

# Computational Astrophysics

ASTR 660, Spring 2020  
計算天文物理

Lecture 8

Boundary Value Problems

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# Class website



[https://kuochuanpan.github.io/courses/109ASTR660\\_CA/](https://kuochuanpan.github.io/courses/109ASTR660_CA/)



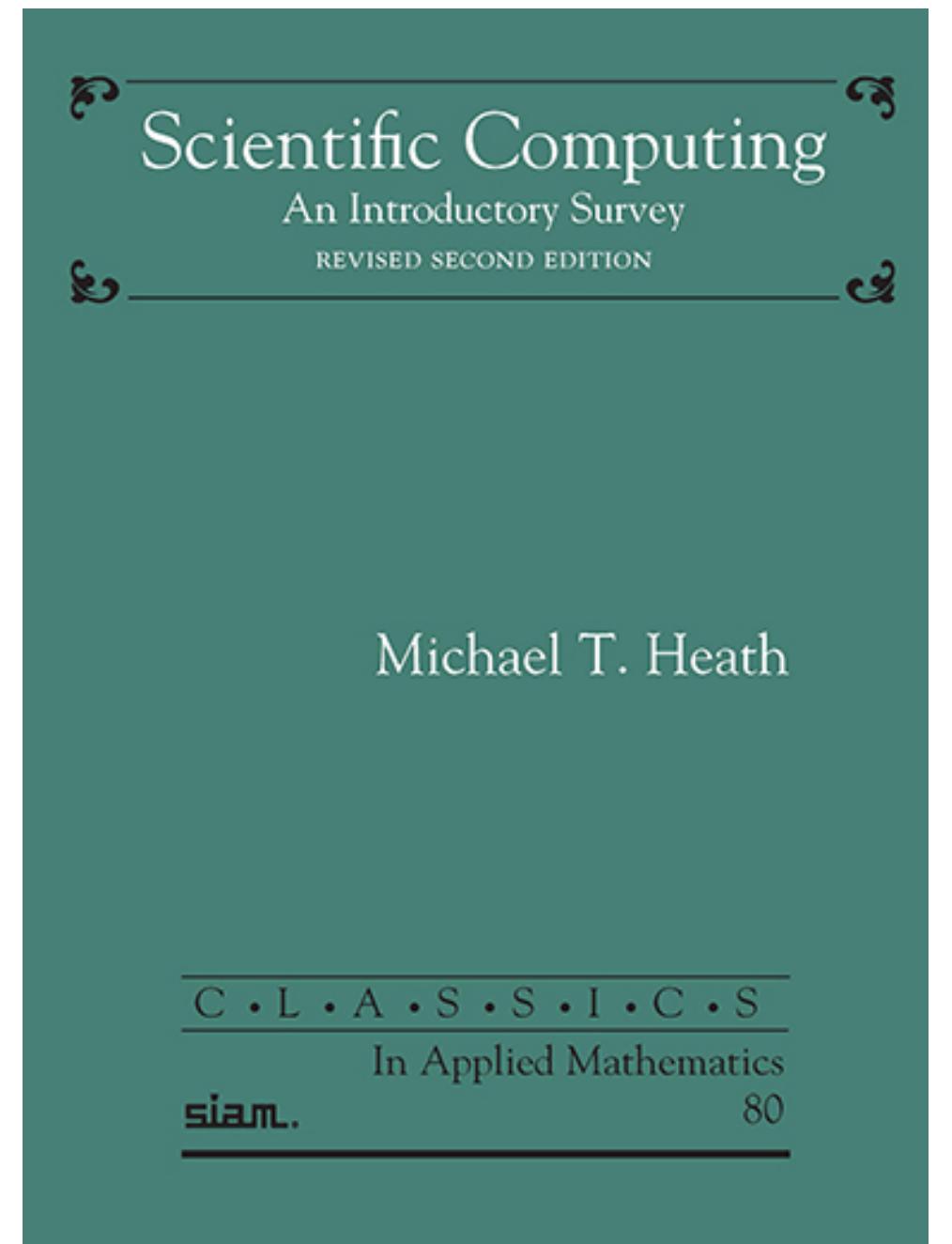
# Plan for today

- Ordinary Differential Equations (ODEs, continues)
- ODE: Boundary Value Problems (BVPs)
- Stellar structure (Polytropes)
- Lab: Polytrope

## Reference:

“Scientific Computing: An introductory survey”, Michael Heath

<https://books.google.com.tw/books?id=f6Z8DwAAQBAJ&hl=zh-TW>





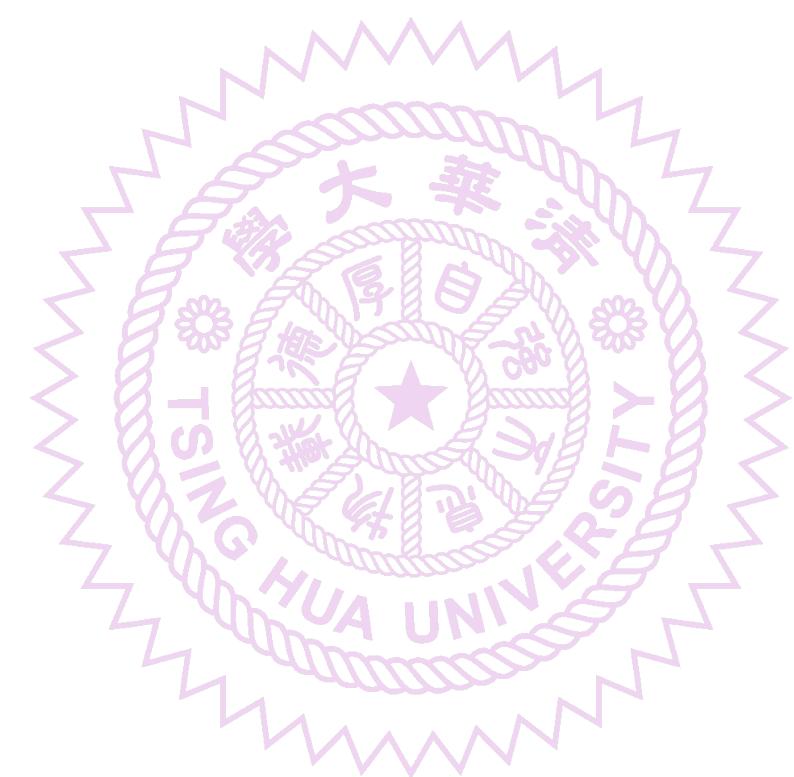
# Boundary Value Problems (BVPs)

# Boundary Value Problem

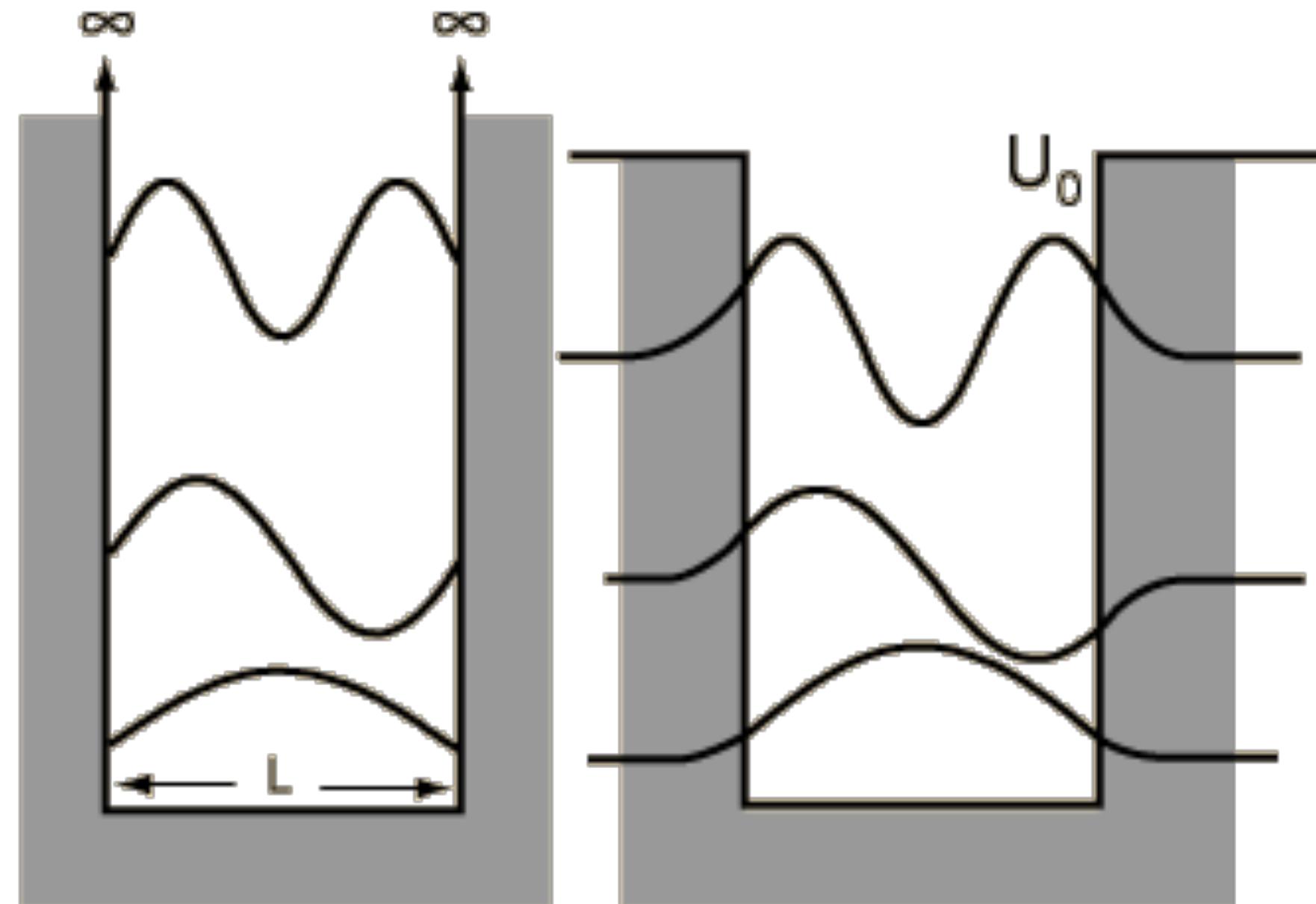


- Side conditions prescribing solution or derivative values at specified points are required to make solution of ODE unique
- For initial value problems, all side conditions are specified at a single point, say at  $t_0$ .
- For **boundary value problem** (BVP), side conditions are specified at more than one point
- K-th order ODE, or equivalent first-order system, requires k side conditions.
- Side conditions are typically specified at endpoints of interval  $[a,b]$ , so we have two-point boundary value problem with **boundary condition** (BC) at  $a$  and  $b$

# Example: Boundary Value Problem

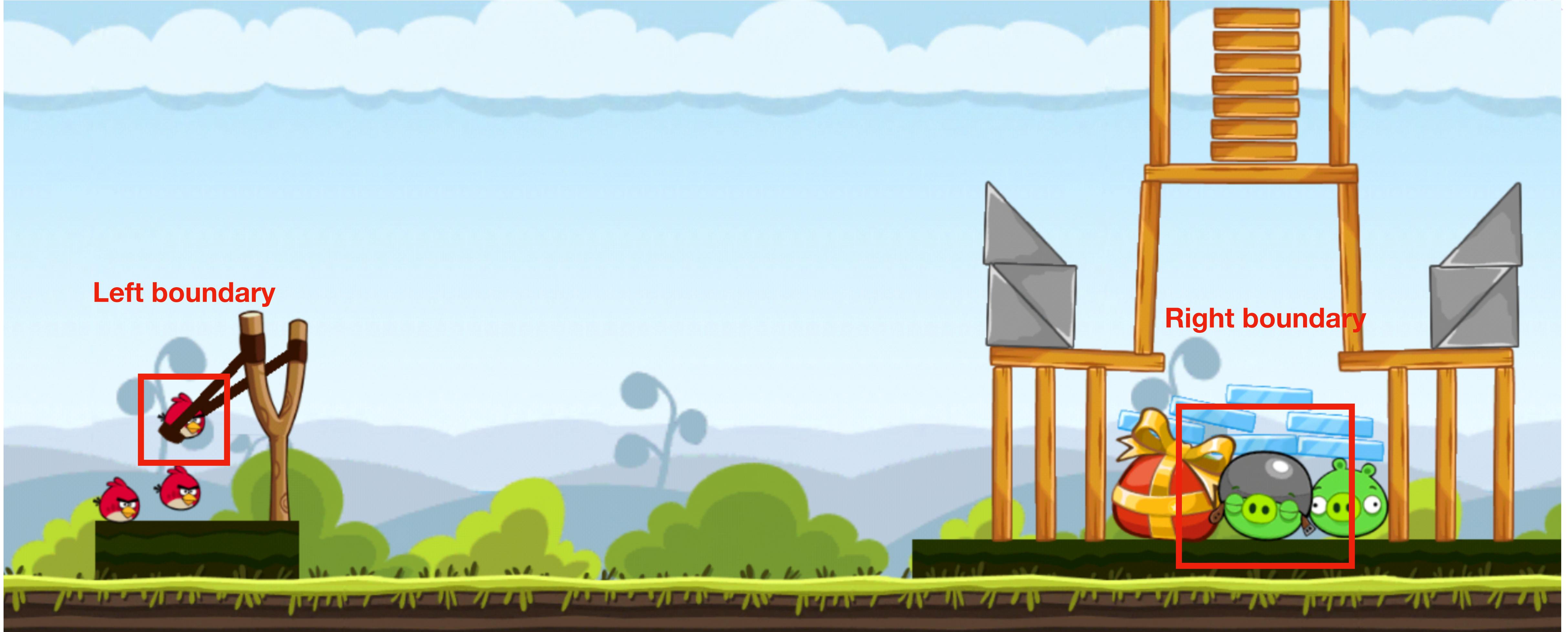
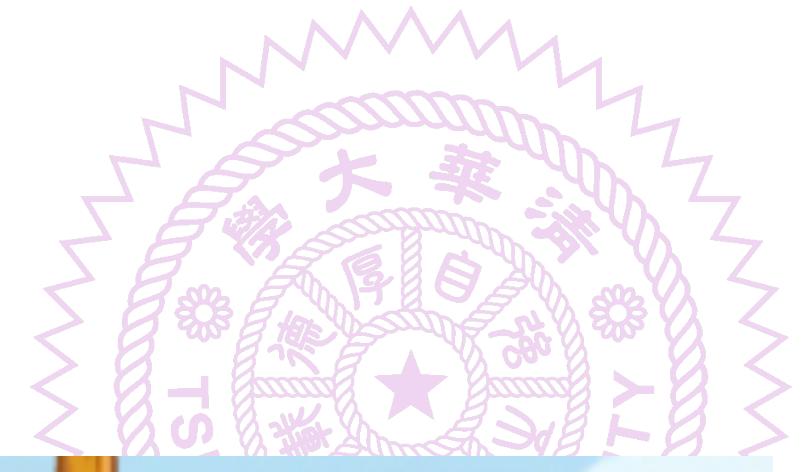


- Schrödinger equation in an infinity potential well



$$-\frac{\hbar^2}{2m} \frac{d^2\varphi}{dx^2} + V(x)\varphi = E\varphi$$

# Example: Boundary Value Problem





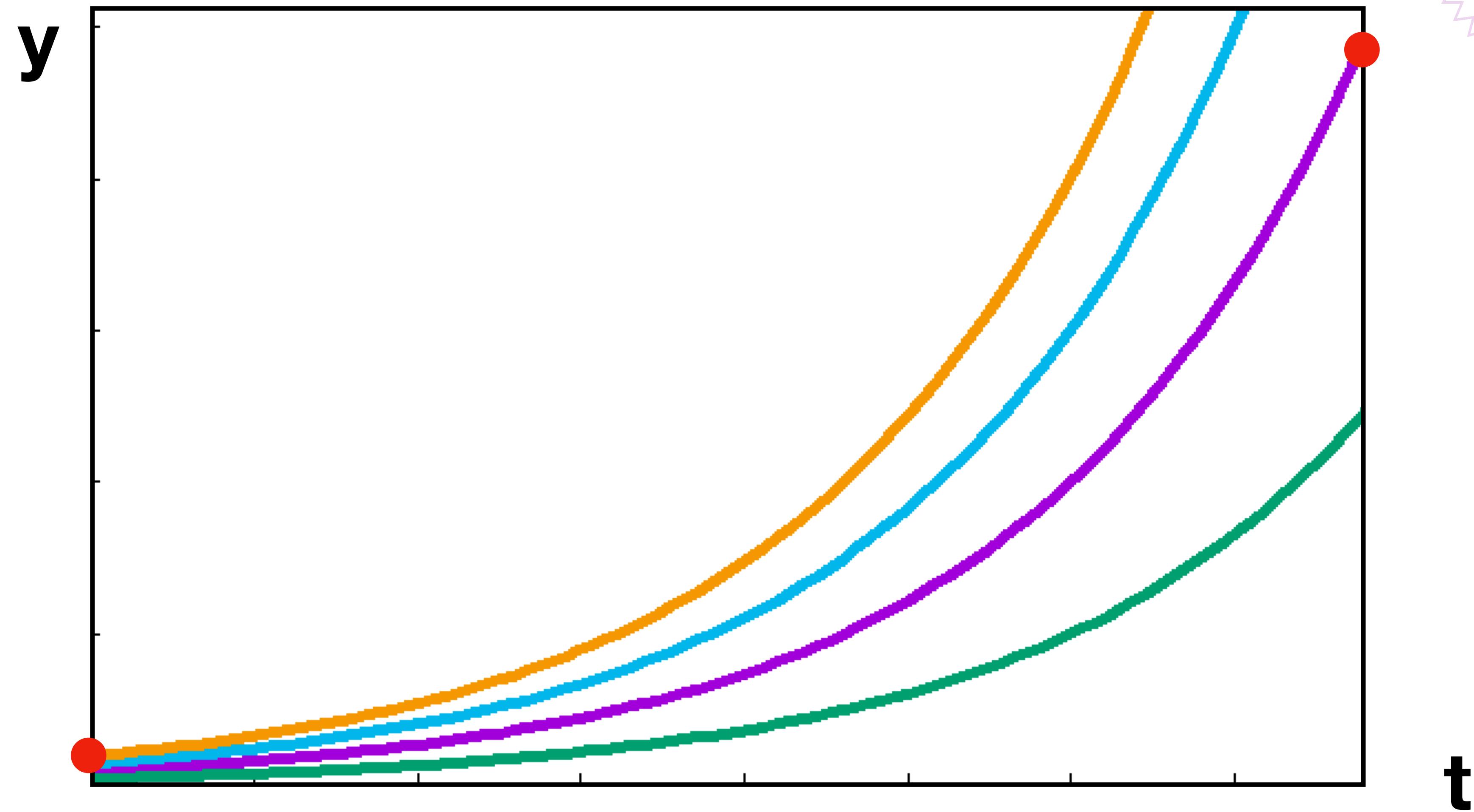
# Shooting method



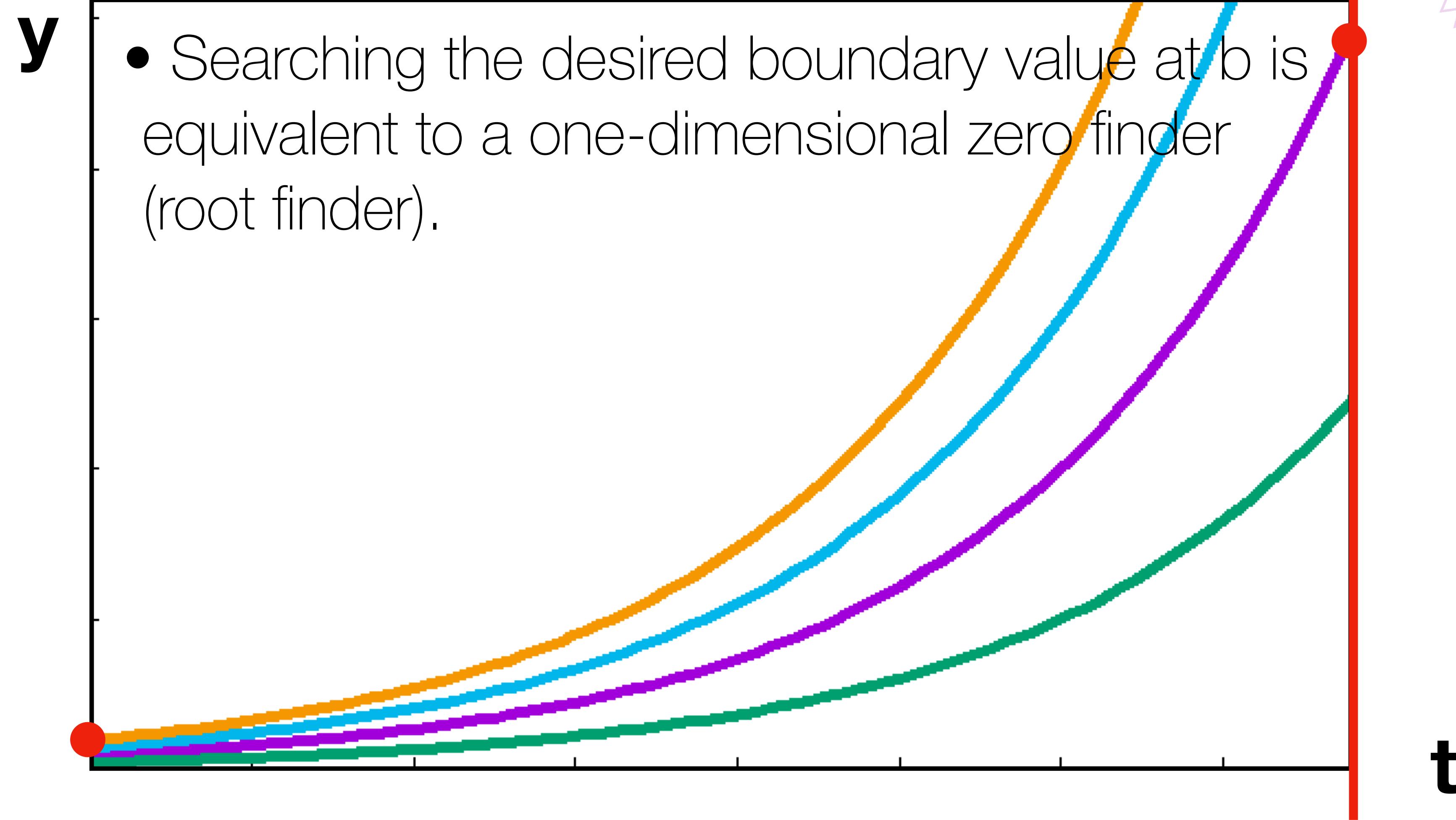
# Shooting method

- If we know the value of  $u(a)$  but lacking  $u'(a)$
- If we also know  $u'(a)$ , then we would have IVP that we could solve be methods discussed in the previous lecture.
- Lacking that information, we could try sequence of increasingly accurate guesses until we find value of  $u'(a)$  such that when we solve resulting IVP, approximate solution value at  $t=b$  matches desired boundary values,  $u(b)$

# Shooting method



# Shooting method



# Exercise: shooting method



- Consider two-point BVP for second-order ODE

$$u'' = 6t \quad 0 < t < 1$$

with BC

$$u(0) = 1 \quad u(1) = 1$$

- For each guess of  $u'(0)$ , we could integrate the resulting IVP using the classical Runge-Kutta method to determine how close we come to hitting desired solution value at  $t=1$

# Exercise: shooting method



- Consider two-point BVP for second-order ODE

$$u'' = 6t \quad 0 < t < 1$$

with BC

$$u(0) = 1 \quad u(1) = 1$$

$$\mathbf{y}'(t) = \begin{bmatrix} y_2 \\ 6t \end{bmatrix}$$

With initial guess  $y_2(0) = 1$



# Exercise: shooting method, part 1

- Solve ODE:  $u'' = 6t$       or       $y'(t) = \begin{bmatrix} y_2 \\ 6t \end{bmatrix}$

with BC       $u(0) = 1$        $u(1) = 1$

- Step 1: Modify the example1.f90 to compute the corresponding IVP with an initial guess of  $y_2(0) = 1$ .
- Step 2: Change the initial guess of  $y_2(0) = -1.5$
- Step 3: Continue vary the initial guess of  $y_2(0)$  until  $u(1)=1$  is satisfied.

# Exercise: shooting method, part 2



- Solve ODE:  $u'' = 6t$       or       $y'(t) = \begin{bmatrix} y_2 \\ 6t \end{bmatrix}$

with BC       $u(0) = 1$        $u(1) = 1$

- Step 1: Copy your example1.f90 to example2.f90
- Step 2: Implement a bisection search for the best  $y_2(0)$
- Step 3: Compare that result with your solution in example1.f90



# Finite Difference Method

# Finite Difference Method



- Finite difference method converts BVP into system of algebraic equations by replacing all derivatives with finite difference approximations
- For example, to solve two-point BVP

$$u'' = f(t, u, u') \quad a < t < b$$

with BC

$$u(a) = \alpha \quad u(b) = \beta$$

We introduce mesh points  $t_i = a + i h$

# Finite Difference Method



- We replace derivatives by finite difference approximations such as

$$u'(t_i) \sim \frac{y_{i+1} - y_{i-1}}{2h}$$
$$u''(t_i) \sim \frac{y_{i+1} - 2y_i + y_{i-1}}{h^2}$$

This yields system of equations

$$\frac{y_{i+1} - 2y_i + y_{i-1}}{h^2} = f(t_i, y_i, \frac{y_{i+1} - y_{i-1}}{2h})$$

to be solved for unknowns  $y_i$

# Finite Difference Method



- For these particular finite difference formulas, system to be solved is **tridiagonal**, which saves on both work and storage compared to general system of equations.
- This is generally true of finite difference methods: they yield sparse systems because each equation involves few variables



# Example: Finite Difference Method

Solve

$$u'' = 6t \quad 0 < t < 1$$

with BC

$$u(0) = 1 \quad u(1) = 1$$

$$\frac{y_{i+1} - 2y_i + y_{i-1}}{h^2} = 6t_i$$

If we take  $h=0.5$ , yielding three points:  $y_0$ ,  $y_1$  and  $y_2$

$$\frac{1 - 2y_1 + 1}{(0.5)^2} = 6t_1 = 3$$

# Example: Finite Difference Method



- In practice, much smaller step size and many more mesh points would be required to achieve acceptable accuracy
- We would therefore obtain system of equations to solve for approximate solution values at mesh points, rather than single equation as in this example

# Exercise: Finite Difference Method



Solve

$$u'' = 6t$$

$$0 < t < 1$$

$$\frac{y_{i+1} - 2y_i + y_{i-1}}{h^2} = 6t_i$$

with BC

$$u(0) = 1$$

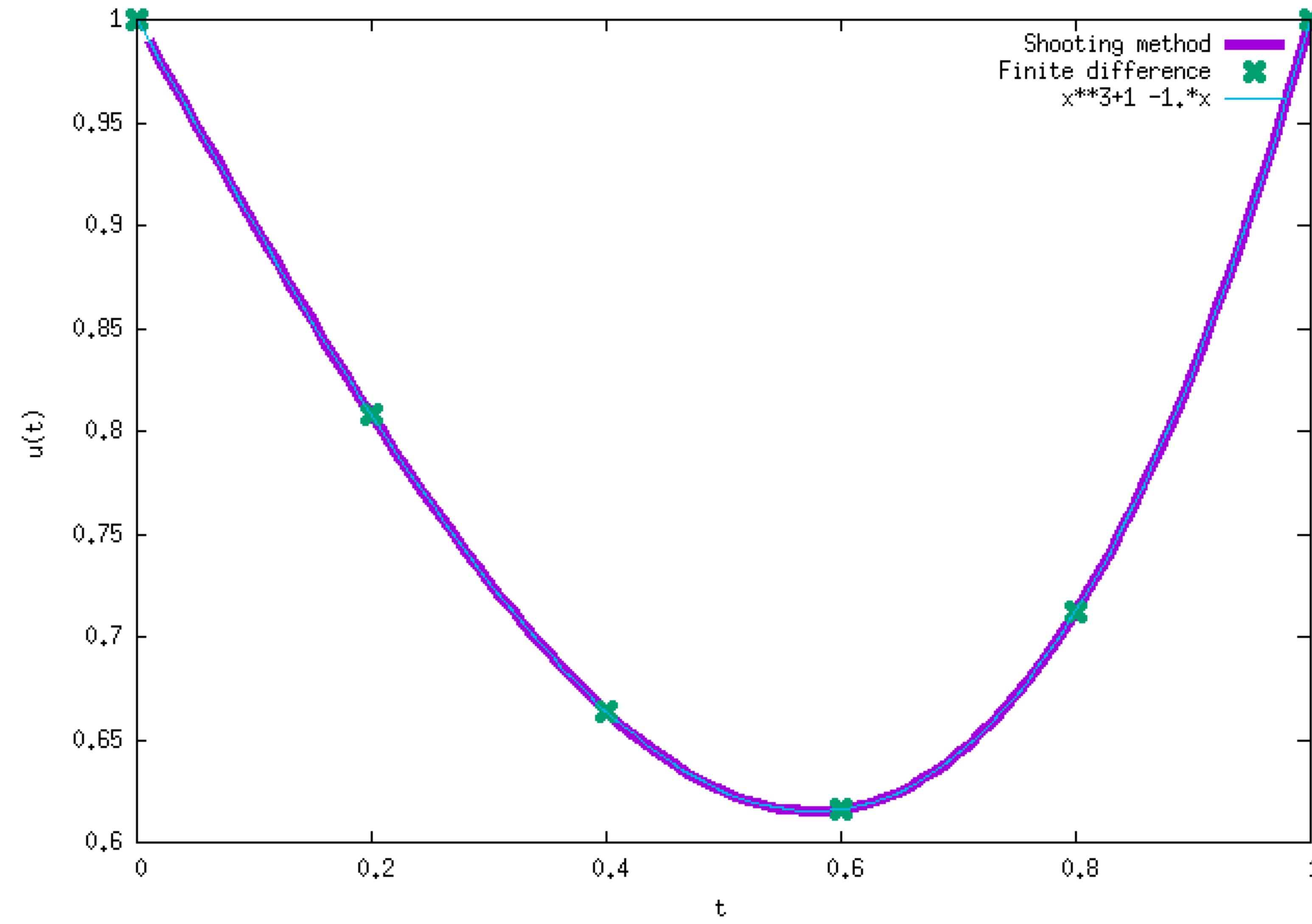
$$u(1) = 1$$

Take  $h=0.2$ , yielding

$$\begin{bmatrix} -2 & 1 & 0 & 0 \\ 1 & -2 & 1 & 0 \\ 0 & 1 & -2 & 1 \\ 0 & 0 & 1 & -2 \end{bmatrix} \begin{bmatrix} y_4 \\ y_3 \\ y_2 \\ y_1 \end{bmatrix} = \begin{bmatrix} 4.8h^2 - 1 \\ 3.6h^2 \\ 2.4h^2 \\ 1.2h^2 - 1 \end{bmatrix}$$

Is a linear system. We know how to solve  $y_i$

# Exercise: Finite Difference Method





# Stellar Structure

# Stellar Structure



$$P = P(\rho, T, X)$$

$$L = L(\rho, T, X)$$

$$\kappa = \kappa(\rho, T, X)$$

$$\epsilon = \epsilon(\rho, T, X)$$

# Stellar Structure



Hydrostatic equation

$$\frac{dP}{dr} = -\rho g = -\rho \frac{GM_r}{r^2}$$

$$dM_r = 4\pi r^2 \rho dr,$$

Mass conservation

$$\frac{dP}{dM_r} = -\frac{GM_r}{4\pi r^4}$$

$$\frac{dr}{dM_r} = -\frac{1}{4\pi r^2 \rho}$$

$$\frac{dL_r}{dM_r} = E_{nuc} - \frac{d\epsilon}{dt} - P \frac{dV}{dt}$$

Energy conservation

# Polytropic EoS and polytopes



- Pressure can be described by

$$P(r) = K\rho^{1+1/n}(r)$$

where  $n$  is the polytropic index (a constant),  
 $K$  is the polytropic constant



# Polytropic EoS and polytopes

- From the hydrostatic equation

$$\frac{dP}{dr} = -\rho g = -\rho \frac{GM_r}{r^2}$$

$$\frac{d}{dr} \left( \frac{r^2}{\rho} \frac{dP}{dr} \right) = -4\pi Gr^2 \rho$$

$$\frac{1}{r^2} \frac{d}{dr} \left( \frac{r^2}{\rho} \frac{dP}{dr} \right) = -4\pi G \rho$$

Which is the Poisson's equation



# Polytropic EoS and polytopes

- Define the dimensionless variable theta by

$$\rho(r) = \rho_c \theta^n(r)$$

The power law for pressure is then

$$P(r) = K \rho_c^{1+1/n} \theta^{n+1}(r) = P_c \theta^{1+n}(r)$$

Substitute these into Poisson's equation

$$\frac{(n+1)P_c}{4\pi G \rho_c^2} \frac{1}{r^2} \frac{d}{dr} \left( r^2 \frac{d\theta}{dr} \right) = -\theta^n$$



# Polytropic EoS and polytopes

- Define the dimensionless radial coordinate  $\xi$

$$r = r_n \xi$$

where we define  $r_n^2 = \frac{(n+1)P_c}{4\pi G \rho_c^2}$

Poisson's equation becomes

$$\frac{1}{\xi^2} \frac{d}{d\xi} \left( \xi^2 \frac{d\theta_n}{d\xi} \right) = -\theta_n^n$$

The Lane-Emden equation



# Polytropic EoS and polytopes

$$\frac{1}{\xi^2} \frac{d}{d\xi} \left( \xi^2 \frac{d\theta_n}{d\xi} \right) = -\theta_n^n$$

The Lane-Emden equation

- If ideal gas EoS  $P = \rho N_A k T / \mu$ , solve for temperature
- Specify K, n and either  $\rho_c$  or  $P_c$ , yields all solutions

$$R = r_n \xi_1 = \left[ \frac{(n+1)P_c}{4\pi G \rho_c^2} \right] \xi_1 \quad T = T_c \theta_n(r)$$

$$P_c = K \rho_c^{1+1/n}$$

$$T_c = K \rho_c^{1/n} \left( \frac{N_A k}{\mu} \right)$$



# Lane-Emden Equation

$$\frac{1}{\xi^2} \frac{d}{d\xi} \left( \xi^2 \frac{d\theta_n}{d\xi} \right) = -\theta_n^n$$

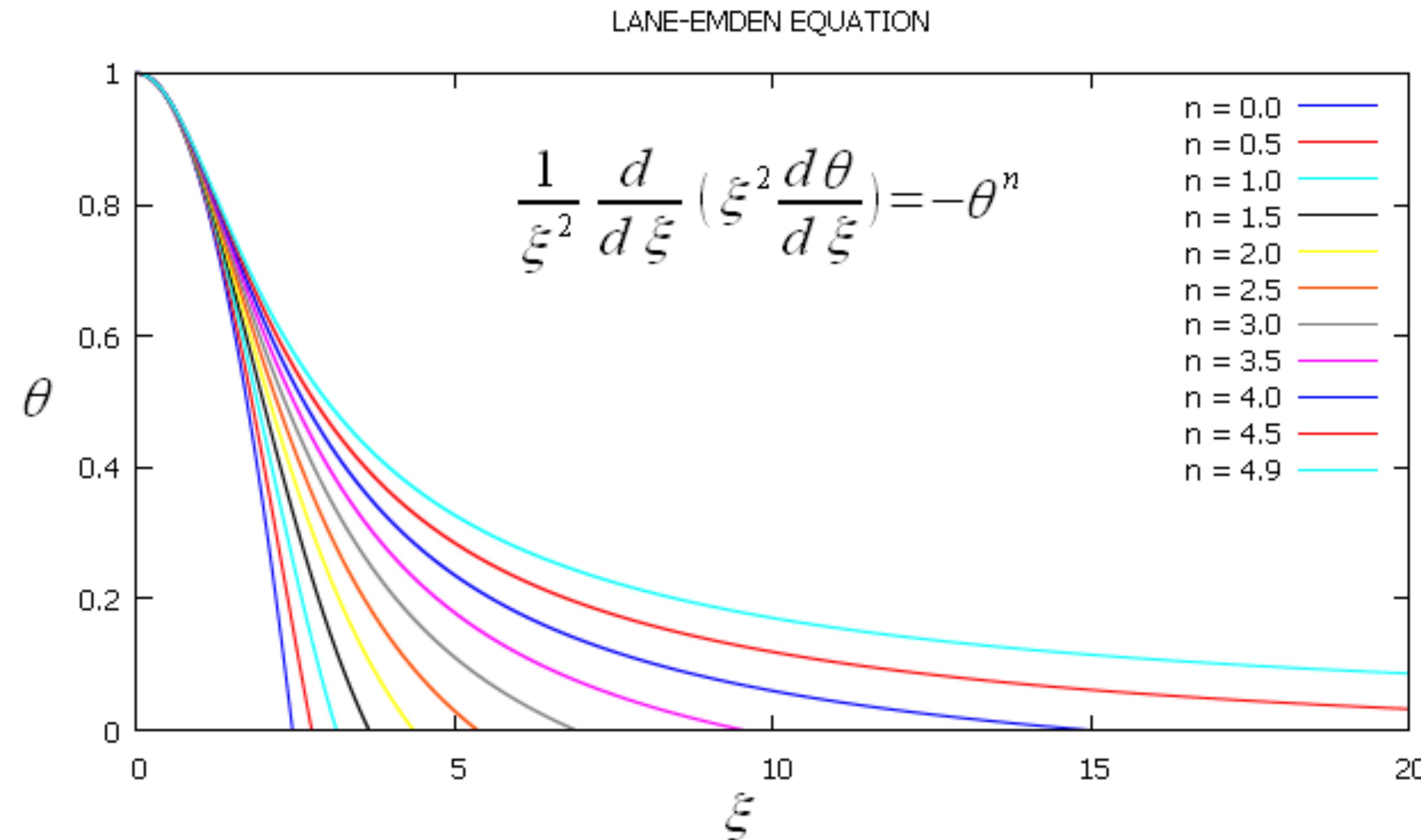
The Lane-Emden equation

- Boundary conditions: density and pressure are 0 at surface. At center, density is central density and pressure is the central pressure  $P_c$

$$\theta_n(0) = 1 \quad \theta'_n(0) = 0 \quad \text{at center}$$

MP !!  $\theta_n(\xi = \xi_1) = 0$  at surface

# Lane-Emden Equation





# Exercise: Polytropes

$$\frac{1}{\xi^2} \frac{d}{d\xi} \left( \xi^2 \frac{d\theta_n}{d\xi} \right) = -\theta_n^n \quad \text{The Lane-Emden equation}$$

$$\theta_n(0) = 1 \quad \theta'_n(0) = 0 \quad \text{at center}$$

$$\theta_n(\xi = \xi_1) = 0 \quad \text{at surface}$$



# Exercise: Polytropes

$$\mathbf{y} = \begin{bmatrix} \theta_n = y_1 \\ \frac{d\theta_n}{d\xi} = y_2 \end{bmatrix} \quad \mathbf{y}' = \begin{bmatrix} y_2 \\ -y_1^n - \frac{2}{\xi}y_2 \end{bmatrix}$$

The Lane-Emden equation

$$\theta_n(0) = 1 \quad \theta'_n(0) = 0 \quad \text{at center}$$

$$\theta_n(\xi = \xi_1) = 0 \quad \text{at surface}$$

- Step1: modify polytrop.f90 and compute the solution with different polytope index ( $n=1, 1.5, 2, 3, 4$ )
- Step2: Visualize the solutions
- Step3: make y-axis in log scale



# Polytropes

- However, the surface of the star could be very sharp
- IVP requires step size adjustment to resolve the envelope of the star
- In realistic cases, BCs of a stellar model might come from the surface region.



# Exercise: Polytropes, Part 2

$$\mathbf{y} = \begin{bmatrix} \theta_n = y_1 \\ \frac{d\theta_n}{d\xi} = y_2 \end{bmatrix} \quad \mathbf{y}' = \begin{bmatrix} y_2 \\ -y_1^n - \frac{2}{\xi}y_2 \end{bmatrix}$$

The Lane-Emden equation

$$\theta_n(0) = 1 \quad \theta'_n(0) = 0 \quad \text{at center}$$

$$\theta_n(\xi = \xi_1) = 0 \quad \text{at surface}$$

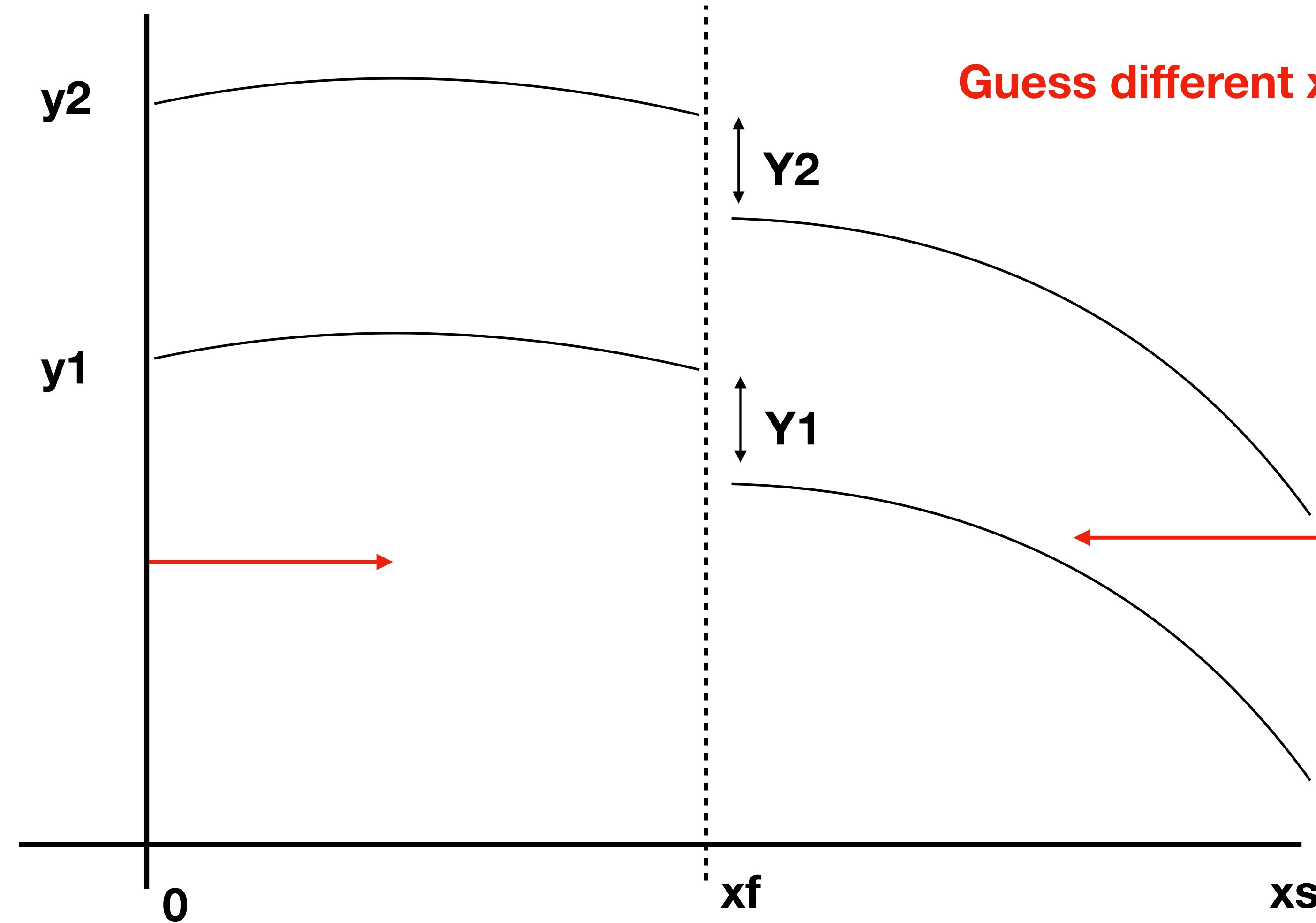
- Step1: copy your poltrope.f90 to polytrop2.f90
- Step2: set the polytrope index  $n = 1$
- Step3: Use shooting method to integrate from the surface (assuming we know  $\xi_1$  at surface is  $\pi$ )



# Lane-Emden Equation

- However, the surface of the star could be very sharp
- A better way to solve Lane-Emden equation is starting from both the center and the surface and try to match at some point -> Fitting method (**multiple shooting**).

# Lane-Emden Equation: Fitting method





# Lane-Emden Equation: Fitting method

$$\begin{aligned} Y1(x_s, y_{2s}) &= y_{1i}(x_f) - y_{1,0}(x_f) \\ Y2(x_s, y_{2s}) &= y_{2i}(x_f) - y_{2,0}(x_f) \end{aligned}$$

$$x_s \rightarrow x_s + \delta x_s$$

$$y_{2s} \rightarrow y_{2s} + \delta y_{(2s)}$$

- Becomes an eigenvalue problem



# Lane-Emden Equation: Fitting method

- Standard eigenvalue problem for second-order ODE has form

$$u'' = \lambda f(y, u, u') \quad a < t < b$$

with BC

$$u(a) = \alpha \quad u(b) = \beta$$

- Scalar lambda is eigenvalue and solution u is corresponding eigenfunction for this two-point BVP

# Summary

- Two points BVP for ODE specifies BC at both endpoints of interval
- Shooting method replaces BVP by sequence of IVPs, with missing initial conditions
- Finite difference method replaces derivatives in ODE by finite difference.

# Problem Set 7



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# Next lecture

- PDE: Hyperbolic systems

