

Lecture 4 (Moved Up): Numerical Integration and Math "Review" (Because you' ve probably forgotten it unless you had me for NWP)

Solving Ordinary Differential Equations in the Stella Environment



Agenda

- Review the Basics of the Taylor Series
- Introduce you to the numerical solving methods used in Stella

Our Goal:

- To make Stella less of a black box and more of a viable learning and working tool.
 - If you don't know what's going on under a GUI, you don't know what's going on period.



What this will require in review

- Review Taylor Expansion
- The Classic Limit Theory Definition of "Derivative"

 Here we will use a very simple timedependant relationship, y=f(y,t)



The Old Derivative's Definition

Let's recall some basic Calc...

$$\frac{dy}{dt} = \lim_{\Delta t \to 0} \frac{y(t_0 + \Delta t) - y(t_0)}{\Delta t}$$



The Derivative's Definition

• So, if Δt is small we can approximate this so that

$$\frac{dy}{dt} \approx \frac{y(t_0 + \Delta t) - y(t_0)}{\Delta t}$$



And playing to to solve for a y(t)?

• So using the simple definition...

$$y(t_0 + \Delta t) = y(t_0) + \frac{dy}{dt} \Delta t$$

- Which would be easy.... if dy/dt is always a constant
 - i.e., dy/dt = 6 and not 6t or $f(t^2)$ or $6\cos(t)$
 - But what do you do when you when you have a higher order dy/dt?



The Taylor Series

$$y(t_0 + \Delta t) = y(t_0) + \frac{dy}{dt} \Delta t$$

• The Taylor Series is, for y(t) at $t=t_0+\Delta t$

$$y(t_{0} + \Delta t) = y(t_{0}) + \sum_{i=1}^{\infty} \left[\frac{(\Delta t)^{i}}{i!} \frac{d^{i}y}{dt^{i}} (t_{0}) \right]$$

$$= y(t_{0}) + \Delta t \frac{dy}{dt} (t_{0}) + \frac{(\Delta t)^{2}}{2!} \frac{d^{2}y}{dt^{2}} (t_{0}) + \frac{(\Delta t)^{3}}{3!} \frac{d^{3}y}{dt^{3}} (t_{0}) + \dots$$

• So... the more you know about the higher order terms of the relationship y(t) as it changes, the better you can project its value onto future values of t from a known starting point, $y(t_0)$ at time t_0



The Taylor Series

FOR YOUR NOTEBOOK!

$$y(t_0 + \Delta t) = y(t_0) + \sum_{i=1}^{\infty} \left[\frac{(\Delta t)^i}{i!} \frac{d^i y}{dt^i} (t_0) \right]$$

$$= y(t_0) + \Delta t \frac{dy}{dt} (t_0) + \frac{(\Delta t)^2}{2!} \frac{d^2 y}{dt^2} (t_0) + \frac{(\Delta t)^3}{3!} \frac{d^3 y}{dt^3} (t_0) + \dots$$

• Get the following... for the 0th through 5th order.

$$y(t) = t^3$$

 $t_0 = 1; \quad y(t_0) = 1; \quad \Delta t = 2; \quad y(t = 3) = ?$

• Answer: 1, 7, 19, 27, 27, 27, 27, 27,



The Taylor Series in **Modeling**

$$y(t) = t^3$$

$$\frac{dy}{dt} = 3t^2$$

$$\frac{dy}{dt} = 3t^2$$

- Applying it to modeling...
- We know that at t=1, $y(t=1)=1^3=1$
- If we knew the perfect solution to y(t), which here is t^3 , then we'd know that $y(t=3)=3^3=27$
- But... we don't know the perfect solution... otherwise modeling would be a perfect and precise art.



The Taylor Series in Modeling

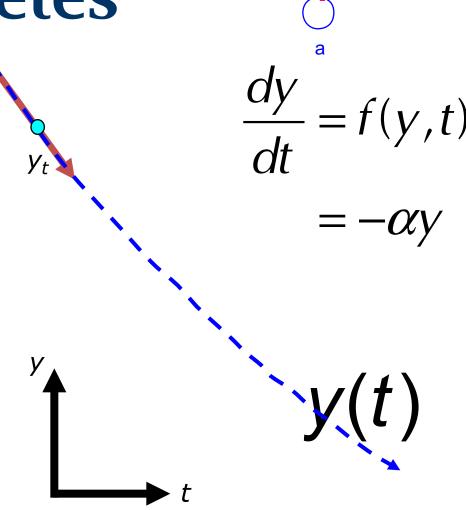
- We DO, however, hopefully know what forces are acting on our system at our current point in time (or how to guess at them)
- Therefore we can approximate the potential changes to our system NOW, and project them into the future, based on what we know NOW assuming that changes in the system are modest over delta-t.
- And the better we can approximate these potential changes, the better the solution!



Spiraling Down to

 $\frac{dy}{dt}(t) = -\alpha y(t)$ Concretes

- Consider the following blue curve. The "perfect" real solution.
- This is the system we want to model.
- We are at the light blue point.
- We want to project the future position of the point along the "perfect" curve.



expense

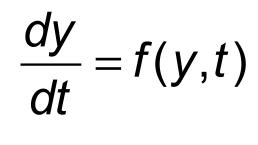


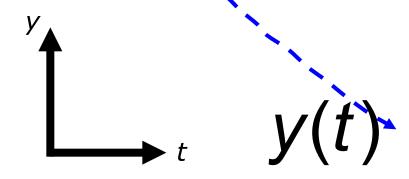
Spiraling Down to Concretes

$$\frac{dy}{dt} = c_1 - c_2 y^4$$

$$\frac{dy}{dt} = -\alpha y$$

- Before proceeding let's sit back and look at a given system.
- We have a parameter we wish to solve (or model): y(t).
- We begin by <u>not</u>
 attacking the formula for
 y and "solve" y





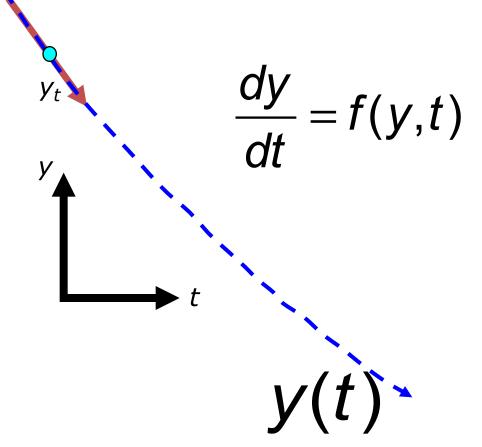


Spiraling Down to Concretes

$$\frac{dy}{dt} = c_1 - c_2 y^4$$

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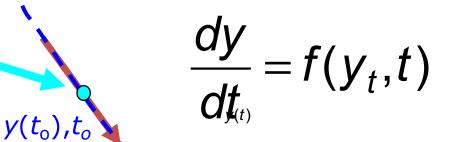
- We begin by <u>not</u> expressing the formula for y
- Rather, we express y in terms of its *changes*.
 - In our example, y could be temperature or internal energy of the earth (or the water depth of a bucket - or the dollars in a trustfund...)
 - Its change is a function of the incoming radiation (c₁) and its ability to offload energy (c₂ T⁴)

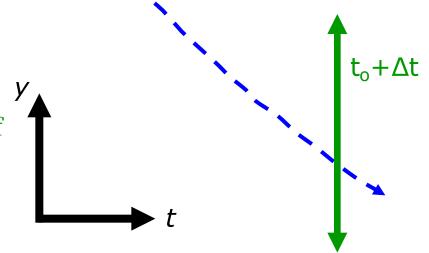




Spiraling Down to Concretes

- "You are here"
 - At the Carolina Blue Dot
 - $-t=t_0$
 - $y = y_0$
 - $dy/dt = f[y(t_0), t_0]$
 - These values we know
- "We want to be there"
 - Green Line
 - to+ Δ t
 - y=?
 - hence the vague line instead of a green dot
 - (the blue line is the ideal but still an unknown!)





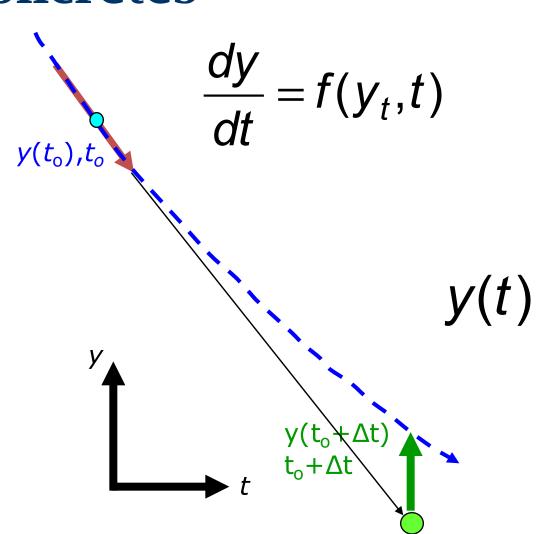


Spiraling Back Down to Concretes

 Notice what happens if you take the first derivative approximation and solve for x(t₀+Δt)

$$y(t_0 + \Delta t) \approx y(t_0) + \Delta t \frac{dy}{dt}(t_0)$$

- ...same as the 1rst-Order approx of the Taylor Series
 - (This is also a variant on the same eq. that I tossed about periodically last time and is what stella "writes" for you)

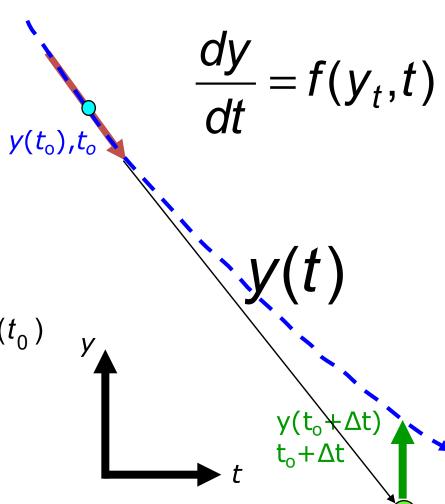




Spiraling Back Down to Concretes

This solution approach is called *Euler's Method*.
 (not to be confused with Euler's theorem or Eulerian expansion) and can serve as a fast solution approach.

$$y(t_0 + \Delta t) \approx y(t_0) + \Delta t \frac{dy}{dt}(t_0)$$





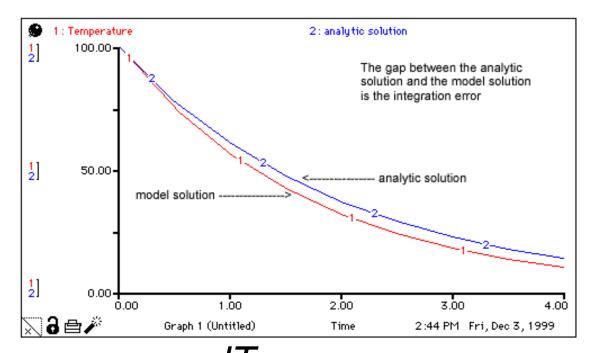
Euler's Method

- Euler's method is clearly very simple and straight forward.
- Unfortunately, it is not the most accurate of methods.
- It is reliant on the **first derivative** at the **point of origin** for a given time step.
 - Analogy: "Driving faster than the range of your headlights"
 - If delta-t is big, you're gonna accrue error.
- Consequently, the smaller the time step the better the prediction, but error is not fully eliminated.



Euler's Method

- ☐ To the side is an other example:
 - Classic Cooling
- ☐ The Blue indicates an analytical solution to the equation.
- ☐ The Red, a Eulerestimated approach.
- ☐ (This works better in a spreadsheet than the previous example)



$$\frac{dI}{dt} = -A_1T$$
$$T = A_2e^{-A_1t}$$



Runge-Kutta Method

- Let's then look at another approach:
- The *Runge Kutta Method*. This method provides superior accuracy than Euler's Method and can handle larger time steps as well.
- But before we proceed, let's change some notation.
- This is mostly to reduce clutter and to make what we do here compatible with most textbooks and refernce manuals like *Numerical Recipes*

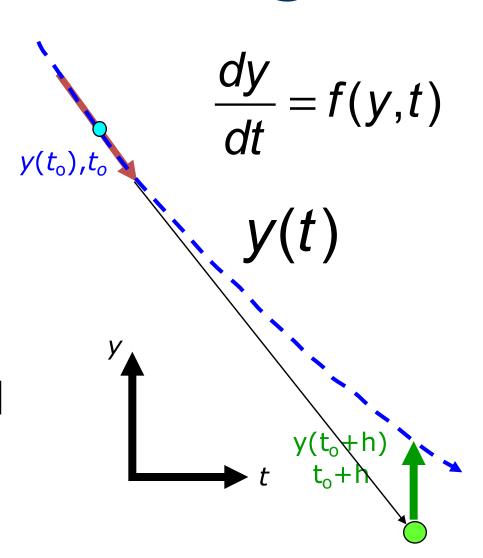


Notation Change

 We're going to use "f" for the first derivative

 $h \equiv \Delta t$

- "f" for "forcings" $\frac{dy}{dt} = f(y,t)$ $y(t_0 + \Delta t) \approx y(t_0) + \Delta t f[y(t_0), t_0]$ also





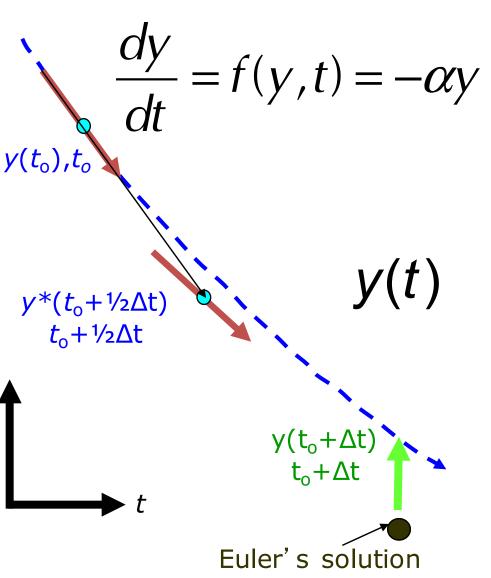
Runge-Kutta Method

• If we were to take an incremental approach, we could define our derivatives at intermediate values between our steps...

• Today, we'll look at the 1^{rst}-, 2nd- and 4rd-Order R-K approaches



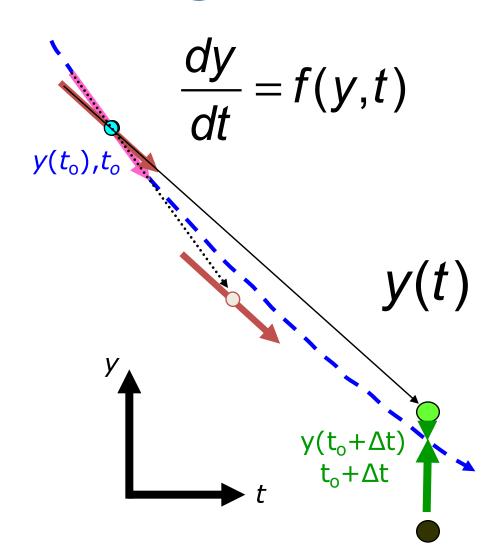
- Let's shoot <u>only</u> half-way first to the mid-point between t and t+Δt
- We now have an intermediate slope between the two points.
- This "midpoint" progression allows us to aim now for the next step.
- But in case we are "off" withy our y estimate, we'll hedge our bets and shoot again from t_o.



For reference



- Using this new slope...
 - Notice, first, that we assume a continuous curve here
- ... We re-cast our line back to t=t+∆t for our final estimate



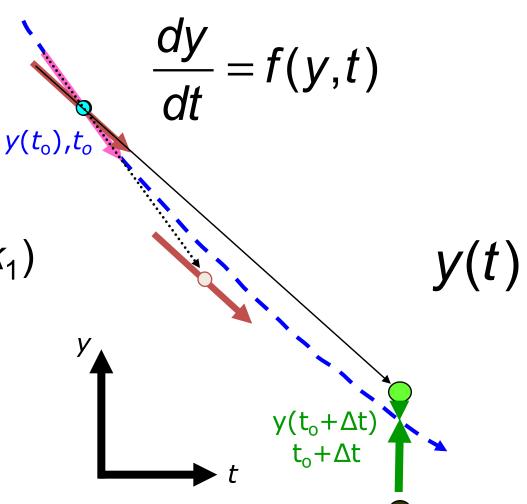


- In the world of perfect forms (aka textbooks):
 - Since Fortran doesn't do pictures!

$$y_{n+1} = y_n + k_2 + O(h^3)$$

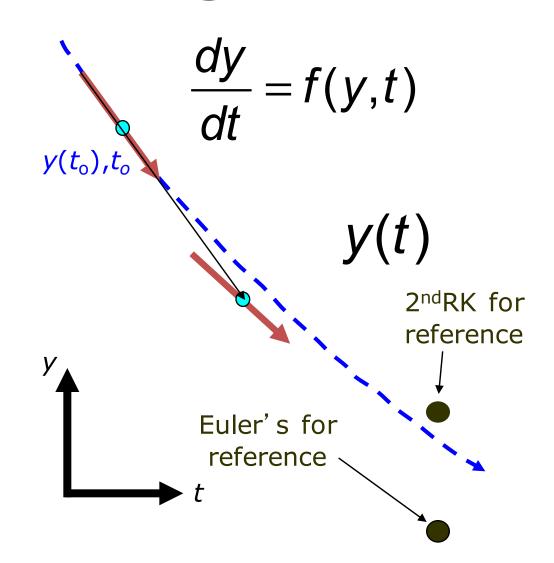
 $k_2 = h f(t_n + \frac{1}{2}h, y_n + \frac{1}{2}k_1)$
 $k_1 = h f(t_n, y_n)$

 This approach is also called the "mid-point" method



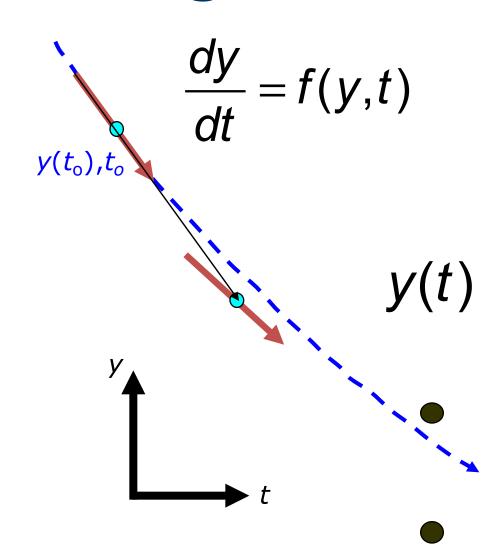


- We can expand this to a 4th-order RK (the first order, btw is the Euler's method (if you go backward from the 2ndorder example, you might be able to see why).
- But for the 4th-order approch we will cast our line more than once (actually, 4 times)



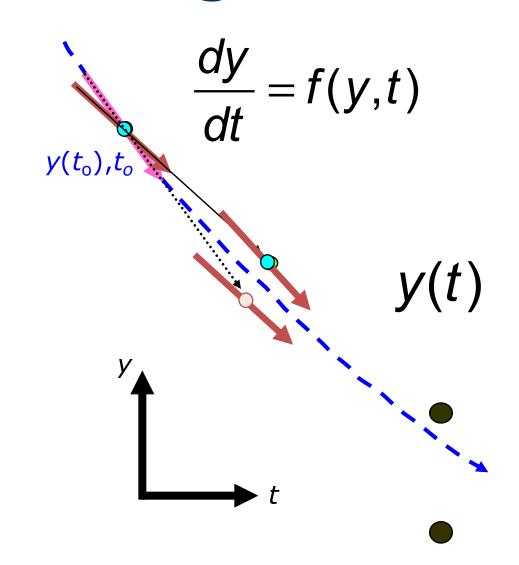


- Cast Number 1
 - The first cast is the same as the 2ndorder.
 - From here, we will cast our second line, not all the way to t+Δt
 - As with the first cast, we'll send it only half way!



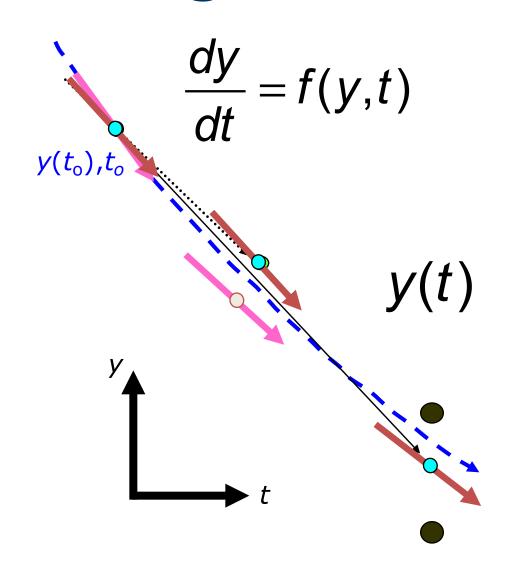


- Cast Number 2
 - Notice that even though this new slope is measured at the same time as the last one. its value can differ from the first "cast"



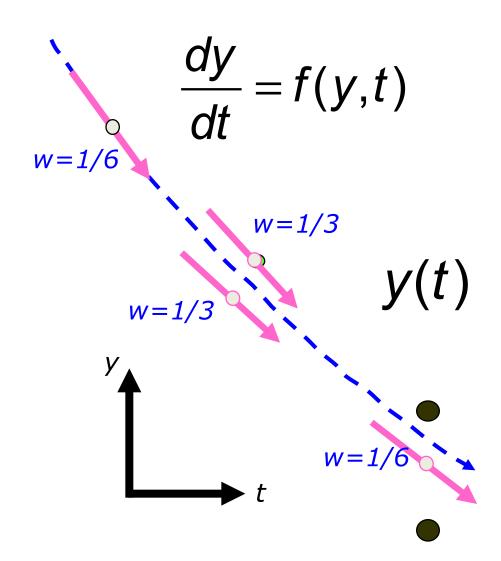


- Cast Number 3
 - Ok, this time we go all the way - but we won't be done yet.
 - Now we have FOUR possible forcings on our system over the full timestep



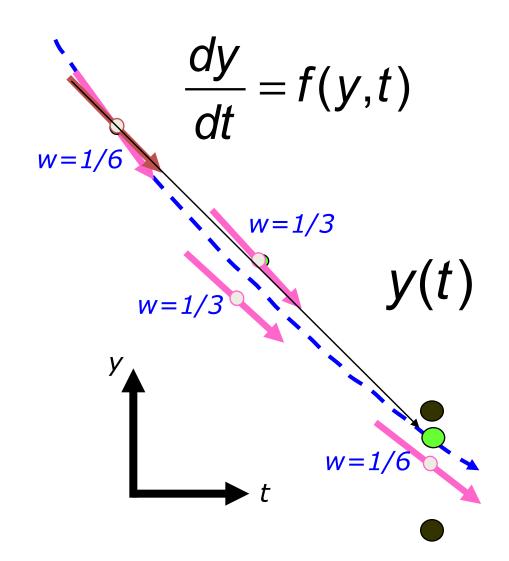


- Creating Cast 4
 - We now have four slopes (possible "forcing trajectories")
 - We'll now take a weighted average of the four noting the larger potential for error for the firstand last-cast slopes

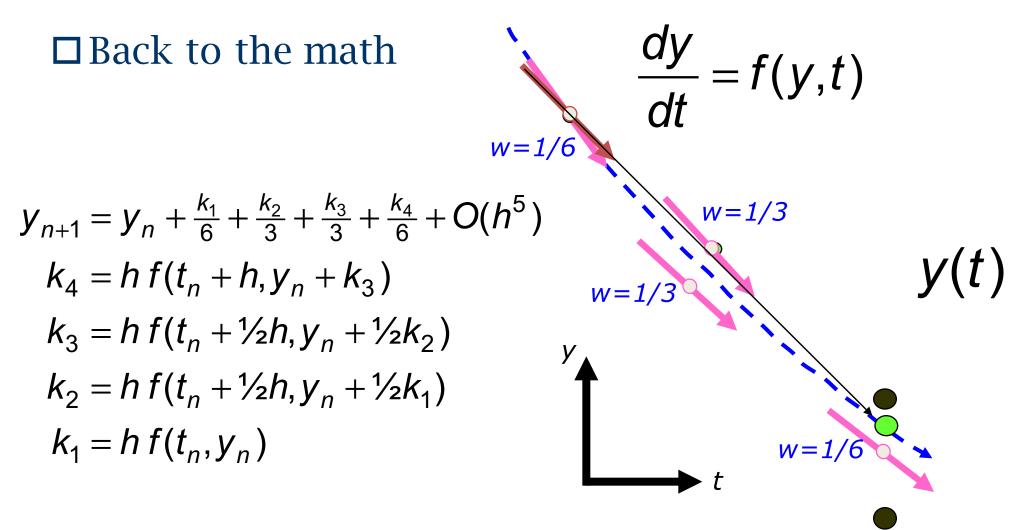




- Creating Cast 4
 - We how have one last cast, this time "aiming" with the weighted mean of the possible forcings.









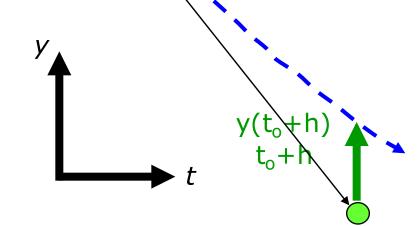
- And for completeness
 - 1^{rst}-Order R-K is just Euler's Method

$$y(t_0), t_0$$

$$\frac{dy}{dt} = f(y,t)$$

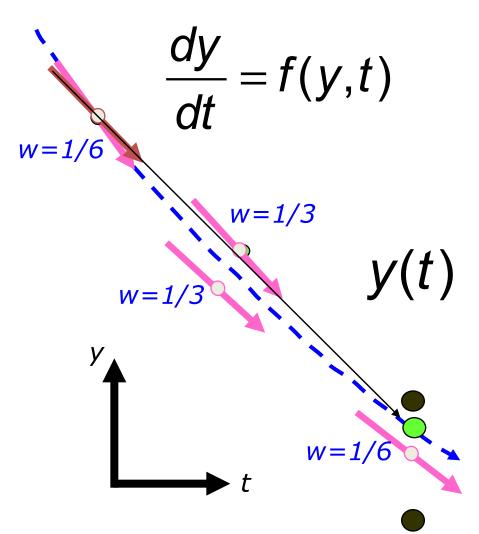
$$y_{n+1} = y_n + k_1 + O(h^2)$$

 $k_1 = h f(t_n, y_n)$



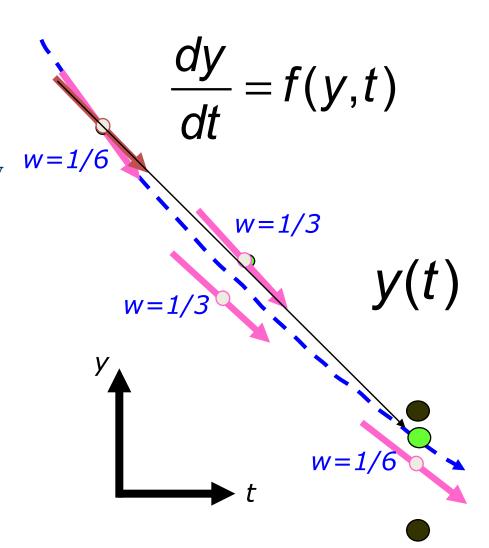


□ Though in this case, we did quite well in going from a Euler approach, through the R-K approaches, higher order approaches are not a guarantee of an improved solution.



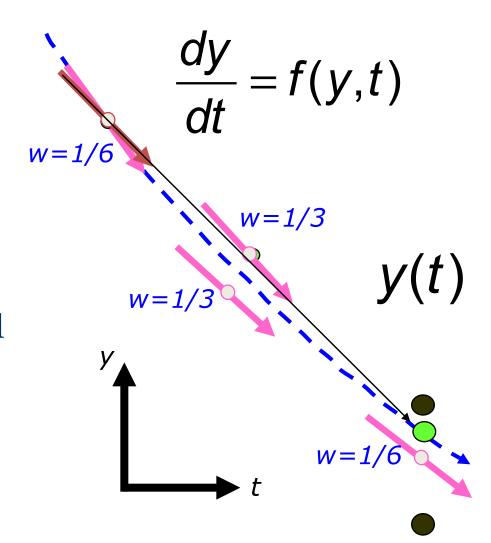


- If you have a smooth continuous function over your time step, R-K should service you fairly well on most occasions if you use the same time step as a Euler approach.
- However, R-K doesn't not liberate you from needing sensible time steps.



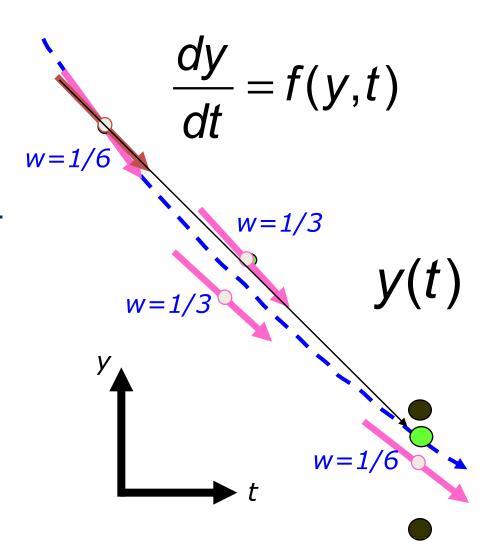


- Stella[™] also doesn't recommend R-K approaches when you use some of their options or when you periodically "shock" your system.
- R-K does provide a good representation of smooth continuous and oscillating systems.





- ☐ But not matter what scheme you use,
- □ ALWAYS DO A TIME-STEP SENSITIVITY TEST!





Let's Play