



What Can Elves Tell Us About Very Strong Lightning?



William Daniels

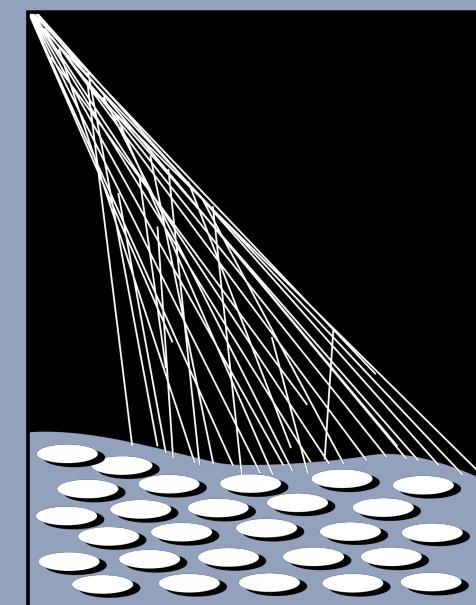
Kevin-Druis Merenda

Lawrence Wiencke

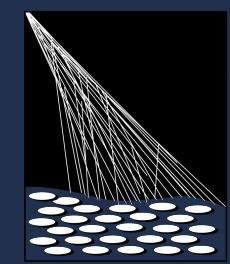
APS April Meeting

Denver, CO

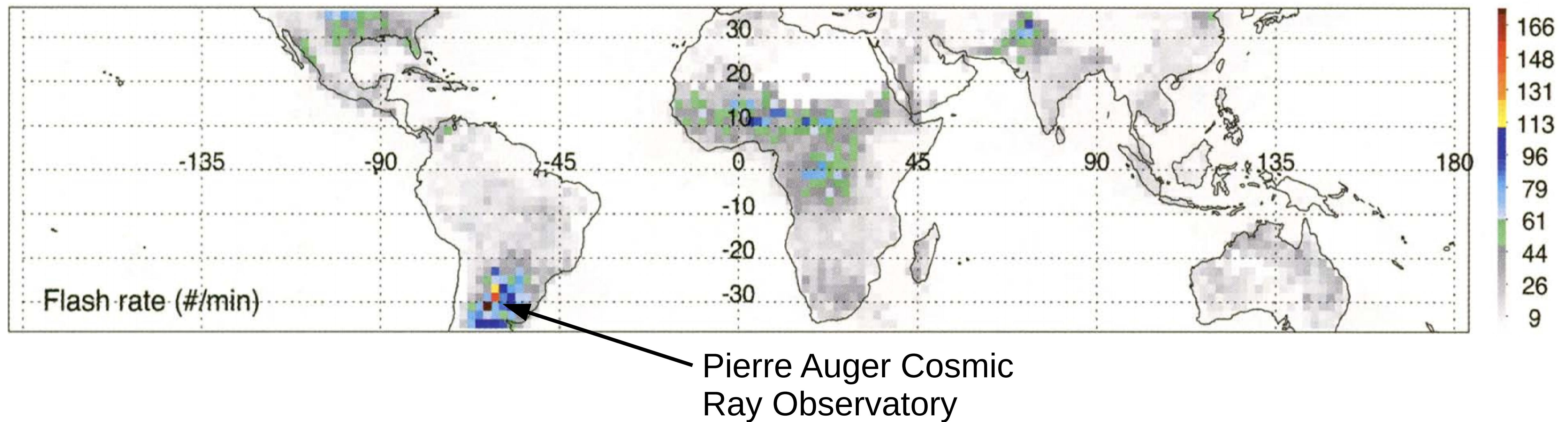
April 14, 2019



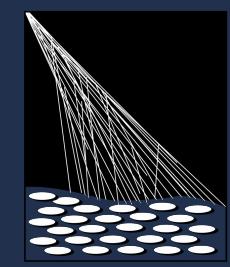
PIERRE
AUGER
OBSERVATORY



- 1) Argentina has some of the world's strongest lightning.
- 2) The Pierre Auger Cosmic Ray Observatory happens to monitor this region.
- 3) Lightning can be dangerous, so it is important to study.

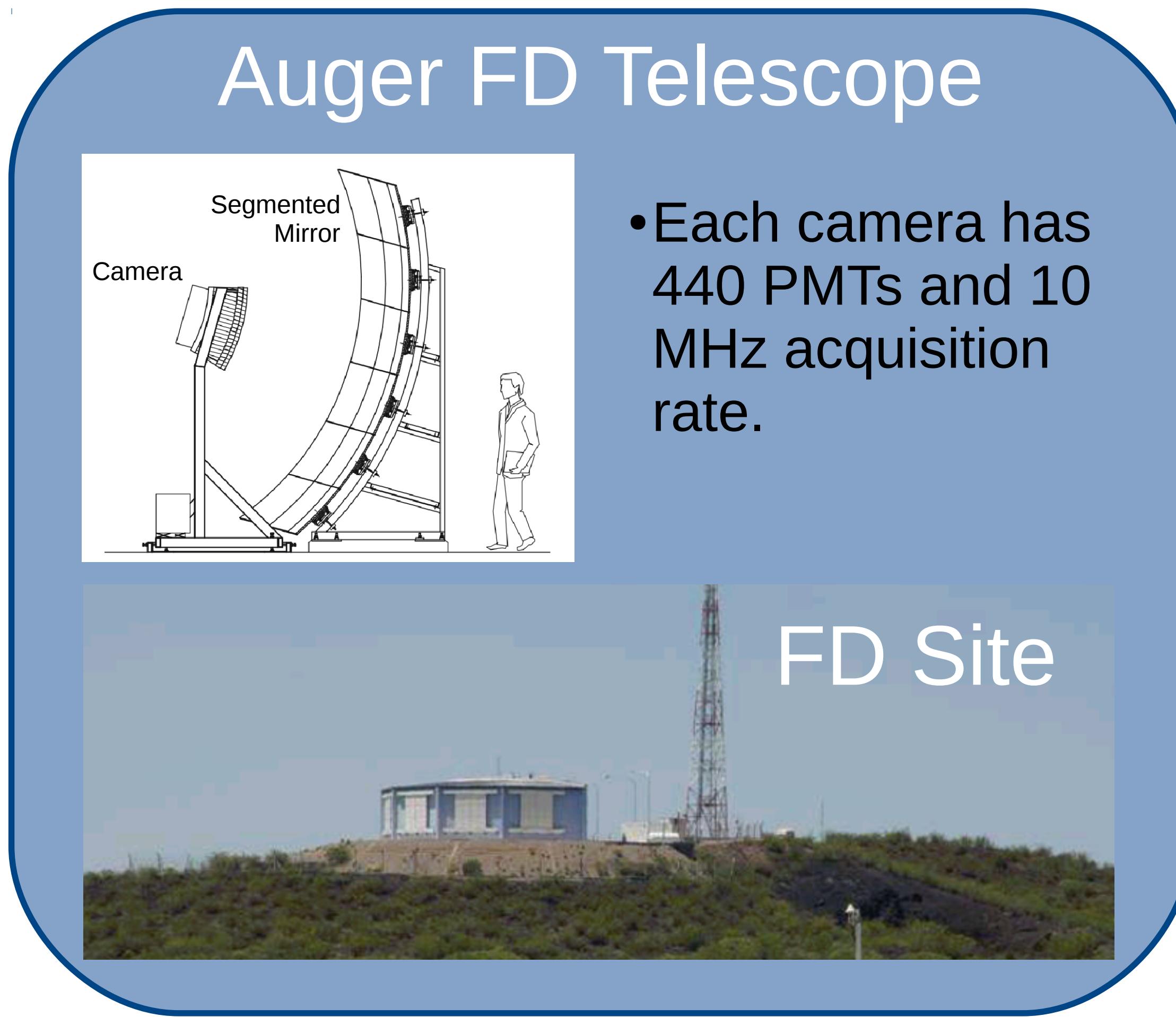


[1] E. J. Zipser et al., "WHERE ARE THE MOST INTENSE THUNDERSTORMS ON EARTH?," Bull. Am. Meteorol. Soc., vol. 87, no. 8, pp. 1057–1072, Aug. 2006.

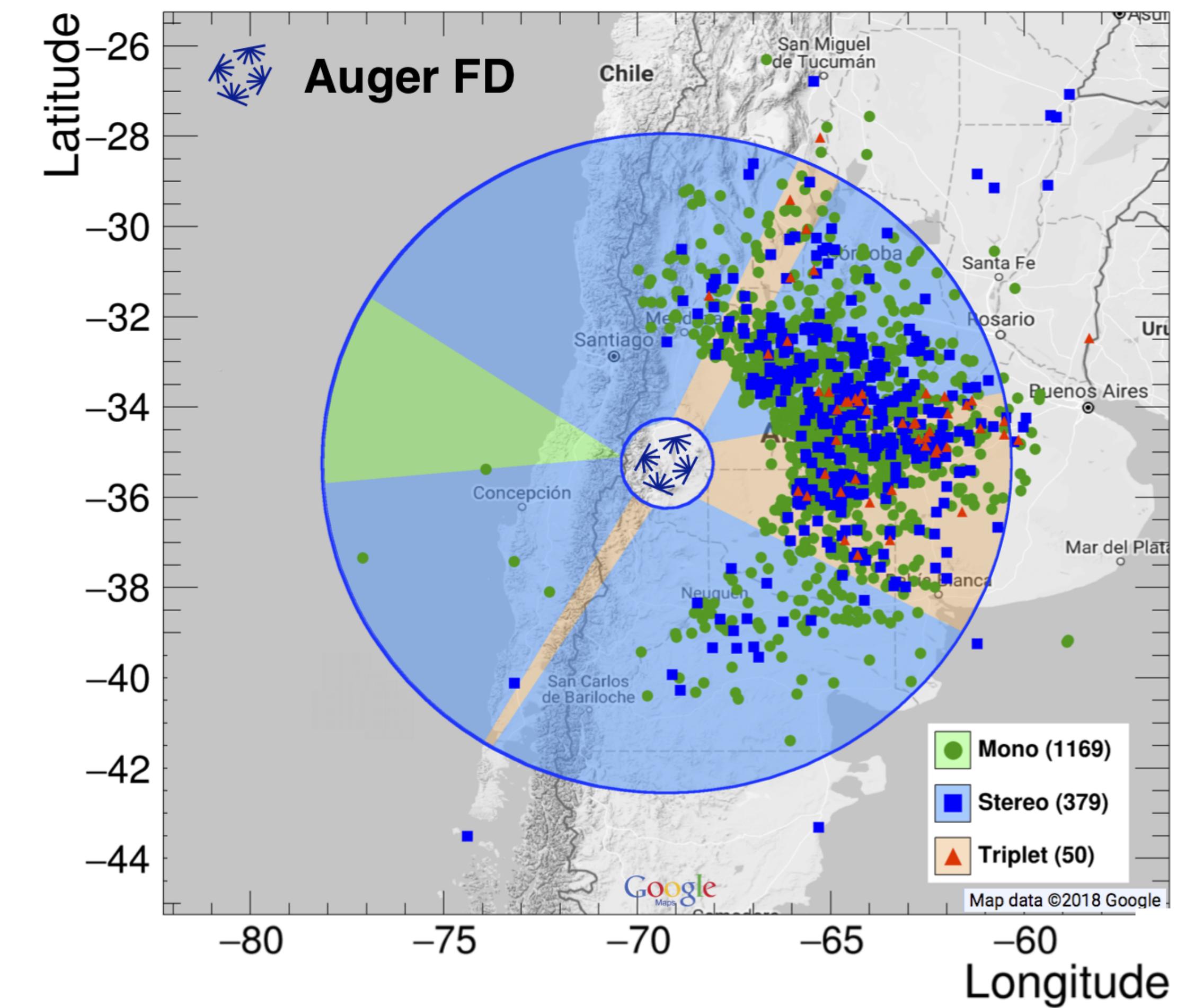


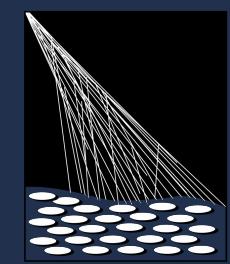
WHAT IS THE AUGER OBSERVATORY?

- The Fluorescence Detector of the Pierre Auger Observatory (Auger FD) records UV fluorescence from cosmic-ray air showers.
- FD made up of four sites with six telescopes each (24 total).



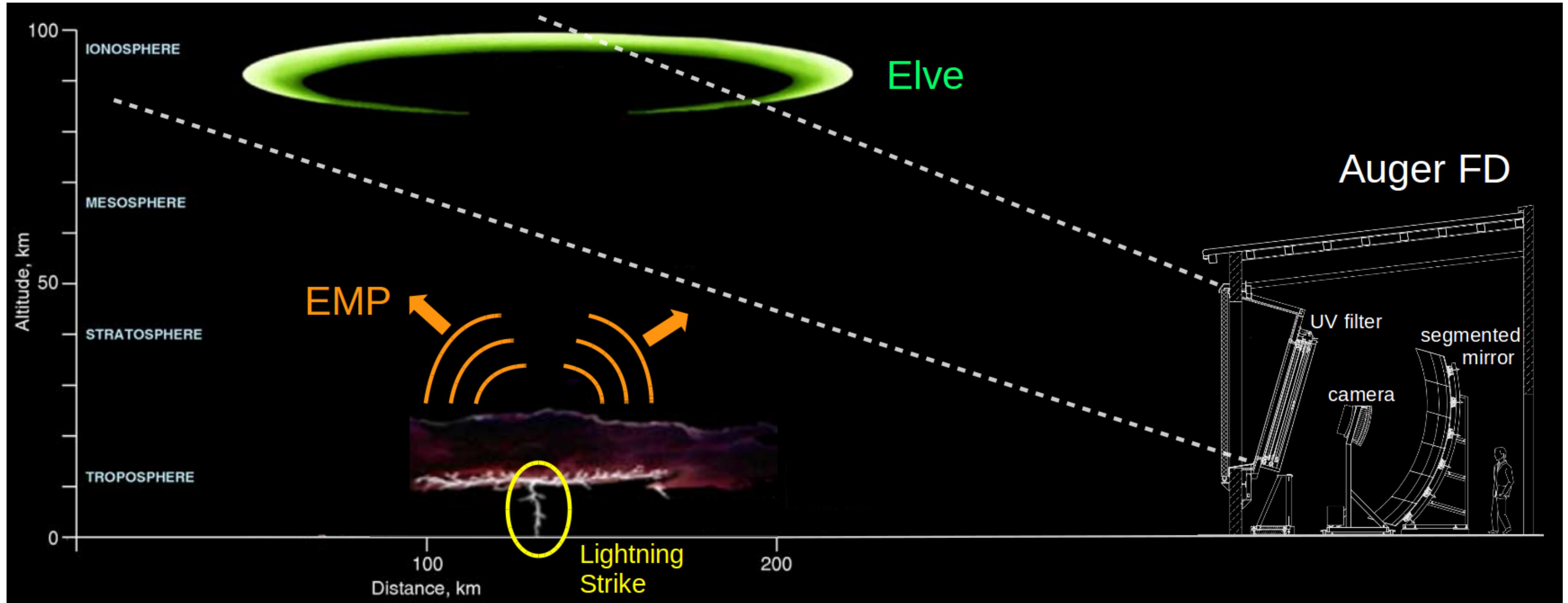
- Each camera has 440 PMTs and 10 MHz acquisition rate.



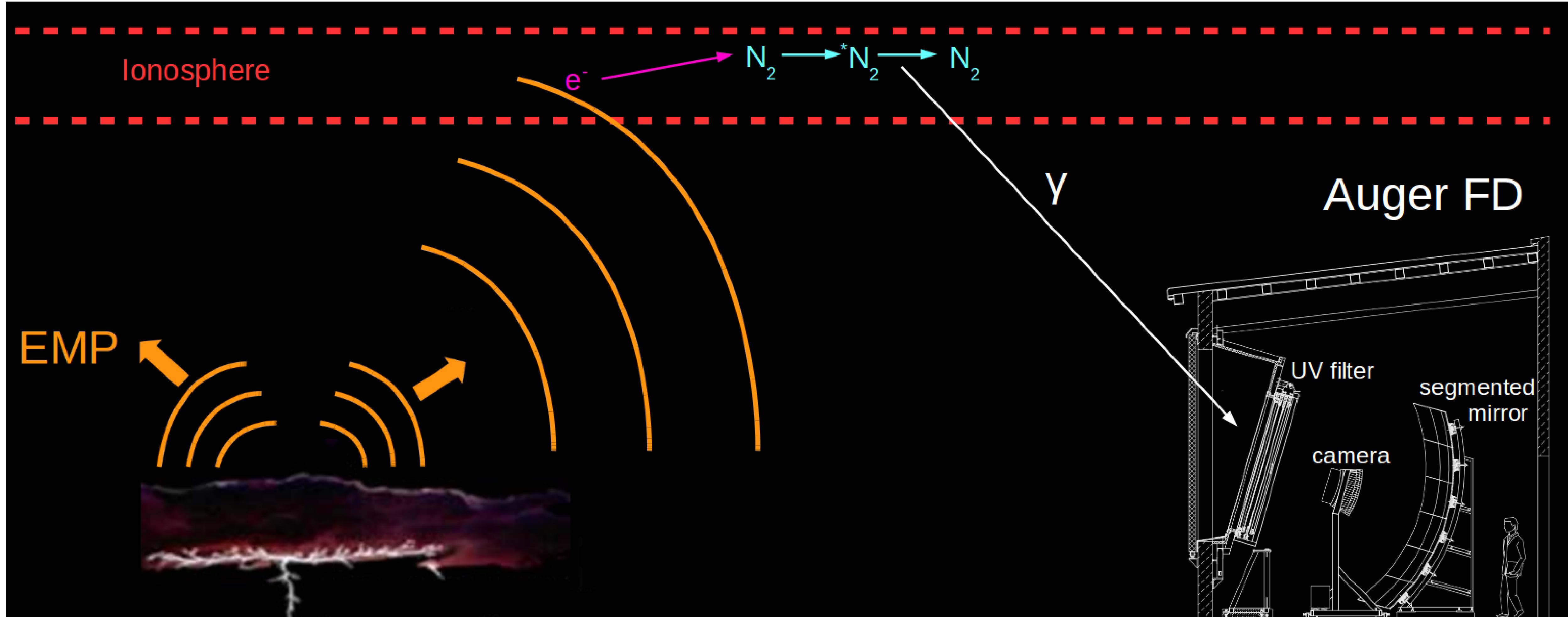
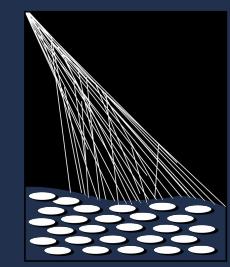


WHAT ARE ELVES?

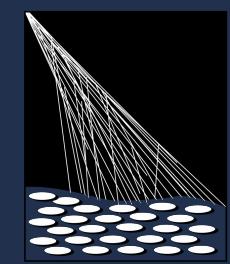
- Elves are a class of transient luminous events that occur in the ionosphere over strong lightning.
- The fast current flow in lightning is modeled as a Hertzian dipole and creates an EMP.
- **Elves** are a result of the interaction between this EMP and the ionosphere.



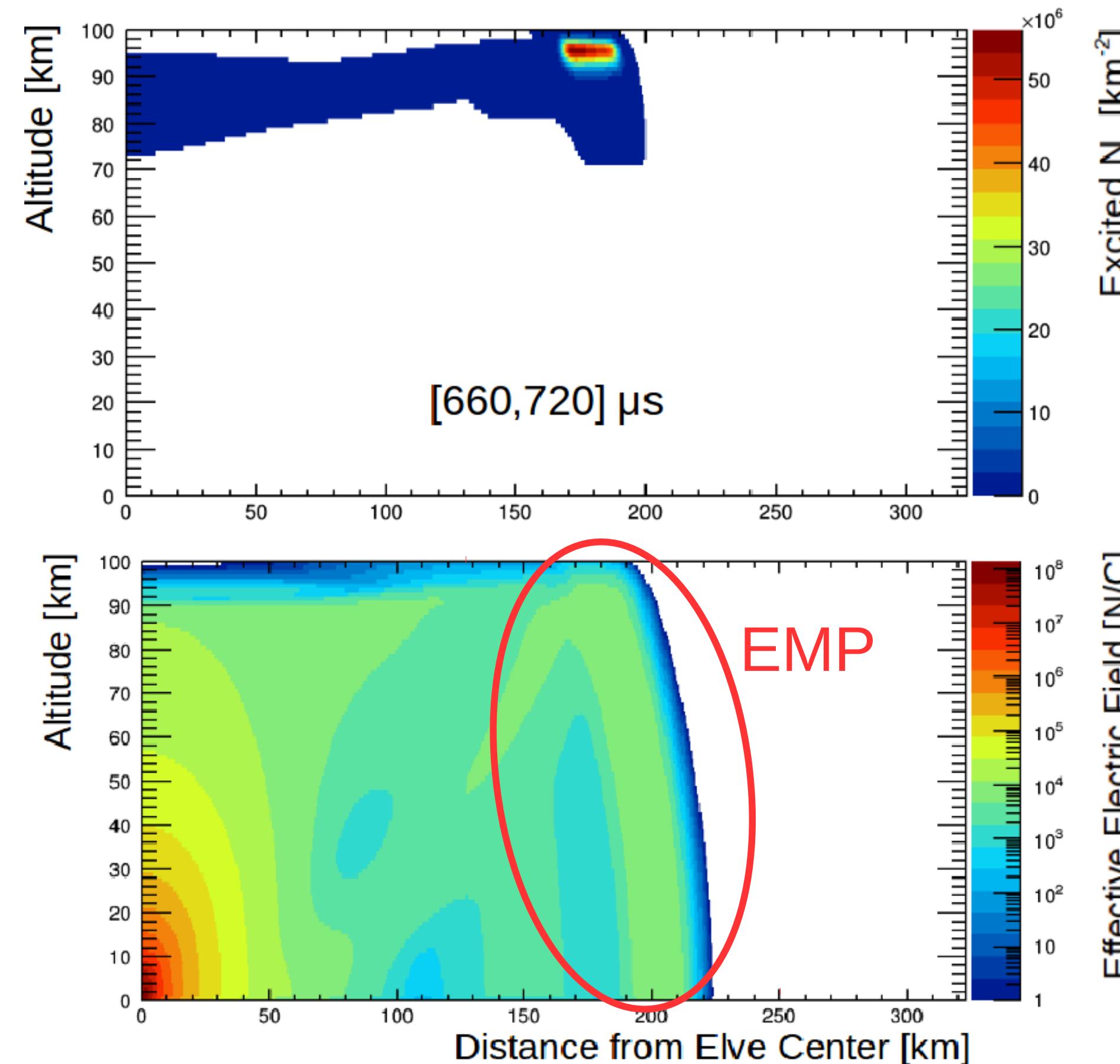
[1] Warrilow, Chrissy. "Transient Luminous Events: Sprites, Jets and Elves Are Mysteries in the Sky (PHOTOS)." The Weather Channel, 27 Aug. 2014, weather.com/news/news/transient-luminous-events-mysteries-sky-20130731.



[1] Warrilow, Chrissy. "Transient Luminous Events: Sprites, Jets and Elves Are Mysteries in the Sky (PHOTOS)." The Weather Channel, 27 Aug. 2014, weather.com/news/news/transient-luminous-events-mysteries-sky-20130731.



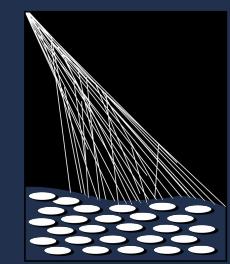
- Models the lightning strike, EMP, and interactions with the ionosphere.
- Numerical propagation of Maxwell's Equations and the Langevin equation.



Maxwell's Equations

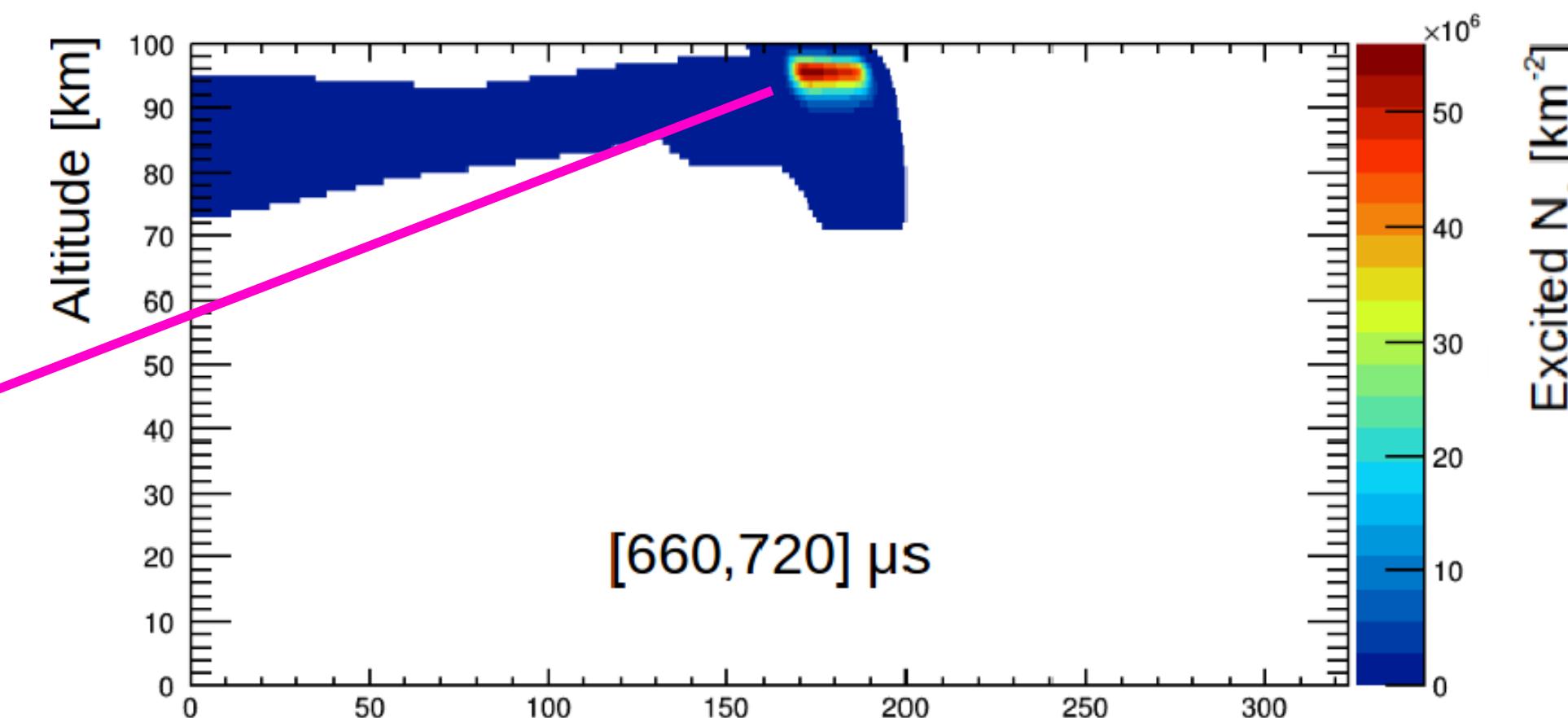
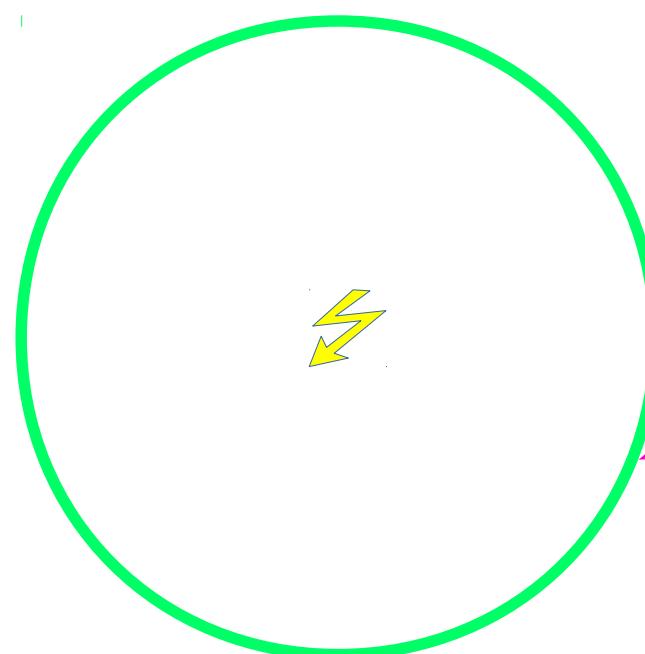
$$\epsilon_0 \frac{\partial \mathbf{E}}{\partial t} = \nabla \times \mathbf{H} - \mathbf{J}_{\text{tot}}$$

$$\mu_0 \frac{\partial \mathbf{H}}{\partial t} = \nabla \times \mathbf{E}$$



- Models the lightning strike, EMP, and interactions with the ionosphere.
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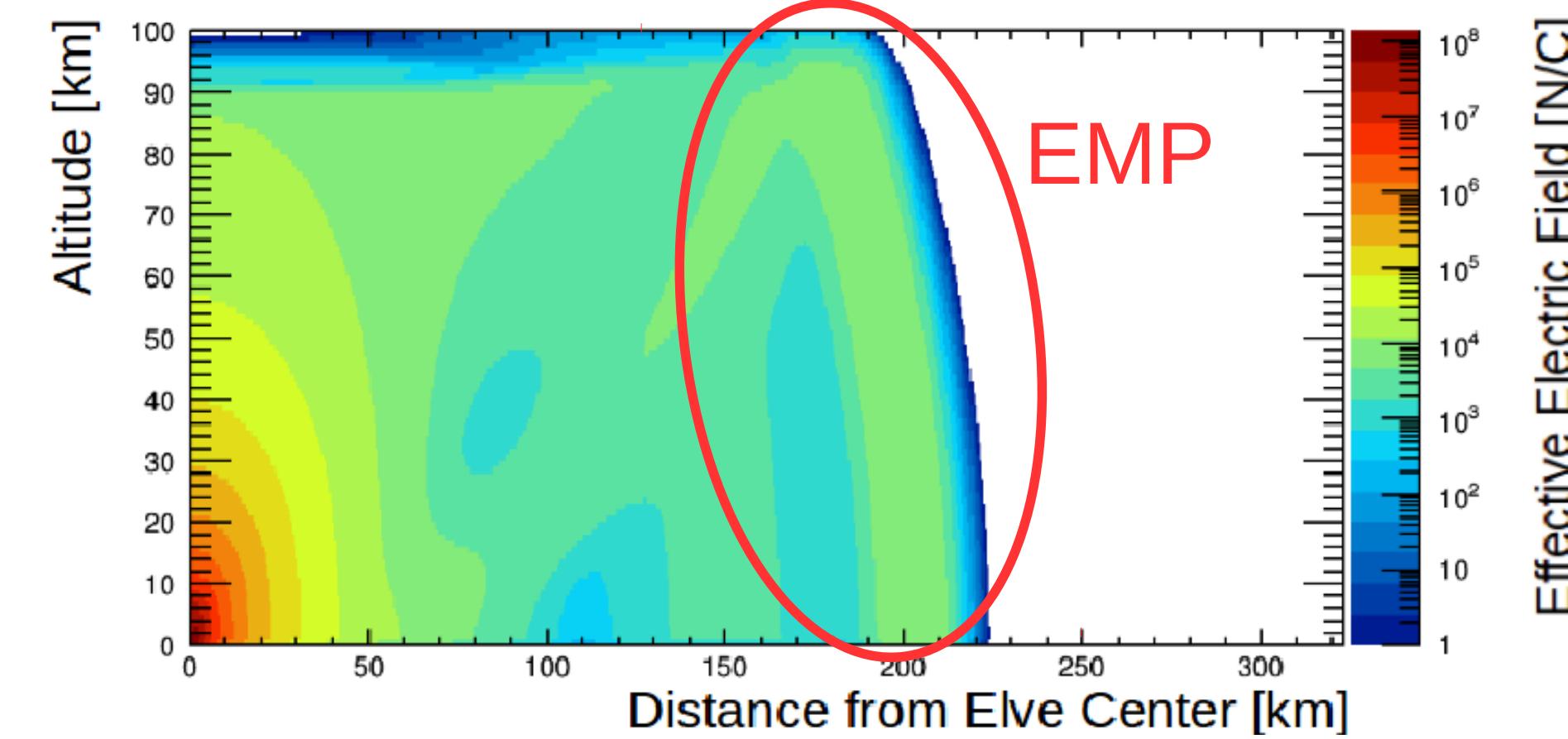
Elve seen from above

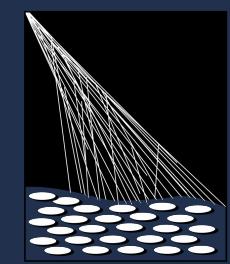


Maxwell's Equations

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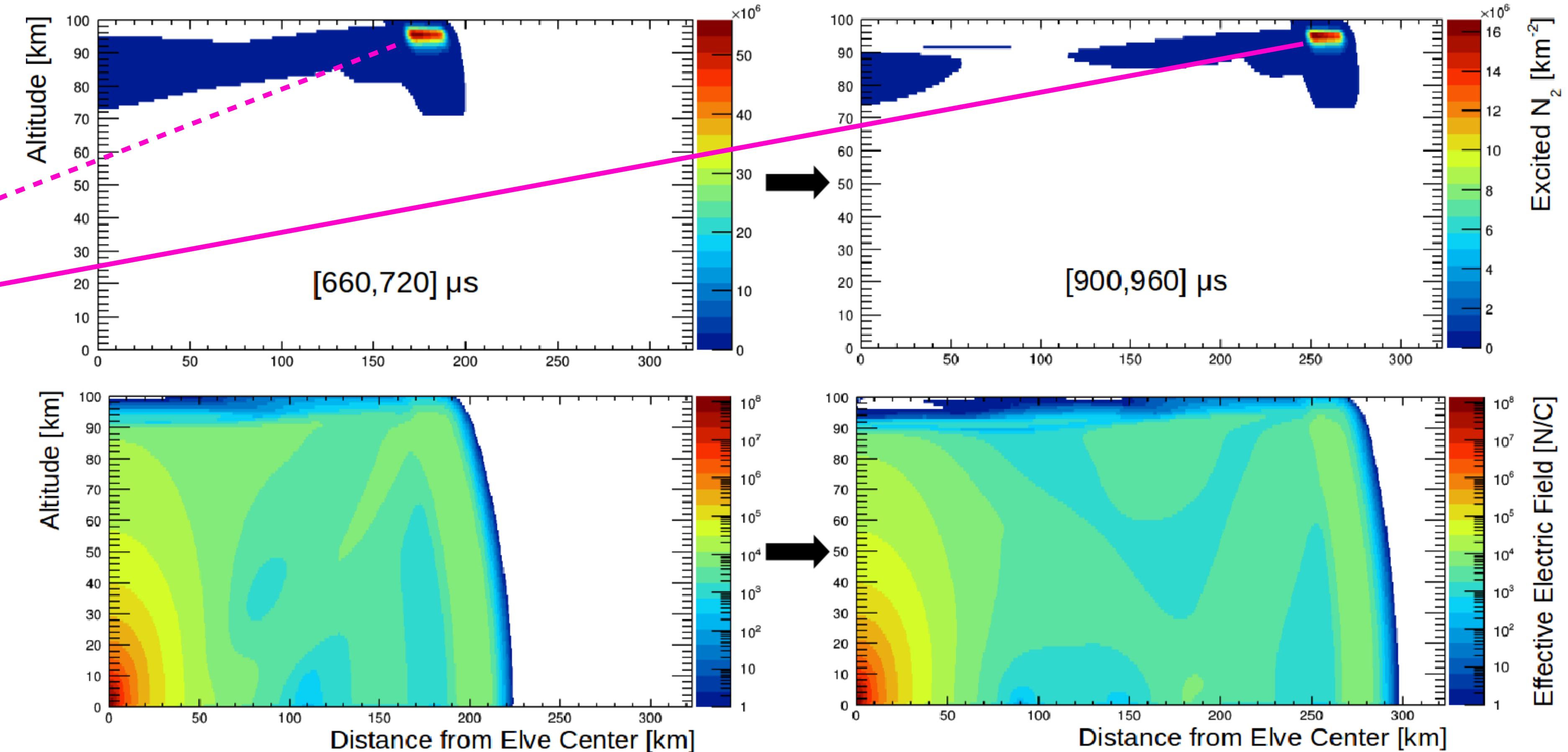
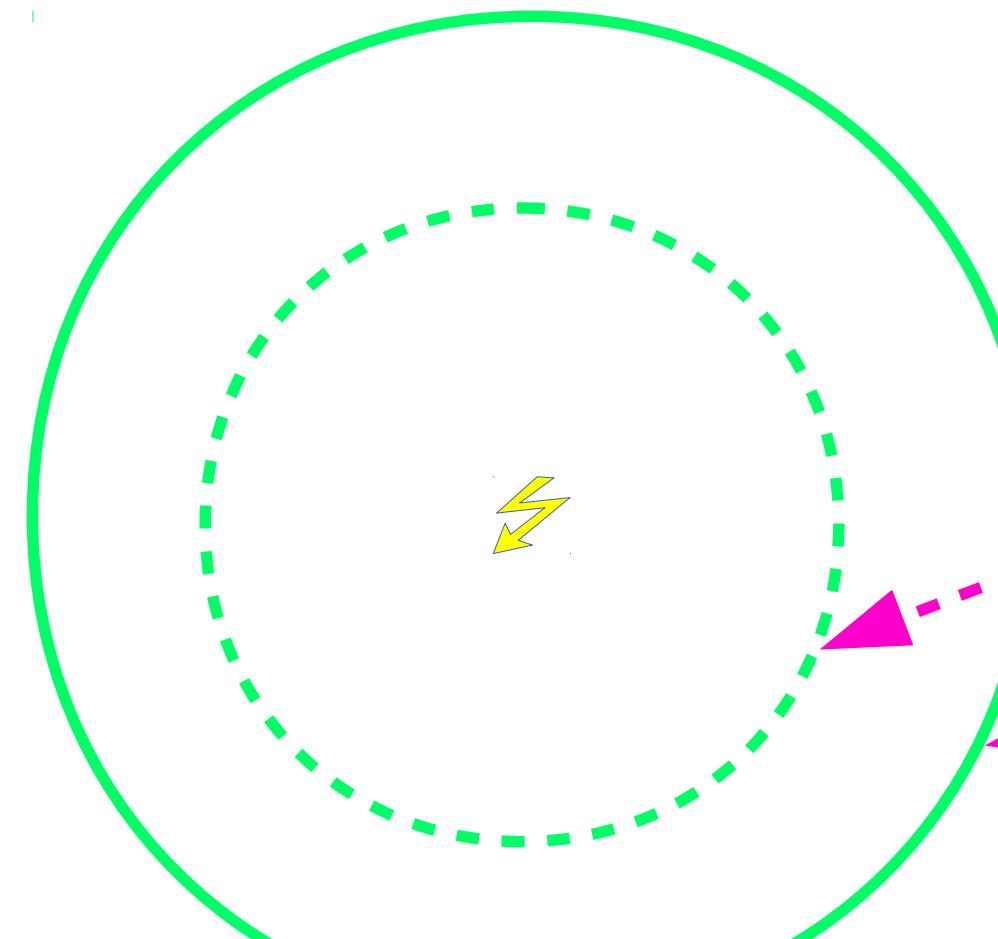




SIMULATION OVERVIEW



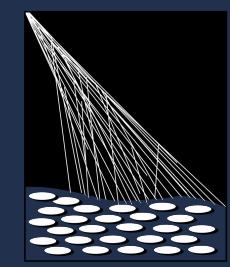
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Maxwell's Equations

$$\epsilon_0 \frac{\partial \mathbf{E}}{\partial t} = \nabla \times \mathbf{H} - \mathbf{J}_{\text{tot}}$$

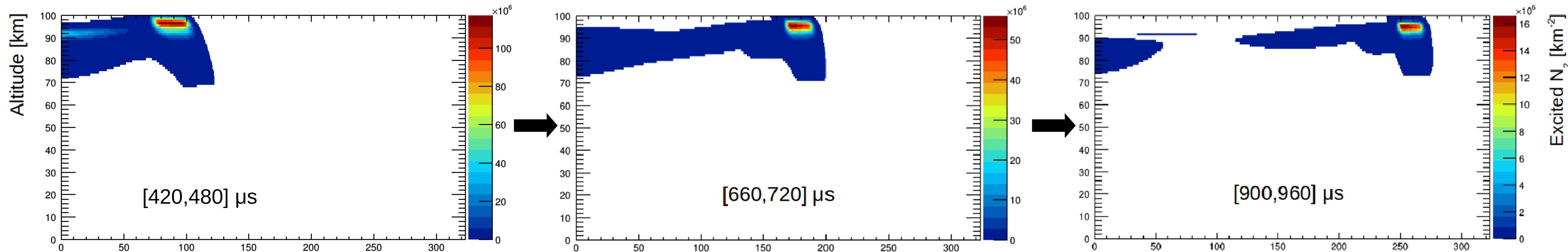
$$\mu_0 \frac{\partial \mathbf{H}}{\partial t} = \nabla \times \mathbf{E}$$



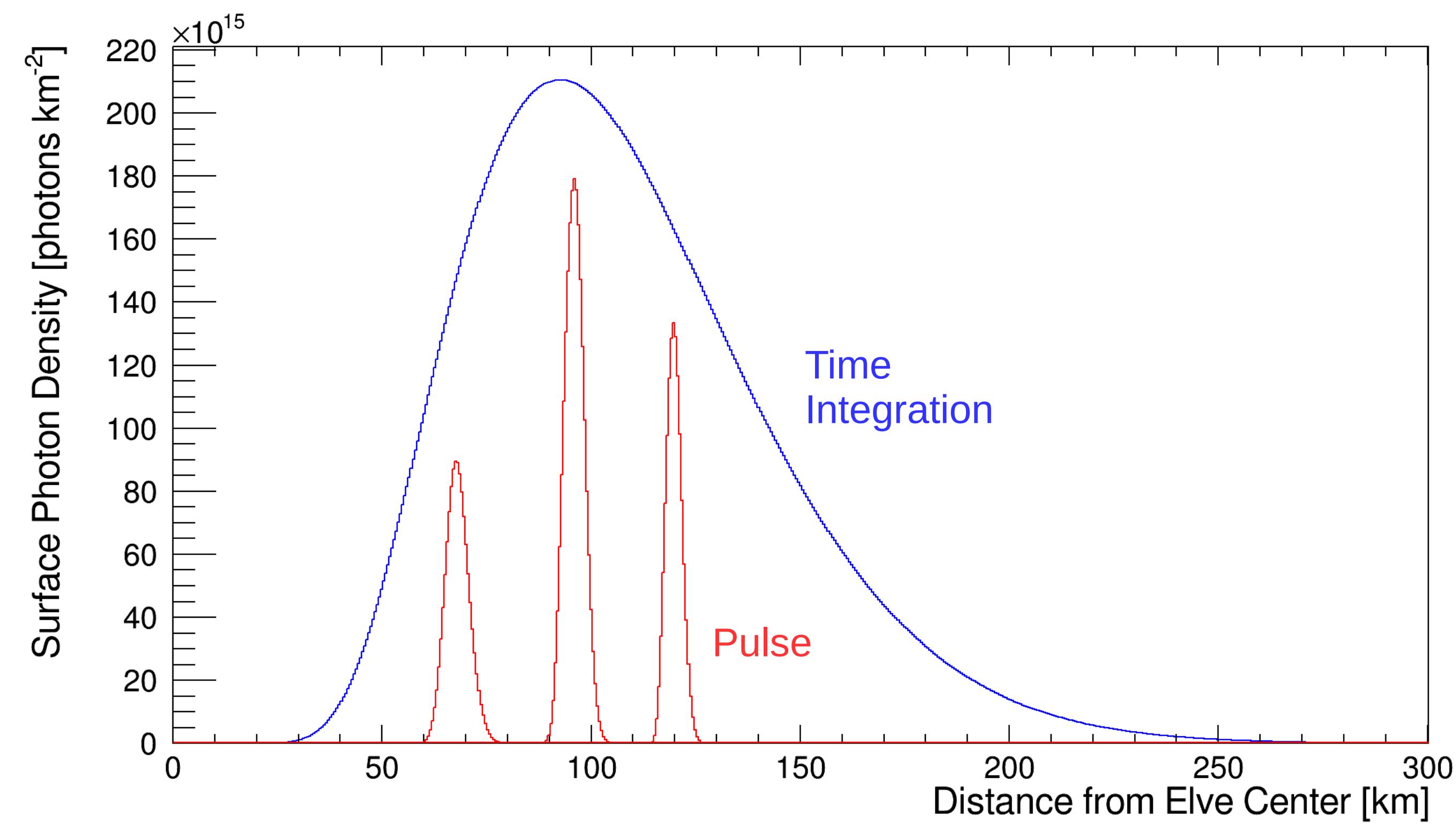
SURFACE PHOTON DENSITY

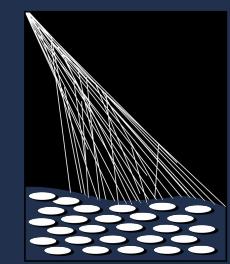


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- Integrate **excited nitrogen** over **altitude** and **time**.
- This gives us **surface photon density**.
- Information reduced to two dimensions.





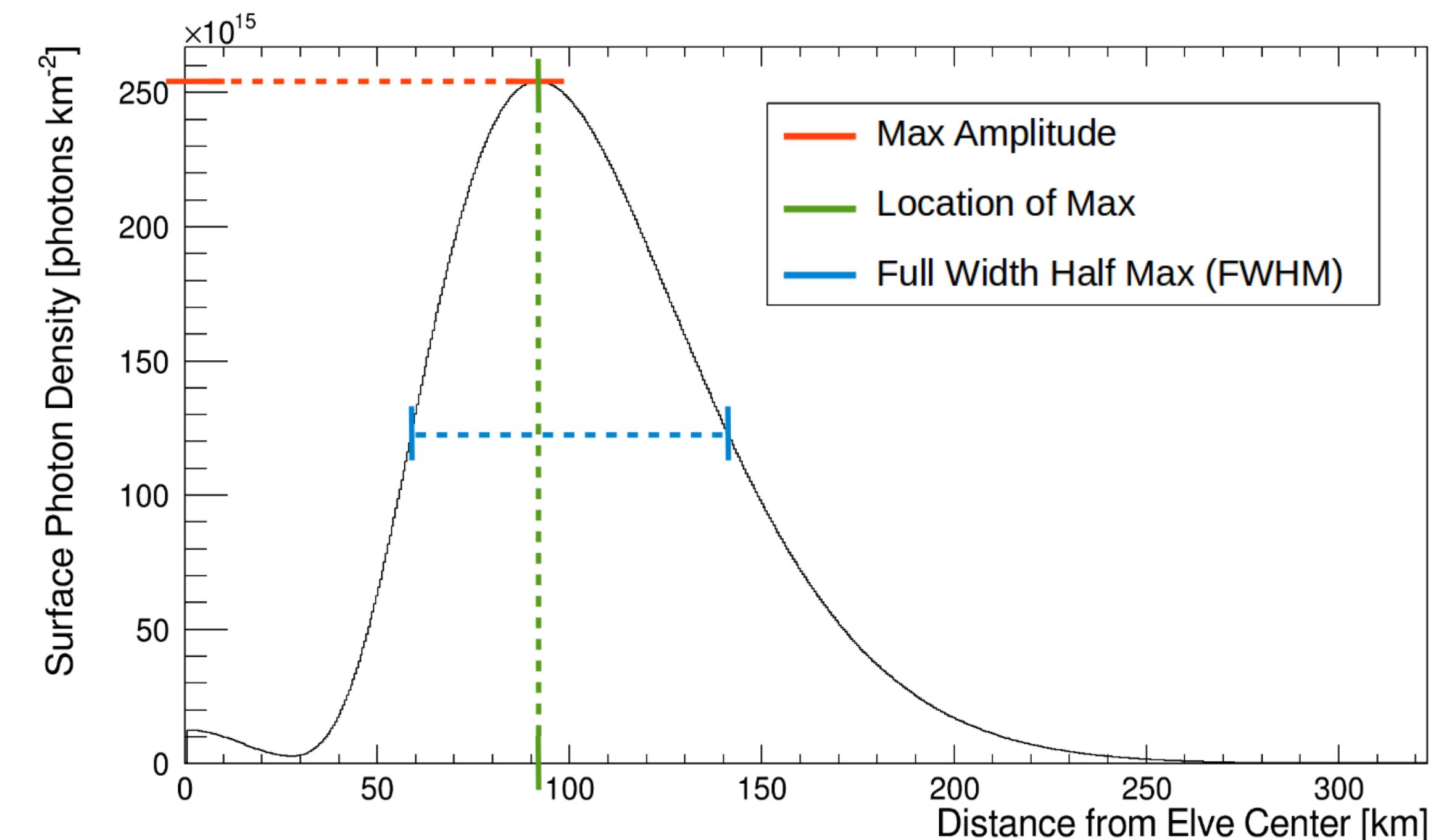
Which lightning parameters affect elve structure the most?

Simulation Input: Lightning and Atmospheric Parameters

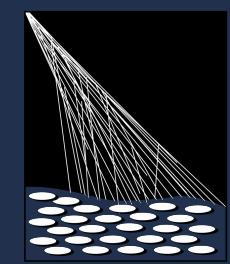
| Parameter | Parameter Name | Nominal Value | Reasonable Range |
|-----------|---------------------------|---------------|------------------|
| I_0 | Peak Current | 100 kA | 30 kA to 300 kA |
| L_{ch} | Channel Length | 6 km | 3 km to 9 km |
| H_{Ion} | Height of Ionosphere Base | 92 km | 80 km to 98 km |

- Range of parameters selected through literature review and simulation study.
- Lightning parameters describe the flow of current in the lightning channel.
- Other parameters studied include: Rise Time, Fall Time, Return Stroke Speed, Continuing Current.

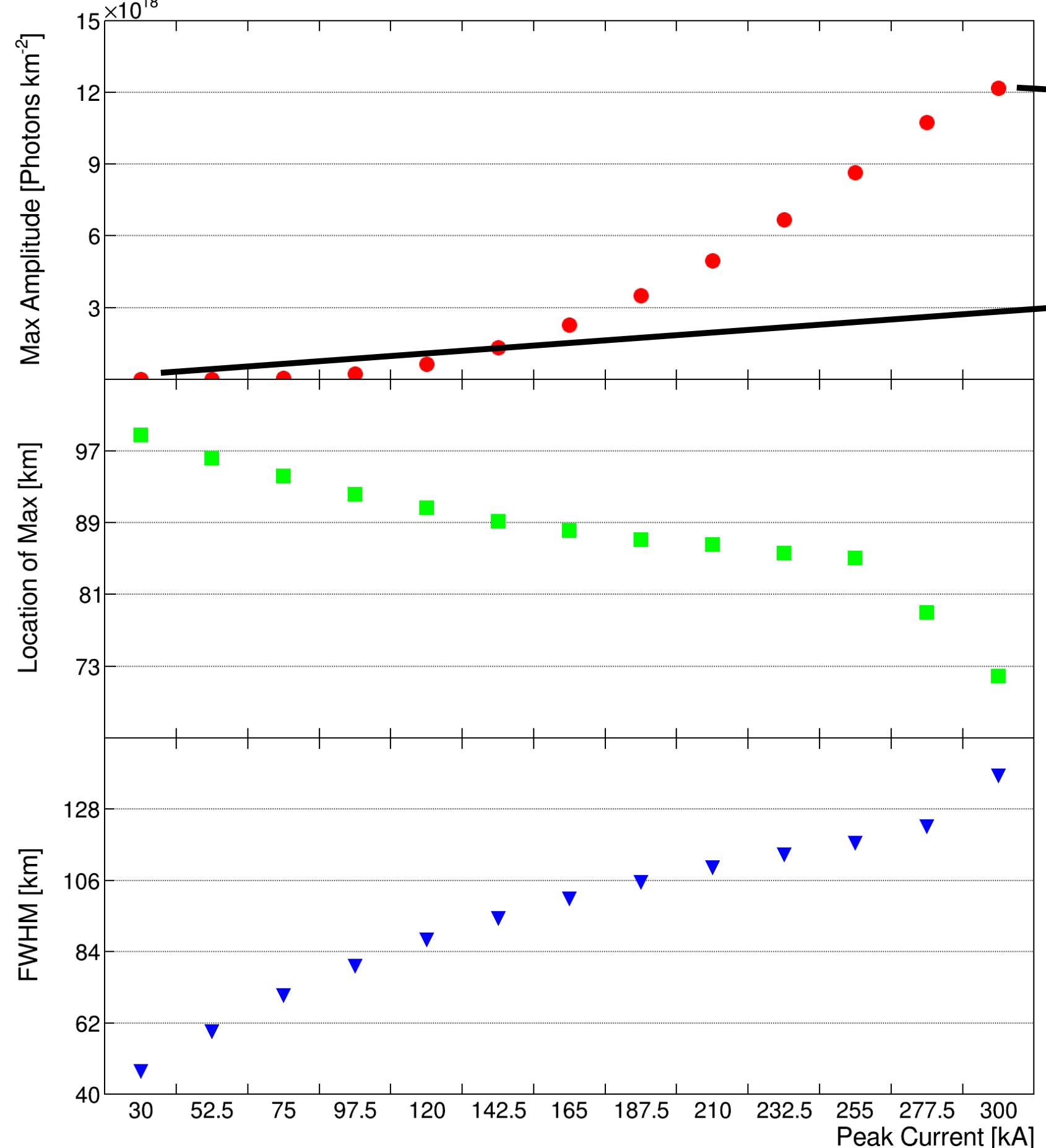
Simulation Output: Elve Parameters



- Want to quantitatively study simulated elves.
- Chosen metrics shown on plot.



Peak Current

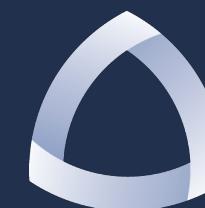
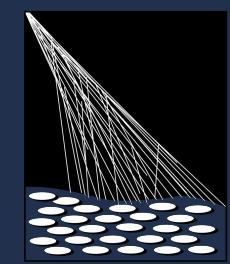


- Varying over peak current while keeping other parameters constant.
- Clearly sensitive to peak current.
- Note the sharp increase in amplitude @ 120 kA.

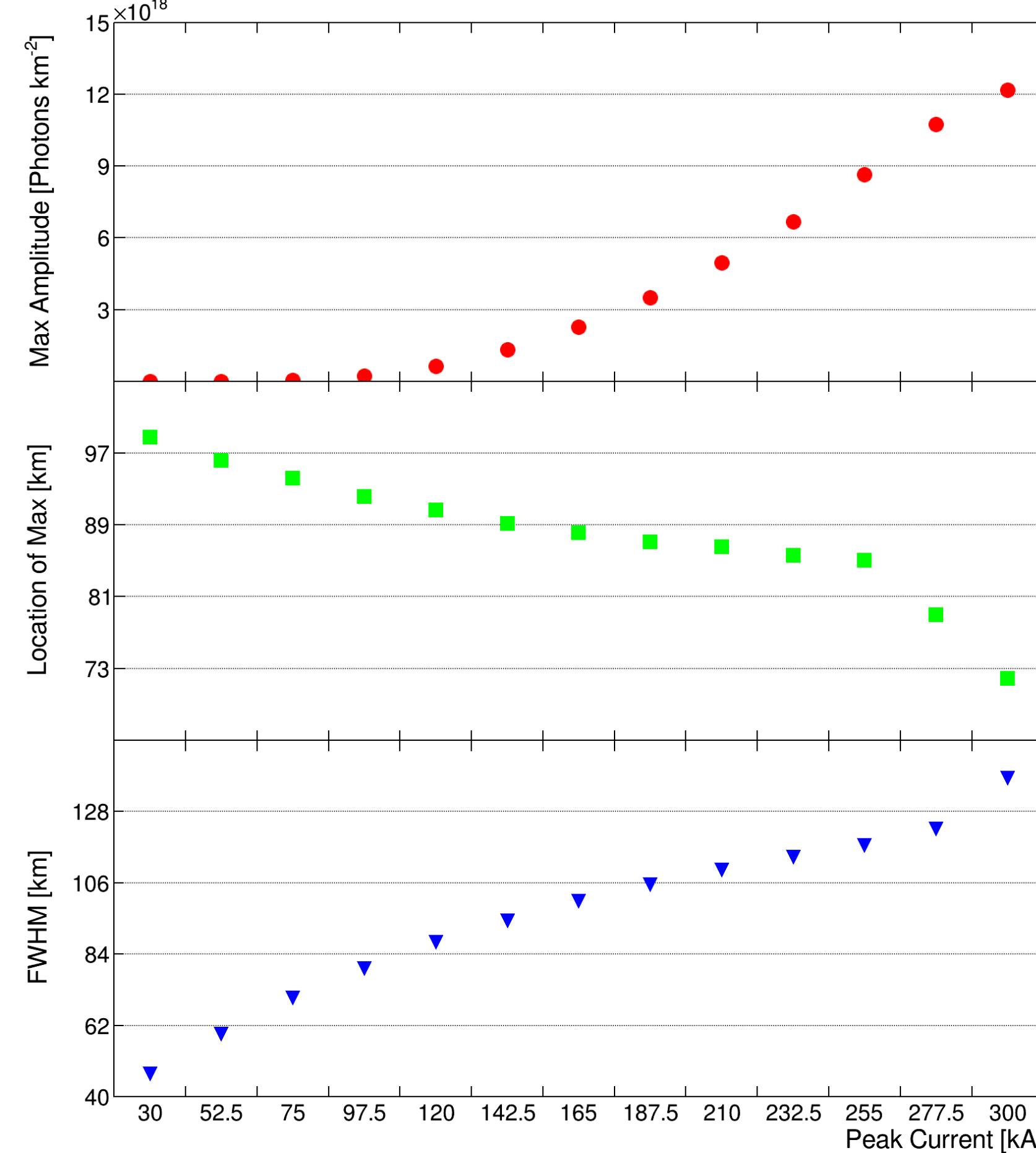
Sensitivity Ratio

Peak Current:

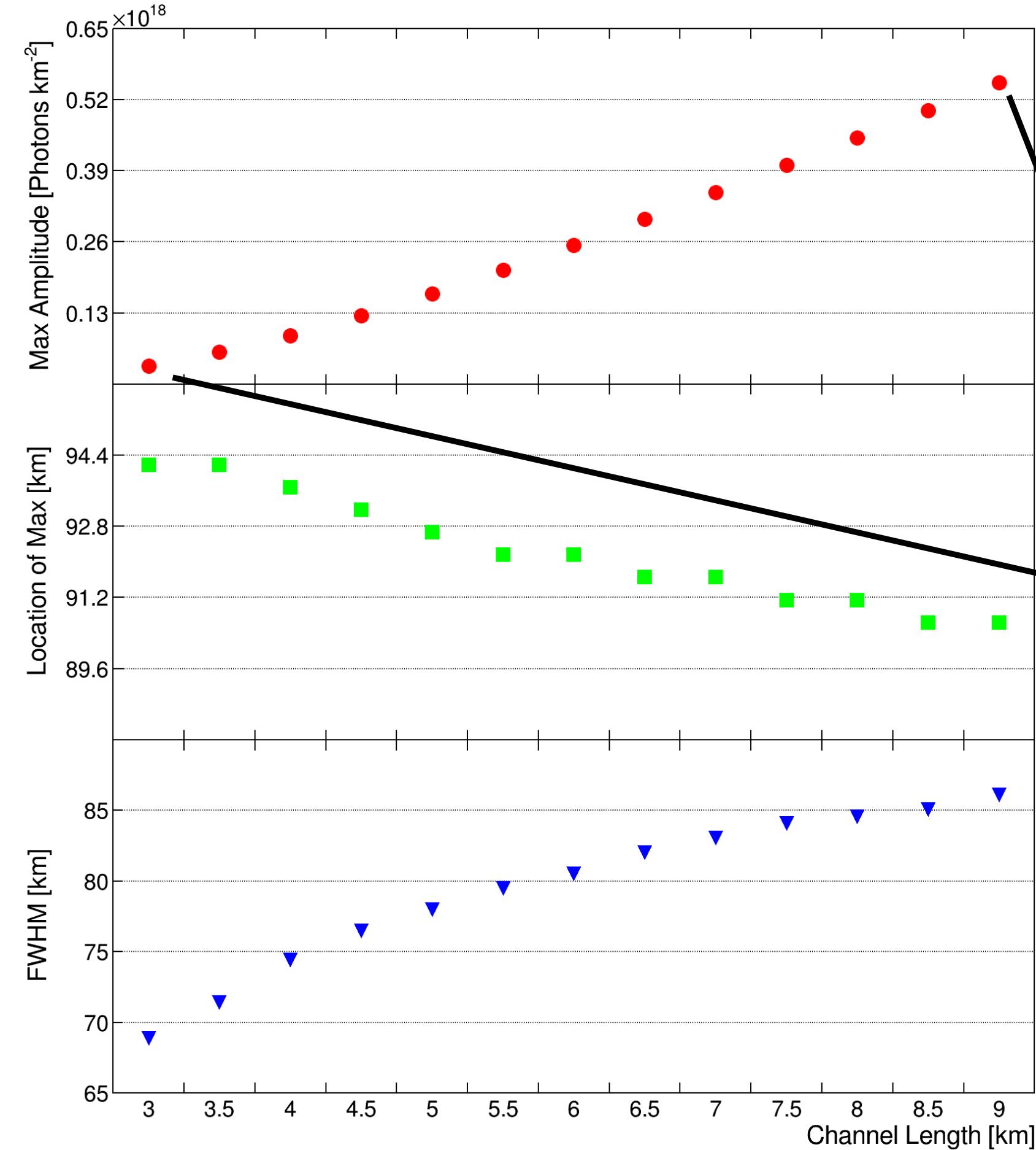
$$\text{Sensitivity Ratio} = \frac{\text{Maximum Value}}{\text{Minimum Value}} \approx \frac{10^{19}}{10^{11}} \approx 10^8$$



Peak Current



Channel Length



Sensitivity Ratio

Peak Current:

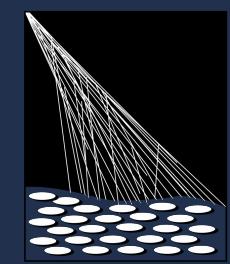
$$\frac{\text{Maximum Value}}{\text{Minimum Value}} \approx \frac{10^{19}}{10^{11}} \approx 10^8$$

Channel Length

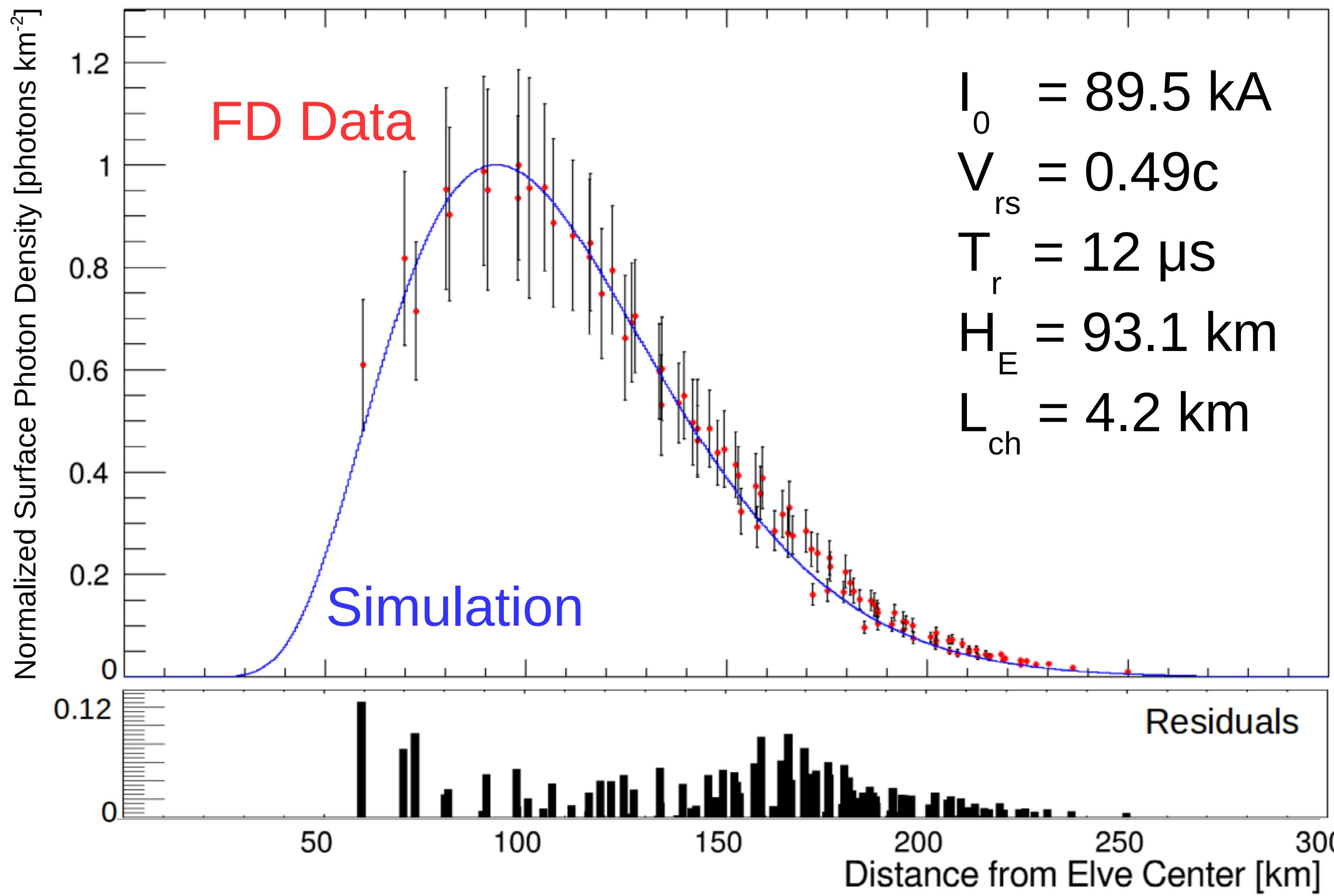
$$\frac{\text{Maximum Value}}{\text{Minimum Value}} \approx \frac{10^{17}}{10^{17}} \approx 1$$

Much more sensitive to Peak Current than Channel Length.

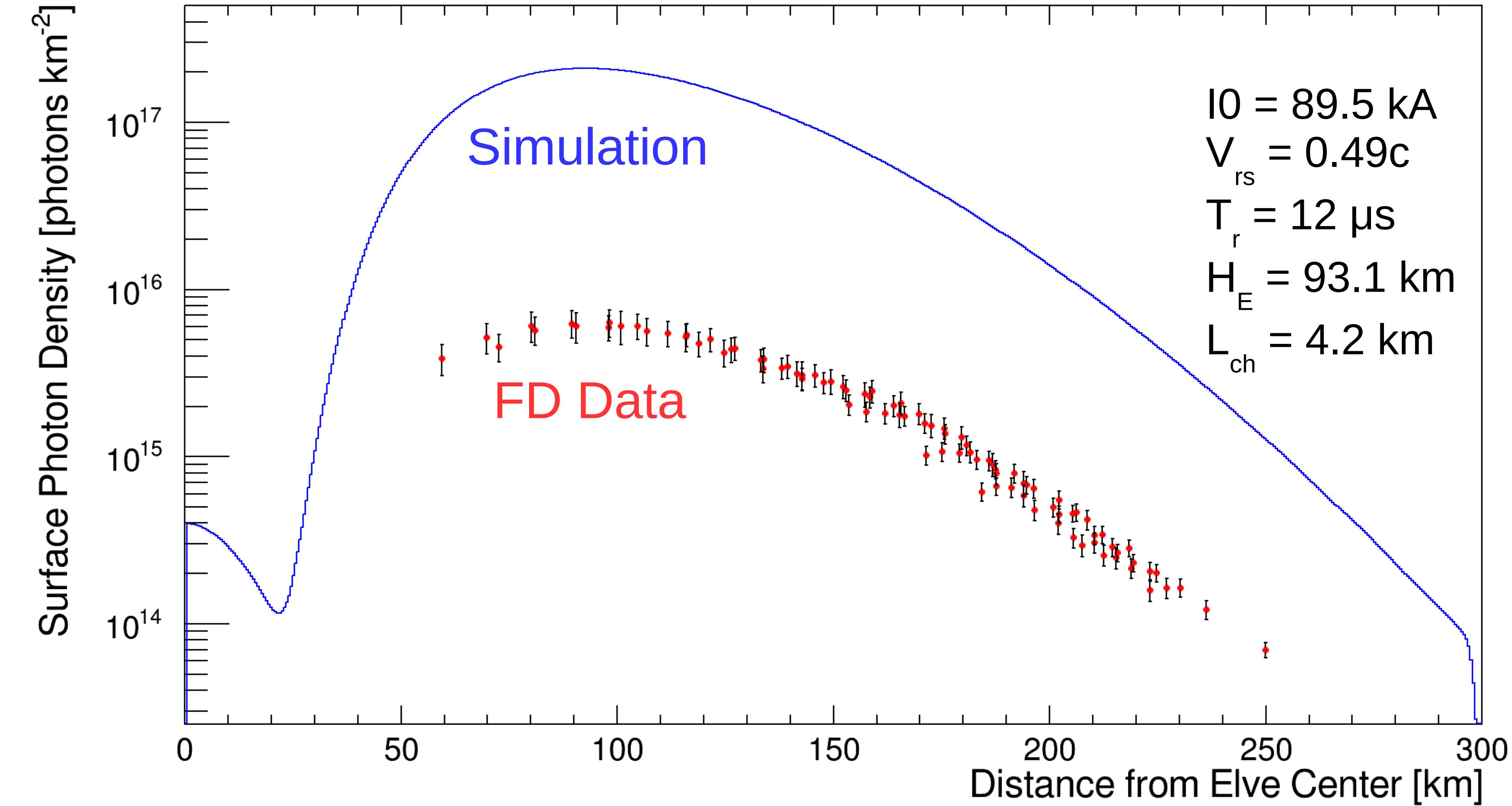
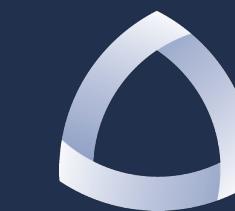
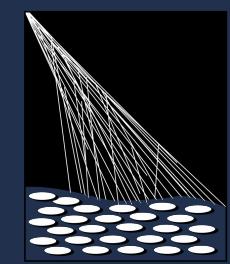
- Varying over channel length while keeping other parameters constant.



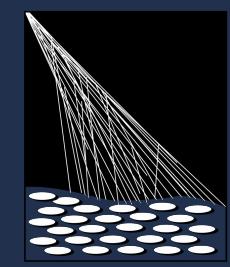
- Sensitivity study + χ^2 minimization \rightarrow attempt to match simulation to data.



- Each red dot is an FD pixel.
- Both plots are **normalized** to their respective maximums.
- Shape of simulation (**Location of Max**, **FWHM**) matches data well.
- Reduced χ^2 with 51 DOF = 3.48311



- Without normalization, **Max Amplitude** is off by ~ 2 orders of magnitude.
- Possible causes:
 - 1) Surface density reconstruction: revise atmospheric attenuation calculations
 - 2) EMP simulation

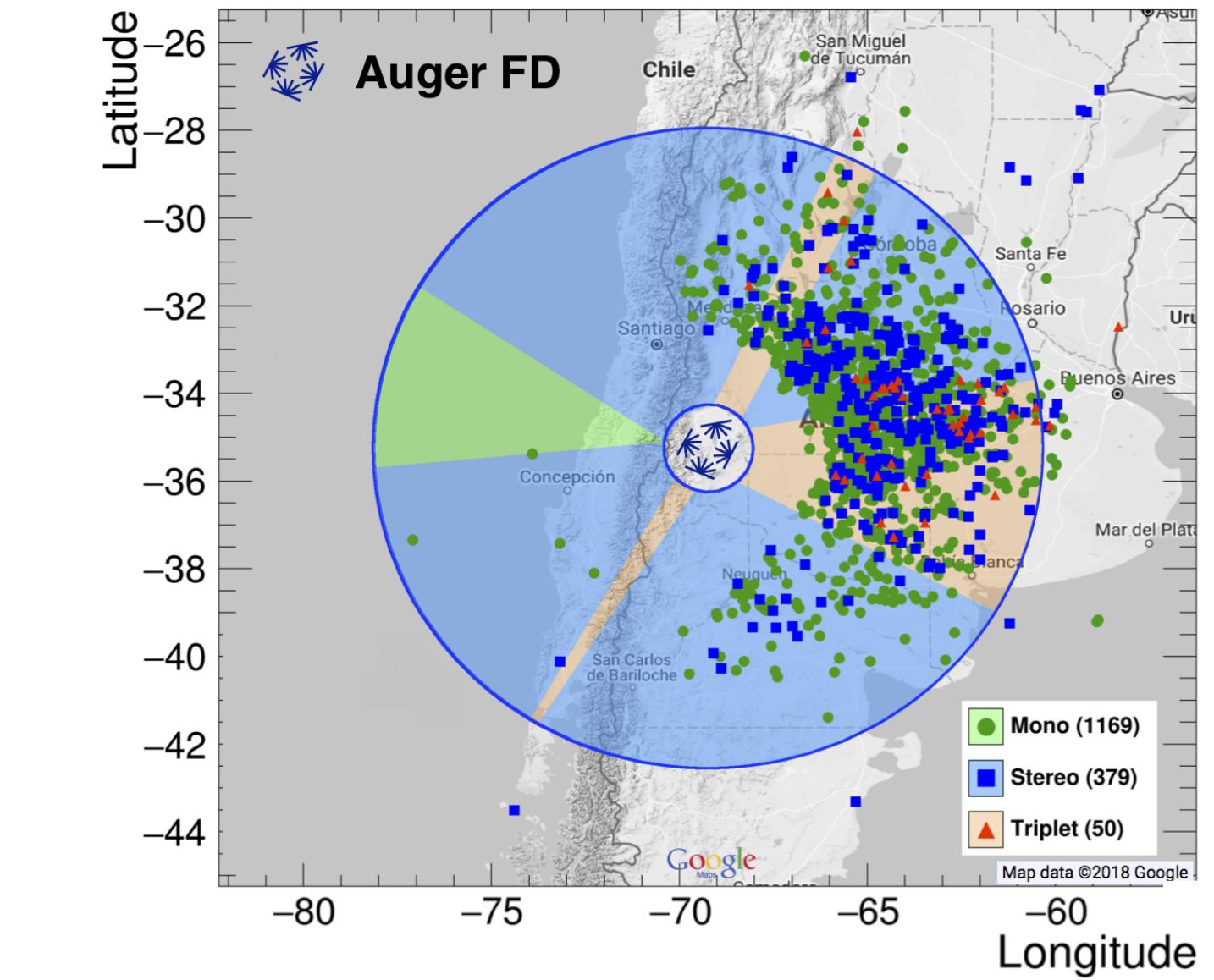
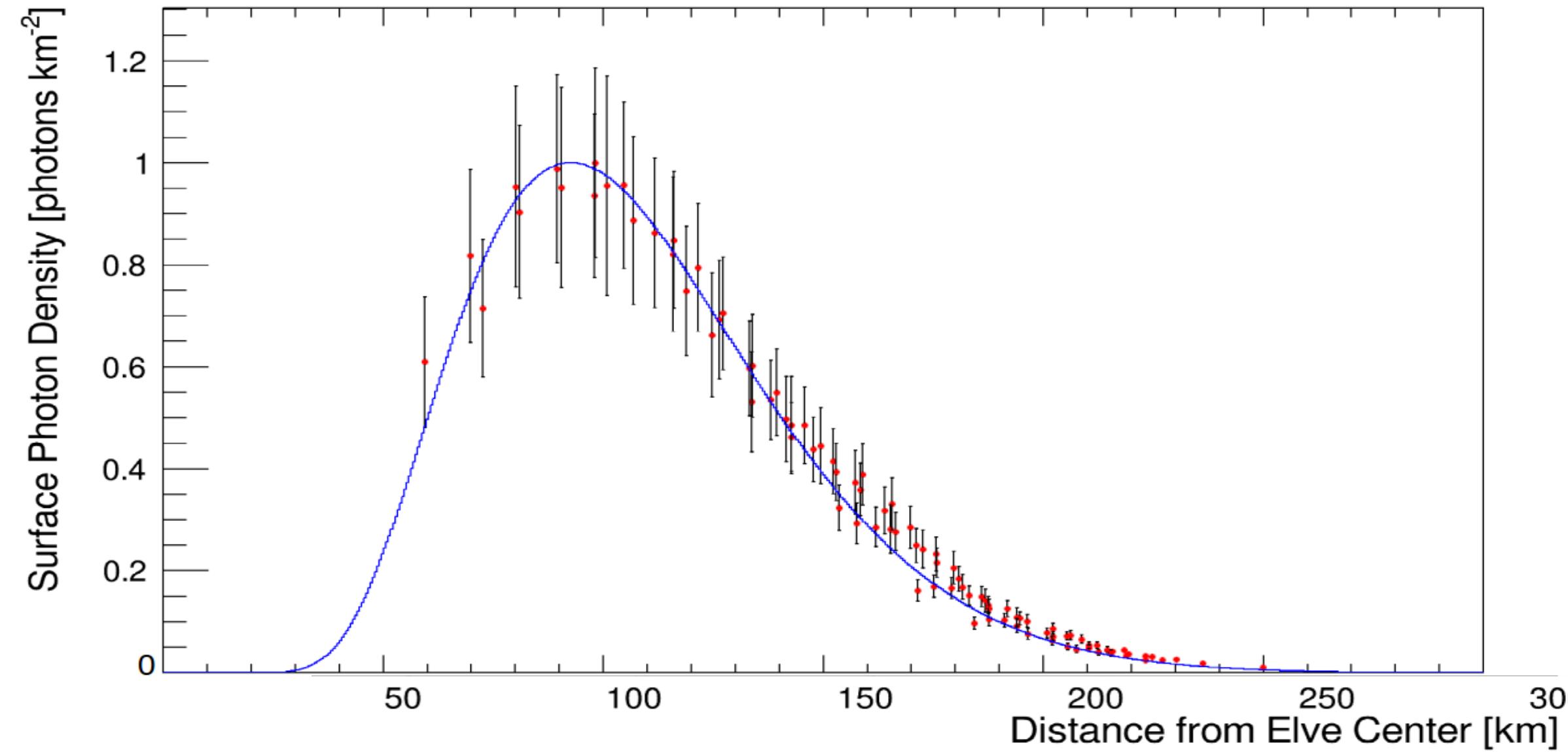


Big Picture Take-Aways

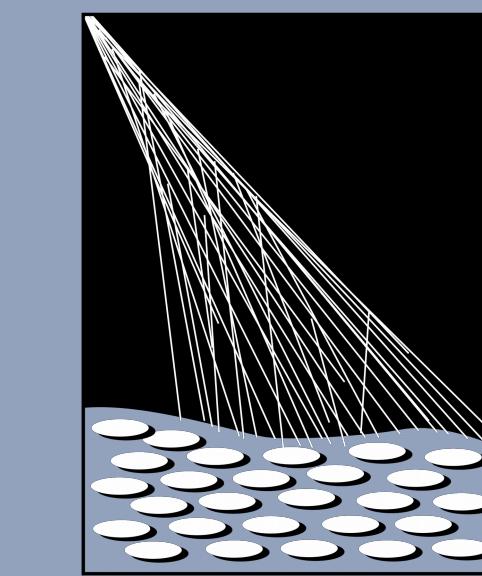
- Lightning can be dangerous, so it is important to study.
- We are using elves to look at lightning in a very novel way.

- Very bright elves means that peak current is greater than 120 kA.
- Better understanding of how lightning affects elve shape and amplitude.
- Simulation can create accurate surface density profiles.
- Problem with amplitude needs to be resolved.

| Parameter | Sensitive to this Parameter? |
|----------------------|------------------------------|
| Peak Current | Very |
| Channel Length | Somewhat |
| Return Stroke Speed | Very |
| Rise Time | Somewhat |
| Fall Time | No |
| Height of Ionosphere | Very |
| Continuing Current | No |



Thank you - Questions?



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