



**UNIVERSITY
OF LONDON**

EC2020 ZB

**BSc DEGREES AND GRADUATE DIPLOMAS IN ECONOMICS, MANAGEMENT,
FINANCE AND THE SOCIAL SCIENCES, THE DIPLOMA IN ECONOMICS AND
SOCIAL SCIENCES AND THE CERTIFICATE IN EDUCATION IN SOCIAL
SCIENCES**

Elements of Econometrics

Thursday 23 May 2019 : 10.00 – 13.00

Time allowed: 3 hours

DO NOT TURN OVER UNTIL TOLD TO BEGIN

Candidates should answer **EIGHT** of the following **TEN** questions: **ALL** of the questions in Section A (8 marks each) and **THREE** questions from Section B (20 marks each). **Candidates are strongly advised to divide their time accordingly.**

Extracts from statistical tables are given after the final question on this paper.

Graph paper is provided at the end of this question paper. If used, it must be detached and fastened securely inside the answer book.

A handheld calculator may be used when answering questions on this paper and it must comply in all respects with the specification given with your Admission Notice. The make and type of machine must be clearly stated on the front cover of the answer book.

If more questions are answered than requested, only the first answers attempted will be counted.

SECTION A

Answer all questions from this section

1. Consider the simple linear regression model

$$Y_i = \beta X_i + u_i, \quad i = 1, \dots, n.$$

We assume that the errors $\{u_i\}_{i=1}^n$ are independent normal random variables with zero mean. The regressor $\{X_i\}_{i=1}^n$ is non-stochastic (fixed under repeated sampling). You suspect that the errors exhibit heteroskedasticity.

- (a) **(3 marks)** Explain what we mean by the concept of heteroskedasticity. Enhance your answer with the help of a graphical illustration.
- (b) **(5 marks)** Derive the variance of the OLS estimator for $\hat{\beta}$ in the presence of heteroskedasticity and explain how a robust (White) standard error for $\hat{\beta}$ can be obtained.

2. Consider the human capital earnings function given by

$$\text{earnings}_i = \beta_0 + \beta_1 \text{educ}_i + \beta_2 \text{exper}_i + \beta_3 \text{exper}_i^2 + u_i, \quad i = 1, \dots, n,$$

where *earnings* denotes the hourly earnings of an individual and *educ* and *exper* denote the years of schooling and experience, respectively. We assume we have obtained a random sample $\{(\text{earnings}_i, \text{educ}_i, \text{exper}_i)\}_{i=1}^n$ from the population. The errors $\{u_i\}_{i=1}^n$ are i.i.d. normal random variables with zero mean and variance σ^2 . We assume independence between the errors and regressors (i.e., we ignore the usual ability bias problem).

- (a) **(2 marks)** Discuss the rationale for including both *exper* and *exper*² in this model. In your answer be explicit about the expected signs for β_2 and β_3 .
- (b) **(2 marks)** Discuss briefly how would you test that *exper* has a significant effect on *earnings*?
- (c) **(4 marks)** Provide a clear interpretation of the parameter β_1 . You are told that you can obtain the estimator for β_1 by running the regression

$$\text{earnings}_i^* = \beta_1 \text{educ}_i^* + e_i, \quad i = 1, \dots, n,$$

where earnings_i^* and educ_i^* are obtained from running a regression of *earnings* (and *educ*) on an intercept, *exper* and *exper*². Explain this statement.

3. Consider the OLS estimator for β in the linear regression model

$$Y_i = \beta X_i + \varepsilon_i, \quad i = 1, \dots, n,$$

where $\{(Y_i, X_i)\}_{i=1}^n$ form an i.i.d. sample from a population and the errors are drawn from an unknown distribution with mean zero and variance σ^2 . You are told that X_i and ε_i are uncorrelated (not necessarily independent therefore!).

- (a) **(4 marks)** Discuss the importance of convergence in probability.
- (b) **(4 marks)** Discuss the importance of convergence in distribution.

4. Consider the following regression model

$$Y_t = \beta_0 + \beta_1 X_t + \varepsilon_t, \quad t = 1, \dots, T.$$

Both $\{Y_t\}_{t=1}^T$ and $\{X_t\}_{t=1}^T$ are difference stationary processes.

- (a) **(4 marks)** Discuss the concept "difference stationarity" and contrast it to the concept "trend stationarity". In your answer make sure you also explain what stationarity means.
- (b) **(4 marks)** How can you test whether the above relation is spurious or cointegrating? In your answer you are expected to explain the difference between a spurious and cointegrating relationship.

5. Consider the simple linear regression model

$$Y_i = \beta_0 + \beta_1 X_i + u_i, \quad i = 1, \dots, n.$$

We assume that the errors $\{u_i\}_{i=1}^n$ are independent random variables with zero mean. The regressor $\{X_i\}_{i=1}^n$ is non-stochastic (fixed under repeated sampling). Under these conditions, the OLS estimator for β_1 , $\hat{\beta}_1$, is unbiased. (You are not asked to derive $\hat{\beta}_1$).

- (a) **(2 marks)** Explain the concept of unbiasedness of an estimator.
- (b) **(4 marks)** Let us consider two other estimators for the slope β_1 :

$$\hat{\beta}_1^\circ = \frac{\sum_{i=1}^n (Z_i - \bar{Z}) Y_i}{\sum_{i=1}^n (Z_i - \bar{Z}) X_i}; \quad \hat{\beta}_1^* = \frac{\sum_{i=1}^n (Z_i - \bar{Z}) Y_i}{\sum_{i=1}^n (Z_i - \bar{Z}) Z_i}$$

where $Z_i = \sqrt{X_i}$ for all i and $\bar{Z} = \frac{1}{n} \sum_{i=1}^n Z_i$. Please indicate whether $\hat{\beta}_1^\circ$ and $\hat{\beta}_1^*$ are unbiased estimators for β_1 . Clearly show your derivations.

- (c) **(2 marks)** Briefly indicate how you would choose between the three estimators, $\hat{\beta}_1$, $\hat{\beta}_1^\circ$ and $\hat{\beta}_1^*$.

SECTION B

Answer three questions from this section.

6. Let us consider the expectations augmented Phillips curve (see also Mankiw, 1994):

$$\text{infl}_t - \text{infl}_t^e = \beta_1 (\text{unem}_t - \mu_0) + e_t,$$

where μ_0 is the natural rate of unemployment (assumed to be constant over time) and infl_t^e is the expected rate of inflation formed in $t - 1$.

This model suggests that there is a tradeoff between unanticipated inflation ($\text{infl}_t - \text{infl}_t^e$) and cyclical unemployment (difference between actual unemployment and the natural rate of unemployment). We assume that e_t (also called supply shock) is an i.i.d. random variable with zero mean.

- (a) **(2 marks)** You are told that expectations are formed as follows

$$\text{infl}_t^e - \text{infl}_{t-1}^e = \lambda (\text{infl}_{t-1} - \text{infl}_{t-1}^e)$$

What name do we give such a process and how should we interpret λ ?

- (b) **(7 marks)** Show that you can rewrite the model as

$$\Delta \text{infl}_t = \gamma_0 + \gamma_1 \text{unem}_t + \gamma_2 \text{unem}_{t-1} + v_t, \quad (6.1)$$

where $\Delta \text{infl}_t = \text{infl}_t - \text{infl}_{t-1}$. Clearly indicate the relation between $(\gamma_0, \gamma_1, \gamma_2)$ and $(\mu_0, \beta_1, \lambda)$ and show that

$$v_t = e_t - (1 - \lambda) e_{t-1}.$$

Hint: If you want you may use the following shorthand notation in your derivations: $y_t = \text{infl}_t$, $y_t^e = \text{infl}_t^e$ and $x_t = \text{unem}_t$.

- (c) **(3 marks)** Discuss what assumptions you would like to make about e_t (the supply shock) that will guarantee consistency of the OLS estimator for (6.1). *Hint:* You may want to give the assumptions you need to make about v_t (composite error term) first.
- (d) **(3 marks)** Show how you can obtain a consistent estimator for λ using your consistent estimates for $(\gamma_0, \gamma_1, \gamma_2)$? Provide a proof of its consistency. [Note: If you did not manage to get an explicit relation between λ and $(\gamma_0, \gamma_1, \gamma_2)$, consider $\lambda = g(\gamma_0, \gamma_1, \gamma_2)$ where $g(\cdot)$ is some continuous function.]
- (e) **(5 marks)** One of the assumptions provided in (c) is rather unreasonable (*Hint:* future unemployment may be related to current supply shocks). Discuss how you could use IV (TSLS) to obtain a consistent estimator for the parameters in (6.1). Discuss what conditions your instruments need to satisfy and propose a suitable instrument.

7. This question is based on "Openness and Inflation: Theory and Evidence", by Romer (*Quarterly Journal of Economics*, 1993). Romer proposed theoretical models of inflation that imply that more "open" countries should have lower inflation. A simple macroeconomic model is

$$infl_i = \beta_0 + \beta_1 open_i + \beta_2 \ln(pcinc_i) + u_{1i} \quad (7.1)$$

$$open_i = \alpha_0 + \alpha_1 infl_i + \alpha_2 \ln(pcinc_i) + \alpha_3 \ln(land_i) + u_{2i}, \quad i = 1, \dots, n \quad (7.2)$$

where $infl$ is the average annual inflation rate (since 1973), $open$ is the average share of imports in gross domestic (or national) product since 1973, $\ln(pcinc)$ is the log of per capita income in US dollars, and $\ln(land)$ is the log of land area in square miles. The variables $\ln(pcinc)$ and $\ln(land)$ are treated as exogenous variables; u_{1i} and u_{2i} are serially uncorrelated disturbances with zero mean, variances σ_1^2 and σ_2^2 and covariance σ_{12} .

- (a) **(5 marks)** Explain the concepts of endogenous versus exogenous explanatory variables and show that $open$ is an endogenous variable in (7.1).
 (b) **(5 marks)** Discuss the identification of each structural form equation and explain what it means to say that an equation is exact identified.

The table below shows the OLS and IV estimation results for equation (7.1). Standard errors are reported in parenthesis.

| | OLS | IV |
|-----------------|--------------------|-------------------|
| <i>constant</i> | 25.109 (15.205) | 26.90 (15.40) |
| <i>open</i> | -0.215 (0.095) | -0.337 (0.144) |
| $\ln(pcinc)$ | 0.018 (1.975) | 0.376 (2.015) |

The IV estimator is obtained by solving the following three conditions (no need to show!):

$$\begin{aligned} \sum_{i=1}^n (infl_i - \hat{\beta}_0^{IV} - \hat{\beta}_1^{IV} open_i - \hat{\beta}_2^{IV} \ln(pcinc_i)) &= 0 \\ \sum_{i=1}^n (infl_i - \hat{\beta}_0^{IV} - \hat{\beta}_1^{IV} open_i - \hat{\beta}_2^{IV} \ln(pcinc_i)) \ln(pcinc_i) &= 0 \\ \sum_{i=1}^n (infl_i - \hat{\beta}_0^{IV} - \hat{\beta}_1^{IV} open_i - \hat{\beta}_2^{IV} \ln(pcinc_i)) \ln(land_i) &= 0 \end{aligned}$$

- (c) **(5 marks)** Describe in detail how you would estimate the parameters of (7.1) using Two Stage Least Squares (TSLS). How would these estimates compare with the reported IV estimates?
 (d) **(5 marks)** Conduct a test to see whether you can find evidence that more "open" countries have lower inflation. Clearly indicate the null and alternative hypothesis, test statistic and its (asymptotic) distribution under the null. Briefly indicate how you could test whether the OLS and IV results are significantly different (so that it matters whether you use the OLS or IV parameter estimates).

8. Let us consider how workplace smoking bans affect the incidence of smoking. Below, we use data on 10,000 US indoor workers from 1991 to 1993 taken from "Do Workplace Smoking Bans Reduce Smoking", by Evans et al. (*American Economic Review*, 1999). Let *smoker* be a dummy variable indicating whether a worker smokes (1=yes, 0=no) and *smkban* a dummy variable indicating whether there is a ban on smoking in the workplace (1=yes, 0=no).

- (a) **(5 marks)** OLS regression results were obtained for the simple linear regression model

$$\widehat{smoker} = \frac{0.290}{(0.007)} - \frac{0.078}{(0.009)} smkban \quad (8.1)$$

$$n = 10000, R^2 = 0.0078, RSS = 1821.59$$

The standard errors are in parentheses. Interpret the parameter estimates of the intercept and the coefficient on *smkban*. Provide the (approximate) 95% confidence interval for the coefficient on *smkban*. How can we use this confidence interval to test the hypothesis that $\beta_{smkban} = 0$?

A further specification was considered that included other characteristics of the worker: the age (in years), gender (male/female), ethnicity (black/hispanic/white), and level of education (*E1*=highschool dropout, *E2*=highschool graduate, *E3*=some college, *E4*=college graduate, *E5* Master degree or above). The following OLS regression results were obtained for this multiple linear regression model

$$\widehat{smoker} = \frac{0.201}{(0.019)} - \frac{0.045}{(0.009)} smkban - \frac{0.033}{(.009)} female - \frac{0.001}{(.0003)} age - \frac{0.027}{(.016)} black \quad (8.2)$$

$$- \frac{0.104}{(.014)} hispanic + \frac{0.310}{(.019)} E1 + \frac{0.224}{(.012)} E2 + \frac{0.156}{(.012)} E3 + \frac{0.042}{(.012)} E4$$

$$n = 10000, R^2 = 0.0526, RSS = 1736.81$$

The standard errors are in parentheses.

- (b) **(3 marks)** Compare the estimates of the coefficient on *smkban* from the simple and multiple regression model in (8.1) and (8.2) and explain why the estimates differ.
- (c) **(4 marks)** Interpret the estimated parameter on *E2* (highschool graduate) in (8.2) and indicate how you can obtain its p-value and what information the p-value provides.
- (d) **(4 marks)** Both the simple and multiple regression model suffer from heteroskedasticity. Explain why. What are the implications of heteroskedasticity for the parameter estimates and the standard errors in (8.1) and (8.2)? What can you do to resolve this problem? Explain your answer.

9. In question 8 we considered the linear regression model to study how workplace smoking bans affect the incidence of smoking. Here we consider the results from applying a probit regression of *smoker* (1=yes, 0=no) on *smkban* (1=yes, 0=no) and the other explanatory variables

```
. probit smoker smkban female age black hispanic E1 E2 E3 E4

Iteration 0:   log likelihood = -5537.1662
Iteration 1:   log likelihood = -5255.1526
Iteration 2:   log likelihood = -5252.349
Iteration 3:   log likelihood = -5252.3489

Probit regression               Number of obs   =    10,000
                                LR chi2(9)         =    569.63
                                Prob > chi2         =    0.0000
                                Pseudo R2          =    0.0514

Log likelihood = -5252.3489
```

| smoker | Coef. | Std. Err. | z | P> z | [95% Conf. Interval] |
|----------|-----------|-----------|--------|-------|----------------------|
| smkban | -.1517626 | .0289268 | -5.25 | 0.000 | -.208458 - .0950671 |
| female | -.1106249 | .0287785 | -3.84 | 0.000 | -.1670298 - .05422 |
| age | -.0042031 | .0011748 | -3.58 | 0.000 | -.0065057 - .0019006 |
| black | -.07969 | .0525369 | -1.52 | 0.129 | -.1826604 .0232804 |
| hispanic | -.3327039 | .0476677 | -6.98 | 0.000 | -.4261308 - .2392769 |
| E1 | 1.094231 | .0714121 | 15.32 | 0.000 | .9542663 1.234197 |
| E2 | .8518588 | .0594747 | 14.32 | 0.000 | .7352906 .9684271 |
| E3 | .6492566 | .0606989 | 10.70 | 0.000 | .530289 .7682241 |
| E4 | .2224747 | .0649939 | 3.42 | 0.001 | .0950891 .3498603 |
| _cons | -.9842425 | .0756055 | -13.02 | 0.000 | -1.132427 - .8360584 |

- (a) **(5 marks)** It is argued that using the probit regression model is better than using the linear probability model when explaining the binary variable *smoker*. Discuss the benefits/drawback of using the Probit model when trying to explain a binary variable.
- (b) **(5 marks)** Explain briefly how the Probit estimates are obtained and discuss the properties of the parameter estimates.

Hint: You may recall that for the Probit model, we will specify

$$\Pr(\text{smoker} = 1) = \Phi(\beta_0 + \beta_1 \text{smkban} + \beta_2 \text{female} + \dots + \beta_8 E3 + \beta_9 E4),$$

where Φ is the standard normal CDF (cumulative distribution function).

- (c) **(5 marks)** Explain how you can estimate the effect of the smoking ban on the probability of smoking for a 50-year old white, college graduated man. You are not expected to use your calculator, clarity of the computations required is enough.
- (d) **(5 marks)** Discuss how you could test the joint significance of the worker's characteristics (age, gender, ethnicity, and level of education) using the likelihood ratio test. Clearly indicate the test statistic, its distribution, the rejection rule and the additional information you would need to implement it.

10. This question is based on "An Empirical Comparison of Alternative Models of the Short-Term Interest Rate", by Chan et al. (*The Journal of Finance*, 1992). In this article, the authors are interested in a regression model for the short-term interest rate r_t given by

$$r_t - r_{t-1} = \beta_0 + \beta_1 r_{t-1} + \varepsilon_t. \quad (10.1)$$

Assume that the errors ε_t are not serially correlated.

- (a) **(5 marks)** Provide sufficient conditions (on β_1) for r_t to be stationary and weakly dependent. Will the OLS estimator for a linear regression of r_t on r_{t-1} in this case be unbiased and consistent? Explain your answer.
- (b) **(5 marks)** A few well-known models in finance hypothesize that $\beta_1 = 0$. Explain how you would test this hypothesis. What name do we give the process r_t when $\beta_1 = 0$.

Suppose you are interested in the following relation between long-term interest rates R_t and short-term interest rates r_t :

$$R_t - r_t = \alpha_0 + \alpha_1 r_t + u_t, \quad (10.2)$$

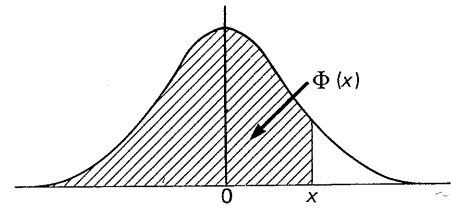
where u_t is an error term that has zero mean.

- (c) **(5 marks)** Let us assume $r_t = \beta_0 + \varepsilon_t$ (that is $\beta_1 = -1$). How would you test whether u_t is serially correlated? Explain your answer.
- (d) **(5 marks)** Let us assume $\beta_1 = 0$. Explaining your answers, what can you say about the OLS estimator for α_0 and α_1 applied to equation (10.2)
- i. if u_t is stationary and weakly dependent?
 - i. if u_t is non-stationary and strongly dependent?

TABLE 4. THE NORMAL DISTRIBUTION FUNCTION

The function tabulated is $\Phi(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^x e^{-\frac{1}{2}t^2} dt$. $\Phi(x)$ is

the probability that a random variable, normally distributed with zero mean and unit variance, will be less than or equal to x . When $x < 0$ use $\Phi(x) = 1 - \Phi(-x)$, as the normal distribution with zero mean and unit variance is symmetric about zero.



| x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ |
|------|-----------|------|-----------|------|-----------|------|-----------|------|-----------|------|-----------|
| 0.00 | 0.5000 | 0.40 | 0.6554 | 0.80 | 0.7881 | 1.20 | 0.8849 | 1.60 | 0.9452 | 2.00 | 0.97725 |
| 0.01 | .5040 | 0.41 | .6591 | 0.81 | .7910 | 1.21 | .8869 | 1.61 | .9463 | 2.01 | .97778 |
| 0.02 | .5080 | 0.42 | .6628 | 0.82 | .7939 | 1.22 | .8888 | 1.62 | .9474 | 2.02 | .97831 |
| 0.03 | .5120 | 0.43 | .6664 | 0.83 | .7967 | 1.23 | .8907 | 1.63 | .9484 | 2.03 | .97882 |
| 0.04 | .5160 | 0.44 | .6700 | 0.84 | .7995 | 1.24 | .8925 | 1.64 | .9495 | 2.04 | .97932 |
| 0.05 | .5199 | 0.45 | .6736 | 0.85 | .8023 | 1.25 | .8944 | 1.65 | .9505 | 2.05 | .97982 |
| 0.06 | .5239 | 0.46 | .6772 | 0.86 | .8051 | 1.26 | .8962 | 1.66 | .9515 | 2.06 | .98030 |
| 0.07 | .5279 | 0.47 | .6808 | 0.87 | .8078 | 1.27 | .8980 | 1.67 | .9525 | 2.07 | .98077 |
| 0.08 | .5319 | 0.48 | .6844 | 0.88 | .8106 | 1.28 | .8997 | 1.68 | .9535 | 2.08 | .98124 |
| 0.09 | .5359 | 0.49 | .6879 | 0.89 | .8133 | 1.29 | .9015 | 1.69 | .9545 | 2.09 | .98169 |
| 0.10 | .5398 | 0.50 | .6915 | 0.90 | .8159 | 1.30 | .9032 | 1.70 | .9554 | 2.10 | .98214 |
| 0.11 | .5438 | 0.51 | .6950 | 0.91 | .8186 | 1.31 | .9049 | 1.71 | .9564 | 2.11 | .98257 |
| 0.12 | .5478 | 0.52 | .6985 | 0.92 | .8212 | 1.32 | .9066 | 1.72 | .9573 | 2.12 | .98300 |
| 0.13 | .5517 | 0.53 | .7019 | 0.93 | .8238 | 1.33 | .9082 | 1.73 | .9582 | 2.13 | .98341 |
| 0.14 | .5557 | 0.54 | .7054 | 0.94 | .8264 | 1.34 | .9099 | 1.74 | .9591 | 2.14 | .98382 |
| 0.15 | .5596 | 0.55 | .7088 | 0.95 | .8289 | 1.35 | .9115 | 1.75 | .9599 | 2.15 | .98422 |
| 0.16 | .5636 | 0.56 | .7123 | 0.96 | .8315 | 1.36 | .9131 | 1.76 | .9608 | 2.16 | .98461 |
| 0.17 | .5675 | 0.57 | .7157 | 0.97 | .8340 | 1.37 | .9147 | 1.77 | .9616 | 2.17 | .98500 |
| 0.18 | .5714 | 0.58 | .7190 | 0.98 | .8365 | 1.38 | .9162 | 1.78 | .9625 | 2.18 | .98537 |
| 0.19 | .5753 | 0.59 | .7224 | 0.99 | .8389 | 1.39 | .9177 | 1.79 | .9633 | 2.19 | .98574 |
| 0.20 | .5793 | 0.60 | .7257 | 1.00 | .8413 | 1.40 | .9192 | 1.80 | .9641 | 2.20 | .98610 |
| 0.21 | .5832 | 0.61 | .7291 | 0.01 | .8438 | 1.41 | .9207 | 1.81 | .9649 | 2.21 | .98645 |
| 0.22 | .5871 | 0.62 | .7324 | 0.02 | .8461 | 1.42 | .9222 | 1.82 | .9656 | 2.22 | .98679 |
| 0.23 | .5910 | 0.63 | .7357 | 0.03 | .8485 | 1.43 | .9236 | 1.83 | .9664 | 2.23 | .98713 |
| 0.24 | .5948 | 0.64 | .7389 | 0.04 | .8508 | 1.44 | .9251 | 1.84 | .9671 | 2.24 | .98745 |
| 0.25 | .5987 | 0.65 | .7422 | 1.05 | .8531 | 1.45 | .9265 | 1.85 | .9678 | 2.25 | .98778 |
| 0.26 | .6026 | 0.66 | .7454 | 0.06 | .8554 | 1.46 | .9279 | 1.86 | .9686 | 2.26 | .98809 |
| 0.27 | .6064 | 0.67 | .7486 | 0.07 | .8577 | 1.47 | .9292 | 1.87 | .9693 | 2.27 | .98840 |
| 0.28 | .6103 | 0.68 | .7517 | 0.08 | .8599 | 1.48 | .9306 | 1.88 | .9699 | 2.28 | .98870 |
| 0.29 | .6141 | 0.69 | .7549 | 0.09 | .8621 | 1.49 | .9319 | 1.89 | .9706 | 2.29 | .98899 |
| 0.30 | .6179 | 0.70 | .7580 | 1.10 | .8643 | 1.50 | .9332 | 1.90 | .9713 | 2.30 | .98928 |
| 0.31 | .6217 | 0.71 | .7611 | 0.11 | .8665 | 1.51 | .9345 | 1.91 | .9719 | 2.31 | .98956 |
| 0.32 | .6255 | 0.72 | .7642 | 0.12 | .8686 | 1.52 | .9357 | 1.92 | .9726 | 2.32 | .98983 |
| 0.33 | .6293 | 0.73 | .7673 | 0.13 | .8708 | 1.53 | .9370 | 1.93 | .9732 | 2.33 | .99010 |
| 0.34 | .6331 | 0.74 | .7704 | 0.14 | .8729 | 1.54 | .9382 | 1.94 | .9738 | 2.34 | .99036 |
| 0.35 | .6368 | 0.75 | .7734 | 1.15 | .8749 | 1.55 | .9394 | 1.95 | .9744 | 2.35 | .99061 |
| 0.36 | .6406 | 0.76 | .7764 | 0.16 | .8770 | 1.56 | .9406 | 1.96 | .9750 | 2.36 | .99086 |
| 0.37 | .6443 | 0.77 | .7794 | 0.17 | .8790 | 1.57 | .9418 | 1.97 | .9756 | 2.37 | .99111 |
| 0.38 | .6480 | 0.78 | .7823 | 0.18 | .8810 | 1.58 | .9429 | 1.98 | .9761 | 2.38 | .99134 |
| 0.39 | .6517 | 0.79 | .7852 | 0.19 | .8830 | 1.59 | .9441 | 1.99 | .9767 | 2.39 | .99158 |
| 0.40 | .6554 | 0.80 | .7881 | 1.20 | .8849 | 1.60 | .9452 | 2.00 | .9772 | 2.40 | .99180 |

TABLE 4. THE NORMAL DISTRIBUTION FUNCTION

| x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ | x | $\Phi(x)$ |
|------|-----------|------|-----------|------|-----------|------|-----------|------|-----------|------|-----------|
| 2.40 | 0.99180 | 2.55 | 0.99461 | 2.70 | 0.99653 | 2.85 | 0.99781 | 3.00 | 0.99865 | 3.15 | 0.99918 |
| 41 | .99202 | 56 | .99477 | 71 | .99664 | 86 | .99788 | 01 | .99869 | 16 | .99921 |
| 42 | .99224 | 57 | .99492 | 72 | .99674 | 87 | .99795 | 02 | .99874 | 17 | .99924 |
| 43 | .99245 | 58 | .99506 | 73 | .99683 | 88 | .99801 | 03 | .99878 | 18 | .99926 |
| 44 | .99266 | 59 | .99520 | 74 | .99693 | 89 | .99807 | 04 | .99882 | 19 | .99929 |
| | | | | | | | | | | | |
| 2.45 | 0.99286 | 2.60 | 0.99534 | 2.75 | 0.99702 | 2.90 | 0.99813 | 3.05 | 0.99886 | 3.20 | 0.99931 |
| 46 | .99305 | 61 | .99547 | 76 | .99711 | 91 | .99819 | 06 | .99889 | 21 | .99934 |
| 47 | .99324 | 62 | .99560 | 77 | .99720 | 92 | .99825 | 07 | .99893 | 22 | .99936 |
| 48 | .99343 | 63 | .99573 | 78 | .99728 | 93 | .99831 | 08 | .99896 | 23 | .99938 |
| 49 | .99361 | 64 | .99585 | 79 | .99736 | 94 | .99836 | 09 | .99900 | 24 | .99940 |
| | | | | | | | | | | | |
| 2.50 | 0.99379 | 2.65 | 0.99598 | 2.80 | 0.99744 | 2.95 | 0.99841 | 3.10 | 0.99903 | 3.25 | 0.99942 |
| 51 | .99396 | 66 | .99609 | 81 | .99752 | 96 | .99846 | 11 | .99906 | 26 | .99944 |
| 52 | .99413 | 67 | .99621 | 82 | .99760 | 97 | .99851 | 12 | .99910 | 27 | .99946 |
| 53 | .99430 | 68 | .99632 | 83 | .99767 | 98 | .99856 | 13 | .99913 | 28 | .99948 |
| 54 | .99446 | 69 | .99643 | 84 | .99774 | 99 | .99861 | 14 | .99916 | 29 | .99950 |
| | | | | | | | | | | | |
| 2.55 | 0.99461 | 2.70 | 0.99653 | 2.85 | 0.99781 | 3.00 | 0.99865 | 3.15 | 0.99918 | 3.30 | 0.99952 |

The critical table below gives on the left the range of values of x for which $\Phi(x)$ takes the value on the right, correct to the last figure given; in critical cases, take the upper of the two values of $\Phi(x)$ indicated.

| | | | | | | | |
|-------|--------|-------|--------|-------|---------|-------|---------|
| 3.075 | 0.9990 | 3.263 | 0.9994 | 3.731 | 0.99990 | 3.916 | 0.99995 |
| 3.105 | 0.9990 | 3.320 | 0.9995 | 3.759 | 0.99991 | 3.976 | 0.99996 |
| 3.138 | 0.9991 | 3.389 | 0.9996 | 3.791 | 0.99992 | 4.055 | 0.99997 |
| 3.174 | 0.9992 | 3.480 | 0.9997 | 3.826 | 0.99993 | 4.173 | 0.99998 |
| 3.215 | 0.9993 | 3.615 | 0.9998 | 3.867 | 0.99994 | 4.417 | 0.99999 |
| | 0.9994 | | 0.9999 | | 0.99995 | | 1.00000 |

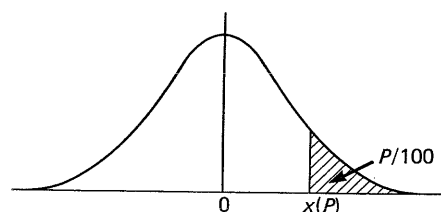
When $x > 3.3$ the formula $1 - \Phi(x) \doteq \frac{e^{-x^2}}{x\sqrt{2\pi}} \left[1 - \frac{1}{x^2} + \frac{3}{x^4} - \frac{15}{x^6} + \frac{105}{x^8} \right]$ is very accurate, with relative error less than $945/x^{10}$.

TABLE 5. PERCENTAGE POINTS OF THE NORMAL DISTRIBUTION

This table gives percentage points $x(P)$ defined by the equation

$$\frac{P}{100} = \frac{1}{\sqrt{2\pi}} \int_{x(P)}^{\infty} e^{-\frac{1}{2}t^2} dt.$$

If X is a variable, normally distributed with zero mean and unit variance, $P/100$ is the probability that $X \geq x(P)$. The lower P per cent points are given by symmetry as $-x(P)$, and the probability that $|X| \geq x(P)$ is $2P/100$.



| P | $x(P)$ | P | $x(P)$ | P | $x(P)$ | P | $x(P)$ | P | $x(P)$ | P | $x(P)$ |
|-----|--------|-----|--------|-----|--------|-----|--------|-----|--------|--------|--------|
| 50 | 0.0000 | 5.0 | 1.6449 | 3.0 | 1.8808 | 2.0 | 2.0537 | 1.0 | 2.3263 | 0.10 | 3.0902 |
| 45 | 0.1257 | 4.8 | 1.6646 | 2.9 | 1.8957 | 1.9 | 2.0749 | 0.9 | 2.3656 | 0.09 | 3.1214 |
| 40 | 0.2533 | 4.6 | 1.6849 | 2.8 | 1.9110 | 1.8 | 2.0969 | 0.8 | 2.4089 | 0.08 | 3.1559 |
| 35 | 0.3853 | 4.4 | 1.7060 | 2.7 | 1.9268 | 1.7 | 2.1201 | 0.7 | 2.4573 | 0.07 | 3.1947 |
| 30 | 0.5244 | 4.2 | 1.7279 | 2.6 | 1.9431 | 1.6 | 2.1444 | 0.6 | 2.5121 | 0.06 | 3.2389 |
| | | | | | | | | | | | |
| 25 | 0.6745 | 4.0 | 1.7507 | 2.5 | 1.9600 | 1.5 | 2.1701 | 0.5 | 2.5758 | 0.05 | 3.2905 |
| 20 | 0.8416 | 3.8 | 1.7744 | 2.4 | 1.9774 | 1.4 | 2.1973 | 0.4 | 2.6521 | 0.01 | 3.7190 |
| 15 | 1.0364 | 3.6 | 1.7991 | 2.3 | 1.9954 | 1.3 | 2.2262 | 0.3 | 2.7478 | 0.005 | 3.8906 |
| 10 | 1.2816 | 3.4 | 1.8250 | 2.2 | 2.0141 | 1.2 | 2.2571 | 0.2 | 2.8782 | 0.001 | 4.2649 |
| 5 | 1.6449 | 3.2 | 1.8522 | 2.1 | 2.0335 | 1.1 | 2.2904 | 0.1 | 3.0902 | 0.0005 | 4.4172 |

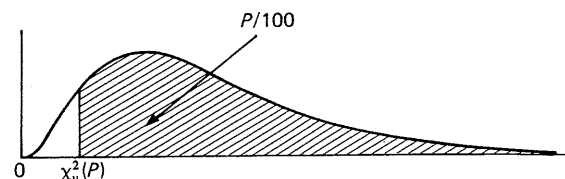
TABLE 8. PERCENTAGE POINTS OF THE χ^2 -DISTRIBUTION

This table gives percentage points $\chi^2_\nu(P)$ defined by the equation

$$\frac{P}{100} = \frac{1}{2^{\nu/2} \Gamma(\frac{\nu}{2})} \int_{\chi^2_\nu(P)}^{\infty} x^{\frac{\nu}{2}-1} e^{-\frac{1}{2}x} dx.$$

If X is a variable distributed as χ^2 with ν degrees of freedom, $P/100$ is the probability that $X \geq \chi^2_\nu(P)$.

For $\nu > 100$, $\sqrt{2X}$ is approximately normally distributed with mean $\sqrt{2\nu-1}$ and unit variance.



(The above shape applies for $\nu \geq 3$ only. When $\nu < 3$ the mode is at the origin.)

| P | 99.95 | 99.9 | 99.5 | 99 | 97.5 | 95 | 90 | 80 | 70 | 60 |
|-----------|----------|----------|----------|----------|----------|----------|---------|---------|--------|--------|
| $\nu = 1$ | 0.003927 | 0.001571 | 0.003927 | 0.001571 | 0.003927 | 0.003927 | 0.01579 | 0.06418 | 0.1485 | 0.2750 |
| 2 | 0.001000 | 0.002001 | 0.01003 | 0.02010 | 0.05064 | 0.1026 | 0.2107 | 0.4463 | 0.7133 | 1.022 |
| 3 | 0.01528 | 0.02430 | 0.07172 | 0.1148 | 0.2158 | 0.3518 | 0.5844 | 1.005 | 1.424 | 1.869 |
| 4 | 0.06392 | 0.09080 | 0.2070 | 0.2971 | 0.4844 | 0.7107 | 1.064 | 1.649 | 2.195 | 2.753 |
| 5 | 0.1581 | 0.2102 | 0.4117 | 0.5543 | 0.8312 | 1.145 | 1.610 | 2.343 | 3.000 | 3.655 |
| 6 | 0.2994 | 0.3811 | 0.6757 | 0.8721 | 1.237 | 1.635 | 2.204 | 3.070 | 3.828 | 4.570 |
| 7 | 0.4849 | 0.5985 | 0.9893 | 1.239 | 1.690 | 2.167 | 2.833 | 3.822 | 4.671 | 5.493 |
| 8 | 0.7104 | 0.8571 | 1.344 | 1.646 | 2.180 | 2.733 | 3.490 | 4.594 | 5.527 | 6.423 |
| 9 | 0.9717 | 1.152 | 1.735 | 2.088 | 2.700 | 3.325 | 4.168 | 5.380 | 6.393 | 7.357 |
| 10 | 1.265 | 1.479 | 2.156 | 2.558 | 3.247 | 3.940 | 4.865 | 6.179 | 7.267 | 8.295 |
| 11 | 1.587 | 1.834 | 2.603 | 3.053 | 3.816 | 4.575 | 5.578 | 6.989 | 8.148 | 9.237 |
| 12 | 1.934 | 2.214 | 3.074 | 3.571 | 4.404 | 5.226 | 6.304 | 7.807 | 9.034 | 10.18 |
| 13 | 2.305 | 2.617 | 3.565 | 4.107 | 5.009 | 5.892 | 7.042 | 8.634 | 9.926 | 11.13 |
| 14 | 2.697 | 3.041 | 4.075 | 4.660 | 5.629 | 6.571 | 7.790 | 9.467 | 10.82 | 12.08 |
| 15 | 3.108 | 3.483 | 4.601 | 5.229 | 6.262 | 7.261 | 8.547 | 10.31 | 11.72 | 13.03 |
| 16 | 3.536 | 3.942 | 5.142 | 5.812 | 6.908 | 7.962 | 9.312 | 11.15 | 12.62 | 13.98 |
| 17 | 3.980 | 4.416 | 5.697 | 6.408 | 7.564 | 8.672 | 10.09 | 12.00 | 13.53 | 14.94 |
| 18 | 4.439 | 4.905 | 6.265 | 7.015 | 8.231 | 9.390 | 10.86 | 12.86 | 14.44 | 15.89 |
| 19 | 4.912 | 5.407 | 6.844 | 7.633 | 8.907 | 10.12 | 11.65 | 13.72 | 15.35 | 16.85 |
| 20 | 5.398 | 5.921 | 7.434 | 8.260 | 9.591 | 10.85 | 12.44 | 14.58 | 16.27 | 17.81 |
| 21 | 5.896 | 6.447 | 8.034 | 8.897 | 10.28 | 11.59 | 13.24 | 15.44 | 17.18 | 18.77 |
| 22 | 6.404 | 6.983 | 8.643 | 9.542 | 10.98 | 12.34 | 14.04 | 16.31 | 18.10 | 19.73 |
| 23 | 6.924 | 7.529 | 9.260 | 10.20 | 11.69 | 13.09 | 14.85 | 17.19 | 19.02 | 20.69 |
| 24 | 7.453 | 8.085 | 9.886 | 10.86 | 12.40 | 13.85 | 15.66 | 18.06 | 19.94 | 21.65 |
| 25 | 7.991 | 8.649 | 10.52 | 11.52 | 13.12 | 14.61 | 16.47 | 18.94 | 20.87 | 22.62 |
| 26 | 8.538 | 9.222 | 11.16 | 12.20 | 13.84 | 15.38 | 17.29 | 19.82 | 21.79 | 23.58 |
| 27 | 9.093 | 9.803 | 11.81 | 12.88 | 14.57 | 16.15 | 18.11 | 20.70 | 22.72 | 24.54 |
| 28 | 9.656 | 10.39 | 12.46 | 13.56 | 15.31 | 16.93 | 18.94 | 21.59 | 23.65 | 25.51 |
| 29 | 10.23 | 10.99 | 13.12 | 14.26 | 16.05 | 17.71 | 19.77 | 22.48 | 24.58 | 26.48 |
| 30 | 10.80 | 11.59 | 13.79 | 14.95 | 16.79 | 18.49 | 20.60 | 23.36 | 25.51 | 27.44 |
| 32 | 11.98 | 12.81 | 15.13 | 16.36 | 18.29 | 20.07 | 22.27 | 25.15 | 27.37 | 29.38 |
| 34 | 13.18 | 14.06 | 16.50 | 17.79 | 19.81 | 21.66 | 23.95 | 26.94 | 29.24 | 31.31 |
| 36 | 14.40 | 15.32 | 17.89 | 19.23 | 21.34 | 23.27 | 25.64 | 28.73 | 31.12 | 33.25 |
| 38 | 15.64 | 16.61 | 19.29 | 20.69 | 22.88 | 24.88 | 27.34 | 30.54 | 32.99 | 35.19 |
| 40 | 16.91 | 17.92 | 20.71 | 22.16 | 24.43 | 26.51 | 29.05 | 32.34 | 34.87 | 37.13 |
| 50 | 23.46 | 24.67 | 27.99 | 29.71 | 32.36 | 34.76 | 37.69 | 41.45 | 44.31 | 46.86 |
| 60 | 30.34 | 31.74 | 35.53 | 37.48 | 40.48 | 43.19 | 46.46 | 50.64 | 53.81 | 56.62 |
| 70 | 37.47 | 39.04 | 43.28 | 45.44 | 48.76 | 51.74 | 55.33 | 59.90 | 63.35 | 66.40 |
| 80 | 44.79 | 46.52 | 51.17 | 53.54 | 57.15 | 60.39 | 64.28 | 69.21 | 72.92 | 76.19 |
| 90 | 52.28 | 54.16 | 59.20 | 61.75 | 65.65 | 69.13 | 73.29 | 78.56 | 82.51 | 85.99 |
| 100 | 59.90 | 61.92 | 67.33 | 70.06 | 74.22 | 77.93 | 82.36 | 87.95 | 92.13 | 95.81 |

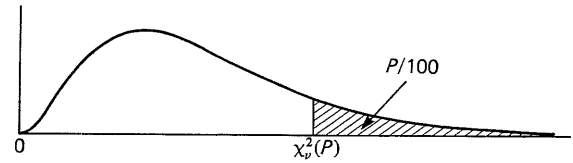
TABLE 8. PERCENTAGE POINTS OF THE χ^2 -DISTRIBUTION

This table gives percentage points $\chi^2_\nu(P)$ defined by the equation

$$\frac{P}{100} = \frac{1}{2^{\nu/2} \Gamma(\frac{\nu}{2})} \int_{\chi^2_\nu(P)}^{\infty} x^{\frac{\nu}{2}-1} e^{-\frac{1}{2}x} dx.$$

If X is a variable distributed as χ^2 with ν degrees of freedom, $P/100$ is the probability that $X \geq \chi^2_\nu(P)$.

For $\nu > 100$, $\sqrt{2X}$ is approximately normally distributed with mean $\sqrt{2\nu-1}$ and unit variance.



(The above shape applies for $\nu \geq 3$ only. When $\nu < 3$ the mode is at the origin.)

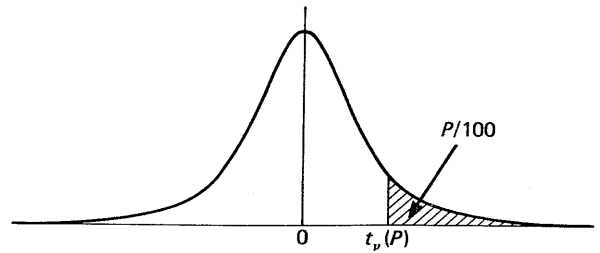
| P | 50 | 40 | 30 | 20 | 10 | 5 | 2.5 | 1 | 0.5 | 0.1 | 0.05 |
|-----------|--------|--------|-------|-------|-------|-------|-------|-------|-------|-------|-------|
| $\nu = 1$ | 0.4549 | 0.7083 | 1.074 | 1.642 | 2.706 | 3.841 | 5.024 | 6.635 | 7.879 | 10.83 | 12.12 |
| 2 | 1.386 | 1.833 | 2.408 | 3.219 | 4.605 | 5.991 | 7.378 | 9.210 | 10.60 | 13.82 | 15.20 |
| 3 | 2.366 | 2.946 | 3.665 | 4.642 | 6.251 | 7.815 | 9.348 | 11.34 | 12.84 | 16.27 | 17.73 |
| 4 | 3.357 | 4.045 | 4.878 | 5.989 | 7.779 | 9.488 | 11.14 | 13.28 | 14.86 | 18.47 | 20.00 |
| 5 | 4.351 | 5.132 | 6.064 | 7.289 | 9.236 | 11.07 | 12.83 | 15.09 | 16.75 | 20.52 | 22.11 |
| 6 | 5.348 | 6.211 | 7.231 | 8.558 | 10.64 | 12.59 | 14.45 | 16.81 | 18.55 | 22.46 | 24.10 |
| 7 | 6.346 | 7.283 | 8.383 | 9.803 | 12.02 | 14.07 | 16.01 | 18.48 | 20.28 | 24.32 | 26.02 |
| 8 | 7.344 | 8.351 | 9.524 | 11.03 | 13.36 | 15.51 | 17.53 | 20.09 | 21.95 | 26.12 | 27.87 |
| 9 | 8.343 | 9.414 | 10.66 | 12.24 | 14.68 | 16.92 | 19.02 | 21.67 | 23.59 | 27.88 | 29.67 |
| 10 | 9.342 | 10.47 | 11.78 | 13.44 | 15.99 | 18.31 | 20.48 | 23.21 | 25.19 | 29.59 | 31.42 |
| 11 | 10.34 | 11.53 | 12.90 | 14.63 | 17.28 | 19.68 | 21.92 | 24.72 | 26.76 | 31.26 | 33.14 |
| 12 | 11.34 | 12.58 | 14.01 | 15.81 | 18.55 | 21.03 | 23.34 | 26.22 | 28.30 | 32.91 | 34.82 |
| 13 | 12.34 | 13.64 | 15.12 | 16.98 | 19.81 | 22.36 | 24.74 | 27.69 | 29.82 | 34.53 | 36.48 |
| 14 | 13.34 | 14.69 | 16.22 | 18.15 | 21.06 | 23.68 | 26.12 | 29.14 | 31.32 | 36.12 | 38.11 |
| 15 | 14.34 | 15.73 | 17.32 | 19.31 | 22.31 | 25.00 | 27.49 | 30.58 | 32.80 | 37.70 | 39.72 |
| 16 | 15.34 | 16.78 | 18.42 | 20.47 | 23.54 | 26.30 | 28.85 | 32.00 | 34.27 | 39.25 | 41.31 |
| 17 | 16.34 | 17.82 | 19.51 | 21.61 | 24.77 | 27.59 | 30.19 | 33.41 | 35.72 | 40.79 | 42.88 |
| 18 | 17.34 | 18.87 | 20.60 | 22.76 | 25.99 | 28.87 | 31.53 | 34.81 | 37.16 | 42.31 | 44.43 |
| 19 | 18.34 | 19.91 | 21.69 | 23.90 | 27.20 | 30.14 | 32.85 | 36.19 | 38.58 | 43.82 | 45.97 |
| 20 | 19.34 | 20.95 | 22.77 | 25.04 | 28.41 | 31.41 | 34.17 | 37.57 | 40.00 | 45.31 | 47.50 |
| 21 | 20.34 | 21.99 | 23.86 | 26.17 | 29.62 | 32.67 | 35.48 | 38.93 | 41.40 | 46.80 | 49.01 |
| 22 | 21.34 | 23.03 | 24.94 | 27.30 | 30.81 | 33.92 | 36.78 | 40.29 | 42.80 | 48.27 | 50.51 |
| 23 | 22.34 | 24.07 | 26.02 | 28.43 | 32.01 | 35.17 | 38.08 | 41.64 | 44.18 | 49.73 | 52.00 |
| 24 | 23.34 | 25.11 | 27.10 | 29.55 | 33.20 | 36.42 | 39.36 | 42.98 | 45.56 | 51.18 | 53.48 |
| 25 | 24.34 | 26.14 | 28.17 | 30.68 | 34.38 | 37.65 | 40.65 | 44.31 | 46.93 | 52.62 | 54.95 |
| 26 | 25.34 | 27.18 | 29.25 | 31.79 | 35.56 | 38.89 | 41.92 | 45.64 | 48.29 | 54.05 | 56.41 |
| 27 | 26.34 | 28.21 | 30.32 | 32.91 | 36.74 | 40.11 | 43.19 | 46.96 | 49.64 | 55.48 | 57.86 |
| 28 | 27.34 | 29.25 | 31.39 | 34.03 | 37.92 | 41.34 | 44.46 | 48.28 | 50.99 | 56.89 | 59.30 |
| 29 | 28.34 | 30.28 | 32.46 | 35.14 | 39.09 | 42.56 | 45.72 | 49.59 | 52.34 | 58.30 | 60.73 |
| 30 | 29.34 | 31.32 | 33.53 | 36.25 | 40.26 | 43.77 | 46.98 | 50.89 | 53.67 | 59.70 | 62.16 |
| 32 | 31.34 | 33.38 | 35.66 | 38.47 | 42.58 | 46.19 | 49.48 | 53.49 | 56.33 | 62.49 | 65.00 |
| 34 | 33.34 | 35.44 | 37.80 | 40.68 | 44.90 | 48.60 | 51.97 | 56.06 | 58.96 | 65.25 | 67.80 |
| 36 | 35.34 | 37.50 | 39.92 | 42.88 | 47.21 | 51.00 | 54.44 | 58.62 | 61.58 | 67.99 | 70.59 |
| 38 | 37.34 | 39.56 | 42.05 | 45.08 | 49.51 | 53.38 | 56.90 | 61.16 | 64.18 | 70.70 | 73.35 |
| 40 | 39.34 | 41.62 | 44.16 | 47.27 | 51.81 | 55.76 | 59.34 | 63.69 | 66.77 | 73.40 | 76.09 |
| 50 | 49.33 | 51.89 | 54.72 | 58.16 | 63.17 | 67.50 | 71.42 | 76.15 | 79.49 | 86.66 | 89.56 |
| 60 | 59.33 | 62.13 | 65.23 | 68.97 | 74.40 | 79.08 | 83.30 | 88.38 | 91.95 | 99.61 | 102.7 |
| 70 | 69.33 | 72.36 | 75.69 | 79.71 | 85.53 | 90.53 | 95.02 | 100.4 | 104.2 | 112.3 | 115.6 |
| 80 | 79.33 | 82.57 | 86.12 | 90.41 | 96.58 | 101.9 | 106.6 | 112.3 | 116.3 | 124.8 | 128.3 |
| 90 | 89.33 | 92.76 | 96.52 | 101.1 | 107.6 | 113.1 | 118.1 | 124.1 | 128.3 | 137.2 | 140.8 |
| 100 | 99.33 | 102.9 | 106.9 | 111.7 | 118.5 | 124.3 | 129.6 | 135.8 | 140.2 | 149.4 | 153.2 |

TABLE 10. PERCENTAGE POINTS OF THE t -DISTRIBUTION

This table gives percentage points $t_\nu(P)$ defined by the equation

$$\frac{P}{100} = \frac{1}{\sqrt{\nu\pi}} \frac{\Gamma(\frac{1}{2}\nu + \frac{1}{2})}{\Gamma(\frac{1}{2}\nu)} \int_{t_\nu(P)}^{\infty} \frac{dt}{(1+t^2/\nu)^{\frac{1}{2}(\nu+1)}}.$$

Let X_1 and X_2 be independent random variables having a normal distribution with zero mean and unit variance and a χ^2 -distribution with ν degrees of freedom respectively; then $t = X_1/\sqrt{X_2/\nu}$ has Student's t -distribution with ν degrees of freedom, and the probability that $t \geq t_\nu(P)$ is $P/100$. The lower percentage points are given by symmetry as $-t_\nu(P)$, and the probability that $|t| \geq t_\nu(P)$ is $2P/100$.



The limiting distribution of t as ν tends to infinity is the normal distribution with zero mean and unit variance. When ν is large interpolation in ν should be harmonic.

| P | 40 | 30 | 25 | 20 | 15 | 10 | 5 | 2.5 | 1 | 0.5 | 0.1 | 0.05 |
|-----------|--------|--------|--------|--------|-------|-------|-------|-------|-------|-------|-------|-------|
| $\nu = 1$ | 0.3249 | 0.7265 | 1.0000 | 1.3764 | 1.963 | 3.078 | 6.314 | 12.71 | 31.82 | 63.66 | 318.3 | 636.6 |
| 2 | 0.2887 | 0.6172 | 0.8165 | 1.0607 | 1.386 | 1.886 | 2.920 | 4.303 | 6.965 | 9.925 | 22.33 | 31.60 |
| 3 | 0.2767 | 0.5844 | 0.7649 | 0.9785 | 1.250 | 1.638 | 2.353 | 3.182 | 4.541 | 5.841 | 10.21 | 12.92 |
| 4 | 0.2707 | 0.5686 | 0.7407 | 0.9410 | 1.190 | 1.533 | 2.132 | 2.776 | 3.747 | 4.604 | 7.173 | 8.610 |
| 5 | 0.2672 | 0.5594 | 0.7267 | 0.9195 | 1.156 | 1.476 | 2.015 | 2.571 | 3.365 | 4.032 | 5.893 | 6.869 |
| 6 | 0.2648 | 0.5534 | 0.7176 | 0.9057 | 1.134 | 1.440 | 1.943 | 2.447 | 3.143 | 3.707 | 5.208 | 5.959 |
| 7 | 0.2632 | 0.5491 | 0.7111 | 0.8960 | 1.119 | 1.415 | 1.895 | 2.365 | 2.998 | 3.499 | 4.785 | 5.408 |
| 8 | 0.2619 | 0.5459 | 0.7064 | 0.8889 | 1.108 | 1.397 | 1.860 | 2.306 | 2.896 | 3.355 | 4.501 | 5.041 |
| 9 | 0.2610 | 0.5435 | 0.7027 | 0.8834 | 1.100 | 1.383 | 1.833 | 2.262 | 2.821 | 3.250 | 4.297 | 4.781 |
| 10 | 0.2602 | 0.5415 | 0.6998 | 0.8791 | 1.093 | 1.372 | 1.812 | 2.228 | 2.764 | 3.169 | 4.144 | 4.587 |
| 11 | 0.2596 | 0.5399 | 0.6974 | 0.8755 | 1.088 | 1.363 | 1.796 | 2.201 | 2.718 | 3.106 | 4.025 | 4.437 |
| 12 | 0.2590 | 0.5386 | 0.6955 | 0.8726 | 1.083 | 1.356 | 1.782 | 2.179 | 2.681 | 3.055 | 3.930 | 4.318 |
| 13 | 0.2586 | 0.5375 | 0.6938 | 0.8702 | 1.079 | 1.350 | 1.771 | 2.160 | 2.650 | 3.012 | 3.852 | 4.221 |
| 14 | 0.2582 | 0.5366 | 0.6924 | 0.8681 | 1.076 | 1.345 | 1.761 | 2.145 | 2.624 | 2.977 | 3.787 | 4.140 |
| 15 | 0.2579 | 0.5357 | 0.6912 | 0.8662 | 1.074 | 1.341 | 1.753 | 2.131 | 2.602 | 2.947 | 3.733 | 4.073 |
| 16 | 0.2576 | 0.5350 | 0.6901 | 0.8647 | 1.071 | 1.337 | 1.746 | 2.120 | 2.583 | 2.921 | 3.686 | 4.015 |
| 17 | 0.2573 | 0.5344 | 0.6892 | 0.8633 | 1.069 | 1.333 | 1.740 | 2.110 | 2.567 | 2.898 | 3.646 | 3.965 |
| 18 | 0.2571 | 0.5338 | 0.6884 | 0.8620 | 1.067 | 1.330 | 1.734 | 2.101 | 2.552 | 2.878 | 3.610 | 3.922 |
| 19 | 0.2569 | 0.5333 | 0.6876 | 0.8610 | 1.066 | 1.328 | 1.729 | 2.093 | 2.539 | 2.861 | 3.579 | 3.883 |
| 20 | 0.2567 | 0.5329 | 0.6870 | 0.8600 | 1.064 | 1.325 | 1.725 | 2.086 | 2.528 | 2.845 | 3.552 | 3.850 |
| 21 | 0.2566 | 0.5325 | 0.6864 | 0.8591 | 1.063 | 1.323 | 1.721 | 2.080 | 2.518 | 2.831 | 3.527 | 3.819 |
| 22 | 0.2564 | 0.5321 | 0.6858 | 0.8583 | 1.061 | 1.321 | 1.717 | 2.074 | 2.508 | 2.819 | 3.505 | 3.792 |
| 23 | 0.2563 | 0.5317 | 0.6853 | 0.8575 | 1.060 | 1.319 | 1.714 | 2.069 | 2.500 | 2.807 | 3.485 | 3.768 |
| 24 | 0.2562 | 0.5314 | 0.6848 | 0.8569 | 1.059 | 1.318 | 1.711 | 2.064 | 2.492 | 2.797 | 3.467 | 3.745 |
| 25 | 0.2561 | 0.5312 | 0.6844 | 0.8562 | 1.058 | 1.316 | 1.708 | 2.060 | 2.485 | 2.787 | 3.450 | 3.725 |
| 26 | 0.2560 | 0.5309 | 0.6840 | 0.8557 | 1.058 | 1.315 | 1.706 | 2.056 | 2.479 | 2.779 | 3.435 | 3.707 |
| 27 | 0.2559 | 0.5306 | 0.6837 | 0.8551 | 1.057 | 1.314 | 1.703 | 2.052 | 2.473 | 2.771 | 3.421 | 3.690 |
| 28 | 0.2558 | 0.5304 | 0.6834 | 0.8546 | 1.056 | 1.313 | 1.701 | 2.048 | 2.467 | 2.763 | 3.408 | 3.674 |
| 29 | 0.2557 | 0.5302 | 0.6830 | 0.8542 | 1.055 | 1.311 | 1.699 | 2.045 | 2.462 | 2.756 | 3.396 | 3.659 |
| 30 | 0.2556 | 0.5300 | 0.6828 | 0.8538 | 1.055 | 1.310 | 1.697 | 2.042 | 2.457 | 2.750 | 3.385 | 3.646 |
| 32 | 0.2555 | 0.5297 | 0.6822 | 0.8530 | 1.054 | 1.309 | 1.694 | 2.037 | 2.449 | 2.738 | 3.365 | 3.622 |
| 34 | 0.2553 | 0.5294 | 0.6818 | 0.8523 | 1.052 | 1.307 | 1.691 | 2.032 | 2.441 | 2.728 | 3.348 | 3.601 |
| 36 | 0.2552 | 0.5291 | 0.6814 | 0.8517 | 1.052 | 1.306 | 1.688 | 2.028 | 2.434 | 2.719 | 3.333 | 3.582 |
| 38 | 0.2551 | 0.5288 | 0.6810 | 0.8512 | 1.051 | 1.304 | 1.686 | 2.024 | 2.429 | 2.712 | 3.319 | 3.566 |
| 40 | 0.2550 | 0.5286 | 0.6807 | 0.8507 | 1.050 | 1.303 | 1.684 | 2.021 | 2.423 | 2.704 | 3.307 | 3.551 |
| 50 | 0.2547 | 0.5278 | 0.6794 | 0.8489 | 1.047 | 1.299 | 1.676 | 2.009 | 2.403 | 2.678 | 3.261 | 3.496 |
| 60 | 0.2545 | 0.5272 | 0.6786 | 0.8477 | 1.045 | 1.296 | 1.671 | 2.000 | 2.390 | 2.660 | 3.232 | 3.460 |
| 120 | 0.2539 | 0.5258 | 0.6765 | 0.8446 | 1.041 | 1.289 | 1.658 | 1.980 | 2.358 | 2.617 | 3.160 | 3.373 |
| ∞ | 0.2533 | 0.5244 | 0.6745 | 0.8416 | 1.036 | 1.282 | 1.645 | 1.960 | 2.326 | 2.576 | 3.090 | 3.291 |

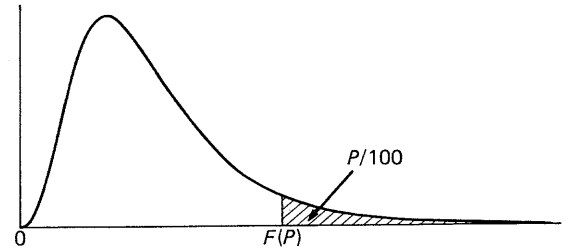
TABLE 12(a). 10 PER CENT POINTS OF THE F-DISTRIBUTION

The function tabulated is $F(P) = F(P|\nu_1, \nu_2)$ defined by the equation

$$\frac{P}{100} = \frac{\Gamma(\frac{1}{2}\nu_1 + \frac{1}{2}\nu_2)}{\Gamma(\frac{1}{2}\nu_1) \Gamma(\frac{1}{2}\nu_2)} \nu_1^{\frac{1}{2}\nu_1} \nu_2^{\frac{1}{2}\nu_2} \int_{F(P)}^{\infty} \frac{F^{\frac{1}{2}\nu_1-1}}{(\nu_2 + \nu_1 F)^{\frac{1}{2}(\nu_1+\nu_2)}} dF,$$

for $P = 10, 5, 2.5, 1, 0.5$ and 0.1 . The lower percentage points, that is the values $F'(P) = F'(P|\nu_1, \nu_2)$ such that the probability that $F \leq F'(P)$ is equal to $P/100$, may be found by the formula

$$F'(P|\nu_1, \nu_2) = 1/F(P|\nu_2, \nu_1).$$

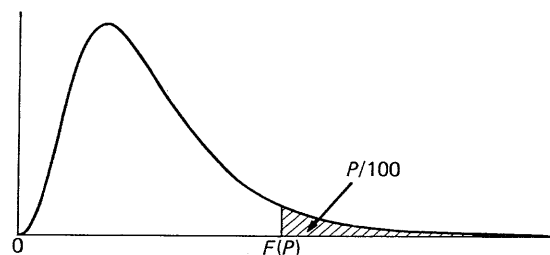


(This shape applies only when $\nu_1 \geq 3$. When $\nu_1 < 3$ the mode is at the origin.)

| $\nu_1 =$ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 10 | 12 | 24 | ∞ |
|-------------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|----------|
| $\nu_2 = 1$ | 39.86 | 49.50 | 53.59 | 55.83 | 57.24 | 58.20 | 58.91 | 59.44 | 60.19 | 60.71 | 62.00 | 63.33 |
| 2 | 8.526 | 9.000 | 9.162 | 9.243 | 9.293 | 9.326 | 9.349 | 9.367 | 9.392 | 9.408 | 9.450 | 9.491 |
| 3 | 5.538 | 5.462 | 5.391 | 5.343 | 5.309 | 5.285 | 5.266 | 5.252 | 5.230 | 5.216 | 5.176 | 5.134 |
| 4 | 4.545 | 4.325 | 4.191 | 4.107 | 4.051 | 4.010 | 3.979 | 3.955 | 3.920 | 3.896 | 3.831 | 3.761 |
| 5 | 4.060 | 3.780 | 3.619 | 3.520 | 3.453 | 3.405 | 3.368 | 3.339 | 3.297 | 3.268 | 3.191 | 3.105 |
| 6 | 3.776 | 3.463 | 3.289 | 3.181 | 3.108 | 3.055 | 3.014 | 2.983 | 2.937 | 2.905 | 2.818 | 2.722 |
| 7 | 3.589 | 3.257 | 3.074 | 2.961 | 2.883 | 2.827 | 2.785 | 2.752 | 2.703 | 2.668 | 2.575 | 2.471 |
| 8 | 3.458 | 3.113 | 2.924 | 2.806 | 2.726 | 2.668 | 2.624 | 2.589 | 2.538 | 2.502 | 2.404 | 2.293 |
| 9 | 3.360 | 3.006 | 2.813 | 2.693 | 2.611 | 2.551 | 2.505 | 2.469 | 2.416 | 2.379 | 2.277 | 2.159 |
| 10 | 3.285 | 2.924 | 2.728 | 2.605 | 2.522 | 2.461 | 2.414 | 2.377 | 2.323 | 2.284 | 2.178 | 2.055 |
| 11 | 3.225 | 2.860 | 2.660 | 2.536 | 2.451 | 2.389 | 2.342 | 2.304 | 2.248 | 2.209 | 2.100 | 1.972 |
| 12 | 3.177 | 2.807 | 2.606 | 2.480 | 2.394 | 2.331 | 2.283 | 2.245 | 2.188 | 2.147 | 2.036 | 1.904 |
| 13 | 3.136 | 2.763 | 2.560 | 2.434 | 2.347 | 2.283 | 2.234 | 2.195 | 2.138 | 2.097 | 1.983 | 1.846 |
| 14 | 3.102 | 2.726 | 2.522 | 2.395 | 2.307 | 2.243 | 2.193 | 2.154 | 2.095 | 2.054 | 1.938 | 1.797 |
| 15 | 3.073 | 2.695 | 2.490 | 2.361 | 2.273 | 2.208 | 2.158 | 2.119 | 2.059 | 2.017 | 1.899 | 1.755 |
| 16 | 3.048 | 2.668 | 2.462 | 2.333 | 2.244 | 2.178 | 2.128 | 2.088 | 2.028 | 1.985 | 1.866 | 1.718 |
| 17 | 3.026 | 2.645 | 2.437 | 2.308 | 2.218 | 2.152 | 2.102 | 2.061 | 2.001 | 1.958 | 1.836 | 1.686 |
| 18 | 3.007 | 2.624 | 2.416 | 2.286 | 2.196 | 2.130 | 2.079 | 2.038 | 1.977 | 1.933 | 1.810 | 1.657 |
| 19 | 2.990 | 2.606 | 2.397 | 2.266 | 2.176 | 2.109 | 2.058 | 2.017 | 1.956 | 1.912 | 1.787 | 1.631 |
| 20 | 2.975 | 2.589 | 2.380 | 2.249 | 2.158 | 2.091 | 2.040 | 1.999 | 1.937 | 1.892 | 1.767 | 1.607 |
| 21 | 2.961 | 2.575 | 2.365 | 2.233 | 2.142 | 2.075 | 2.023 | 1.982 | 1.920 | 1.875 | 1.748 | 1.586 |
| 22 | 2.949 | 2.561 | 2.351 | 2.219 | 2.128 | 2.060 | 2.008 | 1.967 | 1.904 | 1.859 | 1.731 | 1.567 |
| 23 | 2.937 | 2.549 | 2.339 | 2.207 | 2.115 | 2.047 | 1.995 | 1.953 | 1.890 | 1.845 | 1.716 | 1.549 |
| 24 | 2.927 | 2.538 | 2.327 | 2.195 | 2.103 | 2.035 | 1.983 | 1.941 | 1.877 | 1.832 | 1.702 | 1.533 |
| 25 | 2.918 | 2.528 | 2.317 | 2.184 | 2.092 | 2.024 | 1.971 | 1.929 | 1.866 | 1.820 | 1.689 | 1.518 |
| 26 | 2.909 | 2.519 | 2.307 | 2.174 | 2.082 | 2.014 | 1.961 | 1.919 | 1.855 | 1.809 | 1.677 | 1.504 |
| 27 | 2.901 | 2.511 | 2.299 | 2.165 | 2.073 | 2.005 | 1.952 | 1.909 | 1.845 | 1.799 | 1.666 | 1.491 |
| 28 | 2.894 | 2.503 | 2.291 | 2.157 | 2.064 | 1.996 | 1.943 | 1.900 | 1.836 | 1.790 | 1.656 | 1.478 |
| 29 | 2.887 | 2.495 | 2.283 | 2.149 | 2.057 | 1.988 | 1.935 | 1.892 | 1.827 | 1.781 | 1.647 | 1.467 |
| 30 | 2.881 | 2.489 | 2.276 | 2.142 | 2.049 | 1.980 | 1.927 | 1.884 | 1.819 | 1.773 | 1.638 | 1.456 |
| 32 | 2.869 | 2.477 | 2.263 | 2.129 | 2.036 | 1.967 | 1.913 | 1.870 | 1.805 | 1.758 | 1.622 | 1.437 |
| 34 | 2.859 | 2.466 | 2.252 | 2.118 | 2.024 | 1.955 | 1.901 | 1.858 | 1.793 | 1.745 | 1.608 | 1.419 |
| 36 | 2.850 | 2.456 | 2.243 | 2.108 | 2.014 | 1.945 | 1.891 | 1.847 | 1.781 | 1.734 | 1.595 | 1.404 |
| 38 | 2.842 | 2.448 | 2.234 | 2.099 | 2.005 | 1.935 | 1.881 | 1.838 | 1.772 | 1.724 | 1.584 | 1.390 |
| 40 | 2.835 | 2.440 | 2.226 | 2.091 | 1.997 | 1.927 | 1.873 | 1.829 | 1.763 | 1.715 | 1.574 | 1.377 |
| 60 | 2.791 | 2.393 | 2.177 | 2.041 | 1.946 | 1.875 | 1.819 | 1.775 | 1.707 | 1.657 | 1.511 | 1.291 |
| 120 | 2.748 | 2.347 | 2.130 | 1.992 | 1.896 | 1.824 | 1.767 | 1.722 | 1.652 | 1.601 | 1.447 | 1.193 |
| ∞ | 2.706 | 2.303 | 2.084 | 1.945 | 1.847 | 1.774 | 1.717 | 1.670 | 1.599 | 1.546 | 1.383 | 1.000 |

TABLE 12(b). 5 PER CENT POINTS OF THE F-DISTRIBUTION

If $F = \frac{X_1/\nu_1}{X_2/\nu_2}$, where X_1 and X_2 are independent random variables distributed as χ^2 with ν_1 and ν_2 degrees of freedom respectively, then the probabilities that $F \geq F(P)$ and that $F \leq F(P)$ are both equal to $P/100$. Linear interpolation in ν_1 and ν_2 will generally be sufficiently accurate except when either $\nu_1 > 12$ or $\nu_2 > 40$, when harmonic interpolation should be used.

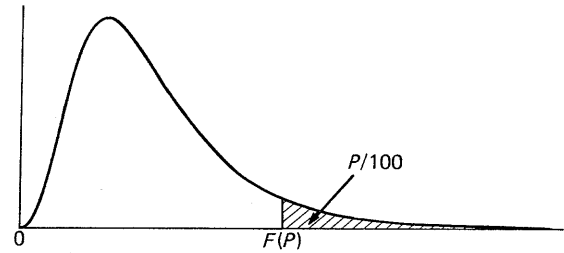


(This shape applies only when $\nu_1 \geq 3$. When $\nu_1 < 3$ the mode is at the origin.)

| $\nu_1 =$ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 10 | 12 | 24 | ∞ |
|-------------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|----------|
| $\nu_2 = 1$ | 161.4 | 199.5 | 215.7 | 224.6 | 230.2 | 234.0 | 236.8 | 238.9 | 241.9 | 243.9 | 249.1 | 254.3 |
| 2 | 18.51 | 19.00 | 19.16 | 19.25 | 19.30 | 19.33 | 19.35 | 19.37 | 19.40 | 19.41 | 19.45 | 19.50 |
| 3 | 10.13 | 9.552 | 9.277 | 9.117 | 9.013 | 8.941 | 8.887 | 8.845 | 8.786 | 8.745 | 8.639 | 8.526 |
| 4 | 7.709 | 6.944 | 6.591 | 6.388 | 6.256 | 6.163 | 6.094 | 6.041 | 5.964 | 5.912 | 5.774 | 5.628 |
| 5 | 6.608 | 5.786 | 5.409 | 5.192 | 5.050 | 4.950 | 4.876 | 4.818 | 4.735 | 4.678 | 4.527 | 4.365 |
| 6 | 5.987 | 5.143 | 4.757 | 4.534 | 4.387 | 4.284 | 4.207 | 4.147 | 4.060 | 4.000 | 3.841 | 3.669 |
| 7 | 5.591 | 4.737 | 4.347 | 4.120 | 3.972 | 3.866 | 3.787 | 3.726 | 3.637 | 3.575 | 3.410 | 3.230 |
| 8 | 5.318 | 4.459 | 4.066 | 3.838 | 3.687 | 3.581 | 3.500 | 3.438 | 3.347 | 3.284 | 3.115 | 2.928 |
| 9 | 5.117 | 4.256 | 3.863 | 3.633 | 3.482 | 3.374 | 3.293 | 3.230 | 3.137 | 3.073 | 2.900 | 2.707 |
| 10 | 4.965 | 4.103 | 3.708 | 3.478 | 3.326 | 3.217 | 3.135 | 3.072 | 2.978 | 2.913 | 2.737 | 2.538 |
| 11 | 4.844 | 3.982 | 3.587 | 3.357 | 3.204 | 3.095 | 3.012 | 2.948 | 2.854 | 2.788 | 2.609 | 2.404 |
| 12 | 4.747 | 3.885 | 3.490 | 3.259 | 3.106 | 2.996 | 2.913 | 2.849 | 2.753 | 2.687 | 2.505 | 2.296 |
| 13 | 4.667 | 3.806 | 3.411 | 3.179 | 3.025 | 2.915 | 2.832 | 2.767 | 2.671 | 2.604 | 2.420 | 2.206 |
| 14 | 4.600 | 3.739 | 3.344 | 3.112 | 2.958 | 2.848 | 2.764 | 2.699 | 2.602 | 2.534 | 2.349 | 2.131 |
| 15 | 4.543 | 3.682 | 3.287 | 3.056 | 2.901 | 2.790 | 2.707 | 2.641 | 2.544 | 2.475 | 2.288 | 2.066 |
| 16 | 4.494 | 3.634 | 3.239 | 3.007 | 2.852 | 2.741 | 2.657 | 2.591 | 2.494 | 2.425 | 2.235 | 2.010 |
| 17 | 4.451 | 3.592 | 3.197 | 2.965 | 2.810 | 2.699 | 2.614 | 2.548 | 2.450 | 2.381 | 2.190 | 1.960 |
| 18 | 4.414 | 3.555 | 3.160 | 2.928 | 2.773 | 2.661 | 2.577 | 2.510 | 2.412 | 2.342 | 2.150 | 1.917 |
| 19 | 4.381 | 3.522 | 3.127 | 2.895 | 2.740 | 2.628 | 2.544 | 2.477 | 2.378 | 2.308 | 2.114 | 1.878 |
| 20 | 4.351 | 3.493 | 3.098 | 2.866 | 2.711 | 2.599 | 2.514 | 2.447 | 2.348 | 2.278 | 2.082 | 1.843 |
| 21 | 4.325 | 3.467 | 3.072 | 2.840 | 2.685 | 2.573 | 2.488 | 2.420 | 2.321 | 2.250 | 2.054 | 1.812 |
| 22 | 4.301 | 3.443 | 3.049 | 2.817 | 2.661 | 2.549 | 2.464 | 2.397 | 2.297 | 2.226 | 2.028 | 1.783 |
| 23 | 4.279 | 3.422 | 3.028 | 2.796 | 2.640 | 2.528 | 2.442 | 2.375 | 2.275 | 2.204 | 2.005 | 1.757 |
| 24 | 4.260 | 3.403 | 3.009 | 2.776 | 2.621 | 2.508 | 2.423 | 2.355 | 2.255 | 2.183 | 1.984 | 1.733 |
| 25 | 4.242 | 3.385 | 2.991 | 2.759 | 2.603 | 2.490 | 2.405 | 2.337 | 2.236 | 2.165 | 1.964 | 1.711 |
| 26 | 4.225 | 3.369 | 2.975 | 2.743 | 2.587 | 2.474 | 2.388 | 2.321 | 2.220 | 2.148 | 1.946 | 1.691 |
| 27 | 4.210 | 3.354 | 2.960 | 2.728 | 2.572 | 2.459 | 2.373 | 2.305 | 2.204 | 2.132 | 1.930 | 1.672 |
| 28 | 4.196 | 3.340 | 2.947 | 2.714 | 2.558 | 2.445 | 2.359 | 2.291 | 2.190 | 2.118 | 1.915 | 1.654 |
| 29 | 4.183 | 3.328 | 2.934 | 2.701 | 2.545 | 2.432 | 2.346 | 2.278 | 2.177 | 2.104 | 1.901 | 1.638 |
| 30 | 4.171 | 3.316 | 2.922 | 2.690 | 2.534 | 2.421 | 2.334 | 2.266 | 2.165 | 2.092 | 1.887 | 1.622 |
| 32 | 4.149 | 3.295 | 2.901 | 2.668 | 2.512 | 2.399 | 2.313 | 2.244 | 2.142 | 2.070 | 1.864 | 1.594 |
| 34 | 4.130 | 3.276 | 2.883 | 2.650 | 2.494 | 2.380 | 2.294 | 2.225 | 2.123 | 2.050 | 1.843 | 1.569 |
| 36 | 4.113 | 3.259 | 2.866 | 2.634 | 2.477 | 2.364 | 2.277 | 2.209 | 2.106 | 2.033 | 1.824 | 1.547 |
| 38 | 4.098 | 3.245 | 2.852 | 2.619 | 2.463 | 2.349 | 2.262 | 2.194 | 2.091 | 2.017 | 1.808 | 1.527 |
| 40 | 4.085 | 3.232 | 2.839 | 2.606 | 2.449 | 2.336 | 2.249 | 2.180 | 2.077 | 2.003 | 1.793 | 1.509 |
| 60 | 4.001 | 3.150 | 2.758 | 2.525 | 2.368 | 2.254 | 2.167 | 2.097 | 1.993 | 1.917 | 1.700 | 1.389 |
| 120 | 3.920 | 3.072 | 2.680 | 2.447 | 2.290 | 2.175 | 2.087 | 2.016 | 1.910 | 1.834 | 1.608 | 1.254 |
| ∞ | 3.841 | 2.996 | 2.605 | 2.372 | 2.214 | 2.099 | 2.010 | 1.938 | 1.831 | 1.752 | 1.517 | 1.000 |

TABLE 12(d). 1 PER CENT POINTS OF THE *F*-DISTRIBUTION

If $F = \frac{X_1/\nu_1}{X_2/\nu_2}$, where X_1 and X_2 are independent random variables distributed as χ^2 with ν_1 and ν_2 degrees of freedom respectively, then the probabilities that $F \geq F(P)$ and that $F \leq F(P)$ are both equal to $P/100$. Linear interpolation in ν_1 or ν_2 will generally be sufficiently accurate except when either $\nu_1 > 12$ or $\nu_2 > 40$, when harmonic interpolation should be used.



(This shape applies only when $\nu_1 \geq 3$. When $\nu_1 < 3$ the mode is at the origin.)

| $\nu_1 =$ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 10 | 12 | 24 | ∞ |
|-------------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|----------|
| $\nu_2 = 1$ | 4052 | 4999 | 5403 | 5625 | 5764 | 5859 | 5928 | 5981 | 6056 | 6106 | 6235 | 6366 |
| 2 | 98.50 | 99.00 | 99.17 | 99.25 | 99.30 | 99.33 | 99.36 | 99.37 | 99.40 | 99.42 | 99.46 | 99.50 |
| 3 | 34.12 | 30.82 | 29.46 | 28.71 | 28.24 | 27.91 | 27.67 | 27.49 | 27.23 | 27.05 | 26.60 | 26.13 |
| 4 | 21.20 | 18.00 | 16.69 | 15.98 | 15.52 | 15.21 | 14.98 | 14.80 | 14.55 | 14.37 | 13.93 | 13.46 |
| 5 | 16.26 | 13.27 | 12.06 | 11.39 | 10.97 | 10.67 | 10.46 | 10.29 | 10.05 | 9.888 | 9.466 | 9.020 |
| 6 | 13.75 | 10.92 | 9.780 | 9.148 | 8.746 | 8.466 | 8.260 | 8.102 | 7.874 | 7.718 | 7.313 | 6.880 |
| 7 | 12.25 | 9.547 | 8.451 | 7.847 | 7.460 | 7.191 | 6.993 | 6.840 | 6.620 | 6.469 | 6.074 | 5.650 |
| 8 | 11.26 | 8.649 | 7.591 | 7.006 | 6.632 | 6.371 | 6.178 | 6.029 | 5.814 | 5.667 | 5.279 | 4.859 |
| 9 | 10.56 | 8.022 | 6.992 | 6.422 | 6.057 | 5.802 | 5.613 | 5.467 | 5.257 | 5.111 | 4.729 | 4.311 |
| 10 | 10.04 | 7.559 | 6.552 | 5.994 | 5.636 | 5.386 | 5.200 | 5.057 | 4.849 | 4.706 | 4.327 | 3.909 |
| 11 | 9.646 | 7.206 | 6.217 | 5.668 | 5.316 | 5.069 | 4.886 | 4.744 | 4.539 | 4.397 | 4.021 | 3.602 |
| 12 | 9.330 | 6.927 | 5.953 | 5.412 | 5.064 | 4.821 | 4.640 | 4.499 | 4.296 | 4.155 | 3.780 | 3.361 |
| 13 | 9.074 | 6.701 | 5.739 | 5.205 | 4.862 | 4.620 | 4.441 | 4.302 | 4.100 | 3.960 | 3.587 | 3.165 |
| 14 | 8.862 | 6.515 | 5.564 | 5.035 | 4.695 | 4.456 | 4.278 | 4.140 | 3.939 | 3.800 | 3.427 | 3.004 |
| 15 | 8.683 | 6.359 | 5.417 | 4.893 | 4.556 | 4.318 | 4.142 | 4.004 | 3.805 | 3.666 | 3.294 | 2.868 |
| 16 | 8.531 | 6.226 | 5.292 | 4.773 | 4.437 | 4.202 | 4.026 | 3.890 | 3.691 | 3.553 | 3.181 | 2.753 |
| 17 | 8.400 | 6.112 | 5.185 | 4.669 | 4.336 | 4.102 | 3.927 | 3.791 | 3.593 | 3.455 | 3.084 | 2.653 |
| 18 | 8.285 | 6.013 | 5.092 | 4.579 | 4.248 | 4.015 | 3.841 | 3.705 | 3.508 | 3.371 | 2.999 | 2.566 |
| 19 | 8.185 | 5.926 | 5.010 | 4.500 | 4.171 | 3.939 | 3.765 | 3.631 | 3.434 | 3.297 | 2.925 | 2.489 |
| 20 | 8.096 | 5.849 | 4.938 | 4.431 | 4.103 | 3.871 | 3.699 | 3.564 | 3.368 | 3.231 | 2.859 | 2.421 |
| 21 | 8.017 | 5.780 | 4.874 | 4.369 | 4.042 | 3.812 | 3.640 | 3.506 | 3.310 | 3.173 | 2.801 | 2.360 |
| 22 | 7.945 | 5.719 | 4.817 | 4.313 | 3.988 | 3.758 | 3.587 | 3.453 | 3.258 | 3.121 | 2.749 | 2.305 |
| 23 | 7.881 | 5.664 | 4.765 | 4.264 | 3.939 | 3.710 | 3.539 | 3.406 | 3.211 | 3.074 | 2.702 | 2.256 |
| 24 | 7.823 | 5.614 | 4.718 | 4.218 | 3.895 | 3.667 | 3.496 | 3.363 | 3.168 | 3.032 | 2.659 | 2.211 |
| 25 | 7.770 | 5.568 | 4.675 | 4.177 | 3.855 | 3.627 | 3.457 | 3.324 | 3.129 | 2.993 | 2.620 | 2.169 |
| 26 | 7.721 | 5.526 | 4.637 | 4.140 | 3.818 | 3.591 | 3.421 | 3.288 | 3.094 | 2.958 | 2.585 | 2.131 |
| 27 | 7.677 | 5.488 | 4.601 | 4.106 | 3.785 | 3.558 | 3.388 | 3.256 | 3.062 | 2.926 | 2.552 | 2.097 |
| 28 | 7.636 | 5.453 | 4.568 | 4.074 | 3.754 | 3.528 | 3.358 | 3.226 | 3.032 | 2.896 | 2.522 | 2.064 |
| 29 | 7.598 | 5.420 | 4.538 | 4.045 | 3.725 | 3.499 | 3.330 | 3.198 | 3.005 | 2.868 | 2.495 | 2.034 |
| 30 | 7.562 | 5.390 | 4.510 | 4.018 | 3.699 | 3.473 | 3.304 | 3.173 | 2.979 | 2.843 | 2.469 | 2.006 |
| 32 | 7.499 | 5.336 | 4.459 | 3.969 | 3.652 | 3.427 | 3.258 | 3.127 | 2.934 | 2.798 | 2.423 | 1.956 |
| 34 | 7.444 | 5.289 | 4.416 | 3.927 | 3.611 | 3.386 | 3.218 | 3.087 | 2.894 | 2.758 | 2.383 | 1.911 |
| 36 | 7.396 | 5.248 | 4.377 | 3.890 | 3.574 | 3.351 | 3.183 | 3.052 | 2.859 | 2.723 | 2.347 | 1.872 |
| 38 | 7.353 | 5.211 | 4.343 | 3.858 | 3.542 | 3.319 | 3.152 | 3.021 | 2.828 | 2.692 | 2.316 | 1.837 |
| 40 | 7.314 | 5.179 | 4.313 | 3.828 | 3.514 | 3.291 | 3.124 | 2.993 | 2.801 | 2.665 | 2.288 | 1.805 |
| 60 | 7.077 | 4.977 | 4.126 | 3.649 | 3.339 | 3.119 | 2.953 | 2.823 | 2.632 | 2.496 | 2.115 | 1.601 |
| 120 | 6.851 | 4.787 | 3.949 | 3.480 | 3.174 | 2.956 | 2.792 | 2.663 | 2.472 | 2.336 | 1.950 | 1.381 |
| ∞ | 6.635 | 4.605 | 3.782 | 3.319 | 3.017 | 2.802 | 2.639 | 2.511 | 2.321 | 2.185 | 1.791 | 1.000 |

Durbin-Watson test statistic d : 1% significance points of d_L and d_U .

| n | $k'=1$ | | $k'=2$ | | $k'=3$ | | $k'=4$ | | $k'=5$ | |
|-----|--------|-------|--------|-------|--------|-------|--------|-------|--------|-------|
| | d_L | d_U | d_L | d_U | d_L | d_U | d_L | d_U | d_L | d_U |
| 15 | 0.81 | 1.07 | 0.70 | 1.25 | 0.59 | 1.46 | 0.49 | 1.70 | 0.39 | 1.96 |
| 16 | 0.84 | 1.09 | 0.74 | 1.25 | 0.63 | 1.44 | 0.53 | 1.66 | 0.44 | 1.90 |
| 17 | 0.87 | 1.10 | 0.77 | 1.25 | 0.67 | 1.43 | 0.57 | 1.63 | 0.48 | 1.85 |
| 18 | 0.90 | 1.12 | 0.80 | 1.26 | 0.71 | 1.42 | 0.61 | 1.60 | 0.52 | 1.80 |
| 19 | 0.93 | 1.13 | 0.83 | 1.26 | 0.74 | 1.41 | 0.65 | 1.58 | 0.56 | 1.77 |
| 20 | 0.95 | 1.15 | 0.86 | 1.27 | 0.77 | 1.41 | 0.68 | 1.57 | 0.60 | 1.74 |
| 21 | 0.97 | 1.16 | 0.89 | 1.27 | 0.80 | 1.41 | 0.72 | 1.55 | 0.63 | 1.71 |
| 22 | 1.00 | 1.17 | 0.91 | 1.28 | 0.83 | 1.40 | 0.75 | 1.54 | 0.66 | 1.69 |
| 23 | 1.02 | 1.19 | 0.94 | 1.29 | 0.86 | 1.40 | 0.77 | 1.53 | 0.70 | 1.67 |
| 24 | 1.04 | 1.20 | 0.96 | 1.30 | 0.88 | 1.41 | 0.80 | 1.53 | 0.72 | 1.66 |
| 25 | 1.05 | 1.21 | 0.98 | 1.30 | 0.90 | 1.41 | 0.83 | 1.52 | 0.75 | 1.65 |
| 26 | 1.07 | 1.22 | 1.00 | 1.31 | 0.93 | 1.41 | 0.85 | 1.52 | 0.78 | 1.64 |
| 27 | 1.09 | 1.23 | 1.02 | 1.32 | 0.95 | 1.41 | 0.88 | 1.51 | 0.81 | 1.63 |
| 28 | 1.10 | 1.24 | 1.04 | 1.32 | 0.97 | 1.41 | 0.90 | 1.51 | 0.83 | 1.62 |
| 29 | 1.12 | 1.25 | 1.05 | 1.33 | 0.99 | 1.42 | 0.92 | 1.51 | 0.85 | 1.61 |
| 30 | 1.13 | 1.26 | 1.07 | 1.34 | 1.01 | 1.42 | 0.94 | 1.51 | 0.88 | 1.61 |
| 31 | 1.15 | 1.27 | 1.08 | 1.34 | 1.02 | 1.42 | 0.96 | 1.51 | 0.90 | 1.60 |
| 32 | 1.16 | 1.28 | 1.10 | 1.35 | 1.04 | 1.43 | 0.98 | 1.51 | 0.92 | 1.60 |
| 33 | 1.17 | 1.29 | 1.11 | 1.36 | 1.05 | 1.43 | 1.00 | 1.51 | 0.94 | 1.59 |
| 34 | 1.18 | 1.30 | 1.13 | 1.36 | 1.07 | 1.43 | 1.01 | 1.51 | 0.95 | 1.59 |
| 35 | 1.19 | 1.31 | 1.14 | 1.37 | 1.08 | 1.44 | 1.03 | 1.51 | 0.97 | 1.59 |
| 36 | 1.21 | 1.32 | 1.15 | 1.38 | 1.10 | 1.44 | 1.04 | 1.51 | 0.99 | 1.59 |
| 37 | 1.22 | 1.32 | 1.16 | 1.38 | 1.11 | 1.45 | 1.06 | 1.51 | 1.00 | 1.59 |
| 38 | 1.23 | 1.33 | 1.18 | 1.39 | 1.12 | 1.45 | 1.07 | 1.52 | 1.02 | 1.58 |
| 39 | 1.24 | 1.34 | 1.19 | 1.39 | 1.14 | 1.45 | 1.09 | 1.52 | 1.03 | 1.58 |
| 40 | 1.25 | 1.34 | 1.20 | 1.40 | 1.15 | 1.46 | 1.10 | 1.52 | 1.05 | 1.58 |
| 45 | 1.29 | 1.38 | 1.24 | 1.42 | 1.20 | 1.48 | 1.16 | 1.53 | 1.11 | 1.58 |
| 50 | 1.32 | 1.40 | 1.28 | 1.45 | 1.24 | 1.49 | 1.20 | 1.54 | 1.16 | 1.59 |
| 55 | 1.36 | 1.43 | 1.32 | 1.47 | 1.28 | 1.51 | 1.25 | 1.55 | 1.21 | 1.59 |
| 60 | 1.38 | 1.45 | 1.35 | 1.48 | 1.32 | 1.52 | 1.28 | 1.56 | 1.25 | 1.60 |
| 65 | 1.41 | 1.47 | 1.38 | 1.50 | 1.35 | 1.53 | 1.31 | 1.57 | 1.28 | 1.61 |
| 70 | 1.43 | 1.49 | 1.40 | 1.52 | 1.37 | 1.55 | 1.34 | 1.58 | 1.31 | 1.61 |
| 75 | 1.45 | 1.50 | 1.42 | 1.53 | 1.39 | 1.56 | 1.37 | 1.59 | 1.34 | 1.62 |
| 80 | 1.47 | 1.52 | 1.44 | 1.54 | 1.42 | 1.57 | 1.39 | 1.60 | 1.36 | 1.62 |
| 85 | 1.48 | 1.53 | 1.46 | 1.55 | 1.43 | 1.58 | 1.41 | 1.60 | 1.39 | 1.63 |
| 90 | 1.50 | 1.54 | 1.47 | 1.56 | 1.45 | 1.59 | 1.43 | 1.61 | 1.41 | 1.64 |
| 95 | 1.51 | 1.55 | 1.49 | 1.57 | 1.47 | 1.60 | 1.45 | 1.62 | 1.42 | 1.64 |
| 100 | 1.52 | 1.56 | 1.50 | 1.58 | 1.48 | 1.60 | 1.46 | 1.63 | 1.44 | 1.65 |

n = number of observations

k' = number of explanatory variables

Durbin-Watson test statistic d : 5% significance points of d_L and d_U .

| n | $k'=1$ | | $k'=2$ | | $k'=3$ | | $k'=4$ | | $k'=5$ | |
|-----|--------|-------|--------|-------|--------|-------|--------|-------|--------|-------|
| | d_L | d_U | d_L | d_U | d_L | d_U | d_L | d_U | d_L | d_U |
| 15 | 1.08 | 1.36 | 0.95 | 1.54 | 0.82 | 1.75 | 0.69 | 1.97 | 0.56 | 2.21 |
| 16 | 1.10 | 1.37 | 0.98 | 1.54 | 0.86 | 1.73 | 0.74 | 1.93 | 0.62 | 2.15 |
| 17 | 1.13 | 1.38 | 1.02 | 1.54 | 0.90 | 1.71 | 0.78 | 1.90 | 0.67 | 2.10 |
| 18 | 1.16 | 1.39 | 1.05 | 1.53 | 0.93 | 1.69 | 0.82 | 1.87 | 0.71 | 2.06 |
| 19 | 1.18 | 1.40 | 1.08 | 1.53 | 0.97 | 1.68 | 0.86 | 1.85 | 0.75 | 2.02 |
| 20 | 1.20 | 1.41 | 1.10 | 1.54 | 1.00 | 1.68 | 0.90 | 1.83 | 0.79 | 1.99 |
| 21 | 1.22 | 1.42 | 1.13 | 1.54 | 1.03 | 1.67 | 0.93 | 1.81 | 0.83 | 1.96 |
| 22 | 1.24 | 1.43 | 1.15 | 1.54 | 1.05 | 1.66 | 0.96 | 1.80 | 0.86 | 1.94 |
| 23 | 1.26 | 1.44 | 1.17 | 1.54 | 1.08 | 1.66 | 0.99 | 1.79 | 0.90 | 1.92 |
| 24 | 1.27 | 1.45 | 1.19 | 1.55 | 1.10 | 1.66 | 1.01 | 1.78 | 0.93 | 1.90 |
| 25 | 1.29 | 1.45 | 1.21 | 1.55 | 1.12 | 1.66 | 1.04 | 1.77 | 0.95 | 1.89 |
| 26 | 1.30 | 1.46 | 1.22 | 1.55 | 1.14 | 1.65 | 1.06 | 1.76 | 0.98 | 1.88 |
| 27 | 1.32 | 1.47 | 1.24 | 1.56 | 1.16 | 1.65 | 1.08 | 1.76 | 1.01 | 1.86 |
| 28 | 1.33 | 1.48 | 1.26 | 1.56 | 1.18 | 1.65 | 1.10 | 1.75 | 1.03 | 1.85 |
| 29 | 1.34 | 1.48 | 1.27 | 1.56 | 1.20 | 1.65 | 1.12 | 1.74 | 1.05 | 1.84 |
| 30 | 1.35 | 1.49 | 1.28 | 1.57 | 1.21 | 1.65 | 1.14 | 1.74 | 1.07 | 1.83 |
| 31 | 1.36 | 1.50 | 1.30 | 1.57 | 1.23 | 1.65 | 1.16 | 1.74 | 1.09 | 1.83 |
| 32 | 1.37 | 1.50 | 1.31 | 1.57 | 1.24 | 1.65 | 1.18 | 1.73 | 1.11 | 1.82 |
| 33 | 1.38 | 1.51 | 1.32 | 1.58 | 1.26 | 1.65 | 1.19 | 1.73 | 1.13 | 1.81 |
| 34 | 1.39 | 1.51 | 1.33 | 1.58 | 1.27 | 1.65 | 1.21 | 1.73 | 1.15 | 1.81 |
| 35 | 1.40 | 1.52 | 1.34 | 1.58 | 1.28 | 1.65 | 1.22 | 1.73 | 1.16 | 1.80 |
| 36 | 1.41 | 1.52 | 1.35 | 1.59 | 1.29 | 1.65 | 1.24 | 1.73 | 1.18 | 1.80 |
| 37 | 1.42 | 1.53 | 1.36 | 1.59 | 1.31 | 1.66 | 1.25 | 1.72 | 1.19 | 1.80 |
| 38 | 1.43 | 1.54 | 1.37 | 1.59 | 1.32 | 1.66 | 1.26 | 1.72 | 1.21 | 1.79 |
| 39 | 1.43 | 1.54 | 1.38 | 1.60 | 1.33 | 1.66 | 1.27 | 1.72 | 1.22 | 1.79 |
| 40 | 1.44 | 1.54 | 1.39 | 1.60 | 1.34 | 1.66 | 1.29 | 1.72 | 1.23 | 1.79 |
| 45 | 1.48 | 1.57 | 1.43 | 1.62 | 1.38 | 1.67 | 1.34 | 1.72 | 1.29 | 1.78 |
| 50 | 1.50 | 1.59 | 1.46 | 1.63 | 1.42 | 1.67 | 1.38 | 1.72 | 1.34 | 1.77 |
| 55 | 1.53 | 1.60 | 1.49 | 1.64 | 1.45 | 1.68 | 1.41 | 1.72 | 1.38 | 1.77 |
| 60 | 1.55 | 1.62 | 1.51 | 1.65 | 1.48 | 1.69 | 1.44 | 1.73 | 1.41 | 1.77 |
| 65 | 1.57 | 1.63 | 1.54 | 1.66 | 1.50 | 1.70 | 1.47 | 1.73 | 1.44 | 1.77 |
| 70 | 1.58 | 1.64 | 1.55 | 1.67 | 1.52 | 1.70 | 1.49 | 1.74 | 1.46 | 1.77 |
| 75 | 1.60 | 1.65 | 1.57 | 1.68 | 1.54 | 1.71 | 1.51 | 1.74 | 1.49 | 1.77 |
| 80 | 1.61 | 1.66 | 1.59 | 1.69 | 1.56 | 1.72 | 1.53 | 1.74 | 1.51 | 1.77 |
| 85 | 1.62 | 1.67 | 1.60 | 1.70 | 1.57 | 1.72 | 1.55 | 1.75 | 1.52 | 1.77 |
| 90 | 1.63 | 1.68 | 1.61 | 1.70 | 1.59 | 1.73 | 1.57 | 1.75 | 1.54 | 1.78 |
| 95 | 1.64 | 1.69 | 1.62 | 1.71 | 1.60 | 1.73 | 1.58 | 1.75 | 1.56 | 1.78 |
| 100 | 1.65 | 1.69 | 1.63 | 1.72 | 1.61 | 1.74 | 1.59 | 1.76 | 1.57 | 1.78 |

n = number of observations

k' = number of explanatory variables

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