

Phy 132 - HW XI-XII solutions

10)

$B = \mu_0 n I$ $n \Rightarrow \text{turns/m}$

a) Flux per turn is then $B \cdot A$

Total # of turns is $n \cdot l$ & total flux is scaled by this, so

$$\Phi = B \cdot A \cdot n l = \mu_0 n^2 I A l$$

~~$V = -L \frac{dI}{dt}$~~ and $V = -\frac{d\Phi}{dt}$, so, since $\Phi \propto I$,

$LI = \Phi$, or $L = \frac{\Phi}{I} = \mu_0 n^2 A l$ \uparrow turns per length!
not total turns

b) $n = 10^4 \text{ m}^{-1}$

$d = .05 \text{ m} \Rightarrow r = .025 \text{ m} \Rightarrow A = \pi r^2 = .00196 \text{ m}^2$

$l = .5 \text{ m}$

so $L = \mu_0 n^2 A l = 4\pi \times 10^{-7} \cdot (10^4)^2 \cdot (.00196) \cdot (.5) = .123 \text{ H}$

If $\frac{dI}{dt} \approx \frac{\Delta I}{\Delta t} = \frac{10 \text{ A}}{.001 \text{ sec}} = 10^4 \text{ A/sec}$

$V = -L \frac{dI}{dt} = -1230 \text{ V!}$

20) $L = 5 \times 10^{-3} \text{ H}$ $V = 5 \text{ V} \Rightarrow \frac{dI}{dt} = \frac{V}{L} = \frac{5 \text{ V}}{5 \times 10^{-3} \text{ H}} = 10^3 \text{ A/sec}$

so $I = \frac{dI}{dt} \cdot t = 10^3 \text{ A/s} \cdot t = 1 \text{ A}$

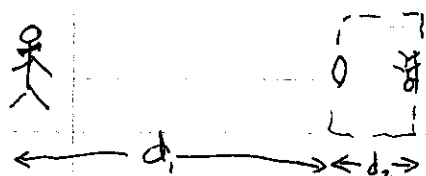
so $t = 10^{-3} \text{ sec}$

3) $\omega = 3 \times 10^{12} \text{ s}^{-1}$ We know $v_{\text{wave}} = \frac{\omega}{k} = c = 3 \times 10^8 \text{ m/s}$

So $k = \frac{\omega}{c}$ and $k = \frac{2\pi}{\lambda}$, so

$$\lambda = \frac{2\pi}{k} = \frac{2\pi c}{\omega} = \frac{2\pi \cdot 3 \times 10^8 \text{ m/s}}{3 \times 10^{12} \text{ s}^{-1}} = \boxed{6.28 \times 10^{-4} \text{ m}}$$

4) $f = 50 \text{ mm}$ $d_2 = 51 \text{ mm}$



$$\frac{1}{f} = \frac{1}{d_1} + \frac{1}{d_2} \Rightarrow \frac{1}{f} - \frac{1}{d_2} = \frac{1}{d_1}$$

$$d_1 = \left(\frac{1}{f} - \frac{1}{d_2} \right)^{-1} = \left(\frac{1}{50} - \frac{1}{51} \right)^{-1} \text{ mm} = 2550 \text{ mm}$$

$$\boxed{d_1 = 2.55 \text{ m}}$$

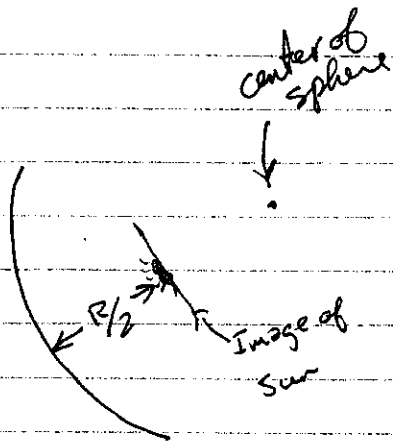
5)

$d_1 = 4 \text{ m}$. We want $M = \frac{2 \text{ cm}}{2 \text{ m}} = 10^{-2} = \frac{d_2}{d_1}$, so

$$d_2 = 10^{-2} d_1 = 4 \text{ cm}$$

What is f ? $\frac{1}{f} = \frac{1}{d_1} + \frac{1}{d_2} \Rightarrow f = \left(\frac{1}{d_1} + \frac{1}{d_2} \right)^{-1} = \left(\frac{1}{.04} + \frac{1}{4} \right)^{-1} \text{ m}$
 $= .0396 \text{ m} = \boxed{39.6 \text{ mm} = f}$

640)



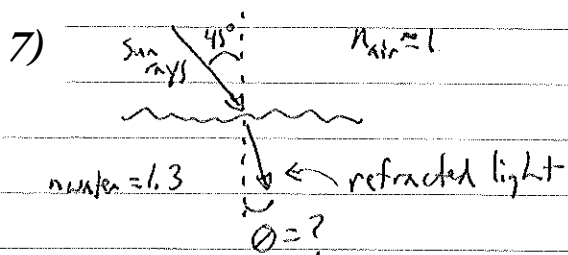
We found $f = \frac{R}{2}$

So $d_1 \approx \infty$ (sun is far away!)

$$\text{So } \frac{1}{d_1} + \frac{1}{d_2} \approx \frac{1}{d_2} = \frac{1}{f}$$

$$\text{So } \boxed{d_2 = f = \frac{R}{2}} = \frac{d}{4} = 2.5 \text{ cm}$$

Session XII, 3



Use Snell's Law to find angle of refraction

$$n_{\text{air}} \sin \theta_{\text{air}} = n_{\text{water}} \sin \theta_{\text{w}}$$

$$\sin \theta_{\text{w}} = \frac{n_{\text{a}}}{n_{\text{w}}} \sin \theta_{\text{a}}$$

$$= \frac{1}{1.3} \frac{\sqrt{2}}{2}$$

$$\boxed{\theta_{\text{w}} \approx 33^\circ}$$

8) for diffraction grating

$$d \sin \theta_n = m \lambda$$

Want to find est. of d -spacing b/t recording pits on CD

We are given that for $\lambda = 500 \text{ nm}$ $m=1$ spot is at 10°

$$d = \frac{m \lambda}{\sin \theta_n} = \frac{(1) 500 \text{ nm}}{\sin(10^\circ)} = \underline{\underline{2.88 \times 10^{-6} \text{ m}}}$$

9) Apply diffraction eqn.

$$d \sin \theta_n = m \lambda$$

Solve for λ

$$\lambda = \frac{d \sin \theta_n}{m} = \frac{(0.15 \text{ nm}) \sin(10^\circ)}{1}$$

$$= 2.6 \times 10^{-11} \text{ m}$$

$2 \sim 2.5 \times 10^4$ times smaller

than light wavelengths!

10.

a. As we saw in Guidebook entry X11.17, for telescope we have d_1 fixed and very large. This means

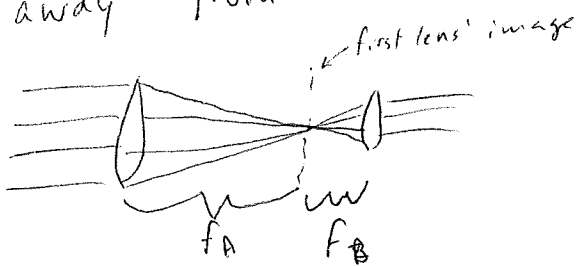
$$\frac{1}{f} = \frac{1}{d_1} + \frac{1}{d_2} \approx \frac{1}{\infty} + \frac{1}{d_2} \approx \frac{1}{d_2} \Rightarrow f \approx d_2. \text{ Thus,}$$

for highest magnification, we want f big.

$$M = \frac{d_2}{d_1} \approx \frac{f}{d_1} \text{ is biggest when } f \text{ is big.}$$

We want biggest f for our first lens,
so want 1cm focal length lens for eyepiece.

b. With infinitely far away object, incoming rays are parallel, and thus are focused at focal length. To get these rays to come out parallel through eye piece, we need the second lens to be its focal length away from the first + focal point.



total separation

$$= f_1 + f_2 = 20\text{cm} + 1\text{cm} = \boxed{21\text{cm}}$$

c. We still want 2nd lens its focal length away from first lens' image. Since we've ~~now~~ brought the ~~image~~ object closer to the first lens, its image is now farther from the lens. Since nothing changes about the second lens, we know it must be two ~~cm~~ farther, that is, $d_2 = f_A + 2\text{cm}$. Using this, we can solve for d_1 . $\frac{1}{f_A} = \frac{1}{d_1} + \frac{1}{d_2} = \frac{1}{d_1} + \frac{1}{f_A + 2\text{cm}} \Rightarrow \frac{1}{d_1} = \frac{1}{f_A} - \frac{1}{f_A + 2\text{cm}}$

$$d_1 = \frac{1}{\frac{1}{20\text{cm}} - \frac{1}{22\text{cm}}} = \boxed{220\text{cm}}$$