

Introduction
<div>► Intro goes here</div>
Pipeline for Forward Modeling Emission from Active Region Cores
<p>We have developed a Python package for forward modeling emission from ARs using ensembles of field-aligned hydrodynamic models. It leverages the full power of the scientific Python stack and relies heavily on the SunPy [2] and Astropy [?] libraries.</p> <div><div><div>1. Fetch observed magnetogram for the desired AR. Fig. ?? shows an HMI magnetogram of NOAA 1109.</div><div>2. Perform a field extrapolation (e.g. PFSS) to derive the three-dimensional vector field \vec{B}</div><div>3. Trace 1000 fieldlines through the extrapolated field, including only closed fieldlines in the range $10 < L < 1000$ Mm.</div><div>4. For each fieldline, run a field-aligned hydrodynamic model. In this case, we'll use the two-fluid EBTEL model described in Barnes et al. [1]</div><div>5. Map T_e and n_e from simulations to 3D field skeleton and calculate emissivity for each selected transition λ_{ij} of element X and charge state k,</div><div>$\varepsilon_{ij}^{X,k} = n_j A_{ij} h c / \lambda_{ij} / n_e \quad [\text{erg cm}^3 \text{ s}^{-1}]$<div>where all atomic data comes from the CHIANTI atomic database [? ?]</div></div><div>6. Integrate the emissivity along the LOS for each transition,</div><div>$I(\lambda_{ij}) = \frac{1}{4\pi} \int_{\text{LOS}} \text{d}h \, 0.83 \text{Ab}(X) f_{X,k} \varepsilon_{ij}^{X,k} n_e^2 \quad [\text{erg cm}^{-2} \text{ s}^{-1} \text{ str}^{-1}]$<div><div>► $f_{X,k}$ calculation includes effects due to nonequilibrium ionization [e.g. ? ?]</div><div>► T_e, n_e functions of h, the distance along the LOS which intersects <i>many</i> loops</div></div><div>7. Convolve with wavelength response function of each channel of each instrument. Here, we synthesize observations from both channels of the Extreme-ultraviolet Imaging Spectrometer (EIS) on <i>Hinode</i>.</div></div></div></div>
Heating Model: Impulsive Heating over a Range of Frequencies
<div>► Describe heating model here</div>
Forward Modeling AIA Intensities
<div><div>► Details about AIA calculation</div><div>► list ions/elements</div><div>► Show equations</div></div>

Simulated AIA Intensities
Computing Timelags Between AIA Channel Pairs
Simulated versus Observed Timelags
Analyzing Observed Pixels with a Random Forest Classifier
Conclusions
<div>► The conclusions go here</div>
References
<div><div>[1] Barnes, W. T., Cargill, P. J., & Bradshaw, S. J. 2016, The Astrophysical Journal, 829, 31</div><div>[2] SunPy Community, Mumford, S. J., Christe, S., et al. 2015, Computational Science and Discovery, 8, 014009</div></div>