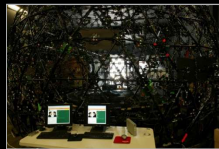


Radiometry and HDR Basics



C0417 – Advanced Computer Graphics: Photographic Image Synthesis

Abhijeet Ghosh

Lecture 02, Jan. 14th 2019

1

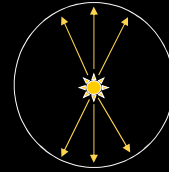
Radiometry & Geometric Optics

- Light transport modeled using geometric or ray optics
 - light as particle, not wave!
 - some exceptions, i.e., polarization
- Basic properties of geometric optics:
 - Linearity
 - Energy conservation

2

Basic Quantities

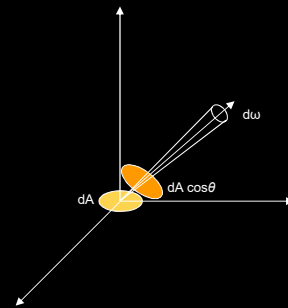
- **Radiant Flux** or **Power** Φ : Energy flowing through a surface per unit time. Units - Joules/second (J/s) or Watts (W).
 - Emission from light sources typically described with flux
- **Irradiance** E : area density of incoming flux (W/m^2)
 - for a sphere of radius r , $E = \Phi / 4\pi r^2$
 - energy received from an isotropic source falls off with squared distance!



3

Basic Quantities

- **Intensity** I : flux density per solid angle [W/sr]
 - $I = d\Phi/d\omega$
 - useful for describing point light sources, with zero area!
- **Radiance** L : radiant flux density per unit area, per unit solid angle [$\text{W}/\text{m}^2\text{sr}$]
 - $L = d^2\Phi/(dA \cos\theta d\omega)$
 - radiance remains constant along a direction!



4

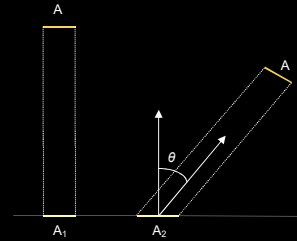
Lambert's Law

- Irradiance E proportional to **cosine** of the angle between light direction \mathbf{l} and surface normal \mathbf{n}

$$E = d\Phi/dA,$$

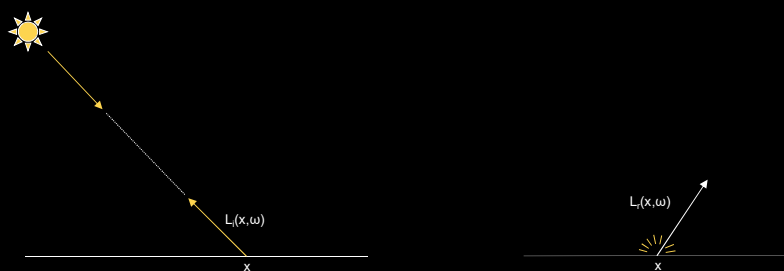
$$\text{hence } E_1 = \Phi/A,$$

$$\text{and } E_2 = \Phi \cos\theta/A.$$



5

Incident and Exitant Radiance

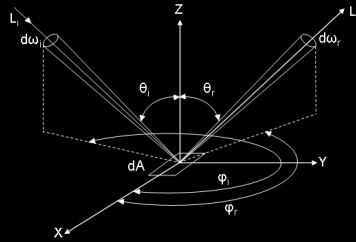


- Incident radiance $L_i(x, \omega)$, due to light arriving from a source
- Exitant radiance $L_r(x, \omega)$, due to reflection from a surface

$$\text{In general } L_i(x, \omega) \neq L_r(x, \omega)$$

6

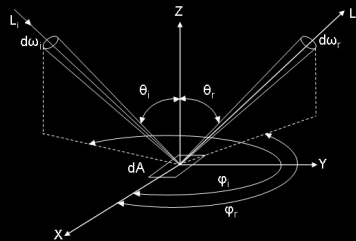
BRDF



- Bidirectional Reflectance Distribution Function [Nicodemus et al. 77]
 - formalizes the reflection of light at a surface!

7

BRDF



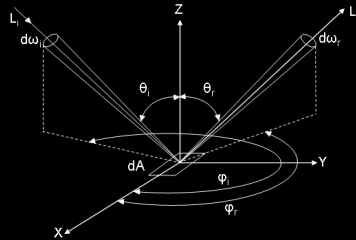
- Defined as the ratio of reflected radiance to incident irradiance:

$$\begin{aligned} f_r(x, \omega_r, \omega_i) &= dL_r(x, \omega_r)/dE_i(x, \omega_i) \\ &= dL_r(x, \omega_r)/(L_i(x, \omega_i) \cos\theta \, d\omega_i). \end{aligned}$$

- the units of a BRDF are inverse steradian [1/sr].

8

BRDF



- Physically based BRDFs have 2 important properties:

Helmholtz Reciprocity: $f_r(x, \omega_r, \omega_i) = f_r(x, \omega_i, \omega_r)$.

and

Energy Conservation: $\int_{\Omega} f_r(x, \omega_r, \omega_i) \cos \theta_i d\omega_i \leq 1$, for all ω_r in Ω .

9

Radiance imaging with cameras



Camera settings:

Shutter speed – 1 sec

Aperture - f/8

gain – ISO 100

ND filters

10

Radiance in the Real World – Dynamic Range



Sony VX2000 video camera

Office interior
Indirect light from window
1/60th sec shutter
f/5.6 aperture
0 ND filters
0dB gain

11

Dynamic Range in the Real World



16 times the light as inside

Outside in the shade
1/1000th sec shutter
f/5.6 aperture
0 ND filters
0dB gain

12

Dynamic Range in the Real World



Outside in the sun
1/1000th sec shutter
f/11 aperture
0 ND filters
0dB gain

64 times the light as inside

13

Dynamic Range in the Real World



Straight at the sun
1/10,000th sec shutter
f/11 aperture
13 stops ND filters
0dB gain

5,000,000 times the light as inside

14

Dynamic Range in the Real World

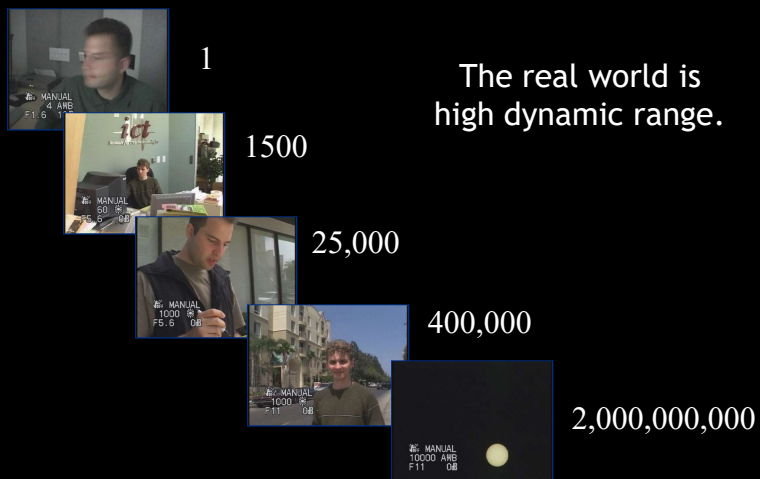


Very dim room
1/4th sec shutter
f/1.6 aperture
0 stops ND filters
18dB gain

1/1500th the light than inside

15

Dynamic Range in the Real World



16

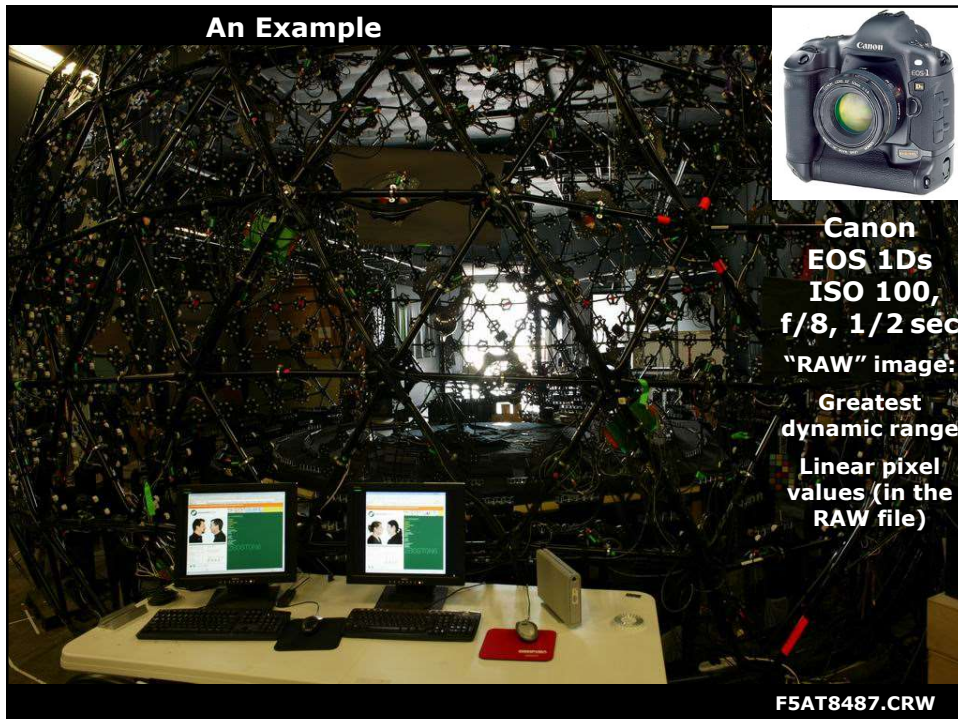
Taking HDR Images



www.debevec.org

17

An Example



18

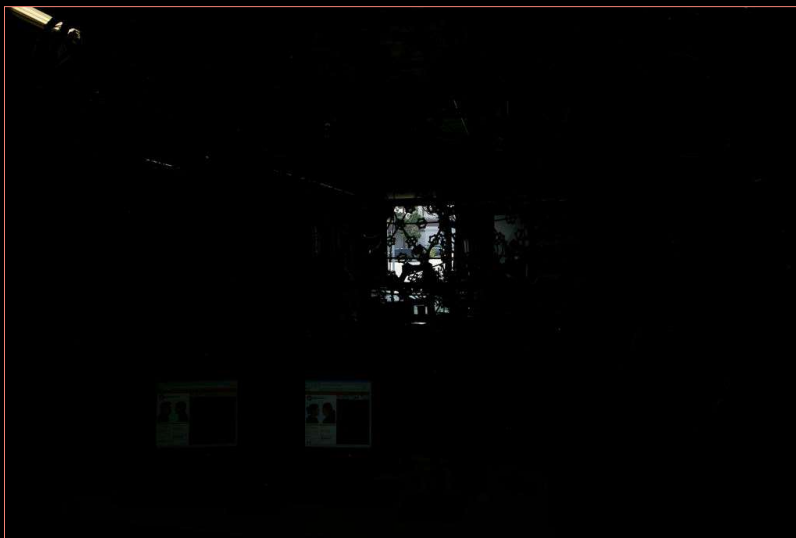
High Dynamic Range Imaging



ISO 100, f/8, 1/1000 sec

19

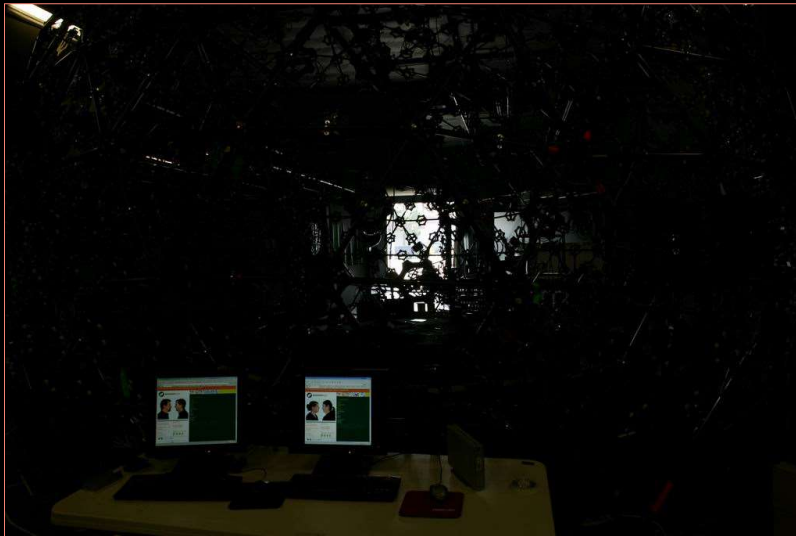
High Dynamic Range Imaging



ISO 100, f/8, 1/125 sec

20

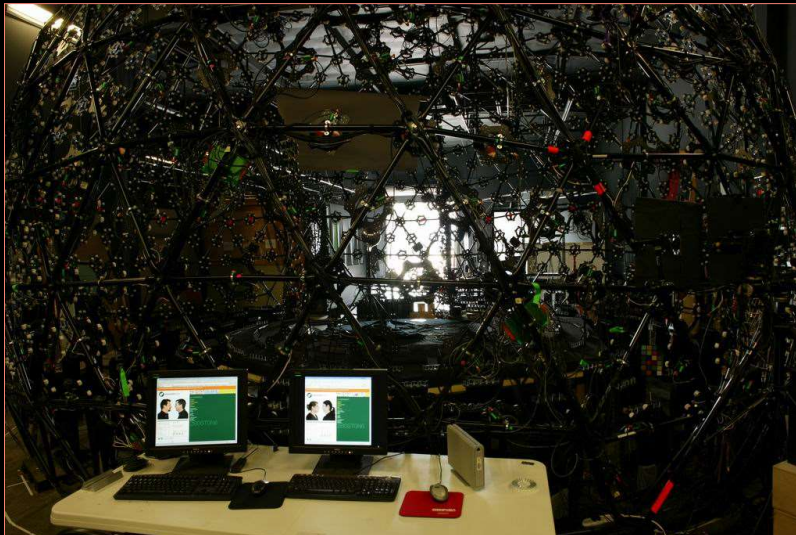
High Dynamic Range Imaging



ISO 100, f/8, 1/15 sec

21

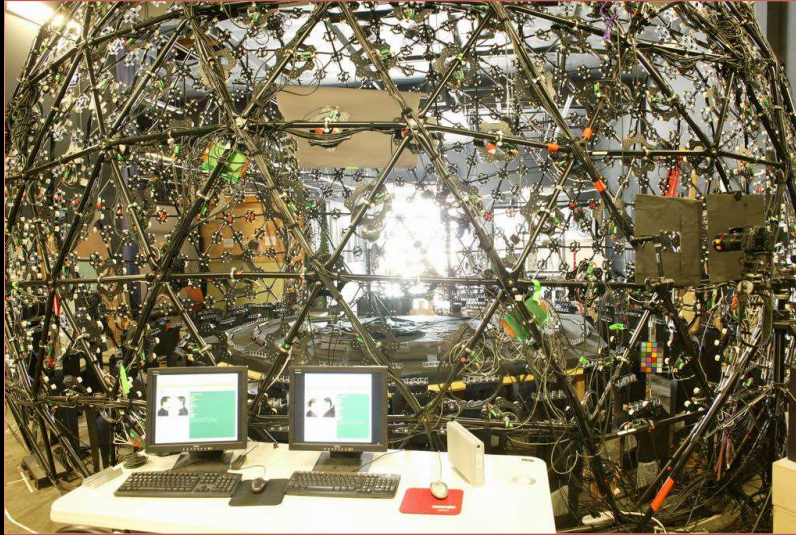
High Dynamic Range Imaging



ISO 100, f/8, 1/2 sec

22

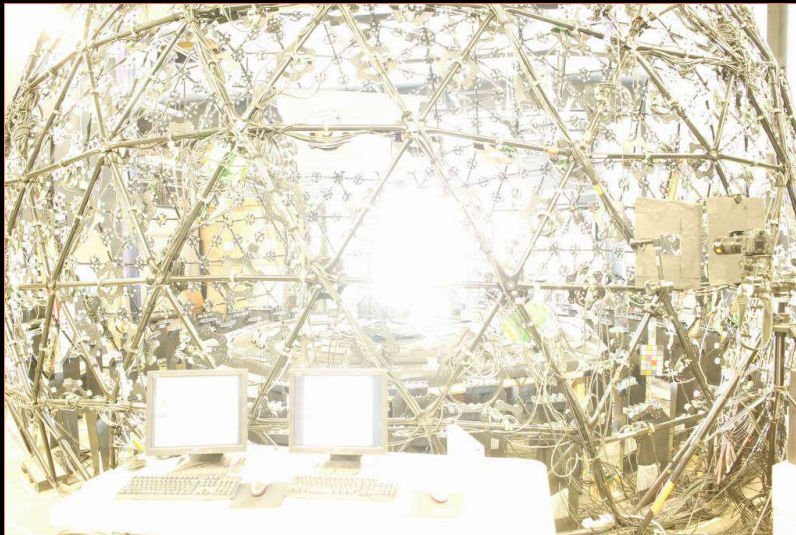
High Dynamic Range Imaging



ISO 100, f/8, 4 sec

23

High Dynamic Range Imaging



ISO 100, f/8, 30 sec

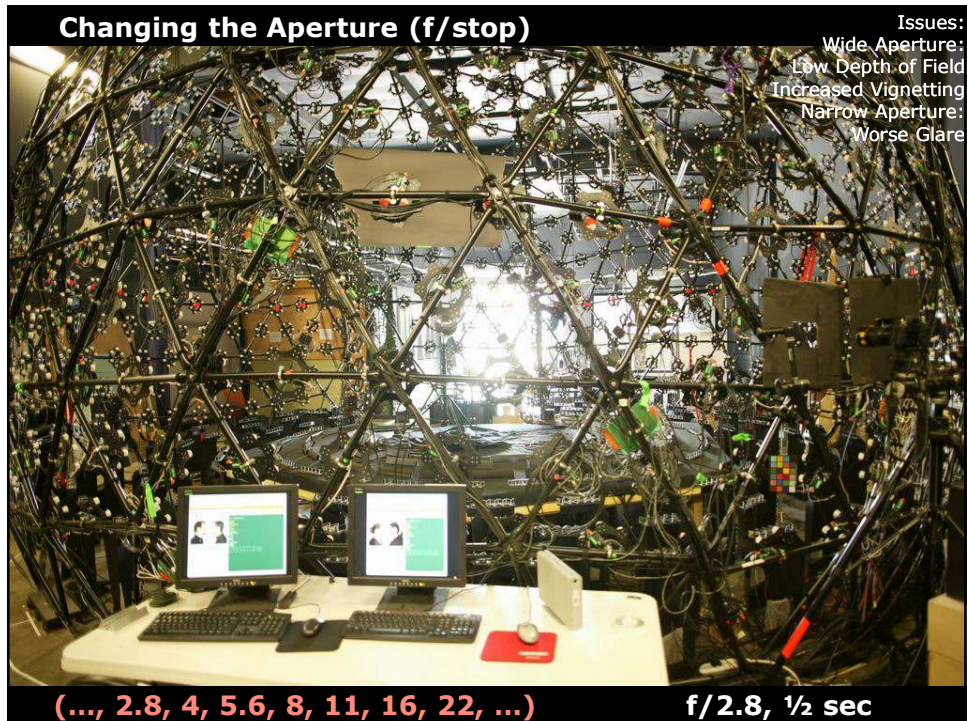
24

High Dynamic Range Imaging

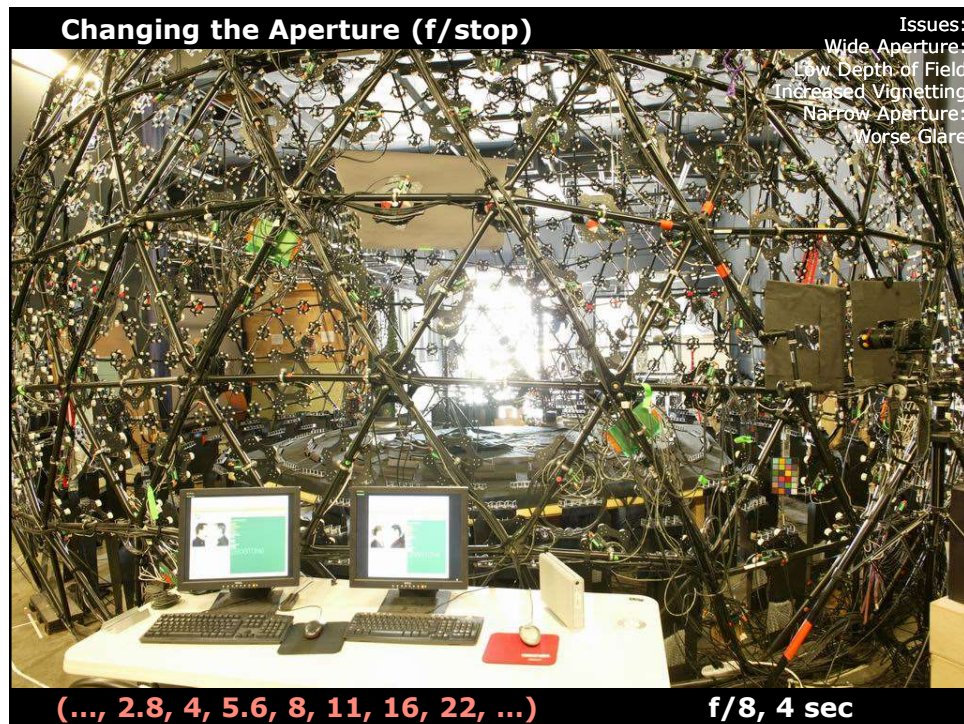


25

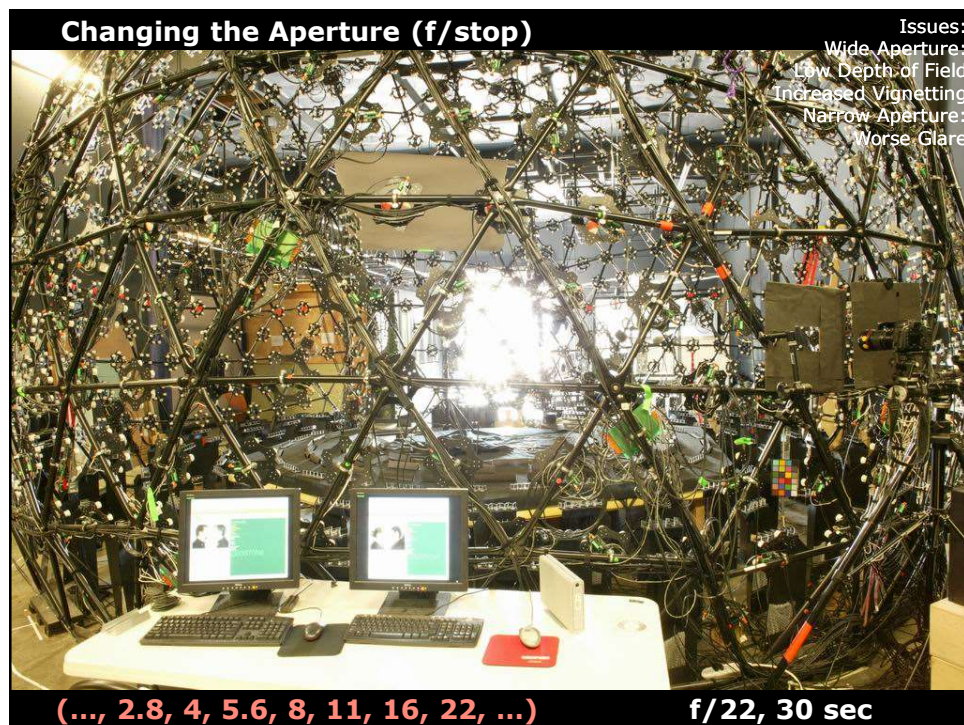
Changing the Aperture (f/stop)



26



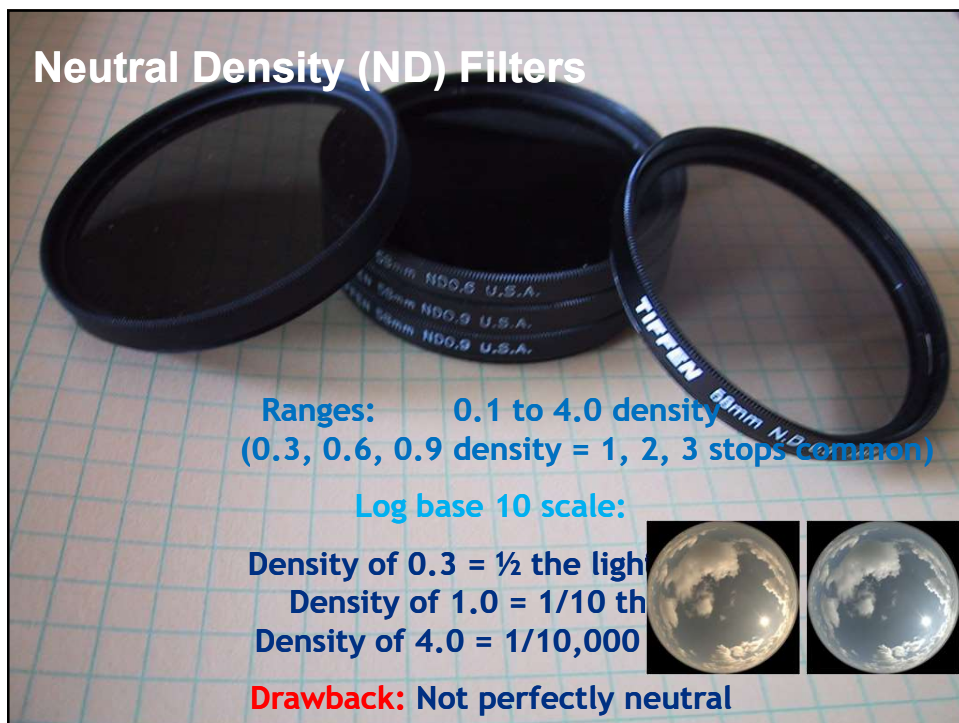
27



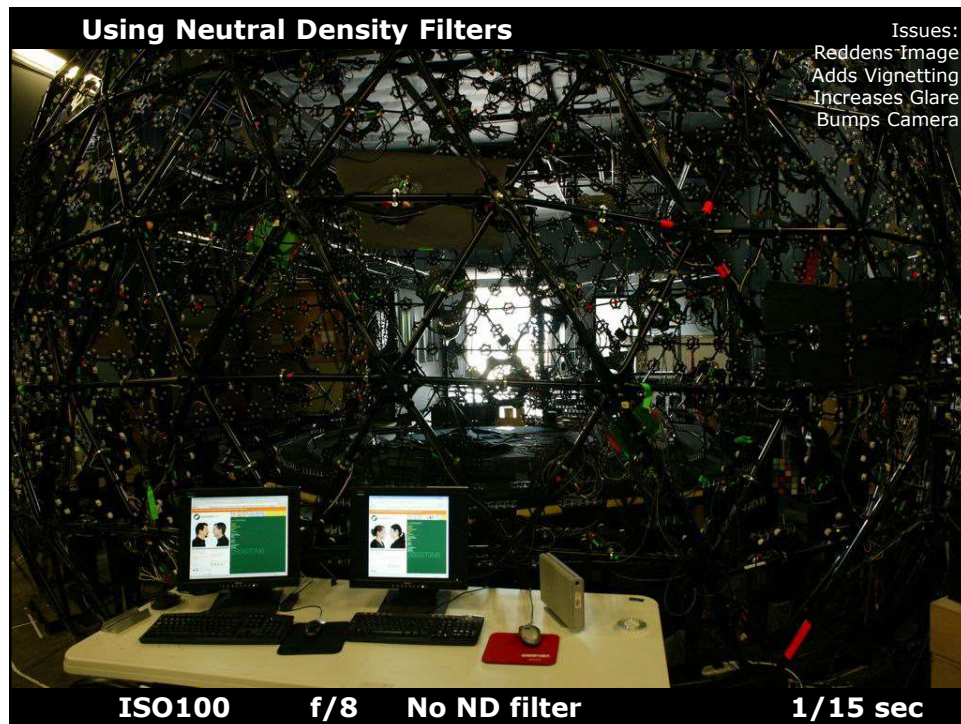
28



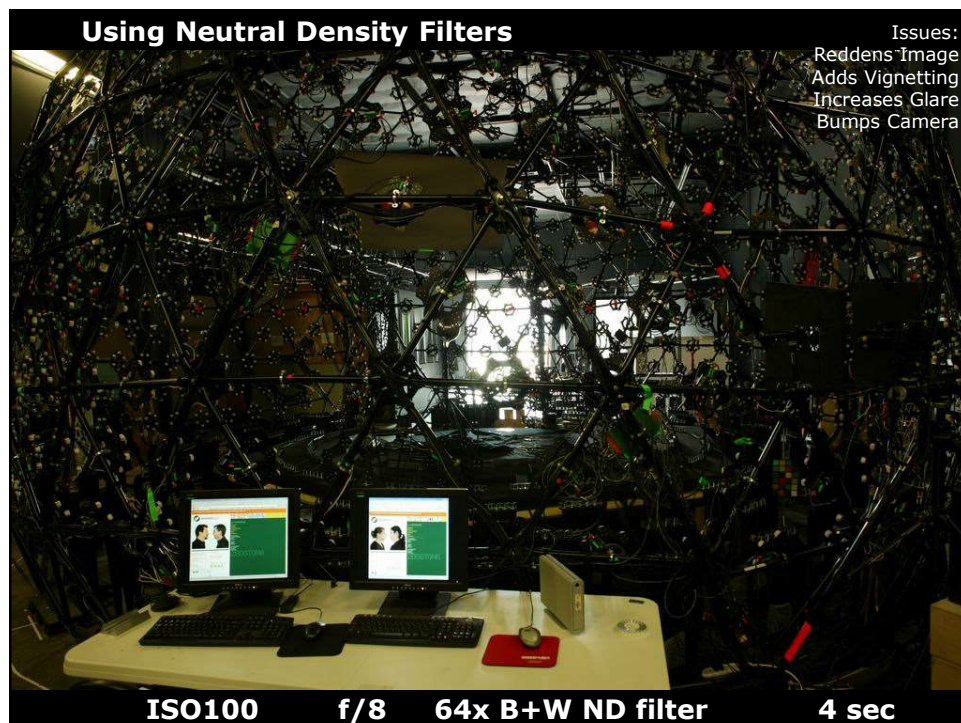
29



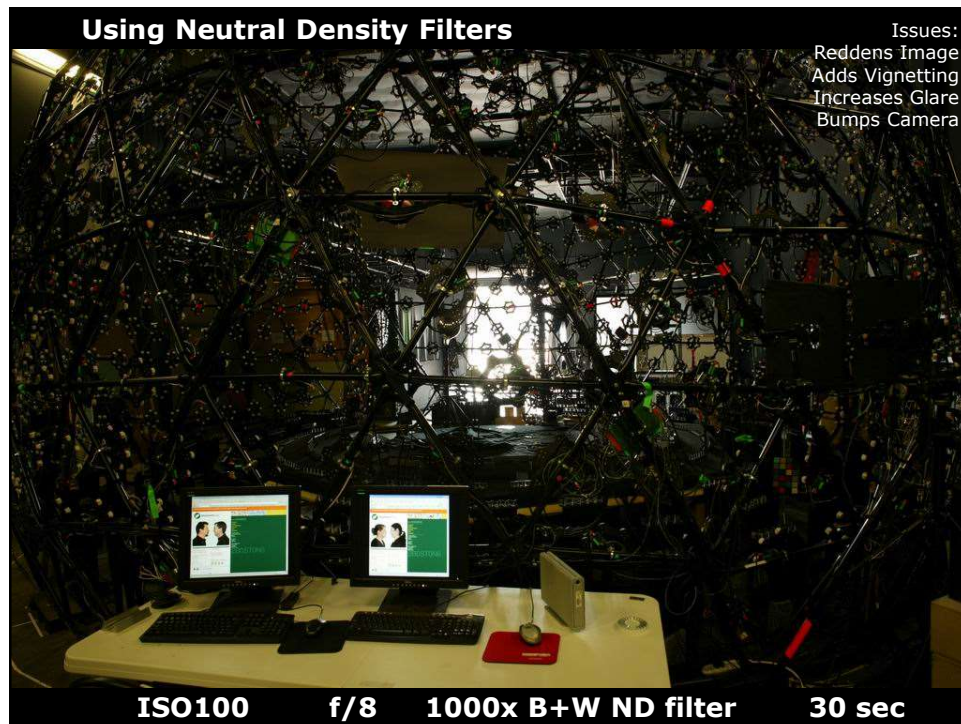
30



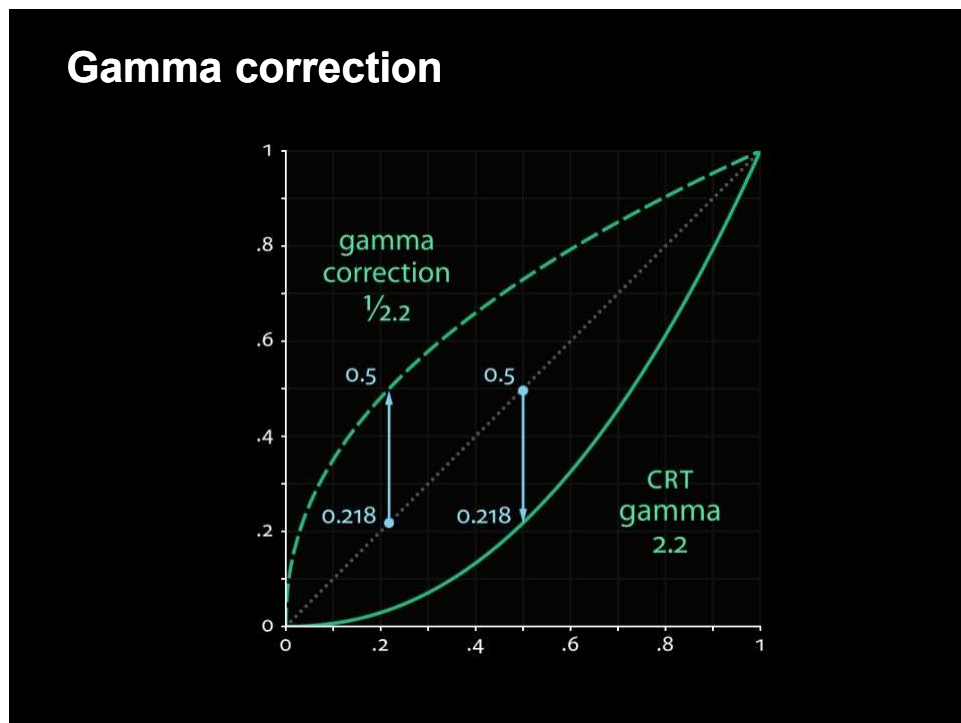
31



32

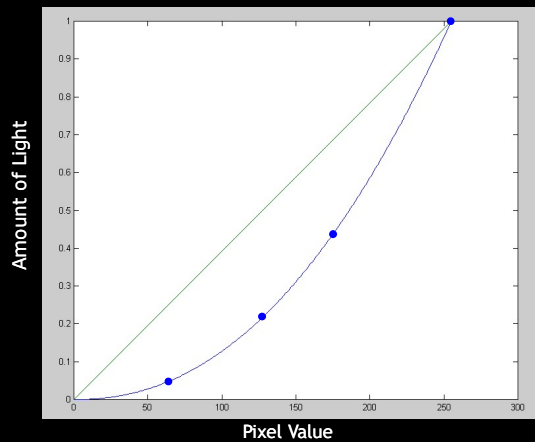


33



34

Gamma 2.2 graph



Implications:

128 is less than $\frac{1}{4}$ as bright as 255

128 is more than 4 times as bright as 64

175 is twice as bright as 128

93 is half as bright as 128

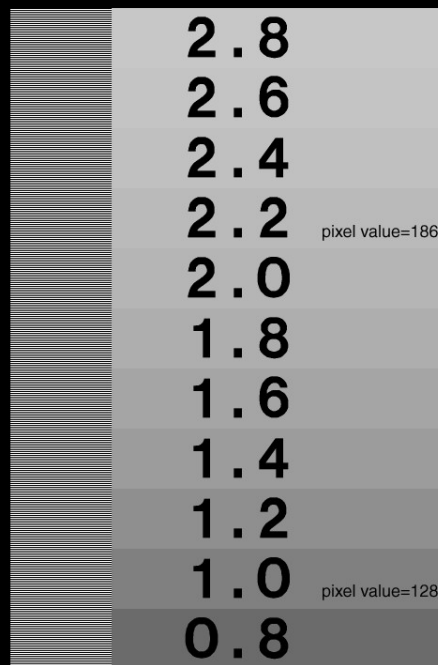
$$128 + 128 = 175$$

$$128 / 2 = 93$$

See also: <http://www.poynton.com/GammaFAQ.html>

35

- Gamma Calibration Chart
- After Greg Ward
- www.debevec.org/gamma.png



36