9 (Divergence product rules)

We have that

- $\nabla \cdot (f\vec{a}) = f(\nabla \cdot \vec{a}) + \vec{a} \cdot (\nabla f),$
- $\nabla \cdot (\vec{a} \times \vec{b}) = \vec{b} \cdot (\nabla \times \vec{a}) - \vec{a} \cdot (\nabla \times \vec{b}).$

Proposition 1.2.10 (Curl product rules)

We have that

- $\nabla \times (f\vec{a}) = f(\nabla \times \vec{a}) + \vec{a} \cdot (\nabla f),$
- $\nabla \times (\vec{a} \times \vec{b}) = (\vec{b} \cdot \nabla)\vec{a} - (\vec{a} \cdot \nabla)\vec{b} + \vec{a}(\nabla \cdot \vec{b}) - \vec{b}(\nabla \cdot \vec{a}).$

§1.2.7 Second Derivatives

We have covered the three possible first derivatives using the del operator. By using the del operator twice, we can construct *five* species of second derivatives.

Since the gradient and curl are vector, we can take the divergence or curl of them. Since the divergence is a scalar, we can only take the gradient of it. Let's consider these possibilities:

Definition 1.2.11. We call the divergence of the gradient of a scalar function T the **Laplacian** of T , denoted $\nabla^2 T$. It is equal to

$$\begin{aligned}\nabla \cdot (\nabla T) &= \left(\hat{x} \frac{\partial}{\partial x} + \hat{y} \frac{\partial}{\partial y} + \hat{z} \frac{\partial}{\partial z} \right) \cdot \left(\hat{x} \frac{\partial T}{\partial x} + \hat{y} \frac{\partial T}{\partial y} + \hat{z} \frac{\partial T}{\partial z} \right) \\ &= \frac{\partial^2 T}{\partial x^2} + \frac{\partial^2 T}{\partial y^2} + \frac{\partial^2 T}{\partial z^2}.\end{aligned}$$

Definition 1.2.12. We can also refer to the **Laplacian** of a vector function, which is simply defined as

$$\nabla^2 \vec{v} := (\nabla^2 v_x)\hat{x} + (\nabla^2 v_y)\hat{y} + (\nabla^2 v_z)\hat{z}.$$

The *curl* of the gradient, however, is always equal to zero:

$$\nabla \times (\nabla T) = \vec{0}.$$

A proof of this by components amounts to the fact that $\frac{\partial^2 T}{\partial x \partial y} = \frac{\partial^2 T}{\partial y \partial x}$.

Remark 1.2.13. There is a simple physical explanation for why the curl of the gradient always vanishes. Imagine T represents the height of some mountain; the gradient tells you which direction walks you upwards. If there was nonzero curl of the gradient, this would amount to the gradient vector going around in a closed loop; however, this would mean there is a closed loop where you're always travelling upwards, which is impossible! You can't end up in the same place but be at a higher height.

In essence, we're combining the fundamental theorems of gradients and curls, which will be introduced in the next section. By the fundamental theorem of curl (Stokes' Theorem), the surface integral of the curl of the gradient of T is equal to the line integral of the gradient of T around the boundary of that surface. Then, the fundamental theorem of gradient

implies that this loop integral equals 0. In symbols,

$$\iint_S \nabla \times (\nabla T) \cdot d\vec{a} = \oint_{\partial S} \nabla T \cdot d\vec{\ell} = 0$$

since $\partial^2 S = 0$.

Like the curl of the gradient, the *divergence* of the curl is always equal to zero:

$$\nabla \cdot (\nabla \times \vec{v}) = 0.$$

This can also be easily checked through the definitions of divergence and curl.

Remark 1.2.14. There is also an explanation of why this is true by combining the fundamental theorems of divergences and curls (also introduced in the next section). By the fundamental theorem of divergence (Gauss's Theorem), the volume integral of the divergence of the curl of \vec{v} is equal to the surface integral of the curl of \vec{v} over the boundary of the volume. Then, the fundamental theorem of curl (Stokes' Theorem) implies that this closed surface integral equals 0. In symbols,

$$\iiint_V \nabla \cdot (\nabla \times \vec{v}) d\tau = \oiint_{\partial V} (\nabla \times \vec{v}) \cdot d\vec{a} = \oint_{\partial^2 V} \vec{v} \cdot d\vec{\ell} = 0$$

since $\partial^2 V = 0$.

Remark 1.2.15. Remarks 1.2.13 and 1.2.14 both rely on the fact that the boundary operator ∂ satisfies $\partial^2 = 0$: the first uses the fact that the boundary of the boundary of a surface (the boundary of a closed loop) is nothing, and the second uses the fact that the boundary of the boundary of a volume (the boundary of a closed surface) is nothing.

Definition 1.2.16. We simply refer to $\nabla(\nabla \cdot \vec{v})$ as the **gradient of the divergence** of \vec{v} . Note that this is *not* equal to the Laplacian.

This quantity seldom appears in physics, so we don't give it a special name. The curl of the curl is also nothing new:

Definition 1.2.17. We also simply refer to $\nabla \times (\nabla \times \vec{v})$ as the **curl of the curl** of \vec{v} .

It turns out that

$$\nabla \times (\nabla \times \vec{v}) = \nabla(\nabla \cdot \vec{v}) - \nabla^2 \vec{v}.$$

Often, this is actually used to *define* the Laplacian of a vector function $\nabla^2 \vec{v}$, since our previous definition relied on Cartesian coordinates $\hat{x}, \hat{y}, \hat{z}$, which is undesirable.

§1.3 Integral Calculus

§1.3.1 Line, Surface, and Volume Integrals

Definition 1.3.1. A **line integral** is an expression of the form

$$\int_{\vec{a}}^{\vec{b}} \vec{v} \cdot d\vec{\ell},$$

where \vec{v} is some vector function and $d\vec{\ell}$ is the infinitesimal displacement vector. These two vectors are dotted along every point on some prescribed path \mathcal{P} from point \vec{a} to point \vec{b} ; these values are summed by the integral to obtain the result.

If \mathcal{P} is a closed loop, i.e. $\vec{a} = \vec{b}$, then we instead write a loop integral

$$\oint \vec{v} \cdot d\vec{\ell}.$$

Usually, the value of the line integral depends on the path taken from \vec{a} to \vec{b} , but there is a certain class of functions where it actually isn't, for all \vec{a} , \vec{b} , and potential paths \mathcal{P} ! We *force* functions with this property **conservative**.

Definition 1.3.2. A **surface integral** is an expression of the form

$$\iint_S \vec{v} \cdot d\vec{a},$$

where \vec{v} is some vector function and $d\vec{a}$ is an infinitesimal patch of area, with magnitude equal to the size of the patch and direction perpendicular to the surface \mathcal{S} that we're integrating over. While the choice of direction of $d\vec{a}$ is arbitrary for open surfaces, if the surface is closed (forming an inside and an outside), we will designate the outward direction as $d\vec{a}$ and write

$$\oiint_S \vec{v} \cdot d\vec{a}$$

instead.

If \vec{v} describes the flow of a fluid, then $\iint_S \vec{v} \cdot d\vec{a}$ represents the total mass per unit time passing through \mathcal{S} , also called the **flux**. This same term appears when talking about surface integrals of electric and magnetic fields.

As with conservative forces, there also exist a certain class of vector functions where the value of a surface integral is independent of the shape of the particular surface chosen, and instead is determined entirely by the boundary line $\partial\mathcal{S}$. Classifying these functions is also an important task.

Definition 1.3.3. A **volume integral** is an expression of the form

$$\iiint_V T d\tau,$$

where T is some scalar function and $d\tau$ is an infinitesimal volume element.

Occasionally, we also encounter the volume integrals of vector functions, which we define by

$$\iiint_V \vec{v} d\tau = \iiint_V (v_x \hat{x} + v_y \hat{y} + v_z \hat{z}) d\tau = \hat{x} \iiint_V v_x d\tau + \hat{y} \iiint_V v_y d\tau + \hat{z} \iiint_V v_z d\tau.$$

Here, since \hat{x} , \hat{y} , and \hat{z} are constants, we took them out of the integral.

§1.3.2 The Fundamental Theorem of Calculus

Recall that for a function $f(x)$ of one variable x , we have that

Theorem 1.3.4 (Fundamental Theorem of Calculus)

$$\int_a^b \left(\frac{df}{dx} \right) dx = f(b) - f(a).$$

Geometrically, this is because $\frac{df}{dx} \cdot dx = df$ is simply the small change in f after walking a small dx ; of course if you walk the whole way from a to b , we get the entire change in f from a to b , which is just $f(b) - f(a)$.

All fundamental theorems presented in this section will have the following form: the integral of some derivative of a function in a region is given by the value of that function at the boundary of the region.

In the fundamental theorem of calculus, we are taking the integral of the standard derivative of f over the region $[a, b]$, and it turns out to be given by the (sum of the) value of the function on the boundary $\partial([a, b]) = \{a, b\}$.

§1.3.3 The Fundamental Theorem for Gradients

Theorem 1.3.5 (Fundamental Theorem for Gradients)

$$\int_{\vec{a}}^{\vec{b}} (\nabla T) \cdot d\vec{\ell} = T(\vec{b}) - T(\vec{a}).$$

The intuition for this is essentially the same as for the fundamental theorem of calculus: the $(\nabla T) \cdot d\vec{\ell}$ on the left hand side give the small change in T as we walk along the path from \vec{a} to \vec{b} ; summing up these changes gives the total difference in T from \vec{a} to \vec{b} .

This gives two important corollaries:

Corollary 1.3.6

Line integrals of gradients are path independent (for forces, what we call *conservative*). In particular,

$$\int_{\vec{a}}^{\vec{b}} (\nabla T) \cdot d\vec{\ell}$$

is independent of the path taken from \vec{a} to \vec{b} .

Corollary 1.3.7

The loop integral of a gradient is equal to zero:

$$\oint (\nabla T) \cdot d\vec{\ell} = 0.$$

§1.3.4 The Fundamental Theorem for Divergences

Theorem 1.3.8 (Fundamental Theorem for Divergences)

$$\iiint_{\mathcal{V}} (\nabla \cdot \vec{v}) d\tau = \oiint_{\partial\mathcal{V}} \vec{v} \cdot d\vec{a}.$$

Note that $\partial\mathcal{V}$ is the closed surface bounding the volume \mathcal{V} .

I may refer to this theorem as either Gauss's Theorem or the divergence theorem.

Recall that if we allow \vec{v} to represent the flow of a fluid, then places of positive divergence $\nabla \cdot \vec{v} > 0$ represent “faucets” where fluid is injected and places of negative divergence $\nabla \cdot \vec{v} < 0$ represent “drains” where fluid is removed. The magnitude of the divergence in these cases represent how powerful the faucets or drains are; therefore, it makes sense that if we sum the divergence up across the entire volume, we get the total amount of fluid leaving (or entering) the volume \mathcal{V} through the surface, i.e. the flux through $\partial\mathcal{V}$.

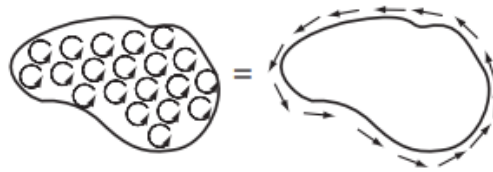
§1.3.5 The Fundamental Theorem for Curls

Theorem 1.3.9 (Fundamental Theorem for Curls)

$$\iint_S (\nabla \times \vec{v}) \cdot d\vec{a} = \oint_{\partial S} \vec{v} \cdot d\vec{\ell}.$$

I may refer to this theorem as Stokes' Theorem.

The intuition for this can come from fluid flow as well: the curl $\nabla \times \vec{v}$ represents how much the fluid swirls in the surface S ; summing up these little swirls gives the total circulation of the line integral around the boundary of the surface ∂S . The following image from Griffiths is a good depiction:



Note that in this theorem, there is ambiguity in picking the direction of $d\vec{a}$ and also the direction to go in the loop integral on the right hand side. However, it suffices to simply be consistent! Use the right hand rule: if your fingers curl around in the direction that $d\vec{\ell}$ travels, your thumb marks the direction of $d\vec{a}$.

Similar to the fundamental theorem for gradients, this gives two important corollaries:

Corollary 1.3.10

Surface integrals of curls are independent of the particular surface chosen. In particular

$$\iint_S (\nabla \times \vec{v}) \cdot d\vec{a}$$

depends only on ∂S .

Corollary 1.3.11

The integral of curls over a closed surface is equal to zero:

$$\oiint (\nabla \times \vec{v}) \cdot d\vec{a} = 0.$$

§1.3.6 Integration by Parts

Recall that the basic technique of integration by parts utilizes the product rule of derivatives:

$$\frac{d}{dx}(fg) = f \left(\frac{dg}{dx} \right) + g \left(\frac{df}{dx} \right).$$

Integrating both sides from a to b with respect to dx gives

$$fg|_a^b = \int_a^b f \left(\frac{dg}{dx} \right) dx + \int_a^b g \left(\frac{df}{dx} \right) dx.$$

Rearranging gives

$$\boxed{\int_a^b f \left(\frac{dg}{dx} \right) dx = - \int_a^b g \left(\frac{df}{dx} \right) dx + fg|_a^b.}$$

Essentially, if you're integrating the product of one function f and the *derivative* of another function g , you can transfer the derivative from g to f at the cost of a boundary term of fg on $\partial([a, b]) = \{a, b\}$.

It turns out that you can exploit the product rules of vector calculus in the same way. For example, take the product rule for the divergence of a scalar times a vector:

$$\nabla \cdot (f\vec{v}) = f(\nabla \cdot \vec{v}) + \vec{v} \cdot (\nabla f).$$

By integrating over a volume, we can abuse Gauss's Law on the left hand side:

$$\iiint_{\mathcal{V}} \nabla \cdot (f\vec{v}) d\tau = \iiint_{\mathcal{V}} f(\nabla \cdot \vec{v}) d\tau + \iiint_{\mathcal{V}} \vec{v} \cdot (\nabla f) d\tau.$$

The left hand side reduces to

$$\oint_{\partial\mathcal{V}} f\vec{v} \cdot d\vec{a},$$

so we can write

$$\boxed{\iiint_{\mathcal{V}} f(\nabla \cdot \vec{v}) d\tau = - \iiint_{\mathcal{V}} \vec{v} \cdot (\nabla f) d\tau + \oint_{\partial\mathcal{V}} f\vec{v} \cdot d\vec{a}.}$$

Notice how this has the same form as vanilla integration by parts: we can transfer a derivative (in this case a divergence) from \vec{v} to f (it turns into a gradient) at the cost of a boundary term.

It turns out that integration by parts is actually one of the most powerful tools in vector calculus; there are two other forms presented in the problem below:

§1.4 Curvilinear Coordinates

§1.4.1 Spherical Coordinates

Recall that instead of Cartesian coordinates, you can also label points in space using **spherical coordinates** (r, θ, φ) . We use physics convention here, where $r = |\vec{r}|$ is the distance to the origin; θ is the **polar angle**, the angle between \vec{r} and the z -axis; and φ is the **azimuthal angle**, the angle around from the x -axis.

We can easily relate these to Cartesian coordinates:

$$x = r \sin \theta \cos \varphi, \quad y = r \sin \theta \sin \varphi, \quad z = r \cos \theta.$$

It turns out that the unit vectors \hat{r} , $\hat{\theta}$, and $\hat{\varphi}$ form an orthonormal basis of \mathbb{R}^3 . Hence you can express any vector \vec{a} as

$$\vec{a} = a_r \hat{r} + a_\theta \hat{\theta} + a_\varphi \hat{\varphi}.$$

It is not too hard to check that

$$\begin{cases} \hat{r} = \sin \theta \cos \varphi \cdot \hat{x} + \sin \theta \sin \varphi \cdot \hat{y} + \cos \theta \cdot \hat{z}, \\ \hat{\theta} = \cos \theta \cos \varphi \cdot \hat{x} + \cos \theta \sin \varphi \cdot \hat{y} - \sin \theta \cdot \hat{z}, \\ \hat{\varphi} = -\sin \varphi \cdot \hat{x} + \cos \varphi \cdot \hat{y}. \end{cases}$$

However, there is a very important caveat about this coordinate system! The unit vectors \hat{r} , $\hat{\theta}$, and $\hat{\varphi}$ are associated with the *particular point* P that is chosen, and they change direction as P moves around! Therefore, unlike \hat{x} , \hat{y} , and \hat{z} , these basis vectors are not constants.

For example, do not combine the spherical coordinates of vectors associated with different points (e.g. if \vec{a} and \vec{b} are unit vectors pointing in opposite directions, even though $\vec{a} = \hat{r}$ in its own coordinate system and $\vec{b} = \hat{r}$ in its own coordinate system, we still have that $\vec{a} + \vec{b} = \vec{0}$, not $2\hat{r}$).

In addition, you need to be careful with differentiating a vector in spherical coordinates, since the unit vectors themselves are functions of position, and need to be differentiated (e.g. $\partial \hat{r} / \partial \theta = \hat{\theta}$). In addition, you cannot take \hat{r} , $\hat{\theta}$, or $\hat{\varphi}$ outside of an integral (e.g. as we did right after Definition 1.3.3 in the definition of the volume integral of a vector function).

Due to the setup of spherical coordinates, infinitesimal elements of length are

$$d\vec{\ell} = dr \cdot \hat{r} + r d\theta \cdot \hat{\theta} + r \sin \theta d\varphi \cdot \hat{\varphi}.$$

We call these respective small steps in the \hat{r} , $\hat{\theta}$, and $\hat{\varphi}$ directions $d\ell_r = dr$, $d\ell_\theta = r d\theta$, and $d\ell_\varphi = r \sin \theta d\varphi$.

Similarly, the infinitesimal elements of volume are also different:

$$d\tau = d\ell_r d\ell_\theta d\ell_\varphi = r^2 \sin \theta dr d\theta d\varphi,$$

the product of the scalings r and $r \sin \theta$ that we applied above to the $\hat{\theta}$ and the $\hat{\varphi}$ directions, respectively.

As you can probably imagine, the expression for surface elements is also different, but it's hard to give a general form since it depends on the orientation of the surface. In general, you can always directly analyze the geometry for any given case, but for some examples:

- If r is constant (e.g. integrating over a sphere), we have

$$d\vec{a} = d\ell_\theta d\ell_\varphi \hat{r} = r^2 \sin \theta d\theta d\varphi \hat{r}.$$

- If the surface lies in the xy plane so that θ is constant, then

$$d\vec{a} = d\ell_r d\ell_\varphi \hat{\theta} = r dr d\varphi \hat{\theta}.$$

Finally, we note that $r \in [0, \infty)$, $\varphi \in [0, 2\pi]$, and $\theta \in [0, \pi]$.

Example 1.4.1

Find the volume of a sphere of radius R .

Solution. Integrate

$$\begin{aligned} V &= \iiint_{\mathcal{V}} d\tau = \int_{r=0}^R \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} (r^2 \sin \theta) d\theta d\varphi dr \\ &= \left(\int_0^R r^2 dr \right) \left(\int_0^{2\pi} d\varphi \right) \left(\int_0^{\pi} \sin \theta d\theta \right) \\ &= \left(\frac{R^3}{3} \right) (2\pi)(2) = \frac{4}{3} \pi R^3, \end{aligned}$$

as expected. □

Alright, that finishes a discussion of the *geometry* of spherical coordinates. However, we still need to translate all the vector derivatives we've already discussed. Brute forcing thing would be quite hard, but Appendix describes a much more efficient method, which happens to treat all coordinate systems at once (I haven't taken notes for this, but maybe will in the future).

In any case, let's cut to the chase and just give the formulas:

Proposition 1.4.2 (Vector derivatives in spherical coordinates)

- Gradient:

$$\nabla T = \frac{\partial T}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial T}{\partial \theta} \hat{\theta} + \frac{1}{r \sin \theta} \frac{\partial T}{\partial \varphi} \hat{\varphi}.$$

- Divergence:

$$\nabla \cdot \vec{v} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 v_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta v_{\theta}) + \frac{1}{r \sin \theta} \frac{\partial v_{\varphi}}{\partial \varphi}.$$

- Curl:

$$\begin{aligned} \nabla \times \vec{v} &= \frac{1}{r \sin \theta} \left[\frac{\partial}{\partial \theta} (\sin \theta v_{\varphi}) - \frac{\partial v_{\theta}}{\partial \varphi} \right] \hat{r} + \frac{1}{r} \left[\frac{1}{\sin \theta} \frac{\partial v_r}{\partial \varphi} - \frac{\partial}{\partial r} (r v_{\varphi}) \right] \hat{\theta} \\ &\quad + \frac{1}{r} \left[\frac{\partial}{\partial r} (r v_{\theta}) - \frac{\partial v_r}{\partial \theta} \right] \hat{\varphi}. \end{aligned}$$

- Laplacian:

$$\nabla^2 T = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial T}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial T}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 T}{\partial \varphi^2}.$$

§1.4.2 Cylindrical Coordinates

Another method of labeling \mathbb{R}^3 is using **cylindrical coordinates** (s, φ, z) ; where we simply project P down to the xy -plane and assign polar coordinates (s, φ) to the projection, then let z be the height off the xy -plane along the z -axis. These relate by

$$x = s \cos \varphi, \quad y = s \sin \varphi, \quad z = z.$$

The unit vectors can be written as

$$\begin{cases} \hat{s} = \cos \varphi \hat{x} + \sin \varphi \hat{y}, \\ \hat{\varphi} = -\sin \varphi \hat{x} + \cos \varphi \hat{y}, \\ \hat{z} = \hat{z}. \end{cases}$$

Infinitesimal elements of length are

$$d\vec{\ell} = ds \cdot \hat{s} + s d\varphi \cdot \hat{\varphi} + dz \cdot \hat{z},$$

where the respective small steps in the \hat{s} , $\hat{\varphi}$, and \hat{z} directions are $d\ell_s = ds$, $d\ell_\varphi = s d\varphi$, and $d\ell_z = dz$.

Hence, the volume element is

$$d\tau = s ds d\varphi dz.$$

Finally, note that $s \in [0, \infty)$, $\varphi \in [0, 2\pi]$, and $z \in (-\infty, \infty)$.

Proposition 1.4.3 (Vector derivatives in cylindrical coordinates)

- Gradient:

$$\nabla T = \frac{\partial T}{\partial s} \hat{s} + \frac{1}{s} \frac{\partial T}{\partial \varphi} \hat{\varphi} + \frac{\partial T}{\partial z} \hat{z}.$$

- Divergence:

$$\nabla \cdot \vec{v} = \frac{1}{s} \frac{\partial}{\partial s}(s v_s) + \frac{1}{s} \frac{\partial v_\varphi}{\partial \varphi} + \frac{\partial v_z}{\partial z}.$$

- Curl:

$$\nabla \times \vec{v} = \left(\frac{1}{s} \frac{\partial v_z}{\partial \varphi} - \frac{\partial v_\varphi}{\partial z} \right) \hat{s} + \left(\frac{\partial v_s}{\partial z} - \frac{\partial v_z}{\partial s} \right) \hat{\varphi} + \frac{1}{s} \left[\frac{\partial}{\partial s}(s v_\varphi) - \frac{\partial v_s}{\partial \varphi} \right] \hat{z}.$$

- Laplacian:

$$\nabla^2 T = \frac{1}{s} \frac{\partial}{\partial s} \left(s \frac{\partial T}{\partial s} \right) + \frac{1}{s^2} \frac{\partial^2 T}{\partial \varphi^2} + \frac{\partial^2 T}{\partial z^2}.$$

§1.5 The Dirac Delta Function

§1.5.1 The Divergence of \hat{r}/r^2

Major apparent paradox: consider

$$\vec{v} = \frac{1}{r^2} \hat{r}.$$

This function has all arrows pointing outwards; if any vector function would have positive divergence, this is it! Yet, if we calculate the divergence anywhere, we get

$$\nabla \cdot \vec{v} = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \cdot \frac{1}{r^2} \right) = 0.$$

What's going on? In fact, it gets even worse if we try applying the divergence theorem. Integrating over a sphere \mathcal{V} of radius R centered at the origin, the surface integral is

$$\begin{aligned}\oint_{\partial\mathcal{V}} \vec{v} \cdot d\vec{a} &= \oint_{\partial\mathcal{V}} \left(\frac{1}{R^2} \hat{r} \right) \cdot (R^2 \sin\theta d\varphi d\theta \cdot \hat{r}) \\ &= \int_0^\pi \int_0^{2\pi} \sin\theta d\varphi d\theta \\ &= 2\pi \cdot \int_0^\pi \sin\theta d\theta = 4\pi.\end{aligned}$$

And yet

$$\iiint_{\mathcal{V}} (\nabla \cdot \vec{v}) d\tau = 0,$$

so what's going on? Are we to believe that the divergence theorem is false?

The solution is where $r = 0$: the divergence must blow up there. It turns out that $\nabla \cdot \vec{v}$ has the wild property that it vanishes everywhere except one point, yet its integral over any volume including that point is 4π . No ordinary function behaves like this, but it turns out we've stumbled on what is known to physicists as the **Dirac delta function**.

§1.5.2 The One-Dimensional Dirac Delta Function

Definition 1.5.1. The **one-dimensional Dirac delta function**, $\delta(x)$ can be pictured as an infinitely high, infinitesimally narrow “spike” at the origin, with area 1. Indeed,

$$\delta(x) = \begin{cases} 0 & \text{if } x \neq 0 \\ \infty & \text{if } x = 0 \end{cases}$$

and

$$\int_{-\infty}^{\infty} \delta(x) dx = 1.$$

If x has units meters, then $\delta(x)$ has units $(\text{meters})^{-1}$.

In reality, $\delta(x)$ is not a function since its value at 0 isn't finite: instead, it's a **generalized function**, or **distribution**. It is a *limit* of a *sequence* of functions, say rectangles with height n and width $1/n$, or isosceles triangles of height n and base $2/n$, centered at the origin (as $n \rightarrow \infty$).

Note that for an ordinary (continuous) function f , we have

$$f(x)\delta(x) = f(0)\delta(x),$$

as their product is 0 whenever $x \neq 0$. Hence

$$\int_{-\infty}^{\infty} f(x)\delta(x) dx = f(0) \int_{-\infty}^{\infty} \delta(x) dx = f(0).$$

Under an integral, the delta function “picks out” the value of $f(x)$ at $x = 0$.

We can shift the spike anywhere with $\delta(x - a)$. Then

$$\boxed{\int_{-\infty}^{\infty} f(x)\delta(x - a) dx = f(a)}.$$

This is so important that I've boxed it.

Example 1.5.2

Evaluate the integral

$$\int_0^3 x^3 \delta(x-2) dx.$$

What if the upper bound was 1 instead?

Solution. The delta function picks out the value of x^3 at 2, which is $\boxed{8}$. If the upper bound was 1 instead, we wouldn't catch the spike since $2 \notin [0, 1]$; hence the integral would be 0. \square

So in fact, you should really think of the delta function as something that is **always intended for use under an integral sign**. In particular, given two expressions D_1, D_2 involving delta function, we say that they're considered equal if

$$\int_{-\infty}^{\infty} f(x) D_1(x) dx = \int_{-\infty}^{\infty} f(x) D_2(x) dx$$

for *all* functions $f(x)$.

Note that this is because if D_1 and D_2 were to differ around the neighborhood of some point $x = x'$, then we could just pick $f(x)$ to have a large spike at x' , and then the integrals wouldn't be equal anymore.

By u -substitution, it can also be shown that

$$\delta(kx) = \frac{1}{|k|} \delta(x).$$

This makes sense, as the factor of k makes it so that you're "passing through" \mathbb{R} k times as fast, so you expect to pick up $\frac{1}{|k|}$ as much of it.

§1.5.3 The Three-Dimensional Delta Function

If we write $\vec{r} = x\hat{x} + y\hat{y} + z\hat{z}$, then we can generalize:

Definition 1.5.3. The **three-dimensional Dirac delta function** is defined as

$$\delta^3(\vec{r}) = \delta(x)\delta(y)\delta(z).$$

It is zero everywhere except the origin, where it blows up, and its volume integral is 1:

$$\iiint_{\mathbb{R}^3} \delta^3(\vec{r}) d\tau = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \delta(x)\delta(y)\delta(z) = 1.$$

In addition, we have that

$$\boxed{\iiint_{\mathbb{R}^3} f(\vec{r}) \delta^3(\vec{r} - \vec{a}) d\tau = f(\vec{a})}.$$

To resolve the paradox at the start of this section, we can note that

$$\nabla \cdot \left(\frac{\hat{r}}{r^2} \right) = 4\pi \delta^3(\vec{r}).$$

More generally,

$$\boxed{\nabla \cdot \left(\frac{\hat{z}}{z^2} \right) = 4\pi \delta^3(\vec{z})},$$

where, recall, $\vec{z} = \vec{r} - \vec{r}'$ for some source \vec{r}' . Also, since $\nabla \left(\frac{1}{z} \right) = -\frac{\vec{z}}{z^2}$, we have that its Laplacian is

$$\nabla^2 \frac{1}{z} = -4\pi\delta^3(\vec{z}).$$

Example 1.5.4

Evaluate the integral

$$J = \iiint_{\mathcal{V}} (r^2 + 2) \nabla \cdot \left(\frac{\hat{r}}{r^2} \right) d\tau,$$

where \mathcal{V} is a ball of radius R centered at the origin.

We present two solutions to this problem: one that demonstrates the Dirac delta function which was just introduced, and another that demonstrates the power of integration by parts.

Solution 1. Rewrite

$$J = \iiint_{\mathcal{V}} (r^2 + 2) \cdot 4\pi\delta^3(\vec{r}) d\tau = 4\pi(0^2 + 2) = \boxed{8\pi}.$$

□

Solution 2. Use integration by parts to transfer the derivative:

$$J = - \iiint_{\mathcal{V}} \nabla(r^2 + 2) \cdot \frac{\hat{r}}{r^2} d\tau + \oint_{\partial\mathcal{V}} (r^2 + 2) \cdot \frac{\hat{r}}{r^2} \cdot d\vec{a}.$$

For the first term, note that the gradient is

$$\nabla(r^2 + 2) = \frac{\partial}{\partial r}(r^2 + 2)\hat{r} = 2r \cdot \hat{r},$$

so

$$- \iiint_{\mathcal{V}} \nabla(r^2 + 2) \cdot \frac{\hat{r}}{r^2} d\tau = -4\pi \int_0^R \frac{2}{r} \cdot r^2 dr = -4\pi R^2.$$

Meanwhile, in the surface integral, we have $d\vec{a} = r^2 \sin\theta d\varphi d\theta \cdot \hat{r}$ and $r = R$, so it equals

$$\int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} (R^2 + 2) d\theta d\varphi = 4\pi(R^2 + 2).$$

Therefore, their sum is

$$J = -4\pi R^2 + 4\pi(R^2 + 2) = \boxed{8\pi},$$

which matches. □

§1.6 The Theory of Vector Fields

§1.6.1 The Helmholtz Theorem

To what extent is a vector function determined by its divergence and curl? If I tell you the divergence of \vec{F} is a certain scalar function D ,

$$\nabla \cdot \vec{F} = D,$$

and the curl of \vec{F} is a certain vector function \vec{C} ,

$$\nabla \times \vec{F} = \vec{C},$$

can you determine \vec{F} ?

Remark 1.6.1. Note here that \vec{C} must satisfy $\nabla \cdot \vec{C} = 0$, since the divergence of any curl vanishes.

And the answer is... **NOPE!** For example, there are many functions with no divergence or curl. To solve such a differential equation (as with any other differential equation), you must be supplied with **boundary conditions**. Generally, we simply require that fields go to zero infinitely far away in any direction (from all charges in the problem); with this information, **Helmholtz theorem** guarantees that the field is uniquely determined by its divergence and curl.

Remark 1.6.2. In some special cases, charges themselves extend to infinity (e.g. electric field of an infinite plane of a certain charge density). In these cases, the normal boundary conditions no longer apply; instead, one must invoke symmetry arguments to determine the fields uniquely.

§1.6.2 Potentials

Recall that the curl of any gradient is equal to zero. Well, it turns out that it works the other way around too.

Definition 1.6.3. If a vector field \vec{F} has a curl that vanishes everywhere, then it can be written as the gradient of a scalar function V that we call its **scalar potential**:

$$\nabla \times \vec{F} = \vec{0} \iff \vec{F} = -\nabla V.$$

Remark 1.6.4. The negative sign is purely for convention, so that forces push you to places of *lower* potential (energy).

This definition works because of the following theorem:

Theorem 1.6.5 (Irrotational fields)

The following are equivalent:

- (1) $\nabla \times \vec{F} = \vec{0}$ everywhere.
- (2) $\int_a^b \vec{F} \cdot d\vec{\ell}$ is independent of path, for any given endpoints.
- (3) $\oint \vec{F} \cdot d\vec{\ell} = 0$ for any closed loop.
- (4) \vec{F} is the gradient of some scalar function: $\vec{F} = -\nabla V$.

The scalar potential is not unique: you can add any constant to V and the gradient will be unaffected.

Definition 1.6.6. If a vector field \vec{F} has divergence that vanishes everywhere, then it can be written as the curl of a vector function \vec{A} that we call its **vector potential**:

$$\nabla \cdot \vec{F} = 0 \iff \vec{F} = \nabla \times \vec{A}.$$

This definition also works because of the following (similar) theorem:

Theorem 1.6.7 (Solenoidal (incompressible) fields)

The following are equivalent:

- (1) $\nabla \cdot \vec{F} = 0$ everywhere.
- (2) $\iint \vec{F} \cdot d\vec{a}$ is independent of surface, for any given boundary line.
- (3) $\oiint \vec{F} \cdot d\vec{a} = 0$ for any closed surface.
- (4) \vec{F} is the curl of some vector function: $\vec{F} = \nabla \times \vec{A}$.

This vector potential is also not unique: the gradient of any scalar function can be added to \vec{A} without affecting the curl.

It turns out that in all cases, a vector field \vec{F} can be written as the gradient of a scalar plus the curl of a vector:

We can always write

$$\vec{F} = -\nabla V + \nabla \times \vec{A}.$$

2 Electrostatics

§2.1 The Electric Field

§2.1.1 Introduction

The fundamental problem in electrodynamics is how **source charges** q_1, q_2, \dots (whose positions are *given* as functions of time) affect a **test charge** Q (whose trajectory is to be *calculated*).

The **law of superposition** helps: the net force \vec{F} on Q is just the vector sum of the forces \vec{F}_i from q_i on Q :

$$\vec{F} = \sum_i \vec{F}_i.$$

Sadly, each individual force \vec{F}_i is not easy to calculate, either. Indeed, it depends not only on the separation distance z , but also *both* their velocities and on the *acceleration* of q . Further, electromagnetic “news” only propagates at the speed of light, so we need to consider *not* what each q_i is doing now, but what it was doing at some slightly earlier time.

So we need to simplify the problem. In **electrostatics**—as the name suggests—we take all the source charges q_i to be static, or stationary (the test charge Q can still move).

§2.1.2 Coulomb's Law

Recall:

Theorem 2.1.1 (Coulomb's Law)

The force of a stationary charge q on a test charge Q that is z away is given by

$$\vec{F} = \frac{1}{4\pi\epsilon_0} \frac{qQ}{z^2} \hat{z}.$$

This is confirmed by experiments.

Definition 2.1.2. The constant here is

$$\epsilon_0 = 8.85 \times 10^{-12} \frac{\text{C}^2}{\text{N} \cdot \text{m}^2},$$

called the **permittivity of free space**.

Notice that the force \vec{F} is repulsive if q and Q have the same sign, and attractive if they have different signs. Indeed, \hat{z} points from q to Q .

§2.1.3 The Electric Field

Now suppose we have n source charges q_1, q_2, \dots, q_n , with respective vectors $\vec{z}_1, \vec{z}_2, \dots, \vec{z}_n$ to a test charge Q . Then, by the superposition principle, we have that

$$\vec{F} = \sum_{i=1}^n \vec{F}_i = \frac{Q}{4\pi\epsilon_0} \sum_{i=1}^n \frac{q_i}{z_i^2} \hat{z}_i.$$

Notice how this quantity is proportional to Q , so we can define another vector field that depends *only* on the source charges:

Definition 2.1.3. We define the **electric field** generated by source charges q_1, \dots, q_n at a point \vec{r} as

$$\vec{E}(\vec{r}) := \frac{1}{4\pi\epsilon_0} \sum_{i=1}^n \frac{q_i}{r_i^2} \hat{z}_i.$$

Then, the force of a test charge Q placed at \vec{r} is $\vec{F} = Q\vec{E}$, by the definition of \vec{E} . Further, since \vec{E} is independent of any test charges Q , we can imagine it as an actual “real” physical field that permeates space, and when test charges Q are placed in the field they get pushed along (or against) the electric field.

§2.1.4 Continuous Charge Distribution

We have only covered the case where the source is a collection of discrete point charges: to cover the continuous case, we need to turn our sum into an integral:

$$\vec{E}(\vec{r}) = \frac{1}{4\pi\epsilon_0} \int \frac{dq}{r^2} \hat{z}.$$

- If the charge is spread over a line, with linear charge density λ , then $dq = \lambda d\ell'$, so

$$\vec{E}(\vec{r}) = \frac{1}{4\pi\epsilon_0} \int \frac{\lambda(\vec{r}')}{r^2} \hat{z} d\ell'.$$

- If the charge is spread over a plane, with surface charge density σ , then $dq = \sigma da'$, so

$$\vec{E}(\vec{r}) = \frac{1}{4\pi\epsilon_0} \int \frac{\sigma(\vec{r}')}{r^2} \hat{z} da'.$$

- If the charge is spread over a volume, with volume charge density ρ , then $dq = \rho d\tau'$, so

$$\boxed{\vec{E}(\vec{r}) = \frac{1}{4\pi\epsilon_0} \int \frac{\rho(\vec{r}')}{r^2} \hat{z} d\tau'}.$$

Often, the final boxed equation is referred to as “Coulomb’s Law”: it is not far from our original formulation, and volume charge distributions are common.

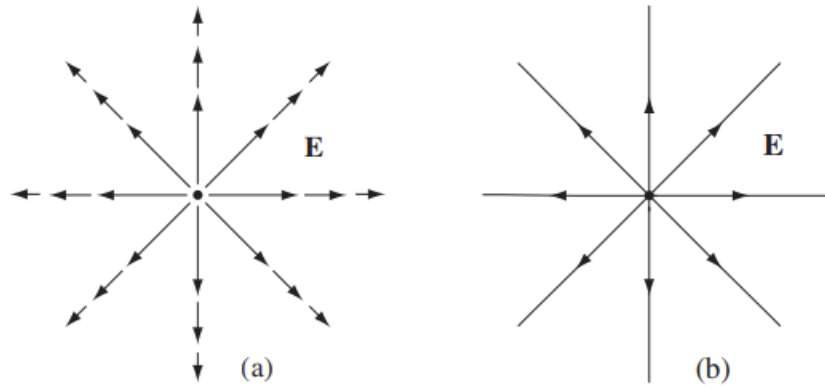
§2.2 Divergence and Curl of Electrostatic Fields

§2.2.1 Field Lines, Flux, and Gauss’ Law

In principle, we are done with electrostatics: just compute the boxed integral above to find the electric field generated by the source charges; we know that a test charge Q will simply receive a force $\vec{F} = Q\vec{E}$ at a point with electric field \vec{E} .

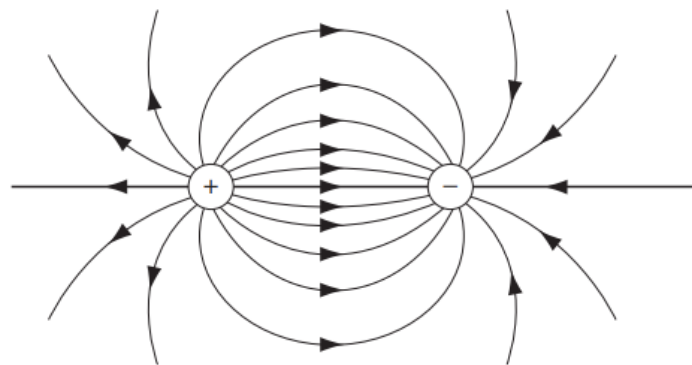
However, that integral is often intractable, so much of electrostatics is actually concerned with **assembling a bag of tools and tricks** to avoid them.

One tool to get a feel for electric fields is through **field lines**. We consider the case of a single point charge q at the origin; shown below are (a) some vectors in the electric field and (b) electric field lines.



It may seem like we lost some information: the strength of the field. However, we actually haven't! It turns out the magnitude of the field is indicated by the **density** of the field lines. However, the picture is actually a bit deceptive when drawn in two dimensions: it seems like the field drops with $1/r$; however, in three dimensions, we do get the correct $1/r^2$.

Note that charges must have a number of field lines emanating/terminating at it proportional to the magnitude of the charge. Further, electric field lines cannot terminate at any points other than charges (else $\nabla \cdot \vec{E} = 0$, something we'll prove later, would be violated): they must start/end at charges or at infinity. Further, field lines cannot cross (otherwise the electric field at that point would have two directions!). Depicted below are the field lines for an electric dipole.



Further, using this model, there's a simple interpretation for the *flux* of \vec{E} through a surface \mathcal{S} :

$$\Phi_E = \iint_{\mathcal{S}} \vec{E} \cdot d\vec{a}$$

is simply the (signed) number of electric field lines passing through \mathcal{S} .

This suggests that the flux through a *closed* surface should measure the total charge inside the surface; in particular, charges outside should have no affect on the flux since any field lines that enter simply exit again.

This is the essence of **Gauss's law**.

Theorem 2.2.1 (Gauss's Law)

For any closed surface \mathcal{S} , we have

$$\oint_{\mathcal{S}} \vec{E} \cdot d\vec{a} = \frac{Q_{\text{enc}}}{\varepsilon_0},$$

where Q_{enc} is the total charge enclosed within \mathcal{S} .

Note that since Newton's law of gravitation also obeys $1/r^2$ fall-off, it will also obey an analogous Gauss's law.

To turn this into differential form, we can rewrite the left hand side using the divergence theorem:

$$\oint_{\mathcal{S}} \vec{E} \cdot d\vec{a} = \iiint_{\mathcal{V}} (\nabla \cdot \vec{E}) d\tau$$

where $\partial\mathcal{V} = \mathcal{S}$; i.e. \mathcal{V} is the interior of \mathcal{S} . In addition, we can write

$$\frac{Q_{\text{enc}}}{\varepsilon_0} = \iiint_{\mathcal{V}} \rho d\tau.$$

Hence.

$$\iiint_{\mathcal{V}} (\nabla \cdot \vec{E}) d\tau = \iiint_{\mathcal{V}} \left(\frac{\rho}{\varepsilon_0} \right) d\tau.$$

However, since this holds true for *any* volume of choice \mathcal{V} , the integrands must also be equal.

Theorem 2.2.2 (Gauss's Law in Differential Form)

We have that

$$\nabla \cdot \vec{E} = \frac{\rho}{\varepsilon_0}.$$

§2.2.2 The Divergence of \vec{E}

Let's calculate the divergence of \vec{E} directly from the boxed equation in Section 2.1.4:

$$\vec{E}(\vec{r}) = \frac{1}{4\pi\varepsilon_0} \iiint_{\mathbb{R}^3} \frac{\hat{\mathbf{z}}}{z^2} \rho(\vec{r}') d\tau'.$$

Since the only dependence on \vec{r} is in $\vec{z} = \vec{r} - \vec{r}'$, and the divergence is linear, we can write

$$\nabla \cdot \vec{E} = \frac{1}{4\pi\varepsilon_0} \iiint_{\mathbb{R}^3} \nabla \cdot \left(\frac{\hat{\mathbf{z}}}{z^2} \right) \rho(\vec{r}') d\tau'.$$

Recall that in Section 1.5.3 we showed that $\nabla \cdot \left(\frac{\hat{\mathbf{z}}}{z^2} \right) = 4\pi\delta^3(\vec{z})$, so

$$\nabla \cdot \vec{E} = \frac{1}{4\pi\varepsilon_0} \iiint_{\mathbb{R}^3} 4\pi\delta^3(\vec{r} - \vec{r}') \rho(\vec{r}') d\tau' = \frac{\rho(\vec{r})}{\varepsilon_0},$$

as desired.

We can recover the original form of Gauss's law by simply integrating over a volume \mathcal{V} .

§2.2.3 Applications of Gauss's Law

Gauss's law is **extraordinarily powerful**, especially when symmetry permits.

Definition 2.2.3. When we apply Gauss's law over a closed surface \mathcal{S} that the boundary of some volume, we generally refer to that surface as a **Gaussian surface**.

Example 2.2.4 (Electric Field of Ball)

Find the field outside a uniformly charged solid ball of radius R and total charge q .

Proof. Suppose we wish to find the field r away from the center: then, draw our Gaussian surface \mathcal{S} as a sphere of radius r that is concentric with our solid charged ball.

Gauss's law tells us that

$$\oiint_{\mathcal{S}} \vec{E} \cdot d\vec{a} = \frac{Q_{enc}}{\epsilon_0} = \frac{q}{\epsilon_0}.$$

Now, by spherical symmetry, the electric fields \vec{E} along \mathcal{S} must all point directly outward (parallel to $d\vec{a}$), and have equal magnitude; hence,

$$E \cdot 4\pi r^2 = \frac{q}{\epsilon_0} \implies \vec{E} = \boxed{\frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \hat{r}}.$$

□

This is actually the exact same as the electric field generated by a point charge q placed at the center of the sphere, which is true by the shell theorem.

In general, there are two types of symmetry: spherical symmetry and cylindrical symmetry. In the first case, make your Gaussian surface a concentric sphere; in the second, make your Gaussian surface a coaxial cylinder.

Exercise 2.2.5. A long cylinder carries a charge density that is proportional to the distance from the axis: $\rho = ks$, for some constant k . Find the electric field inside this cylinder.

Exercise 2.2.6 (Electric Field of Plane). An infinite plane carries a uniform surface charge σ . Find its electric field.

Exercise 2.2.7 (Parallel Plate Capacitor). Two infinite parallel planes carry equal but opposite uniform charge densities $\pm\sigma$. Find the field in each of the three regions: (i) to the left of both, (ii) between them, (iii) to the right of both.

§2.2.4 The Curl of \vec{E}

We can calculate the curl of \vec{E} in the simple case of a point charge at the origin:

$$\vec{E} = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \hat{r}.$$

If we calculate the line integral

$$\int_{\vec{a}}^{\vec{b}} \vec{E} \cdot d\vec{\ell}$$

of this field from point \vec{a} to point \vec{b} , remember that we can write

$$d\vec{\ell} = dr\hat{r} + r d\theta\hat{\theta} + r \sin\theta d\varphi\hat{\varphi}.$$

Hence,

$$\vec{E} \cdot d\vec{\ell} = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} dr.$$

Hence

$$\int_{\vec{a}}^{\vec{b}} \vec{E} \cdot d\vec{\ell} = \frac{1}{4\pi\epsilon_0} \int_{\vec{a}}^{\vec{b}} \frac{q}{r^2} dr = \frac{1}{4\pi\epsilon_0} \left(\frac{q}{r_a} - \frac{q}{r_b} \right),$$

where r_a, r_b are the distances of \vec{a}, \vec{b} from the origin, respectively.

Therefore, the line integral between two points \vec{a} and \vec{b} doesn't depend on the particular path chosen; so the loop integral of \vec{E} is also equal to zero:

$$\oint \vec{E} \cdot d\vec{\ell} = 0$$

for any closed loop. Therefore, by Stokes' Theorem, we have that $\nabla \times \vec{E} = \vec{0}$.

Since ∇ is linear, this also applies for any set of discrete point charges by the superposition principle.

§2.3 Electric Potential

§2.3.1 Introduction to Potential

Electric fields aren't just any old vector fields. They are in a very special class of vector functions: ones that have no curl anywhere, i.e. $\nabla \times \vec{E} = \vec{0}$. However, because of this, we can use Theorem 1.6.5 to reduce the *vector* problem of finding \vec{E} to the *scalar* problem.

Definition 2.3.1. Since $\nabla \times \vec{E} = \vec{0}$, there exists some scalar field V such that $\vec{E} = -\nabla V$. We call V the **electric potential**.

If we agree on some standard reference point \mathcal{O} in the plane, then it turns out the function

$$V(\vec{r}) := - \int_{\mathcal{O}}^{\vec{r}} \vec{E}(\vec{r}') \cdot d\vec{\ell}.$$

Since \vec{E} is path-independent, this integral depends only on \vec{r} . In particular:

Proposition 2.3.2

The potential *difference* between any two points \vec{a} and \vec{b} is simply

$$V(\vec{b}) - V(\vec{a}) = - \int_{\vec{a}}^{\vec{b}} \vec{E} \cdot d\vec{\ell}.$$

This is equivalent to $\vec{E} = -\nabla V$ by the fundamental theorem of gradients:

$$V(\vec{b}) - V(\vec{a}) = \int_{\vec{a}}^{\vec{b}} (\nabla V) \cdot d\vec{\ell}.$$

§2.3.2 Comments on Potential

The name. The name “potential” invariably reminds you of “potential energy.” There is a connection, but they are completely different terms that should have different names. A surface over which the potential is constant is called an **equipotential**.

Advantage of the potential formulation. If you know V , you can easily get \vec{E} just by taking the gradient:

$$\vec{E} = -\nabla V.$$

Yet \vec{E} is a *vector* quantity while V is a *scalar* quantity. How can *one* function V contain all the information for the *three* components of \vec{E} .

The reason, as it turns out, is that \vec{E} is a very special type of function, being irrotational. Indeed, the condition $\nabla \times \vec{E} = 0$, in components, translates to

$$\frac{\partial E_x}{\partial y} = \frac{\partial E_y}{\partial x}, \quad \frac{\partial E_z}{\partial y} = \frac{\partial E_y}{\partial z}, \quad \frac{\partial E_x}{\partial z} = \frac{\partial E_z}{\partial x}.$$

The reference point. There is an ambiguity in defining V , based on the choice of reference point \mathcal{O} . However, altering the reference point amounts to just shifting by a constant.

In this sense, potential is a lot like altitude: if I asked you the height of something, you'd probably give it to me in terms of its height above sea level, but it would be just as valid to measure with respect to New York, or Greenwich, or wherever. Shifting by a constant doesn't matter; rather, the only quantity that actually matters is the *difference* in altitude between any two points.

Nonetheless, there is a natural spot for \mathcal{O} in electrostatics: and that is a point infinitely far from any charge. So, we set the zero of potential at infinity. Once again, this convention fails if the charge itself extends to infinity; in these cases, the potential often blows up. The remedy is to pick some other point in space as the reference point. Also note that in real life, we never have to deal with this issue.

Potential obeys the superposition principle. Since the electric field obeys this principle and ∇ is linear, so should the potential. However, it's even better now since it's an *ordinary* sum of scalars, not vectors.

Units of potential. As electric fields are newtons per coulomb, potential is newton-meters per coulomb, or joules per coulomb. We call this unit a **volt**.

Example 2.3.3

Find the potential inside and outside a spherical shell of radius R that carries a uniform surface charge, and total charge q .

Solution. Recall that the electric field generated by the sphere is

$$\vec{E}(\vec{r}) = \begin{cases} \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \hat{r} & \text{if } r > R, \\ 0 & \text{otherwise.} \end{cases}$$

For points outside of the sphere we integrate

$$V(r) = - \int_{\infty}^{\vec{r}} \vec{E} \cdot d\vec{\ell} = - \frac{1}{4\pi\epsilon_0} \int_{\infty}^r \frac{q}{r'^2} dr' = \frac{1}{4\pi\epsilon_0} \frac{q}{r} \quad (r > R).$$

For points inside the sphere, note that $\nabla V = \vec{0}$ inside, so we simply have

$$V(r) = \frac{1}{4\pi\epsilon_0} \frac{q}{R} \quad (r < R).$$

□

§2.3.3 Poisson's Equation and Laplace's Equation

Recall that we can write $\vec{E} = -\nabla V$. However, also recall that we've determined that

$$\nabla \cdot \vec{E} = \frac{\rho}{\varepsilon_0} \quad \text{and} \quad \nabla \times \vec{E} = \vec{0}.$$

Substituting in V gives

$$\nabla \cdot (\nabla V) = -\frac{\rho}{\varepsilon_0}.$$

Recall that the divergence of the gradient is the *Laplacian*. This gives **Poisson's Equation**.

Theorem 2.3.4 (Poisson's Equation)

We have

$$\nabla^2 V = -\frac{\rho}{\varepsilon_0}.$$

In regions with no charge, this reduces to **Laplace's Equation**,

$$\nabla^2 V = 0.$$

What about the fact that \vec{E} is irrotational? Sadly, it doesn't give anything, since the curl of a gradient (of V) should always equal $\vec{0}$ anyway.

§2.3.4 The Potential of a Localized Charge Distribution

Usually, we want to find \vec{E} from V , and V from the charge distribution ρ . Sadly, Poisson's Equation is the wrong way around: we can take the Laplacian of V to get ρ . Is there some way we can “invert” Poisson's equation?

Recall that a point charge q at the origin generates a potential

$$V(r) = \frac{1}{4\pi\varepsilon_0} \frac{q}{r}.$$

It turns out that we can generalize this to arbitrary charge distributions:

$$V(\vec{r}) = \frac{1}{4\pi\varepsilon_0} \int \frac{1}{z} dq.$$

For volume charge distributions, $dq = \rho \cdot d\tau'$, so

$$V(\vec{r}) = \frac{1}{4\pi\varepsilon_0} \iiint \frac{\rho(\vec{r}')}{z} d\tau'.$$

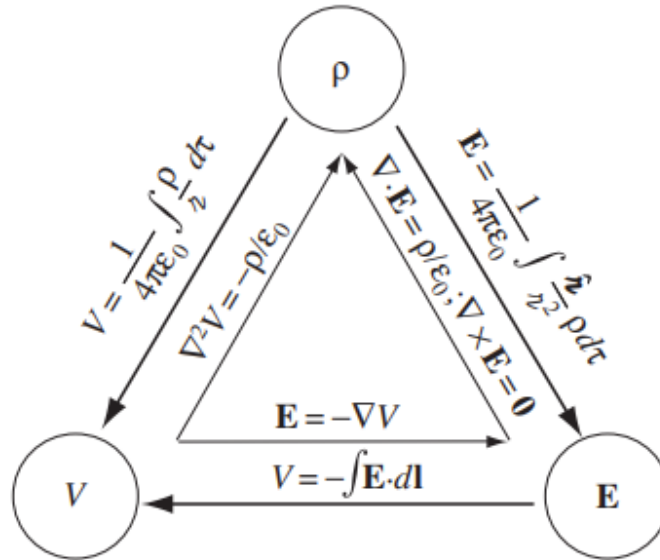
This is a solution to Poisson's equation for a localized charge distribution. Similarly, for surface and line charge distributions we get potentials

$$V = \frac{1}{4\pi\varepsilon_0} \int \frac{\lambda(\vec{r}')}{z} d\ell' \quad \text{and} \quad V = \frac{1}{4\pi\varepsilon_0} \iint \frac{\sigma(\vec{r}')}{z} da'.$$

Recall once again that this section depends on the fact that the reference point is at infinity.

§2.3.5 Boundary Conditions

We have covered the three fundamental quantities in electrostatics: ρ , \vec{E} , and V . We have also derived all six formulas interrelating them, summarized below. All six of these formulas followed from two experimental observations: (1) the principle of superposition and (2) Coulomb's law.



Now onto the point of this section: what happens to \vec{E} and V when you cross a *boundary*, i.e. a surface with charge density σ .

It is simple to determine this for \vec{E} using Gauss's law: simply draw a really thin wafer of area A and thickness ϵ containing the surface; then

$$\oiint_S \vec{E} \cdot d\vec{a} = \frac{Q_{\text{enc}}}{\epsilon_0} = \frac{\sigma A}{\epsilon_0}.$$

Even if there are other electric charges in the area, the flux through the sides of the pillbox is 0 as we take $\epsilon \rightarrow 0$. Hence, the difference in the normal component of the electric field below and the *normal* component (obtained after dotting with $d\vec{a}$) of the electric field above is $\frac{\sigma}{\epsilon_0}$. Note that this is true for a single infinitely flat plane with surface charge density σ as well: the electric fields are both of magnitude $\frac{\sigma}{2\epsilon_0}$, but pointing in opposite directions.

However, the *tangential* component of \vec{E} is *always* continuous. After all, the loop integral of \vec{E} is always 0, so if we considered a small loop that is the along the side of the pillbox (so length ℓ parallel to the surface and length ϵ through the surface), then we see that the tangential components must be equal (since their difference is 0).

Hence, we can summarize:

$$\boxed{\vec{E}_{\text{above}} - \vec{E}_{\text{below}} = \frac{\sigma}{\epsilon_0} \hat{n}},$$

where \hat{n} is a unit vector pointing from below to above.

On the other hand, the potential is continuous across *any* boundary, since the difference in V is given by the path integral of \vec{E} , but since the path length shrinks to zero, so too does the integral:

$$V_{\text{above}} - V_{\text{below}} = - \int_{\text{below}}^{\text{above}} \vec{E} \cdot d\vec{\ell} = 0.$$

The *gradient* of V keeps the discontinuity though, so

$$\nabla V_{\text{above}} - \nabla V_{\text{below}} = -\frac{\sigma}{\varepsilon_0} \hat{n}.$$

More conveniently, we can consider the \hat{n} component of both sides by taking the dot product with \hat{n} , recalling that the dot product of the gradient of V with a unit vector gives the rate of change of V along the direction of that unit vector:

$$\frac{\partial V}{\partial n} = (\nabla V) \cdot \hat{n};$$

in this case, we call it the **normal derivative** of V , i.e. the rate of change of V in the direction perpendicular (normal) to the surface. Hence, we can write

$$\boxed{\frac{\partial V_{\text{above}}}{\partial n} - \frac{\partial V_{\text{below}}}{\partial n} = -\frac{\sigma}{\varepsilon_0}}.$$

§2.4 Work and Energy in Electrostatics

§2.4.1 The Work It Takes to Move a Charge

Suppose we have a *stationary* configuration of source charges, and want to move a test charge Q from \vec{a} to \vec{b} . How much work will we have to do?

Well, the force at any point is given by $\vec{F} = Q\vec{E}$; so you need to exert $-Q\vec{E}$. Hence the needed work is

$$W = \int_{\vec{a}}^{\vec{b}} \vec{F} \cdot d\vec{\ell} = -Q \int_{\vec{a}}^{\vec{b}} \vec{E} \cdot d\vec{\ell} = Q[V(\vec{b}) - V(\vec{a})].$$

As $\nabla \times \vec{E} = \vec{0}$, this answer is independent of the path from \vec{a} to \vec{b} : therefore, we call the electrostatic force *conservative*.

Hence,

$$V(\vec{b}) - V(\vec{a}) = \frac{W}{Q}.$$

This gives us another view of potential:

The potential difference between points \vec{a} and \vec{b} is equal to the work per unit charge required to carry a particle from \vec{a} to \vec{b} .

In particular, if you want to bring Q in from far away and stick it at \vec{r} , you must do work $W = Q \cdot V(\vec{r})$ (as we set our reference point at infinity). In this sense, potential is *potential energy per unit charge* of a single particle moving within a field generated by stationary source charges.

§2.4.2 The Energy of a Point Charge Distribution

But what we want to assemble an entire *collection* of point charges? How much work would we need?

The first charge q_1 takes zero work: there's nothing to help or fight against. Bringing in q_2 will cost $q_2 V_1(\vec{r}_2)$, where V_1 is the potential due to q_1 , which is then evaluated at \vec{r}_2 :

$$W_2 = \frac{1}{4\pi\varepsilon_0} \frac{q_1 q_2}{r_{12}},$$

where z_{12} is the distance from \vec{r}_1 to \vec{r}_2 . Moving in q_3 then takes

$$W_3 = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_3}{z_{13}} + \frac{1}{4\pi\epsilon_0} \frac{q_2 q_3}{z_{23}},$$

and so on. Continuing this pattern, the total work necessary is

$$W = \frac{1}{4\pi\epsilon_0} \sum_{i=1}^n \sum_{j=i+1}^n \frac{q_i q_j}{z_{ij}} = \frac{1}{8\pi\epsilon_0} \sum_{i=1}^n \sum_{j \neq i}^n \frac{q_i q_j}{z_{ij}}.$$

This is equal to

$$\frac{1}{2} \sum_{i=1}^n q_i \left(\sum_{j \neq i}^n \frac{1}{4\pi\epsilon_0} \frac{q_j}{z_{ij}} \right) = \frac{1}{2} \sum_{i=1}^n q_i V(\vec{r}_i). \quad (2.1)$$

This is exactly half of what you'd expect if you just added the work of placing any point charge into the configuration (holding everything else and their generated potential fixed).

This value represents the energy stored in the system of charges (avoiding the annoying term “*potential*” energy).

§2.4.3 The Energy of a Continuous Charge Distribution

For a volume charge density ρ , this becomes

$$W = \frac{1}{2} \iiint_{\mathbb{R}^3} \rho V d\tau.$$

There is actually a wonderful way to rewrite this equation to get rid of ρ and V in place of \vec{E} . First, use Gauss's law to convert ρ to \vec{E} :

$$\rho = \epsilon_0 \nabla \cdot \vec{E} \implies W = \frac{\epsilon_0}{2} \iiint_{\mathbb{R}^3} (\nabla \cdot \vec{E}) V d\tau.$$

Now, use integration by parts to transfer the derivative from \vec{E} to V :

$$W = \frac{\epsilon_0}{2} \left[- \iiint_{\mathbb{R}^3} \vec{E} \cdot (\nabla V) d\tau + \oint_{\partial\mathbb{R}^3} V \vec{E} \cdot d\vec{a} \right].$$

Wait a second... what's $\partial\mathbb{R}^3$? Well... if we expand the volume containing charge over and over then the surface integral will drop to 0 since area grows with r^2 but V falls off with $1/r$ and \vec{E} falls off with $1/r^2$.

Now we can use $\nabla V = -\vec{E}$ to get

$$\boxed{W = \frac{\epsilon_0}{2} \iiint_{\mathbb{R}^3} E^2 d\tau}. \quad (2.2)$$

The energy stored in an electric field per unit volume, then, is $\frac{\epsilon_0 E^2}{2}$.

§2.4.4 Comments on Electrostatic Energy

A perplexing “inconsistency”. The equation

$$W = \iiint \frac{\epsilon_0 E^2}{2} d\tau$$

seems to imply that the energy of a stationary charge distribution is *always positive*. Yet the energy of two equal but opposite charges a distance z apart is apparently negative quantity $-(1/4\pi\epsilon_0)(q^2/z)$, according to Equation 2.1.

What's going on? Well, *both are correct*. The problem here is that the energy in $-(1/4\pi\epsilon_0)(q^2/z)$ **doesn't take into account the energy necessary to create the point charges** (it supposes the point charges are already given, we're just assembling them), while Equation 2.2 does.

This is good, in fact, as the energy of a point charge is actually *infinite*. By Equation 2.2,

$$\begin{aligned} W &= \frac{\epsilon_0}{2} \iiint_{\mathbb{R}^3} \frac{1}{(4\pi\epsilon_0)^2} \cdot \frac{q^2}{r^4} d\tau \\ &= \frac{q^2}{32\pi^2\epsilon_0} \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} \int_{r=0}^{\infty} \frac{1}{r^4} \cdot r^2 \sin\theta dr d\theta d\varphi. \\ &= \frac{q^2}{8\pi\epsilon_0} \int_0^{\infty} \frac{1}{r^2} dr = \infty. \end{aligned}$$

So Equation 2.2 is more *complete* in giving the total energy stored in the charge configuration. Yet for point charges, as we've seen, we definitely shouldn't include this, so Equation 2.1 assumes the point charges are already made (e.g. electrons), and cannot be taken apart. So the energy inside them doesn't matter.

Where is the energy stored? Is the energy stored in the charge, or in the field? Presently, this is an unanswerable question: we can compute *what* the total energy is, and several ways to compute it, but worrying *where* it is is irrelevant. In radiation theory and general relativity, it is useful to regard the energy as stored in the field with energy density $\frac{\epsilon E^2}{2}$ per unit volume.

However, in electrostatics, we can just as well say that the energy is stored in the charge, with density $\frac{1}{2}\rho V$.

The superposition principle. Because electrostatic energy is *quadratic* in the fields, it does *not* obey a superposition principle. The energy of a compound system is *not* the sum of the energies of its parts considered equally, as there are cross terms:

$$W_{\text{tot}} = \frac{\epsilon_0}{2} \iiint (\vec{E}_1 + \vec{E}_2)^2 d\tau = W_1 + W_2 + \epsilon_0 \iiint \vec{E}_1 \cdot \vec{E}_2 d\tau.$$

For example, if you double the charge everywhere, you quadruple the total energy.

§2.5 Conductors

§2.5.1 Basic Properties

Definition 2.5.1. In an **insulator**, such as glass or rubber, electrons are not free to move around: they are held closely to particular atoms by short leashes.

Definition 2.5.2. In a metallic **conductor**, on the other hand, one or more electrons per atom are free to roam. (In liquid conductors, ions move).

From this definition, we can derive several electrostatic properties:

$\vec{E} = \vec{0}$ inside a conductor. The general idea is that if not, then the free charges inside the conductor would move to generate an *opposing electric field* that cancels out any external one. This movement of charges occurs practically instantaneously.

$\rho = 0$ inside a conductor. Follow from Gauss's Law: $\rho = \epsilon_0 \nabla \cdot \vec{E}$.

Any net charge resides on the surface. Where else could the net charge be?

Remark 2.5.3. This particular condition is pretty surprising: charges do push each other so it's natural for them to spread out, but placing them all on the boundary seems like a waste of the space inside (shouldn't you at least sprinkle a little charge in the middle?). Nonetheless, it is true that minimizing the energy stored in the system is best done by moving all the charge to the surface.

A conductor is an equipotential. Since the path integral of \vec{E} is 0, as \vec{E} itself is always $\vec{0}$.

\vec{E} is perpendicular to the surface, just outside a conductor. Otherwise charge on the conductor would immediately move around to cancel out any tangential component of the electric field. Charge cannot move in the normal direction, of course, since it is confined to the conductor.

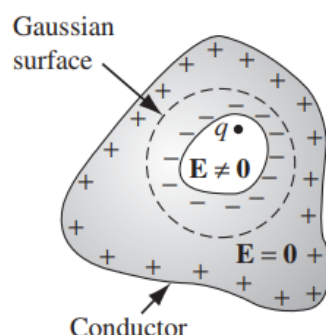
§2.5.2 Induced Charges

Placing a $+q$ charge near an uncharged (!) conductor will cause them to attract. This is because the $+q$ charge will pull negative charges closer to it and repel positive charges farther away. Since the negative charge in the conductor is closer to the $+q$ charge, we will see a net attractive force.

While the field inside the actual material making up a conductor must be zero, there can be an electric field within any hollow cavities closed off within a conductor.

However, it turns out that charges within a cavity in a conductor are *electrically isolated from the outside world*: no fields penetrate the conductor, but you can still detect the presence of a charge inside due to charge present on the outer surface of the conductor.

Let's investigate closely what happens: if we have a point charge $+q$ inside a hollow cavity in a conductor, then negative charges will rush to the boundary of the hollow cavity (the inside surface of the conductor), while positive charges will be pushed outward to the outside surface of the conductor.



If we take a Gaussian surface running entirely through the conductor that encloses the hollow cavity, we know the flux through this surface must be zero (as the electric field along the entire surface is zero). Therefore, the charge enclosed by this Gaussian surface is zero; hence, the inside surface of the conductor surrounding the hollow cavity develops a charge $-q$. Since the conductor started neutral, its outside surface must develop a charge $+q$.

Example 2.5.4

An uncharged spherical conductor centered at the origin has a cavity of some weird shape carved out of it in the inside. Somewhere within the cavity is a charge q . What is the field outside the sphere?

Proof. It turns out the answer depends neither on the shape of the cavity nor the location of the charge! Wild, isn't it.

Indeed, it turns out that any imbalance from the position of q or the shape of the cavity is balanced out by the negative charge developed on the inner surface, and the charge q spread on the outside of the sphere is uniform! Hence, the answer is

$$\vec{E} = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \hat{r}.$$

□

Remark 2.5.5. The argument here isn't conclusive. There are *three* fields at work here: \vec{E}_q from the point charge, \vec{E}_{induced} from the charge of $-q$ on the inside surface of the conductor, and $\vec{E}_{\text{leftover}}$ from the charge of q on the outside surface of the conductor.

We cannot supply a fully rigorous argument yet, but for now, take note that there *exists* a way of distributing the $-q$ on the inner surface so that \vec{E}_q and \vec{E}_{induced} cancel at any exterior points. Indeed, the same cavity (with the point charge inside) could've been carved out of an insanely large conductor, and the positive charges out the outside just would be too far away to affect the inside, so the first two fields would have to do all the work of cancellation by themselves.

Yet how do we know that this is actually what happens? Well, it turns out that due to some uniqueness theorems of Chapter 3, electrostatics has the property that there is precisely one way of distributing the charge on a conductor so that the field inside is zero. Having found a single *possible* way, we are guaranteed that no alternative exists.

If there is no charge inside the cavity, then we can argue that $\vec{E} = 0$; indeed, if there was some field line beginning and ending on the cavity wall, integrating across it would violate the fact that all points on the conductor have equal potential. This is the principle behind grounded **Faraday cages**, which will shield out stray electric fields. In practice, the conductor doesn't even have to be solid; something like chicken wire often suffices.

§2.5.3 Surface Charge and the Force on a Conductor

Since the field inside a conductor is zero, and all the charge is concentrated on the surface of the conductor, the boundary conditions from Section 2.3.5 require that the field immediately outside the conductor is

$$\vec{E} = \frac{\sigma}{\epsilon_0} \hat{n},$$

where \hat{n} is the unit vector normal to the surface of the conductor. In terms of potential, we get

$$\frac{\partial V}{\partial n} = -\frac{\sigma}{\epsilon_0}$$

for the value of the directional derivative of V in the direction of the normal \hat{n} . Hence, if you can determine \vec{E} or V , you can calculate the surface charge on a conductor.

In the presence of an electric field, a surface charge will experience a force per unit area of $\vec{f} = \sigma \vec{E}$. But the electric field is *discontinuous* at a surface charge, so what do we use for \vec{E} ? The answer is the average of the electric field below and above:

$$\vec{f} = \sigma \vec{E}_{\text{average}} = \frac{1}{2} \sigma (\vec{E}_{\text{above}} + \vec{E}_{\text{below}}).$$

All this average does is remove any contribution of the surface itself to the electric field (as obviously the patch cannot exert a force on itself).

This argument applies to *any* surface charge; in the particular case of a conductor, the field inside is zero and outside is $\frac{\sigma}{\epsilon_0} \hat{n}$; hence the force per area is

$$\vec{f} = \frac{\sigma^2}{2\epsilon_0} \hat{n}.$$

This amounts to an outwards **electrostatic pressure** on the surface, drawing the conductor into the field, regardless of the sign of σ . We can also write the pressure as

$$P = \frac{\epsilon_0 E^2}{2}.$$

§2.5.4 Capacitors

Suppose we have two conductors, one with charge $+Q$ on it and one with charge $-Q$ on it. Since V is constant over conductors, the potential difference between these two conductors is unambiguous:

$$V = V_+ - V_- = - \int_{(-)}^{(+)} \vec{E} \cdot d\vec{\ell}.$$

We don't know how the charges get distributed, or what the exact field looks like, but we do know that \vec{E} and V should both be proportional to Q . After all, \vec{E} is given by Coulomb's law

$$\vec{E} = \frac{1}{4\pi\epsilon_0} \iiint_{\mathbb{R}^3} \frac{\rho}{r^2} \hat{r} d\tau,$$

so if you double ρ everywhere, you double \vec{E} as well.

Remark 2.5.6. There is a slight caveat here: how do we know that, say, doubling Q actually doubles ρ everywhere. Maybe the charge would actually move around, so that in some places ρ is quadrupled while in others maybe it is halved, so that the total charge Q is still doubled. Well, it turns out that doubling Q does in fact double ρ everywhere, but the proof of this will have to wait for Chapter 3.

Since \vec{E} is proportional to Q , and $\vec{E} = -\nabla V$, V is also proportional to Q .

Definition 2.5.7. The constant of proportionality between Q and V is called the **capacitance** of the arrangement:

$$C := \frac{Q}{V}.$$

Capacitance is a purely geometrical quantity, determined entirely by the sizes, shapes, and separation between the two conductors. Capacitance is measured in **farads**, which is a coulomb-per-volt. This is generally inconveniently large: microfarads and picofarads are more convenient.

Note that since Q and V are both *defined* to be positive, C is always positive.

Remark 2.5.8. Sometimes, you will hear people speak of the capacitance of a *single* conductor. In this case, the “second conductor” is an imaginary spherical shell of infinite radius surrounding the conductor. It contributes nothing to the field, so the capacitance is just given by Q/V where V is the potential with ∞ as the reference point.

Example 2.5.9 (Capacitance of parallel-plate capacitor)

Find the capacitance of two metal surfaces of area A held a distance of $d \ll A$ apart.

Solution. When $d \ll A$, the field between the two plates can be approximated as $Q/A\epsilon_0$ from the positive plate to the negative plate. Therefore, the potential difference is

$$V = - \int_{(-)}^{(+)} \vec{E} \cdot d\vec{\ell} = \frac{Qd}{A\epsilon_0}.$$

Hence,

$$C = \frac{Q}{V} = \boxed{\frac{A\epsilon_0}{d}}.$$

□

Example 2.5.10

Find the capacitance of two concentric spherical metal shells, with radii a and b .

Solution. Place a charge $+Q$ on the inner sphere of radius a , and $-Q$ on the outer sphere of radius b . Then the field between the two spheres is

$$\vec{E} = \frac{1}{4\pi\epsilon_0} \frac{Q}{r^2} \hat{r}.$$

Therefore, the potential difference is

$$V = \int_{(+)}^{(-)} \vec{E} \cdot d\vec{\ell} = \frac{Q}{4\pi\epsilon_0} \int_a^b \frac{1}{r^2} dr = \frac{Q}{4\pi\epsilon_0} \left(\frac{1}{a} - \frac{1}{b} \right).$$

Hence,

$$C = \frac{Q}{V} = \boxed{4\pi\epsilon_0 \cdot \frac{ab}{b-a}}.$$

□

In order to charge up a capacitor, you must remove electrons from the positive plate and carry them to the negative plate. Doing so fights against the electric field. How much work does it take to charge up the capacitor to a final amount Q ?

Proposition 2.5.11

The amount of work necessary to charge a capacitor up to a final amount Q is given by

$$W = \frac{Q^2}{2C} = \frac{1}{2}CV^2.$$

Proof. Indeed, if we've already charged the capacitor up by q , then the potential difference is q/C . Hence, the work required to move it is

$$dW = \frac{q}{C}dq.$$

The total work, then, is

$$W = \int_0^Q \frac{q}{C}dq = \frac{1}{2} \frac{Q^2}{C}.$$

Since $Q = CV$, we often write this as

$$W = \frac{1}{2}CV^2.$$

□