Experiment 25 - Balmer Lines of Hydrogen and Deuterium

Peak-finding function

To make the calibration process easier, we write a quick algorithm to find peaks in a reasonably-restricted dataset (in the broad spectra, finding peaks quickly becomes intractable because of noise and complex features in the spectra not corresponding to Balmer lines), and a second function to find the full-width at half maximum in restricted datasets with a single peak.

```
In[⊕]:= (* Finds peaks and displays plots for visual confirmation. Selects
      for peaks that survive Gaussian blurring \sigma to account for noise *)
     peakEstimate[filename_, \sigma_] :=
      Module[{data, rawData, peaks, listPeaks, interFun, p1, p2}, (data = Import[filename];
        rawData = Table[If[data[[i]][[1]] == "#", Null, data[[i]]], {i, 1, Length[data]}] /.
          Null → Sequence[]; (* removing comments and headers from data *)
        peaks = FindPeaks[rawData[[All, 2]], \sigma]; (* we select for peaks
         surviving a Gaussian blurring with \sigma to account for noise *)
        listPeaks = Table[rawData[[Floor[peaks[[i]][[1]]]]], {i, 1, Length[peaks]}];
        interFun = Interpolation[rawData];
        p1 = ListLogPlot[listPeaks,
          PlotRange → {{rawData[[1]][[1]], rawData[[Length[rawData]]][[1]]},
             {0.9 Min[rawData[[All, 2]]], All}}, FrameLabel → {"Wavelength (nm)", "Intensity"}];
        p2 = LogPlot[interFun[x], {x, rawData[[1]][[1]], rawData[[Length[rawData]]][[1]]},
           FrameLabel → {"Wavelength (nm)", "Intensity"}];
        {listPeaks, Show[p1, p2, ImageSize → 315]})]
```

```
\[ \ln[\pi] := (* Finds full-width at half-maximum for a selected peak,
     assuming the section in the file has a single peak as only
      feature. This should be sufficient for the calibration data. *)
     fwhm[filename_, peakLocation_] := Module[
       {data, rawData, interFun, halfMax, lowHalf, highHalf, x, y}, (data = Import[filename];
        rawData = Table[If[data[[i]][[1]] == "#", Null, data[[i]]], {i, 1, Length[data]}] /.
          Null → Sequence[]; (* removing comments and headers from data *)
        interFun = Interpolation[rawData];
        halfMax = Min[rawData[[All, 2]]] + (peakLocation[[2]] - Min[rawData[[All, 2]]]) / 2;
        lowHalf = x /. FindRoot[interFun[x] == halfMax,
            {x, peakLocation[[1]] + 0.1, peakLocation[[1]], rawData[[Length[rawData]]][[1]]}];
        highHalf = y /. FindRoot[interFun[y] == halfMax,
            {y, peakLocation[[1]] - 0.1, rawData[[1]][[1]], peakLocation[[1]]}];
        Abs[highHalf - lowHalf])]
```

Calibrating the Monochromator using NIST data for the Hglines

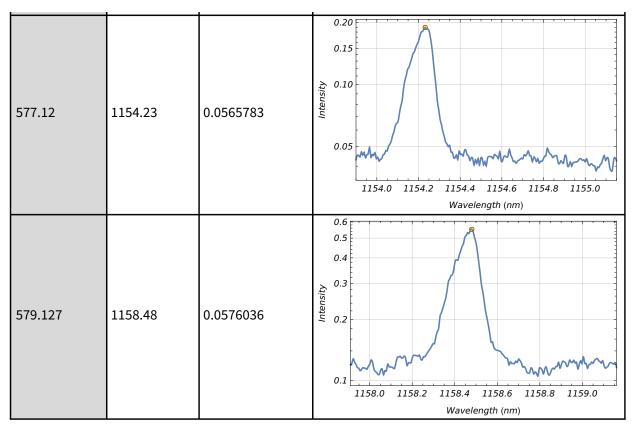
We calibrate the monochromator (note that we will be doing all observations in second order, therefore we calibrate our second order observations with the actual lines in first order) and expect a linear fit with slope ~2. We have to go up to a Gaussian smoothing with σ =28 to get rid of unwanted peaks, at which point the functions accurately finds all desired line positions. (See table below, where we show a plot of the peaks and the estimated line positions for visual confirmation that the function is performing as desired.)

We estimate the uncertainties in peak position by $\frac{\text{FWHM}}{2.355}$, where FWHM is the full-width at half-maximum. We cannot determine an accurate, independent line position uncertainty for each line, so we use the width of the peaks as a rough measure of the uncertainty in the line positions. If our lines were somehow convolved with a Gaussian scattering function as a result of our sources of uncertainty, we would expect the standard deviation to be a measure of the uncertainty in our line position, hence our choice of 2.355*FWHM. (Of course, we have no idea as to whether the sources of uncertainty act by any stretch in a Gaussian manner on our measurements, but this is the simplest assumption to make and we have no information to help us make a better guess.) This should at least provide an upper bound on the error in peak location.

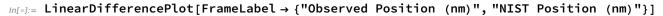
```
In[*]:= calList = {"365119.dat", "365588.dat", "366432.dat",
        "404770.dat", "435955.dat", "546226.dat", "577120.dat", "579227.dat"};
In[@]:= (* We pull the expected wavelength from
      NIST data from the comment added to each file *)
     nist[filename_] := Module[{data}, (data = Import[filename];
        data[[1]][[2]])
```

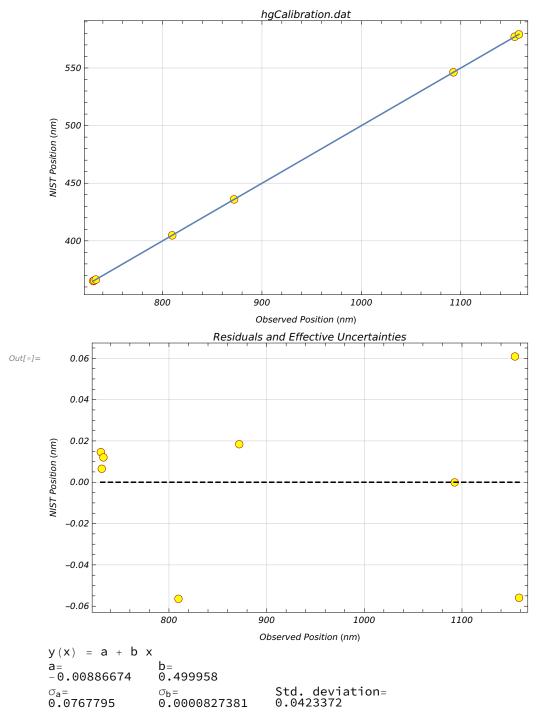
```
In[*]:= hgPeaks = {#, nist[#], Flatten[peakEstimate[#, 28][[1]], 1][[1]],
           Flatten[peakEstimate[#, 28][[1]], 1][[2]], peakEstimate[#, 28][[2]]} & /@ calList;
ln[\cdot]:= unc = \frac{1}{2.355} fwhm[#[[1]], {#[[3]], #[[4]]}] & /@ hgPeaks;
In[*]:= calTable = Join[{{"NIST
      Position (nm)", "Observed
     Position (nm)", "Estimated
     Uncertainty (nm)", "Plot of Peak"}}, Table[{hgPeaks[[i]][[2]],
           hgPeaks[[i]][[3]], unc[[i]], hgPeaks[[i]][[5]]}, {i, 1, Length[hgPeaks]}]];
In[*]:= Grid[calTable, Alignment → Left, Frame → All,
      ItemStyle → "Text", Background → {{LightGray, None}, {Gray, None}}]
```

NIST Position (nm)	Observed Position (nm)	Estimated Uncertainty (nm)	Plot of Peak
365.119	730.287	0.0593201	0.10 0.05 729.6 729.8 730.0 730.2 730.4 730.6 Wavelength (nm)
365.588	731.242	0.0586332	0.50 0.20 0.05 0.05 0.02 731.0 731.2 731.4 731.6 731.8 732.0 732.2 Wavelength (nm)



```
In[*]:= (* Having confirmation that our peak-finding function performs
       satisfactorily in finding estimates of line positions from our data,
    we use our observed line positions along with the NIST data to calibrate the
      monochromator for observations in 2<sup>nd</sup> order. We simply export the dataset
      above and perform a linear fit without uncertainties using CurveFit. *)
In[*]:= calSet = Join[{{"S NIST_Position(nm)", "Observed_Position(nm)"}},
        Drop[calTable[[All, {1, 2}]], 1]];
In[@]:= Export["hgCalibration.dat", calSet];
In[*]:= With[ {name = SystemDialogInput["FileOpen",
          {DataFileName, {"data files" -> {"*.dat", "*.mca"}, "all files" -> {"*"}}}]},
      If[ name =!= $Canceled,
       LoadFile[name]]
     ]
In[*]:= SwitchXXandYY[]
     xx and yy have been switched (so have sx and sy).
In[*]:= LinearFit[]
```





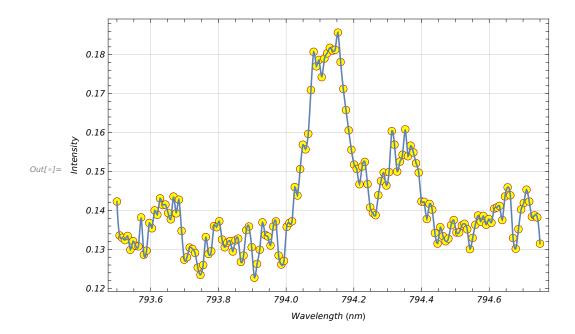
We observe that the maximum deviation from our linear fit is 0.06 nm, and our maximum fractional deviation from the fit is ~0.007% (at the 809.738 nm data point). This is a rather small fractional error, and along with the spread of the deviations from the fit, indicates that a linear fit is adequate to calibrate the monochromator for observations in second order.

The slope of 0.499958 ± 0.000083 is consistent with the expected value of 0.5 for second order diffraction, so it is difficult to tell whether the index of refraction of air in the lab or some other source of error affected our calibration in any way. However, for our purposes a slope within error of 0.5 indicates that any such source of error has a negligible effect on our calibration.

Locating the Balmer lines

Processing data points

```
In[@]:= (* First, we have to process our data for the line observed at 794nm,
     as it was not obtained using a single scan. We simply
      average over multiple scans to obtain a reasonable signal. *)
     Module[{data, rawData, avg}, (data = Import["Balmer/h794.dat"];
        rawData = Table[If[data[[i]][[1]] == "#", Null, data[[i]]], {i, 1, Length[data]}] /.
          Null → Sequence[]; (* removing comments and headers from data *);
        avg = Mean /@ GatherBy[rawData, First];
        Export["Balmer/avg794.dat", avg])];
In[@]:= data2 = Import["Balmer/avg794.dat"];
     before1 = ListPlot[data2, FrameLabel → {"Wavelength (nm)", "Intensity"}];
     beforeFun = Interpolation[data2];
     before2 = Plot[beforeFun[x], {x, data2[[1]][[1]], data2[[Length[data2]]][[1]]},
        FrameLabel → {"Wavelength (nm)", "Intensity"}];
     Show[before1, before2]
```



```
In[*]:= (* The resulting data is nevertheless very noisy,
      making it very difficult to reliably identify a peak in the
       interpolated function. We therefore perform a moving average with a
       window of 5 data points to make it possible to identify peaks in our data,
      and restrict our data to our Balmer lines of interest. *)
In[@]:= data3 = MovingAverage[data2[[All, 2]], 5];
      data4 = Partition[Riffle[data2[[All, 1]], data3], 2];
      data5 = Select[data4, 793.9 < #[[1]] < 794.5 &];</pre>
      afterFun = Interpolation[data5];
      after1 = ListPlot[data5, FrameLabel → {"Wavelength (nm)", "Intensity"}];
      after2 = Plot[afterFun[x], {x, data5[[1]][[1]], data5[[Length[data5]]][[1]]},
         FrameLabel → {"Wavelength (nm)", "Intensity"}];
      Show[
       after1,
       after2]
        0.18
        0.17
Out[•]= |
        0.16
        0.15
        0.14
        0.13
          793.9
                    794.0
                               794.1
                                         794.2
                                                    794.3
                                                              794.4
                                      Wavelength (nm)
```

In[@]:= Export["Balmer/movingAvg794.dat", data5];

 $ln[\cdot \cdot] :=$ (* Because of the noise in the data (recall that the hydrogen lamp gets dim very fast), we have to restrict our dataset to our Balmer peaks of interest in order for our function to be able to pick our peaks. We also perform a moving average as before to get a better peak location, as some of the peaks can be displaced by noise near the peaks. *)

In[@]:= SetDirectory["Balmer"];

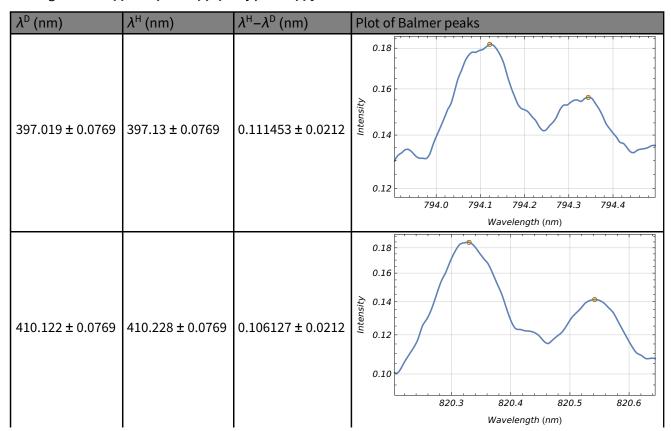
```
In[*]:= trim[filename , range ] :=
      Module[{data, rawData, avgData, avgList, trimmed}, (data = Import[filename];
        rawData = Table[If[data[[i]][[1]] == "#", Null, data[[i]]], {i, 1, Length[data]}] /.
           Null → Sequence[]; avgData = MovingAverage[rawData[[All, 2]], 5];
        avgList = Partition[Riffle[rawData[[All, 1]], avgData], 2];
        trimmed = Select[avgList, range[[1]] < #[[1]] < range[[2]] &];</pre>
        Export[StringReplace["trim%", "%" → filename], trimmed])]
In[*]:= trim["h820.dat", {820.2, 820.65}];
     trim["h868.dat", {868.0, 868.5}];
     trim["h972.dat", {972.05, 972.65}];
     trim["h1312.dat", {1312.2, 1313.1}];
```

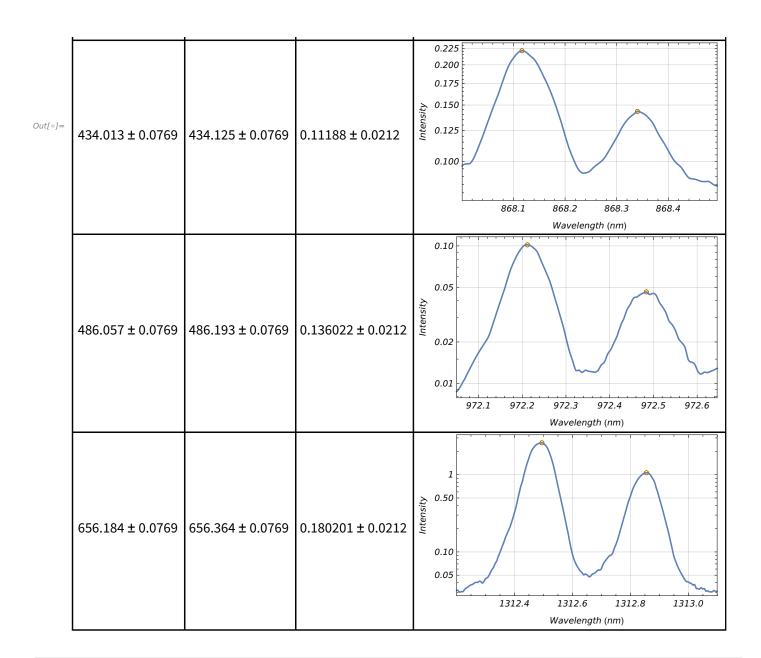
Balmer peaks and Deuterium shift

We can now use our data to locate the Balmer lines, and convert the observed values to line positions with uncertainties using our calibration.

```
In[*]:= balmer =
        {"movingAvg794.dat", "trimh820.dat", "trimh868.dat", "trimh972.dat", "trimh1312.dat"};
In[*]:= obs = peakEstimate[#, 5] & /@ balmer;
      (* We define our calibration function,
     such that it returns line positions along with propagated
        uncertainties. Since it is difficult to obtain anything more than an
        estimated upper bound on the uncertainties in observed line positions -
       because we only have one sweep for each peak - we use the FWHM on the Hydrogen peaks,
     as we had for the Hg-line calibration data. Using our FWHM function,
     we observe that the FWHM for all of our Balmer lines are very similar at ~0.1 nm,
     corresponding to a standard deviation of ~0.042 nm,
     very close to the standard deviation we have obtained for the calibration. \star)
     calFun[observation_] :=
      \left\{a + b \text{ observation, } \operatorname{Sqrt}\left[\operatorname{siga}^2 + \operatorname{observation } b \left(\left(\frac{0.1}{2.355 \text{ observation}}\right)^2 + \left(\frac{\operatorname{sigb}}{b}\right)^2\right)\right]\right\}
      (* However, our fit parameter a simply represents a systematic offset. Therefore,
     it does not affect the shift in the Balmer lines,
     since \lambda^D and \lambda^H are measured in a single scan for each data point. We can then
       obtain the corrected shift using the slope of the calibration function only,
     such that the uncertainty in the shift is only determined by the
       uncertainty in the fit parameter b. We can see, as shown below,
     that this entails that we know the separation with greater
       accuracy than we know the individual line positions. *)
     calShift[obs\D_, obs\H_] := (obs\H - obs\D) \{b, b \text{Sqrt}\left[\left(\frac{0.1}{2.355 \text{(obs\H - obs\D)}}\right)^2 + \left(\frac{\text{sigb}}{\text{b}}\right)^2\right]\}
```

```
In[*]:= balmerLines = Module[{d0bs, h0bs},
         Table[{calFun[d0bs = obs[[i]][[1]][[1]]], calFun[h0bs = obs[[i]][[1]][[2]][[1]]],
           calShift[dObs, hObs], obs[[i]][[2]]}, {i, 1, Length[obs]}]];
     balmerGrid = Join[\{\{"\lambda^D (nm)", "\lambda^H (nm)", "\lambda^H - \lambda^D (nm)", "Plot of Balmer peaks"\}\}, Table[
          \{StringReplace["%1 \pm %2", {"%1"} \rightarrow ToString[NumberForm[balmerLines[[i]]][[1]][[1]], 6]], \}
              "%2" → ToString[NumberForm[balmerLines[[i]][[1]][[2]], 3]]}],
           StringReplace["%1 \pm \%2", \{"\%1" \rightarrow ToString[NumberForm[balmerLines[[i]][[2]][[1]], 6]], \\
              "\$2" \rightarrow ToString[NumberForm[balmerLines[[i]][[2]][[2]], 3]]\}],
           StringReplace["%1 ± %2", {"%1" → ToString[NumberForm[balmerLines[[i]][[3]][[1]], 6]],
              "%2" → ToString[NumberForm[balmerLines[[i]][[3]][[2]], 3]]}],
           balmerLines[[i]][[4]]}, {i, 1, Length[balmerLines]}]];
     Grid[balmerGrid, Alignment → Left, Frame → All, ItemStyle → "Text",
      Background → {{None, None}, {Gray, None}}]
```



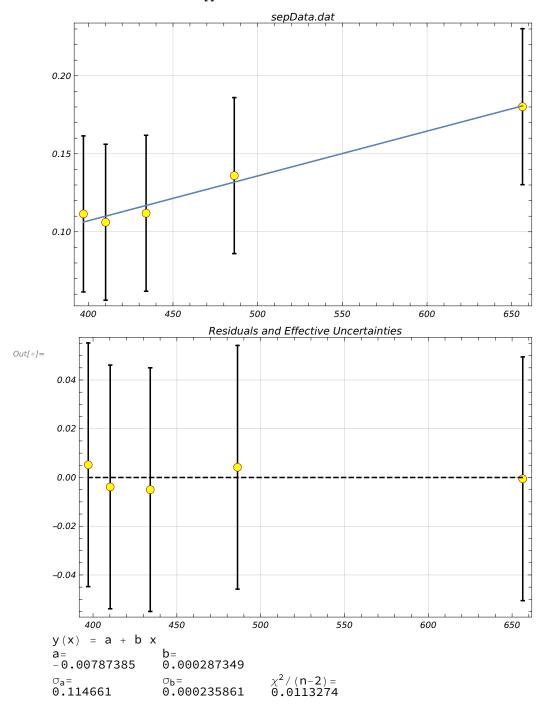


Line separations and Hydrogen line wavelengths

```
In[*]:= (* We export the hydrogen line wavelength
      and line separation data for analysis in CurveFit. *)
    hdData = Join[{{"S lambdaH(nm)", "shift(nm)", "sigmaShiftH(nm)", "sigmaLambdaH"}}, Table[
         {balmerLines[[i]][[2]][[1]], balmerLines[[i]][[3]][[1]], balmerLines[[i]][[3]][[2]],
          balmerLines[[i]][[2]][[2]]], {i, 1, Length[balmerLines]}]];
    Export["sepData.dat", hdData];
```

```
In[@]:= With[ {name = SystemDialogInput["FileOpen",
              {DataFileName, {"data files" -> {"*.dat", "*.mca"}, "all files" -> {"*"}}}]},
        If[ name =!= $Canceled,
         LoadFile[name]]
       ]
In[•]:= LinearFit[]
       n = 5
       y(x) = a + b x
       Fit of (x,y) (unweighted)
                           b=
0.000287349
       a=
-0.00787385
                                                 Std. deviation= 0.00532107
       \sigma_a = 0.0122031
                           σ<sub>b</sub>=
0.0000251021
       Fit of (x, y \pm \sigma_y)
                           b=
0.000287349
       a=
-0.00787385
                                               \chi^2/(n-2) = 0.0113274
                           <sub>Ob</sub>=
0.000235861
       \sigma_a = 0.114661
       Fit of (x \pm \sigma_x, y \pm \sigma_y)
                           b=
0.000287349
       a=
-0.00787385
                                               \chi^2/(n-2) = 0.0113274
       \sigma_a = 0.114661
                           \sigma_{b}^{=} 0.000235861
```

In[*]:= LinearDifferencePlot[]



We observe a large error in our fit parameters for our weighted fit, and a $\tilde{\chi}^2$ well below 1, indicating that, as expected, the FWHM estimate of uncertainty in peak location was a large underestimate of the accuracy of the monochromator. We therefore use the parameters of the unweighted fit in the following analysis. From equation (11), we have (where b is the fit parameter for our fit of $\lambda^H - \lambda^D$ against λ^H):

$$\frac{m_e}{m_p} = 2 b$$

```
In[\bullet]:= ratio = 2 \{b, 0.0000251\}
Out[\bullet] = \{0.000574698, 0.0000502\}
```

Therefore we estimate $\frac{m_e}{m_p} = (5.75 \pm 0.50) \times 10^{-4}$. This is within error of the NIST value of 5.466×10^{-4} .

Estimating Avogadro's number

From the above, we estimate the proton rest energy as:

$$In[*]:= ep = \frac{0.511}{ratio[[1]]} (* in MeV *)$$
 $Out[*]= 889.163$

We then estimate the mass of the hydrogen atom as:

$$ln[\cdot] := mh = \frac{(ep + 0.511) * 10^6 * 1.602 * 10^{-19}}{(3 * 10^8)^2} (*in kg *)$$

 $\textit{Out[•]}{=}~1.58362 \times 10^{-27}$

If a mole of H_2 masses 2g, we can then estimate Avogadro's number as:

$$ln[*]:= na = \frac{0.002}{2 \text{ (mh)}} (* in mol^{-1} *)$$

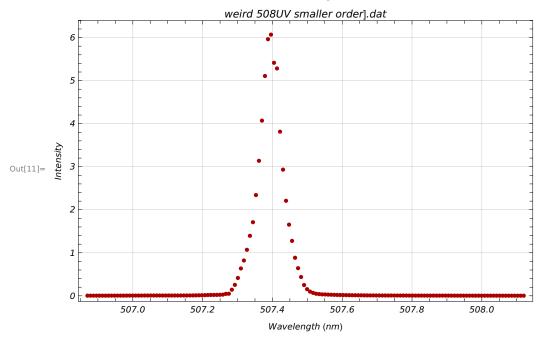
$$Out[*]:= 6.31465 \times 10^{23}$$

which is reasonably close to the NIST value of 6.033×10^{23} mol⁻¹.

Optional observation: line at 508 nm

```
In[10]:= With[ {name = SystemDialogInput["FileOpen",
          {DataFileName, {"data files" -> {"*.dat", "*.mca"}, "all files" -> {"*"}}}]},
      If[ name =!= $Canceled,
       LoadFile[name]]
     ]
```

ln[11]:= LinearDataPlot[FrameLabel \rightarrow {"Wavelength (nm)", "Intensity"}]



We observe this feature at 507.4 nm, which seems to disappear if we use the clear filter. Since the clear filter should let 507.4 nm light through (cyan), we deduce that we are observing a lower wavelength feature in higher order. In fact this corresponds to an observation in second order of a strong line in the NIST database at 253.48 nm.

We can observe the feature in fourth order:

```
In[7]:= With[ {name = SystemDialogInput["FileOpen",
         {DataFileName, {"data files" -> {"*.dat", "*.mca"}, "all files" -> {"*"}}}]},
     If[ name =!= $Canceled,
      LoadFile[name]]
    ]
```

ln[9]:= LinearDataPlot[FrameLabel \rightarrow {"Wavelength (nm)", "Intensity"}]

