# Assignment – 1, 2, 3, 4, 5

Course: SC-374

Computational and Numerical Methods

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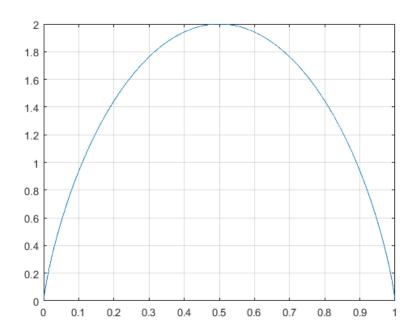
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#### **The Binary Search and Information Entropy**

Average information content is given by the formula,  $\langle I \rangle = -k \sum i Pi log 2 \mathbb{Z}^{n} Pi$ , in which k is constant and Pi is the probability of an event.

(a) For a two-outcome problem (eg. a coin toss), Show that  $\langle I \rangle$  peaks at P=12.

$$fp = \langle I \rangle = -k \ (plog2p+1-plog2p-1-p)$$
 $f'p = -k \ (log2p+1-1-log2p-1-p)$ 
 $f'p = 0 \ for \ maximize,$ 
 $log2p-1-p$ 
 $p=1-p$ 
 $p=1-p$ 



(b) Apply a very small perturbation as  $P = \frac{1}{2} + e$ , in which  $e < < \frac{1}{2}$ . Show that perturbation approach  $< I > = a - be^2$ , where a = k and b = (4k)/ln2.

 $\langle I \rangle = -k \left( \varepsilon + 12 \ln \varepsilon + 12 + 12 - \varepsilon \ln \varepsilon + 12 - \varepsilon \right)$ 

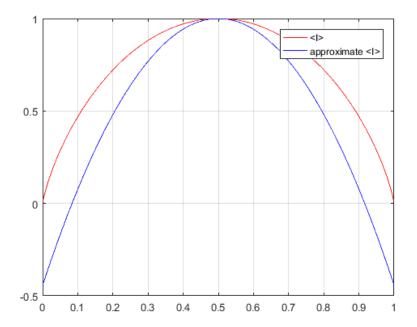
$$\langle I \rangle = -k \left( \varepsilon + 12 \ln \frac{10}{10} 2\varepsilon + 12 + 12 - \varepsilon \ln \frac{10}{10} - 2\varepsilon + 12 \right)$$

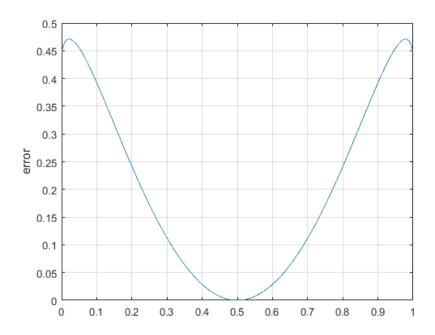
$$\langle I \rangle = -k \left( \varepsilon + 122 \varepsilon \ln \frac{10}{10} 2 - 1 + 12 - \varepsilon - 2\varepsilon \ln \frac{10}{10} 2 - 1 \right)$$

$$\langle I \rangle = -k 4\varepsilon 2 \ln \frac{10}{10} 2 - 1$$

$$\langle I \rangle = k - 4k \ln 2\varepsilon 2$$

(c) plot for both the actual function and the approximate function together and then compare the graph for closeness on the line. For plotting choose k=1.





#### **An Astrophysical Inflow**

In the problem of spherically symmetric astrophysical accretion, interstellar fluid matter (a very thin gas) travels a great distance (almost from infinity) along radial lines and falls on to a massive star (or a neutron star or even a black hole) located at the origin of coordinates. The star can be treated as a point-like particle, and the rate of the fluid flow (matter flowing in unit time) on to it is given as

$$m.=\pi G2M2\rho \infty cs3 \infty 25-3\gamma 5-3\gamma 2\gamma -2$$

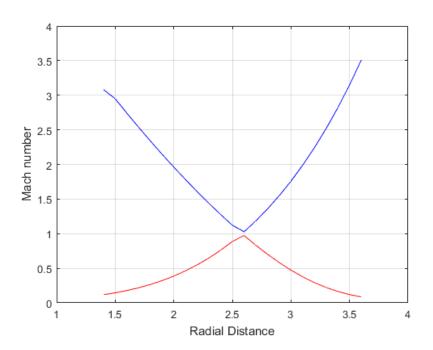
in which G is Newton's universal gravitational constant, M is the mass of the central astrophysical object,  $\rho \infty$  is the constant density of the gas at infinity,  $cs(\infty)$  is the speed of sound at infinity, and  $\gamma$  is a dimensionless number called the polytropic exponent  $(1 \le \gamma \le 5/3)$ . The velocity of the fluid flow v, as a function of the radial distance from the centre r, is given by the equation

$$fv,r=v22+n\mu.vr21n-GMr-ncs2\infty=0$$

$$\mu.=m.4\pi\rho\infty cs2n$$

$$n=1\gamma-1$$

Solve Eq.(1) by the bisection method to find v(r), using the values  $M = 2 \times 1030$  kg,  $cs(\infty) = 10$  km s-1,  $\rho\infty = 10$ -21 kg m-3 and n = 2.5. These values are typical of accretion of the interstellar medium on to a star. Each value of r in Eq.(1) will give a set of two real and physical roots of v. The plot of v(r) is shown in Fig. 1, in which v is scaled as the Mach number, v(r)/cs(r). Obtain a similar plot.

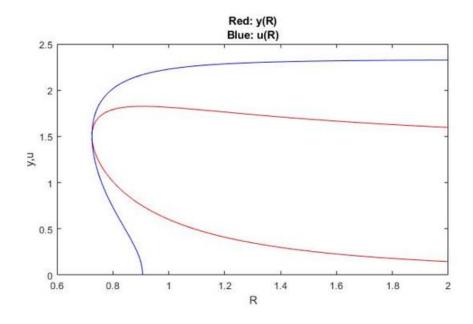


## **A Nuclear Outflow**

High-energy impacts and collisions among elementary particles can result in an outflow of nuclear fluid. The rescaled equations of the steady outflow are,

$$xyR^2 = 1$$
,  $y^2 + 3x^2 - 4x = B$ ,

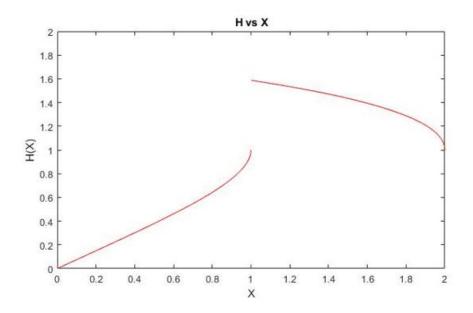
the velocity of an acoustic wave in the nuclear matter is u = x(3x - 2). On the same graph now plot R versus u.



#### The Hydraulic Jump

In a hydraulic jump, the height of a flowing liquid increases abruptly, without any pumping action. In a steady one-dimensional liquid flow, the rescaled equation of the flow is 4H - H4 = 3 (X - D), in which H is the flow height, X is the distance, and D is a constant.

- (a) Restrict your study to the range  $X \ge 0$ , and first analyse all the implications of dH/dX.
- (b) Analytically solve the quartic equation  $H \equiv H(X)$ . For the condition X = H = 0, plot X along the horizontal axis and H along the vertical axis of a graph. On the same graph, repeat the plotting exercise for H = 1 when X = 2



# The Lienard system

Apply the fourth-order Runge-Kutta method on this system. Solve separately for the two initial conditions  $(\phi, \psi) = (0.95, 0), (1.05, 0)$ . Plot  $\phi$  versus  $\psi$ .

