

*Veritatem inquirenti, semel in vita de omnibus,
quantum fieri potest, esse dubitandu:*

*In order to seek truth, it is necessary once in the course of our life, to
doubt, as far as possible, of all things.*

- Descartes, Rene, *Principles of Philosophy*

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Chapter 1

Introduction

“In the beginning the Universe was created. This has made a lot of people very angry and been widely regarded as a bad move”

- Douglas Adams, *The Restaurant at the End of the Universe*

A Preface on Humanity and the Climate

The development of humanity is not unlike the chirography of an Aristotelian tragedy. It starts with a simple/primitive species cradling a noble cause - to improve their chances of survival. Here the protagonist (humankind) develops a fatal flaw: an insecurity and latent distruction of their home due to a sudden rise to power. Having acknowleged this flaw, we now strive to imporve our understanding of the universe, correct past mistakes and stem the tide of inevitable change.

With tragedy being an imitation not of humanity, but of action and life, happiness and misery, it is only expected that such a comparison to our current affairs should stir feelings of catharsis when exploring our need for research and scientific advancement. It is with that I begin this thesis with the begining of the planet, its atmosphere and consequently the beginning of humankind.

1.1 Whence

This section describes the intial formation of an atmosphere, how this led to life, and ultimately the human race.

1.1.1 Formation of the Atmosphere

4.5 billion years ago the earth was part of a disk of dust and gas orbiting our sun. As these gasses move about, resonant drag instability led to the clumping of dust particles, Hopkins and Squire [2018]; Woo [2018]. As these ‘clumps’ become denser, other forces come in to play, further increasing the size - eventually forming the hot mix of gas and solid which became the Earth.

As it cooled, this Earth becan to accumelate an atmosphere of primordial gasses from the vollatile componenets of the gas cloud. These gasess were then supplemented through outgassing (volcanic eruptions). At this point in time oxygen was not only absent in the atmopshere, but also had many siks within the Earths anoxidised crust. It was not until oxygenic photosynthesis (Peretó [2011]) that the concentrations of oxygen in the atmosphere started to increase. Eventually the development of multicellular cyanobacteria¹ resulted in biologically induced oxygen accumelating in the atmosphere, University of Zurich [2013]. This led to the most significant climate event in the planets history: the Great Oxigenation Event (2.5 billion years ago), Planavsky et al. [2014]. This increase of oxygen allowed oragnisms to become larger and more active, eventually resulting in the human race.

¹The phylum of phtosynthetic prokaryotic (cells not containing a distinct nucleus) bacteria - e.g. blue-green algae

1.1.2 Rise of the Homo Spiens ('Wise Man')

About x million years ago there were many varieties of the homo genus. With the development of the human brain, energy transfer changed. A larger brain required more fuel, and therefore with the development of cooking² humans were able to increase their...

This led to the first know source of indoor air pollution.

With the this increased capability, a language capable of communicating information, allowing for the ability to not only hunt larger prey but also. Ability to metaphorical, allowed fruther knowplege transfer , cvave paings and metaphoirical for people over 150

REFERENCES TO OTHER CHAPTERS... - vis - accounting via metaphors - and an interest in science, and atmosphere

1.2 Motivation

Climate change and air quality have been prominent features of many recent current affair stories[LONG LIST OF REF animals pollution , etc.]. Anthropogeninc in fluence as air qulaity pollicy kingxx

1.3 Ozone and its role

1.4 Changing Climate

The main removal

1.4.1 HOx Cycle

1.5 Air Pollution

VOC

²The first known case of indoor air pollution

1.6 Modelling the Earth

1.6.1 Global Models

Earth Simulation Models (ESM) are models which predict past or future interactions of the planetary system. They represent our foremost understanding of the complex interplay between land-surface (geosphere), ocean (hydrosphere), ice (cryosphere) and the air (atmosphere). These act as a surrogate to manual experimentation, which is just not possible on the global scale, and enable us to access changes, situations and predict the future.

Earth simulation models often have to address many problems relating to computational efficiency.

1.6.2 The Atmosphere

1.6.2.1 Types of model

3d transport emission/deposition chemical and physical processes

1.6.2.2 Chemical Mechanisms

The atmosphere consists of thousands of species, with tens of thousands of reactions between them. Since manual experimentation is not possible on both the macro (every location on earth) or micro (every atom) scale for the Earth System any prediction of these requires the use of a numerical model.

These models represent real world reactions

In modelling these we can describe their rate of production and loss with respect to the species they react with.

1.6.3 The model development cycle

Scientific understanding is the product of many cycles of trial and error, Figure 1.1. In atmospheric chemistry we start with a hypothesis or a question, e.g. will changing X have a negative response on Y. We then construct a theoretical model to represent the chemistry within. This chemistry is updated to reflect the rates and reactions that have been recorded in laboratory/chamber experiments. This cycle is then repeated until the model and real-world observations produce a comparable result.

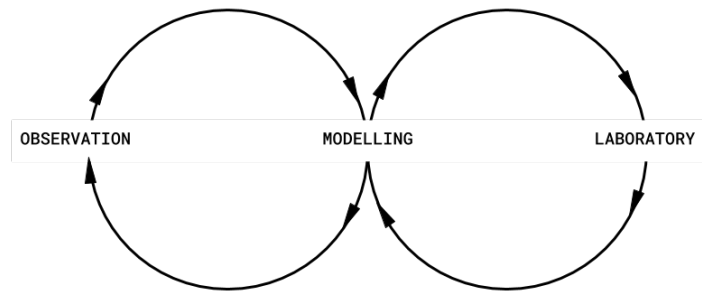


Figure 1.1: **The scientific development cycle.** This shows the iterative nature between modelling, observation and laboratory experimentation

ESM

A series of box models.

1.6.4 The Dynamically Simple Model of Atmospheric Chemical Complexity

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