

# utils.py

---

## Contents

- [utils.py](#)
  - [Contents](#)
  - [1. GetSFR](#)
    - [Parameters](#)
    - [Returns](#)
  - [massFunction](#)
    - [Parameters](#)
    - [Local](#)
    - [Global](#)
    - [Returns](#)
    - [Warning](#)
  - [ImportOriData](#)
    - [Parameters](#)
    - [Returns](#)
    - [Examples](#)
  - [DELallrbwind](#)
    - [Parameters](#)
    - [Returns](#)
    - [Criterion](#)
    - [Warning](#)
  - [Beamingfactor](#)
    - [Parameters](#)
    - [Returns](#)
  - [OBSplot](#)
    - [Parameters](#)
    - [Returns](#)
  - [plot\\_Lx\\_mesaNS](#)
    - [Parameters](#)
    - [Returns](#)
    - [Warning](#)
  - [plot\\_Lx\\_b\\_mesa](#)
  - [plot\\_Lx\\_NNsum\\_gedian](#)
    - [Parameters](#)
    - [Returns](#)
  - [plot\\_lx\\_tb\\_gedian\\_paper](#)
    - [Parameters](#)
    - [Returns](#)
  - [plot\\_mass\\_tb\\_gedian\\_paper](#)
    - [Parameters](#)
    - [Returns](#)
  - [SETplot](#)

## GetSFR

Calculate the Binary Star forming rate of this program

### Parameters

SFR: float, original star forming rate obtained by observation. MminMINE: float or double, minimum higher initial mass. alpha: float or double, the exponent of mass function, which is different according to different version of mass function. qav: float, average mass ratio (q) of popbin program (default: 0.535) binary: int, Whether your SFR is for binary or all stars. ONLY 0 or 1! MminSFR: float, the minimum mass taken into account while calculating SFR.

### Returns

Sb: float, Binary Star forming rate of this program

## massFunction

Calculate the birth rate of particular binary system.

### Parameters

#### Local

mass: Double or float, the **higher** initial mass of binaries system.

#### Global

AllmassNUM: int, total mass number of Popbin. M1MAX: float or double, maximum higher initial mass M1MIN: float or double, minimum higher initial mass SFR: float or double, binary star forming rate qNUM: int, total q (mass ratio) number of Popbin. aNUM: int, total a (separation) number of Popbin. alpha: float or double, the exponent of mass function, which is different according to different version of mass function.

### Returns

rate: float, the birth rate of particular binary system.

### Warning

This function only support initial mass heavier than 1 solar mass! Otherwise the birth rate will be None!

## ImportOriData

Import original data generated by Popbin\_mine.f in ../Fortran0415, and delete some unrelated data. Finally output two array in the structured form.

Please make sure your data files saved in

`{DataFolder}/{OutName}/all_Lx37erg_z0.01_rb_{OutName}.out`

### Parameters

NumParallel : int, the same as the setting in popbin program, (default: 1) DataFolder : string, the data saving folder of original output of popbin program. OutName : string, the same as the setting in popbin program, which is the version of popbin program. SaveNPY : bool, whether this program save the all\_rb and all\_wind array in NPY form in the same file folder. LoadNPY : bool, whether this program load the all\_rb and all\_wind array in NPY form in the same file folder. Please set it to **False** in the first time. NPY form can speed up the loading progress significantly.

## Returns

RLselected: 2-D array, the meaning of each column are as following. The output array have been selected by **DELallrbwind** function. Windseleted: 2-D array, the meaning of each column are as following. The output array have been selected by **DELallrbwind** function.

Column	0	1	2	3	4	5	6	7	8	9
all_wind_1/WindSelected	m1	m2	Lx	ecc0	a	tb0	mt1	mt2	kw	kw2
all_rb_1/RLselected	m1	m2	Lx	ecc0	a	tb0	mt1	mt2	kw	kw2
Column	10	11	12	13	14	15	16	17	18	19
all_wind_1/WindSelected	mx	mx2	eccx	tbx	f	nnn	uuu	m_tr	i	delta T
all_rb_1/RLselected	mx	mx2	eccx	tbx	f	m_transferrate	b_iso	m_tr	i	delta T

## Examples

```
#!/usr/bin/env python3
from utils import *
all_rb_1,all_wind_1 = ImportOriData(
    NumParallel      = 8,
    DataFolder       = '..\Data',
    OutName          = '0415',
    SaveNPY          = True,
    LoadNPY         = False)
all_rb_1,all_wind_1 = ImportOriData(
    NumParallel      = 1,
    DataFolder       = '..\Data',
    OutName          = '0415',
    SaveNPY          = False,
    LoadNPY         = True)
```

## DELallrbwind

Caculate the Binary Star forming rate of this program

## Parameters

all\_rb: 2-D array, original data of Roche-lobe binaries obtained from Fortran0415 `popbin_mine.f` program.  
 all\_wind: 2-D array, original data of Wind Roche-lobe binaries obtained from Fortran0415 `popbin_mine.f` program.  
 minLx: float or double, minimum luminosity of selection. (default: 39) CHAbeamingFlag=False ,  
 eddfac=1e4, DELtmax = 200, MaxDuration = 1 CHAbeamingFlag: bool, Whether changing beaming mode.  
 ABANDONED! (default: False) eddfac: double, Maximum edding factor of selection (default: 1e4) DELtmax:  
 float, Maximum evolution time of selection (default: 200 for Ring Galaxies) MaxDuration: float, Maximum  
 stage duration of selection. In other words, the duration between two row of original bcm array in bse may be  
 longer than several million year which is unreliable for rapid mass transfer. (default: 1)

## Returns

RLselected: 2-D array, Selected array Windseleted: 2-D array, Selected array

## Criterion

We delete following data which follows those criterion.

- less than Maximum evolution time of selection
- mass1 is smaller than 2Msun, which is unstable.
- orbit is 0
- $\Delta T < 0$
- $\Delta T > \text{MaxDuration}$
- Lx is lower than Lxmin

## Warning

Please never set eddfac smaller than  $10^4$  or set CHAbeamingFlag to True untill you understand this part of code!

## Beamingfactor

Different mode for beaming. The larger b is, the realistic the result is (maybe).

## Parameters

b: float, beaming factor. beaming: int, mode for caculation of beaming factor. The larger b is, the realistic the result is (maybe).(Default: 1)

- 0 无修正：数目不做调整
- 1 直接乘以b：数目乘以b=观测的数目
- 2 方法二：像我昨天下午发的修正，乘以 $(8*b)^{0.5/\pi}$
- 3 方法三，调整了一些估计项 $\arccos(1-b)*2/\pi$

## Returns

Beaming factor

## OBsplot

Plot the observation result of Wolter 2018 in seven Ring Galaxies. [2.6,2.5,2.3,20,8.0,4.1,4] |num|Ring Galaxies|SFR| | :---: | :---: | |1|AM0644x3 |2.6| |2|Arp148x1 |2.5| |3|Arp143 |2.3| |4|Cartwheel |20.0| |5|N922 |8.0| |6|Arp 147-Ring |4.1| |7|Arp 284 |4.0|

## Parameters

logFlag: {bool,0,1}, whether return a log10(num) result or not. obs: (local) 1-d array, the original observation data.

## Returns

This function return a tuple obsort: array, the luminosity of each ULXs in seven Ring Galaxies in order. num: 1-D array, the number of ULXs more luminous than each luminosity of obsort.

```
num = np.linspace(len(obsort),1,len(obsort))
```

## plot\_Lx\_mesaNS

Different mode for beaming. The larger b is, the realistic the result is (maybe).

## Parameters

beaming: float, beaming factor. tmax: float, maximum evolution time for bianries. Number\_XLFpoint: int, the length of the array while plotting XLF. minLx: float, minimum luminosity of XLF maxLx: float, maximum luminosity of XLF fileFolder: path, the saving path of mesa data

## Returns

NNsumwind: array, the number of predicted ULXs via Roche-Lobe overflow by mesa program which is more luminous than particular luminosity. NNsumrb: array, the number of predicted ULXs via Wind Roche-Lobe overflow by mesa program which is more luminous than particular luminosity.

## Warning

make sure that there is Lx\_N1207.npy and hanchen\_first\_neutron1124.out in your file folder!

## ~~plot\_Lx\_b\_mesa~~

Abandoned!

## plot\_Lx\_NNsum\_gedian

Generate data for plotting the X-ray Luminosity function(XLF).

## Parameters

data: 2-D array, in the same form with all\_rb\_1 or all\_wind\_1, and could be part of them. beaming: float, beaming factor. Number\_XLFpoint: int, the length of the array while plotting XLF. minLx: float, minimum luminosity of XLF maxLx: float, maximum luminosity of XLF

## Returns

(NNsum,Lx\_x,np.log10(NNsum)) NNsum: 1-D array, the number of predicted ULXs which is more luminous than particular luminosity (Lx\_x) obtained from the input data. Lx\_x: 1-D array, particular luminosity (Lx\_x)

```
Lx_x=np.linspace(minLx,maxLx,Number_XLFpoint)
```

## plot\_lx\_tb\_gedian\_paper

Generate data for plotting the distribution of luminosity-orbit plane.

### Parameters

data\_ori: 2-D array, in the same form with all\_rb\_1 or all\_wind\_1, and could be part of them. beaming: float, beaming factor. Number\_Lx: int, the length of the array related to luminosity. Number\_orb: int, the length of the array related to orbit. my\_title: string, title of this figure Limit\_cri: number, the minimum distribution for showing in the plot. For example, if Limit\_cri=1000 and the maximum distribution is 1. then the plot will not display distribution less than 1/1000.

### Returns

NN: 2-D array, the distribution number of each panel.

## plot\_mass\_tb\_gedian\_paper

Generate data for plotting the distribution of mass-orbit plane.

### Parameters

data\_ori: 2-D array, in the same form with all\_rb\_1 or all\_wind\_1, and could be part of them. beaming: float, beaming factor. Number\_mass: int, the length of the array related to mass. Number\_orb: int, the length of the array related to orbit. my\_title: string, title of this figure Limit\_cri: number, the minimum distribution for showing in the plot. For example, if Limit\_cri=1000 and the maximum distribution is 1. then the plot will not display distribution less than 1/1000.

### Returns

NN: 2-D array, the distribution number of each panel.

## ~~SETplot~~

Abandoned!