### Example 13.7.3:

We are given the following table of data that gives the number of astronomical observations which fall into a given velocity interval.

Intervals of Velocities	0,	ρ̂ <sub>jo</sub>	ê,
(-80, -70)	1)	.000)	
(-70, -60)	1 2 2 2 2 8	.001	
(-60, -50)	2 15	.009 \ .224	17.92
-50, -40)	2	.163	
-40, -30)	24	.284	22.72
-30, -20) -20, -10)	26	.283	22.64
-10, 0)	11)	.154)	
0, 10)		.046 209	16.72
	1	.008	10.72
	1)	.001 )	
		4.000	80.00
(10, 20) (20, 30)	2 15 15 180	.008	

Where  $o_j$  represents the number of observed occurrences in the interval  $\{a_j, a_{j+1}\}^1$ . A test for normality proceeds as follows; assume  $X_i \sim N[\mu, \sigma]$ . Let  $x_{ij}$  denote the  $i^{th}$  observation in cell (velocity interval) j. We have  $\sum_i x_{i,j} = o_j$ . Being as we are testing for a normal distribution the central limit theorem is not strictly necessary, but for reasons which will soon become apparent it is required for the Chi-squared test. So we set  $z = \frac{x-\mu}{2}$  which gives the probability density function (pdf) of the  $j^{th}$  cell,

$$p_j = \left[ \frac{1}{\sqrt{2\pi\sigma}} \int_{z_j}^{z_{j+1}} e^{-\frac{1}{2}z^2} dz \right]^{o_j} = \left[ \Phi(z_{j+1}) - \Phi(z_j) \right]^{o_j}$$

Where  $\Phi$  denotes the integration of the standard normal. We've integrated over the cell  $\{a_j, a_{j+1}\}$  because we are only concerned with whether a given observation occurred anywhere within this interval. By taking this integral to the power  $o_j$  we've assumed observations within each cell are independent of one another, which is to say the fact one observation was made in this cell had no relation to the fact that some other observation was also made in it.

In order to perorm a Chi-squared goodness of fit test it is of use to relate this pdf to the multinomial distribution. In fact we've already done the bulk of the work required to do this by integrating over the cells in which observations are contained. This has the effect of reducing the expected values of the cells to  $np_j$  where  $n = \sum_j o_j$ .

We now make use of the multinational expansion on the sum of probabilities:

$$(p_0 + p_1 + \dots + p_c) = \frac{n!}{o_1! o_2! \dots o_c!} pq * p2 * \dots * pc$$
(1)

<sup>&</sup>lt;sup>1</sup>The table suggests that  $\{a_j, a_{j+1}\}$  is a velocity interval, but considering the math to follow, unit analysis requires that  $a_j$  must be taken to represent a boundary of o – the continuous equivalent of  $o_j$ , i.e.  $a_j$  is a unitless number. Considering the variable of interest is actually  $\frac{o_1-np}{\sqrt{npq}}$  (c.f. eq 2) there is not neccessarily a contradiction. Dividing numerator and denominator by t we infer the central limit theorem is being applied to the number of observations per unit time (hence the title of the graph – frequencies for...).

Where c represents the number of cells (c = 10 in our case). The multinomial distribution has expected values  $E[o_j] = \mu_j = np_j$  and variance values  $\sigma_j = np_j(1 - p_j) = np_jq_j$ .

### Chi-squared test

Now consider for just a moment the case of c=1 which is to say there is only one cell or velocity interval in which observations are made. We perform n trials in which we attempt to observe an asteroid moving with the range of velocities that define this cell. There are only two outcomes; an asteroid was observed to moving at some velocity within the given velocity range of our cell on trial i (successful trial), or an asteroid was not observed to move within the given velocity range on trial i (failed trial), i.e. these are Bernoulli trials. If  $x_i \in 0, 1$  is our Bernoulli variable, we have  $\sum_i x_i = o_1$  as the total number of successful trials, and  $o_2 = n - o_1$  unsuccessful trials. The probability of success is  $p_1 = p$ .

Making the central limit approximation on a sum of n Bernoulli trials gives for the integration variable,

$$z = \frac{\sum_{i} x_{i} - n\mu}{\sqrt{n\sigma}} = \frac{o_{1} - np}{\sqrt{npq}}$$
 (2)

Now it is well known that A) for sufficiently large n a variable of the form of equation 2 is distributed as N[0,1], and B) for  $Z \sim N[0,1]$  we have  $z^2 \sim \mathcal{X}^2(1)$  where  $\mathcal{X}^2(v)$  denotes the Chi-squared distribution with v degrees of freedom. We can decompose  $z^2$  as follows,

$$z^{2} = \frac{(o_{1} - np)^{2}}{np(1 - p)} = \frac{(o_{1} - np)^{2}}{np(1 - p)} + \frac{(o_{1} - np)^{2}}{n(1 - p)}$$

$$= \frac{(o_{1} - np)^{2}}{np} + \frac{((n - o_{1}) - n(1 - p))^{2}}{n(1 - p)}$$

$$= \sum_{i=1}^{2} \frac{(o_{i} - np)^{2}}{np}$$
(3)

where  $p_1 = p$  is again the probability of a successful trial and  $p_2 = 1 - p$  is the probability of a failed trial. In terms of expected values  $e_1$ ,  $e_2$  for  $o_1$ ,  $o_2$ ,

$$z^{2} = \sum_{i=1}^{2} \frac{(o_{i} - e_{j})^{2}}{e_{j}} \sim \mathcal{X}^{2}(1)$$
(4)

So, for two distinct outcomes we got only one degree of freedom. In general, for k different outcomes we have that,

$$\sum_{i=1}^{k} \frac{(o_i - np_j)^2}{np_j} \sim \mathcal{X}^2(k-1)$$
 (5)

This is known as the Pearson Chi-squared statistic

Question; what would the corresponding normal variable(s) look like for the case of k > 2? Attempting the above decomposition for even the relatively simple case of k = 3 is difficult. There are however, a number of advanced theorems which show that equation 8 holds. Reference [1] provides no less than seven different proofs.

Likewise, there are subtleties in calculating the degrees of freedom. Noting that a sum of Chi-squared variables is distributed as a Chi-squared statistic with degrees of freedom equal to the sum of the degrees of freedom of the summed variables  $(\sum_{i}^{c} \mathcal{X}^{2}(v_{i}) \sim \mathcal{X}^{2}(\sum_{i}^{c} v_{i}))$ , from the above analysis it is tempting to say that c cells with k possible outcomes would produce c(k-1) degrees of freedom. In fact we'll have for our total degrees of freedom c-1. Why? A detailed proof can be found in  $[1]^{2}$  We will settle for noting that if an asteroid was not observed in the velocity interval of interest, then implicitly we've assumed there was an asteroid, it just was moving at some other velocity, ergo, it must exist in another cell! So the number of different outcomes k is not in all cases unrelated to the number of cells c.

If there are any parameters which need to be estimated then we'll have to subtract a degree of freedom for each variable estimated. Being as we've not been given values for the parameters  $\mu$ ,  $\sigma$  we'll need to estimate these parameters and so we will subtract 2 giving c-1-2=c-3 degrees of freedom.

Lastly, note that it sometimes occurs that cells need to be pooled together because a criteria of the Pearson Chi-squared test is that the expected values are  $e_j > 5$ . In such cases we of course count the number of grouped cells when calculating the degrees of freedom. Being as  $e_j = np_j$  the criteria that  $e_j > 5$  amounts to  $p_j > 5/n$ .

## 1 Estimating $\mu, \sigma$

We need to estimate the parameters  $\mu, \sigma$  by finding the maximum likelihood estimates (MLE's). Taking the natural log of equation 1 and differentiating with respect to the variables  $\mu$  and  $\sigma$  then setting the two resulting equations to zero gives,

$$\sum_{j} o_{j} \frac{\Phi'(z_{j+1}) - \Phi'(z_{j})}{\Phi(z_{j+1}) - \Phi(z_{j})} = 0$$
 (6)

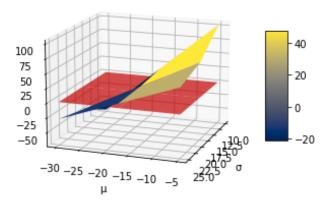
$$\sum_{j} o_{j} \frac{z_{j+1} \Phi'(z_{j+1}) - z_{j} \Phi'(z_{j})}{\Phi(z_{j+1}) - \Phi(z_{j})} = 0$$
(7)

where  $\Phi'$  denotes the derivative of an integral, so it is simply the standard normal distribution evaluated at its given argument. Equations 6 and 7 are solved via the two dimensional Newton-Raphson algorithm contained in raphson.py. The Newton-Raphson method is among the more primitive root-finding algorithms, yet with a little trial and error so as to [graphically or otherwise] guess the vicinity of a root, it serves its purpose in most cases.

First we take a guess. As described in the code contained in  $Ch\_test\_for\_normality\_interval\_data.py$  we first run the program with PLOT = True and RAPHSON = False. The graph of equation 6 gives a rough idea of what the values of  $\mu$ ,  $\sigma$  are. We ascertain this estimate by looking for where the three dimensional plane is zero. After some trial and error with what range of values over which we ought to plot  $\mu$ ,  $\sigma$  (and also different viewing angles) we get the following plot,

<sup>&</sup>lt;sup>2</sup>see Sixth Proof: Generic induction with De Moivre-Laplace theorem.

#### μ MLE function



So we estimate  $\mu, \sigma = -20, 5$ . Now we turn PLOT = False, RAPHSON = True and run the code again with this estimate to see if we can get an even better estimate from the Newton-raphson algorithm contained in raphson.py.

After some trial and error we find that [integral] equations 6 and 7 are very touchy, so we reluctantly increase the integration steps to 1/10,000 and we set the step size for calculating derivatives in the Newton-Raphson algorithm to 1/10,000 as well. Noting that equations 6 and 7 are not just integrated equations, but are a sum of integrated equations, and furthermore that we ought to expect the Newton-Raphson algorithm to run  $\tilde{5}$ -20 times before finding a root, we let the program run while we take a coffee break. We come back the program has estimated the values of  $\mu$ ,  $\sigma$  as,

$$\mu, \sigma \approx -19.867115036443728, 11.2626087946973$$

Which are not too far off from the reported values in  $[2]^3$  which gives  $\mu, \sigma \approx -21.3, 12.7$ .

These estimates enter into the Chi test via the calculation of the probabilities. We get for our Chi-squared statistic,

$$\sum_{j=1}^{k} \frac{(o_i - np_j)^2}{np_j} = 1.302 \tag{8}$$

The calculated value in [2] is 1.22.

Finally, we need a critical value to which we can compare our result to. It turns out we need to pool a number of cells in this example and we are left with c=4 cells. With degrees of freedom v=c-1-2=1 we have for a critical value  $\mathcal{X}^2(1) \approx 3.927$ . As noted in the code, the Chi-squared distribution explodes at the origin when v=1 so this is a very touchy integration in which we again need to make very small and time consuming integration steps. The result given in [2] are  $\mathcal{X}^2(1)=3.84$ .

Recalling that we originally set out to test  $Ho: X_i \sim N[\mu, \sigma^2]$ . We reject Ho If our test statistic is greater than the critical value. In our case 1.302 < 3.927 so we do not reject this hypothesis.

<sup>&</sup>lt;sup>3</sup>See top of page 457

# References

- [1] Eric Benhamou, Valentin Melot Seven proofs of the Pearson Chi-squared independence test and its graphical interpretation. AMS 1991 subject classification: 62E10, 62E15
- [2] Bain, Engelhardt Introduction to Probability and Mathematical Statistics [second ed.]. Brooks/ Cole, Belmont, CA