

# Summary of Current Work - Vaidya Black Hole in $f(R)$ Gravity

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This is a summary of work done by J. P. Hou, NUAA. One main topic is included:

- 1) We investigate the thermodynamics properties of Vaidya black holes in  $f(R)$  gravity

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The action of f(R) gravity is

$$S = \frac{1}{16\pi G} \int d^4s \sqrt{-g} f(R) + S_{matter} \quad (.1)$$

Varying the action, one gets the equation of gravitational field

$$f_R R_{\mu\nu} - \frac{1}{2} f g_{\mu\nu} + g_{\mu\nu} \square f_R = 8\pi G T_{\mu\nu} \quad (.2)$$

## I. STATIC

For simplicity, we firstly consider the static situation. A general form of static spherical symmetric spacetime can be written as

$$ds^2 = -F(r)dt^2 + \frac{1}{F(r)}dr^2 + b^2(r)(d\theta^2 + \sin^2\theta d\phi^2) \quad (I.1)$$

Assuming the metric (I.1) is non-extreme black holes, that is,  $F(r)|_{r=r_+} = 0$  while  $F'(r)|_{r=r_+} \neq 0$ , where  $r = r_+$  is the horizon. From the definition

$$\kappa = \frac{1}{2\sqrt{-h}} \partial_a (\sqrt{-h} h^{ab} \partial_b r) \quad (I.2)$$

one could obtain the temperature of the black hole

$$T = \frac{\kappa}{2\pi}|_{r=r_+} = \frac{1}{4\pi} F'(r_+) b'(r_+) \quad (I.3)$$

The components of (.2) with metric (I.1) are

$$T_t^t = \frac{1}{2b} (-bf - 2f_R b' F' + 4F b' f'_R - b f_R F'' - 2b F f''_R) \quad (I.4)$$

$$T_r^r = \frac{1}{2b} (-bf - 2f_R b' F' + 4F b' f'_R - b f_R F'' - 4F f_R b'') \quad (I.5)$$

Here, ' is derivative with respect to  $r$ . Note that  $F(r_+) = 0$ , thus

$$T_t^t|_{r=r_+} = T_r^r|_{r=r_+} = \frac{1}{2b} (-bf - 2f_R b' F' + 4F b' f'_R - b f_R F'')|_{r=r_+} \quad (I.6)$$

It is obviously that  $T_r^r|_{r=r_+}$  can be treated as pressure, as a result

$$4\pi G b^2 P dr_+ = \frac{b}{4G} (-bf - 2f_R b' F' + 4F b' f'_R - b f_R F'') dr_+ \quad (I.7)$$

In f(R) gravity, the entropy is

$$S = \frac{\pi b^2}{G} f_R \quad (I.8)$$

Inserting  $TdS$  into (I.7)

$$PdV = TdS - \frac{1}{4G}(b'b^2(f - R) + 2b')dr_+ \quad (\text{I.9})$$

To investigate this equation further, we calculate the Misner-Shape energy. One can define a Kodama vector in a spherical symmetric spacetime and it is apparent that it satisfies the following constraints

$$\nabla_\mu \nabla_\nu f_R \nabla^\mu K^\nu = 0 \quad (\text{I.10})$$

Then, one may calculate the Misner-Shape energy with

$$dE_{eff} = dE_{eff} = A\Phi_a dx^a + WdV = A(t, r)dt + B(t, r)dr \quad (\text{I.11})$$

where  $A = V_{n-2}^k b^{n-2}$  and  $V = V_{n-2}^k b^{n-1}/(n-1)$  are area and volume of  $(n-2)$  dimension sphere with radius  $r$ , and

$$A(t, r) = 0, \quad B(t, r) = A(FT_{rr} + 2W) \quad (\text{I.12})$$

Inserting  $T_{rr}$  into the equation above, one has

$$dE_{eff} = \frac{b^2}{2G}(\frac{1}{2}f + f_R \frac{b'}{b}F' - 2F \frac{b'}{b}f'_R - \frac{1}{2}F'f'_R + \frac{1}{2b}f_R F'' - Ff''_R)b'dr_+ \quad (\text{I.13})$$

After integrated, the Misner-Shape energy is

$$E_{eff} = \frac{b}{2G}((1 - h^{ab}\partial_a b \partial_b b)f_R + \frac{b^2}{6}(f - f_R R) - bh^{ab}\partial_a f_R \partial_b b) \quad (\text{I.14})$$

$$- \frac{1}{2G} \int dr (-\frac{1}{2}b^2 b'^2 F' + b - b'^2 Fb - \frac{1}{6}b^3 R - b^2 b'' F)f'_R \quad (\text{I.15})$$

Then, one may get  $dE_{eff}|_{r=r_+}$  readily

$$dE_{eff} = \frac{b^2}{2G}(\frac{1}{2}f + f_R \frac{b'}{b}F' - 2F \frac{b'}{b}f'_R - \frac{1}{2}F'f'_R + \frac{1}{2b}f_R F'' - Ff''_R)b'dr_+ \quad (\text{I.16})$$

Finally, we have the thermodynamics relation at the horizon

$$dE = TdS - PdV + \frac{1}{4G}((b^2 R - 2)(1 - f_R) - b^2 F' f'_R)b'dr_+ \quad (\text{I.17})$$

## II. DYNAMIC

The metric we chosen is

$$ds^2 = -F(t, r)dt^2 + 2dtdr + r^2(d\theta^2 + \sin^2\theta d\phi^2) \quad (\text{II.1})$$

Some the non-zero component of the motion equation are

$$8\pi GT_{tt} = \frac{1}{2r}(frF + 2f_R FF_{,r} - 4F^2 f_{R,r} - rFF_{,r}f_{R,r} - 2rF^2 f_{R,rr} - 2f_R F_{,t} - rf_{R,r}F_{,r} - 4Ff_{R,t} + rF_{,r}f_{R,t} - 4rFf_{R,tr} - 2rf_{R,tt}) \quad (\text{II.2})$$

$$8\pi GT_{tr} = \frac{1}{2G}(-fr - 2f_R F_{,r} + 4Ff_{R,r} + rF_{,r}f_{R,r} - f_R rF_{,rr} + 2rFf_{R,rr} + 4f_{R,t} + 2rf_{R,tr}) \quad (\text{II.3})$$

$$8\pi GT_{rr} = -f_{R,rr} \quad (\text{II.4})$$

$$8\pi GT_{\theta\theta} = \frac{1}{2}(2f_R - fr^2 - f_R F - 2f_R rF_{,r} + 2rFf_{R,r} + 2rf_{R,t} + 2r^2 F_{,r}f_{R,r} + 2r^2 Ff_{R,rr} + 4r^2 f_{R,tr}) \quad (\text{II.5})$$

$$8\pi GT_{\phi\phi} = \frac{1}{2}(2f_R - fr^2 - 2f_R F - 2f_R rF_{,r} + 2rFf_{R,r} + 2rf_{R,t} + 2r^2 F_{,r}f_{R,r} + 2r^2 Ff_{R,rr} + 4r^2 f_{R,tr})\sin^2\theta \quad (\text{II.6})$$

We follow the aforementioned steps to investigate its thermodynamics. However, constraint of Kodama vector dose not satisfy automatically. Now, the Kodama vector is simply

$$k^\mu = \left(\frac{\partial}{\partial t}\right)^\mu \quad (\text{II.7})$$

and the constraint turns out to be

$$-2f_R F_{,t} - 4rf_{R,r}F_{,t} - r^2 f_{R,rr}F_{,t} - 4rf_R F_{,tr} - 2r^2 f_{R,r}F_{,tr} - r^2 f_R F_{,trr} = 0 \quad (\text{II.8})$$

Solving the differential equation, one obtains

$$f_R F_{,t} = \frac{C_1}{r} + \frac{C_2}{r^2} \quad (\text{II.9})$$

Assuming this holds, we can still apply the unified first law

$$\begin{aligned} \delta Q &= A\Phi_a dx^2 \\ &= A(T_{tt} + FT_{tr})dt \\ &= \frac{r_A}{4G}(r_A f_{R,t} F_{,r} - 2f_R F_{,t} - r_A f_{R,r} F_{,t} - 2r_A f_{R,tt})dt \end{aligned} \quad (\text{II.10})$$

Notice that,  $F(t, r_A(t)) = 0$ , and thus,  $r_{A,t} = -F_{,t}/F_{,r}$ , is used. While now,  $T$  takes the form  $F_{,r}/(4|p_i)$  and  $S$  are the same as the static situation since it has nothing to do with metric. Thus,

$$TdS = \frac{r_A}{4G}(r_A f_{R,t} F_{,r} - 2f_R F_{,t} - r_A f_{R,r} F_{,t})dt \quad (\text{II.11})$$