# Digital Image Processing Notebook

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# 1 ELEC70078 Digital Image Processing

### 1.1 Image Acquisition

An image is a projection of a 3D scene into a 2D projection plane. An image can be defined as a function of two variables (x, y) as:

$$f(x,y): \mathbb{R}^2 \to \mathbb{R}$$

where, for each position (x, y) in the projection plane, f(x, y) defines the light intensity at this point.

### 1.2 Sampling and Quantization

The analogue image needs to be converted into a digital image by sampling and quantization (Digitize).

Obviously, the higher the sample rate, the better the quality. The more quantization levels, the better the quality.

# 1.3 Four Main Parts of Digital Image Processing

- 1. Image Transforms
- 2. Image Enhancement
- 3. Image Restoration
- 4. Image Compression

#### 1.4 Natural Signals and Compression

The natural signals are sparse, and that's why they can be compressed.

# 2 2D Discrete Fourier Transform (DFT)

#### 2.1 1D DFT

The generic form of a one-dimensional signal transform:

$$g(u) = \sum_{x=0}^{N-1} T(u, x) f(x), \text{ for } 0 \le u \le N-1$$

T(u,x) is a function of u,x called the **forward transformation kernel**.

$$\mathbf{g} = T \cdot \mathbf{f}$$

For DFT, we have:

$$T(u,x) = e^{-j2\pi \frac{ux}{N}}$$

The generic form of inverse transform:

$$f(x) = \frac{1}{N} \sum_{u=0}^{N-1} I(u, x) g(u), \text{ for } 0 \le x \le N-1$$

I(u, x) is called the **inverse transformation kernel**.

$$\mathbf{f} = I \cdot \mathbf{g}, \quad \mathbf{g} \in \mathbb{R}^N$$

$$I = \frac{1}{N} e^{j2\pi \frac{ux}{N}} = \frac{1}{N} T^* = T^{-1}$$

### 2.2 Matrix Properties

Recall:

$$\frac{1}{N}T^* = T^{-1}, \quad T^*T = N \cdot I, \quad T^TT = N \cdot I$$

This implies that the matrix is **orthogonal**.

### 2.3 2D Transform

The generic form of a 2D transform:

$$g(u,v) = \sum_{x=0}^{M-1} \sum_{u=0}^{N-1} T(u,x)T(v,y)f(x,y)$$

### 2.4 Separable and Symmetric Transforms

A 2D transform is **separable** if:

$$T(u, x, v, y) = T_1(u, x)T_2(v, y)$$

A 2D transform is **symmetric** if:

$$T(u,x) = T_2(v,y) = T(x,y)$$

$$f(x,y) = \frac{1}{MN} \sum_{u=0}^{M-1} \sum_{v=0}^{N-1} I(u,x)I(v,y)g(u,v)$$

### 2.5 Separable Transform Expansion

$$g(u,v) = \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} T_1(u,x) T_2(v,y) f(x,y)$$
$$= \sum_{y=0}^{N-1} T_2(v,y) \left( \sum_{x=0}^{M-1} T_1(u,x) f(x,y) \right)$$
$$= \sum_{y=0}^{N-1} T_2(v,y) F(u,y)$$

Where F(u, y) is the **intermediate image**.

## 3 Properties of DFT

### 3.1 Energy Preservation (Parseval's Theorem)

$$||g||^2 = ||f||^2$$

For images, this can be written as:

$$\sum_{x=0}^{M-1} \sum_{y=0}^{N-1} |f(x,y)|^2 = \sum_{u=0}^{M-1} \sum_{v=0}^{N-1} |g(u,v)|^2$$

### 3.2 Energy Compaction

Transformed image values close to the origin correspond to **low frequency**.

Most energy of the transformed image is concentrated in a small area close to the origin.

Low Frequency  $\leftrightarrow$  Smooth Surface

High Frequency  $\leftrightarrow$  Edge

# 4 The 2D Discrete Cosine Transform (DCT)

#### 4.1 1D DCT

The one-dimensional DCT is defined as:

$$C(u) = a(u) \sum_{x=0}^{N-1} f(x) \cos \left[ \frac{(2x+1)u\pi}{2N} \right], \quad 0 \le u \le N-1$$

where the scaling factor a(u) is:

$$a(u) = \begin{cases} \sqrt{\frac{1}{N}}, & \text{if } u = 0\\ \sqrt{\frac{2}{N}}, & \text{if } u = 1, 2, \dots, N - 1 \end{cases}$$

#### **Inverse Transform**

The inverse 1D DCT is given by:

$$f(x) = \sum_{u=0}^{N-1} a(u)C(u)\cos\left[\frac{(2x+1)u\pi}{2N}\right]$$

The signal is projected onto **real sinusoids** instead of complex exponentials.

### 4.2 2D DCT

The two-dimensional DCT is:

$$C(u,v) = a(u)a(v) \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} f(x,y) \cos\left[\frac{(2x+1)u\pi}{2M}\right] \cos\left[\frac{(2y+1)v\pi}{2N}\right]$$

Inverse 2D DCT

$$f(x,y) = \sum_{u=0}^{M-1} \sum_{v=0}^{N-1} a(u)a(v)C(u,v)\cos\left[\frac{(2x+1)u\pi}{2M}\right]\cos\left[\frac{(2y+1)v\pi}{2N}\right]$$

#### 4.3 Remarks

- DCT projects signals onto real cosine bases.
- DCT has better **energy compaction** properties compared to the DFT.
- It is considered the real-valued version of the DFT.

## 5 The Walsh-Hadamard Transform

### 5.1 Binary Representation

Each integer x can be represented in binary as:

$$x = (b_{n-1}(x) \ b_{n-2}(x) \ \dots \ b_0(x))$$

**Example:** For x = 6

$$b_2(x) = 1$$
,  $b_1(x) = 1$ ,  $b_0(x) = 0$ 

#### 5.2 Definition

The Walsh transform is defined as:

$$W(u) = \frac{1}{\sqrt{N}} \sum_{x=0}^{N-1} f(x) \prod_{i=0}^{n-1} (-1)^{b_i(x) \cdot b_{n-1-i}(u)}$$

or equivalently:

$$W(u) = \frac{1}{\sqrt{N}} \sum_{x=0}^{N-1} f(x)(-1)^{\sum_{i=0}^{n-1} b_i(x) \cdot b_{n-1-i}(u)}$$

Where:

$$N = 2^n, \quad b_i(x) \in \{0, 1\}$$

#### 5.3 Inverse Transform

$$f(x) = \frac{1}{\sqrt{N}} \sum_{u=0}^{N-1} W(u) \prod_{i=0}^{n-1} (-1)^{b_i(x) \cdot b_{n-1-i}(u)}$$

or equivalently:

$$f(x) = \frac{1}{\sqrt{N}} \sum_{u=0}^{N-1} W(u)(-1)^{\sum_{i=0}^{n-1} b_i(x) \cdot b_{n-1-i}(u)}$$

### 5.4 Matrix Property

We can show:

$$I = N \cdot T^{-1} = T^T$$

This implies that we can easily get the inverse matrix.

### 6 2D Walsh and Hadamard Transforms

#### 6.1 2D Walsh Transform

$$W(u,v) = \frac{1}{N} \sum_{x=0}^{N-1} \sum_{y=0}^{N-1} f(x,y) \prod_{i=0}^{n-1} (-1)^{b_i(x)b_{n-1-i}(u) + b_i(y)b_{n-1-i}(v)}$$

Inverse transform:

$$f(x,y) = \frac{1}{N} \sum_{u=0}^{N-1} \sum_{v=0}^{N-1} W(u,v) \prod_{i=0}^{n-1} (-1)^{b_i(x)b_{n-1-i}(u) + b_i(y)b_{n-1-i}(v)}$$

#### 6.2 Hadamard Transform

The Hadamard transform is similar to the Walsh transform:

$$H(u,v) = \frac{1}{N} \sum_{x=0}^{N-1} \sum_{y=0}^{N-1} f(x,y) (-1)^{\sum_{i=0}^{n-1} [b_i(x)b_{n-1-i}(u) + b_i(y)b_{n-1-i}(v)]}$$

### 6.3 Domain Transformation and Sequencing

The Walsh and Hadamard transforms move data from one domain to another (e.g., from spatial to frequency). We interpret frequency as the number of zero-crossings or sign changes in the basis vector. This is referred to as **sequency**.

Ordered Basis Example (Hadamard Matrix):

Rows correspond to increasing sequency from 0 to 7

# 7 Karhunen–Loève Transform (KLT)

#### 7.1 Eigenvalues and Eigenvectors

Given a covariance matrix C, the KLT relies on solving:

$$C\mathbf{e} = \lambda \mathbf{e}$$

where:

- e: eigenvector
- $\lambda$ : eigenvalue

#### 7.2 Image Vector Representation

Suppose we have n images sampled at the same location (x, y), forming a vector:

$$\vec{X}_{x,y} = \begin{bmatrix} x_1(x,y) \\ x_2(x,y) \\ \vdots \\ x_n(x,y) \end{bmatrix}$$

#### 7.3 Mean Vector and Covariance Matrix

The mean vector:

$$\vec{m}_x(x,y) = \mathbb{E}[\vec{X}(x,y)] = \begin{bmatrix} m_1(x,y) & m_2(x,y) & \dots & m_n(x,y) \end{bmatrix}^T$$

Mean value for component i:

$$m_i = \frac{1}{MN} \sum_{k=0}^{M-1} \sum_{l=0}^{N-1} x_i(k, l)$$

Covariance matrix:

$$\operatorname{Cov}(\vec{X}) = \mathbb{E}[(\vec{X} - \vec{m}_x)(\vec{X} - \vec{m}_x)^T]$$
 of size  $n \times n$ 

### 7.4 Covariance Matrix Properties

- The covariance matrix is symmetric and positive definite.
- $\mathbf{x}^T C \mathbf{x} > 0$  for all non-zero real vectors  $\mathbf{x}$ .

#### 7.5 Transform Matrix Construction

Let matrix A be formed by stacking the eigenvectors of the covariance matrix C:

$$A = \begin{bmatrix} e_1^T \\ e_2^T \\ \vdots \\ e_n^T \end{bmatrix}, \quad \text{with corresponding eigenvalues } \lambda_1 \geq \lambda_2 \geq \cdots \geq \lambda_n$$

Transformation:

$$\mathbf{y} = A(\mathbf{x} - \vec{m}_x)$$

It's called the Karhunen-Loève Transform.

#### 7.6 Proof of Zero-Mean Output

$$\vec{m}_{u} = \mathbb{E}[\mathbf{y}] = 0$$

**Proof:** 

$$\vec{m}_y = \mathbb{E}[\mathbf{y}]$$

$$= \mathbb{E}[A(\mathbf{x} - \vec{m}_x)]$$

$$= A\mathbb{E}[\mathbf{x} - \vec{m}_x]$$

$$= A(\mathbb{E}[\mathbf{x}] - \vec{m}_x)$$

$$= A(\vec{m}_x - \vec{m}_x) = 0$$

### 7.7 Covariance in Transformed Domain

$$Cov(\mathbf{y}) = ACov(\mathbf{x})A^T$$

Matrix A and its transpose:

$$A = \begin{bmatrix} e_1^T \\ e_2^T \\ \vdots \\ e_n^T \end{bmatrix}, \quad A^T = [e_1 \ e_2 \ \dots \ e_n]$$

Diagonalization:

$$Cov(\mathbf{x})A^T = C_x[e_1 \dots e_n] = [\lambda_1 e_1 \dots \lambda_n e_n]$$

$$A\text{Cov}(\mathbf{x})A^T = \begin{bmatrix} e_1^T \\ \vdots \\ e_n^T \end{bmatrix} [\lambda_1 e_1 \dots \lambda_n e_n] = \begin{bmatrix} \lambda_1 & 0 & \dots & 0 \\ 0 & \lambda_2 & \dots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \dots & \lambda_n \end{bmatrix}$$

This is a diagonal matrix of variances—uncorrelated components.

### 7.8 Reconstruction and Compression

**Inverse KLT:** 

$$\mathbf{x} = A^T \mathbf{y} + \vec{m}_x$$

Keeping only top K components:

$$\hat{\mathbf{x}} = A^T \mathbf{y}_K + \vec{m}_x$$

### 7.9 Mean Square Error (MSE)

$$MSE = \frac{1}{n} \sum_{i=1}^{n} ||x_i - \hat{x}_i||^2 = ||x - \hat{x}||^2 = \sum_{j=K+1}^{n} \lambda_j$$

### 7.10 Interpretation

The Karhunen–Loève Transform allows for decorrelation and energy compaction:

- Most of the signal energy is concentrated in the first few transformed coefficients.
- Enables compression by retaining only significant components.

**Illustration:** Compressing correlated images by concentrating information into a smaller number of principal components.

# 8 Image Enhancement

#### 8.1 Classification

Image enhancement techniques can be classified into:

- Spatial-domain methods (operate on (x, y))
- Frequency-domain methods (operate on (u, v))

#### 8.2 Intensity Transformations

Let s = T(r), where:

- r: input grey level at pixel f(x,y)
- s: output grey level at pixel g(x,y)

#### Image Negatives

$$s = T(r) = L - 1 - r$$

#### Piecewise Linear Transformation

Often used to stretch or compress intensity ranges.

### 8.3 Log Transformation

$$s = c \cdot \log(1+r)$$

Enhances the dynamic range of low intensity values.

#### 8.4 Power-Law Transformations

$$s = T(r) = c \cdot r^{\gamma}$$

Used for gamma correction:

- $\gamma > 1$ : compress higher intensities, darkens image
- $\gamma < 1$ : expand higher intensities, brightens image

### 8.5 Gray Level Slicing

Highlights specific ranges of grey values:

- Preserve intensities in a certain range
- Suppress others or keep them unchanged

### 8.6 Bit-Plane Slicing

Assume each pixel is represented by 8 bits. Separate an image into its binary layers:

- MSB (Most Significant Bit) layers hold visually significant data
- LSB (Least Significant Bit) layers capture fine details

## 9 Histogram Equalization

Consider an image with intensity levels  $r_k \in [0, L-1]$ , and image size  $M \times N$ .

### 9.1 Histogram Definitions

- The number of pixels with intensity  $r_k$  is  $n_k$ .
- The histogram of the image is the function:  $h(r_k) = n_k$ .
- The normalized histogram is:

$$p(r_k) = \frac{n_k}{MN}$$

This represents the probability of a pixel having grey value  $r_k$ .

#### 9.2 Image Contrast Types

- Low contrast: narrow histogram.
- High contrast: widely spread histogram.
- Dark image: histogram biased to low intensity values.
- Bright image: histogram biased to high intensity values.

#### 9.3 Transformation Function

We want a transformation s = T(r), with T'(r) > 0, satisfying:

- Monotonically increasing
- One-to-one mapping
- $0 \le T(r) \le L 1$  for  $0 \le r \le L 1$

### 9.4 Continuous Case

Let  $p_r(r)$  be the probability density function (PDF) of input intensities.

We want to find a transformation s = T(r) such that  $p_s(s)$  is **uniformly distributed**.

$$p_s(s) ds = p_r(r) dr \Rightarrow p_s(s) = p_r(r) \frac{dr}{ds}$$

To make  $p_s(s) = \frac{1}{L-1}$ , we integrate:

$$\int_0^s p_s(s) \, ds = \int_0^r p_r(r) \, dr = \frac{1}{L-1} \int_0^s ds = \frac{s}{L-1} \Rightarrow s = (L-1) \int_0^r p_r(r) \, dr$$

### 9.5 Discrete Version

$$s_k = (L-1) \sum_{j=0}^{k} p(r_j)$$

### 9.6 Effect of Histogram Equalization

- Enhances image contrast
- Makes the light distribution more uniform
- More visually pleasing for human perception

# 9.7 Example (Illustrative)

A dark image (e.g., most pixels at intensity 1) is transformed to use a broader intensity range (e.g., intensity 1 to 3), improving visibility and spreading the histogram.

### 9.8 Histogram Specification (Matching)

We seek a transformation z = T(r) such that the output image has a specified probability density function  $p_z(z)$ .

Step 1: Equalize  $p_r(r)$ 

$$s = (L-1) \int_0^r p_r(r) dr = T(r)$$

Step 2: Equalize  $p_z(z)$ 

$$s = (L-1) \int_0^z p_z(t) dt = G(z)$$

Step 3: Invert to get z

$$z = G^{-1}(s) = G^{-1}(T(r)), \text{ or } s = T(r) = G(z)$$

### Conceptual Flow

$$r \xrightarrow{T(r)} s \xrightarrow{G^{-1}(s)} z$$

#### Interpretation:

- Adjusts the histogram to match a specified shape (e.g., bimodal, Gaussian).
- Useful for style matching or enhancing specific contrast profiles.

### 9.9 Local Histogram Equalization

Instead of applying histogram equalization globally, apply it to small regions to enhance local details.

#### Advantages:

- Enhances contrast in small areas.
- Preserves edge and texture information better.

#### **Process:**

- Slide a window (e.g., 3x3 or 5x5) across the image.
- Apply histogram equalization to each window.
- Combine results to form the final image.

# 10 Spatial Filters in Image Processing

#### 10.1 Model with Noise

$$g(x,y) = f(x,y) + \eta(x,y)$$

Where  $\eta(x,y)$  is white noise.

White Noise: A sequence of uncorrelated random variables with zero mean and finite variance. If it follows a normal distribution with zero mean, it's called *white Gaussian noise*.

#### 10.2 Filter Kernels and Convolution

#### Convolution:

$$g(x,y) = \sum_{k=-K}^{K} \sum_{l=-L}^{L} w(k,l) f(x+k,y+l)$$

Box Filter (Low-Pass):

$$\frac{1}{9} \begin{bmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{bmatrix} = \frac{1}{3} \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix} \cdot \frac{1}{3} \begin{bmatrix} 1 & 1 & 1 \end{bmatrix}$$

This filter is **separable**.

#### 10.3 Gaussian Filter

#### Gaussian Kernel:

$$w(x,y) = \frac{1}{2\pi\sigma^2} e^{-\frac{x^2+y^2}{2\sigma^2}}$$

- Smooths the image.
- Gaussian in spatial domain remains Gaussian in frequency domain.
- Acts as a low-pass filter.

#### 10.4 Order-Statistic Filters

Median Filter:

$$g(x,y) = \text{median}\{f(x+i,y+j)\}, \quad i = -N,...,N, \ j = -M,...,M$$

- Effective for removing "salt and pepper" noise.
- Non-linear filtering.

#### 10.5 High-Pass Filtering

$$\frac{1}{9} \begin{bmatrix} -1 & -1 & -1 \\ -1 & 8 & -1 \\ -1 & -1 & -1 \end{bmatrix}$$

#### Properties:

- The sum of coefficients must be zero.
- Emphasizes edges and suppresses smooth areas.

### 10.6 High-Boost Filtering

Formula:

High-boost image = 
$$A \cdot I_{\text{orig}} - I_{\text{low}} = (A - 1) \cdot I_{\text{orig}} + I_{\text{high}}$$

**Application:** Enhances edges more strongly than high-pass filtering.

## 11 Edge Detection

Edges are abrupt changes in pixel intensity.

## 11.1 Gradient and Derivatives

Change rate:

$$\frac{\partial f}{\partial x} = \lim_{\Delta x \to 0} \frac{f(x + \Delta x) - f(x)}{\Delta x}$$

Discrete approximation ( $\Delta x = 1$ ):

$$\frac{\partial f}{\partial x} \approx f(x+1) - f(x)$$

#### 11.2 Second Derivative (Laplacian)

$$\frac{\partial^2 f}{\partial x^2} = f(x+1) - 2f(x) + f(x-1)$$

### 11.3 Gradient Magnitude

$$\nabla f(x,y) = \begin{bmatrix} \frac{\partial f}{\partial x} \\ \frac{\partial f}{\partial y} \end{bmatrix}$$
$$\|\nabla f(x,y)\|_2 = \sqrt{\left(\frac{\partial f}{\partial x}\right)^2 + \left(\frac{\partial f}{\partial y}\right)^2} \quad \text{(Euclidean norm)}$$
$$\|\nabla f(x,y)\|_1 \approx \left|\frac{\partial f}{\partial x}\right| + \left|\frac{\partial f}{\partial y}\right| \quad \text{(Faster to compute)}$$

### 11.4 Laplacian Operator

$$\nabla^2 f = \frac{\partial^2 f}{\partial x^2} + \frac{\partial^2 f}{\partial y^2}$$

Used to detect areas of rapid intensity change.

## 12 Image Restoration

### 12.1 Degradation Model

$$g(i,j) = H[f(i,j)] + n(i,j)$$

For Linear Time (or Space) Invariant (LTI/LSI) systems:

$$g(i,j) = f(i,j) * h(i,j) + n(i,j)$$

Where \* denotes convolution.

### 12.2 Types of Degradation Functions

• Uniform motion blur:

$$h(i,j) = \begin{cases} \frac{1}{(2L+1)^2}, & -L \le i, j \le L \\ 0, & \text{otherwise} \end{cases}$$

• Atmospheric turbulence:

$$h(i,j) = e^{-k(i^2 + j^2)^{5/6}}$$

• Out-of-focus blur:

$$h(i,j) = \begin{cases} \frac{1}{\pi R^2}, & i^2 + j^2 \le R^2 \\ 0, & \text{otherwise} \end{cases}$$

#### 12.3 Restoration Quality Metrics

• BSNR (Blurred Signal-to-Noise Ratio):

BSNR = 
$$10 \log_{10} \left( \frac{1}{MN} \sum_{i,j} [E\{f(i,j)\} - \bar{f}(i,j)]^2 / \sigma_n^2 \right)$$

• ISNR (Improvement SNR):

ISNR = 
$$10 \log_{10} \left( \frac{\sum_{i,j} [f(i,j) - g(i,j)]^2}{\sum_{i,j} [f(i,j) - \hat{f}(i,j)]^2} \right)$$

#### 12.4 Lexicographic Ordering

Transform image matrix to a 1D vector  $\mathbf{y} = H\mathbf{f} + \mathbf{n}$ . Matrix H represents the Toeplitz-block structure.

#### 12.5 Inverse Filtering

$$\hat{F}(u,v) = \frac{G(u,v)}{H(u,v)}, \quad f(i,j) = \mathrm{IDFT}[\hat{F}(u,v)]$$

Noise Amplification Problem:

$$\hat{F}(u,v) = \frac{F(u,v)H(u,v) + N(u,v)}{H(u,v)} = F(u,v) + \frac{N(u,v)}{H(u,v)}$$

#### 12.6 Modified Inverse Filter

$$H'(u,v) = \begin{cases} \frac{1}{H(u,v)}, & \text{if } |H(u,v)| \ge \varepsilon \\ 0, & \text{otherwise} \end{cases}$$

## 12.7 Constrained Least Squares (CLS) Restoration

Objective:

$$\min \|y - Af\|^2 + \lambda \|Cf\|^2$$

Solution:

$$\hat{F}(u,v) = \frac{H^*(u,v)}{|H(u,v)|^2 + \lambda |C(u,v)|^2} G(u,v)$$

### 12.8 Wiener Filter

Assume:

$$\mathbb{E}[f] = \mathbb{E}[y] \Rightarrow \text{Linear estimator: } \hat{f} = W\mathbb{E}[y]$$

Error independent of distortion:  $\mathbb{E}[(f - \hat{f})(y - \mathbb{E}[y])^T] = 0$ 

Wiener Solution in Frequency Domain:

$$W(u,v) = \frac{H^*(u,v)S_{ff}(u,v)}{|H(u,v)|^2S_{ff}(u,v) + S_{nn}(u,v)}$$

If H(u, v) = 1:

$$W(u,v) = \frac{S_{ff}(u,v)}{S_{ff}(u,v) + S_{nn}(u,v)} = \frac{\text{SNR}}{\text{SNR} + 1}$$

- If SNR  $\gg 1$ :  $W(u, v) \approx 1$
- If SNR  $\ll 1$ :  $W(u,v) \approx \text{SNR}$

# 13 Image Compression

#### 13.1 Overview

Compression reduces data size in images, video, or audio. It involves two processes:

- Encoding (compress)
- **Decoding** (reconstruct)

Compression Ratio:

$$CR = \frac{\text{Input size}}{\text{Output size}} \quad \text{e.g., } \frac{64 \text{kbps}}{4 \text{kbps}} = 16$$

#### 13.2 Lossless vs. Lossy Compression

- Lossless:  $\hat{f} = f$
- Lossy:  $SNR_{\hat{f}} < SNR_f$

SNR (Signal-to-Noise Ratio):

$$SNR = \frac{signal\ energy}{noise\ energy}$$

## 14 Information Theory Basics

#### 14.1 Self-Information

$$I(s_i) = \log_2\left(\frac{1}{p_i}\right) = -\log_2 p_i$$

### 14.2 Entropy

$$H(S) = \sum_{i} p_i I(s_i) = -\sum_{i} p_i \log_2 p_i$$
 (bits/symbol)

#### 14.3 Source Model

- Source: sequence of symbols from a finite alphabet.
- DMS (Discrete Memoryless Source): symbols generated independently.

### 14.4 Average Length

$$L_{avg} = \sum_{i=1}^{n} p_i \cdot l_i$$
 (length of codewords)  
$$L_{avg} \ge H(S)$$

# 15 Huffman Coding

- Prefix code, optimal for a given probability model.
- Binary tree representation.
- $H(S) \leq L_{avg} < H(S) + 1$

**Decoding:** Traverse the tree based on bits.

# 16 Coding Extensions and Redundancy

#### 16.1 Nth Extension

Encoding blocks of N symbols improves efficiency:

$$H(S) \le \frac{L_{avg}}{N} < H(S) + \frac{1}{N}$$

### 16.2 Redundancy

Redundancy = 
$$L_{avg} - H(S)$$

# 17 Differential Coding

Difference Encoding:

$$g(x,y) = f(x,y) - f(x-1,y)$$
 (horizontal) or  $f(x,y) - f(x,y-1)$ 

**Prediction Residual:** 

$$r = y - x$$
, where  $y = f(a, b, c)$ 

# 18 JPEG Compression Pipeline

- 1. Divide image into  $8 \times 8$  blocks.
- 2. Apply 2D DCT.
- 3. Quantize DCT coefficients.
- 4. Use zig-zag scan to serialize.
- 5. Apply:
  - Differential coding (DC)
  - Run-length coding (AC)
  - Huffman encoding

### 18.1 DCT and Quantization

$$Z_{i,j} = \left\lfloor \frac{Y_{i,j}}{Q_{i,j}} \right\rfloor$$

# 18.2 Differential Coding (DC Coefficient)

$$\Delta DC = DC_i - DC_{i-1}$$

## 18.3 Run-Length Coding (AC Coefficients)

- Use zig-zag order to group zeros.
- Encode using (run, value) pairs.
- Each pair Huffman encoded.

#### 18.4 JPEG Final Structure

 $Original \rightarrow DCT \rightarrow Quantization$ 

DC: differential + Huffman, AC: zig-zag + run-length + Huffman