

$\rho(r) = \rho_0$

When r is much larger than r_0

Density decreases with distance in air

Initialize simulation and put in gravity, should settle down to it

Make plot of this with spherical coordinates

But look at how to actually run shock tube

Fluid

Mass density = ρ

Velocity components 3

Energy density (E)

Solve for five numbers at every grid point in computational grid

Have to give initial values

Just do one r and 1 r dependence on cartesian

Use 1 grid point θ and 1 in ϕ

Message passing interface

Make short file of steps

Read and look at history output file

- ten numbers at time stops

- momentum, energy, mass over whole grid

Look at slices

- output from code at one time

- what is value of five basic quantities at every grid point

Go and review python dictionaries

- access columns by name of column

- name of dictionary ['name of column']

Make sure to ask questions

Overtime, tidal friction tries to make them spin up or spin down to rotate once per orbit

- also tries to make it much more circular

ANGULAR MOMENTUM IS CONSERVED AND GIVEN TO ANOTHER

- from earth to moon

Angular momentum

$$L = mvr$$

R goes to positive power

If earth spins down, increased angular momentum of moons orbit \rightarrow radius increases

Out of earth's spin and transfer it to orbit

Know this because laser shoots to moon, rate is speed of light, distance can be solved by knowing the period

Angular momentum = $r \times v$

Thumb is direction of th

Angular momentum = $r^{1/2}$

Angular momentum in the solar system in the orbit of Jupiter

All mass goes to center, all angular momentum goes outside

Mass spiraling into center because you're sharing it out to larger radial disks