Fast Estimation of Matrix Exponential Spatial Models

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Abstract

The matrix exponential spatial specification (MESS) is an alternative to the spatial autoregressive-type (SAR-type) specifications with several attractive properties. The spatial dependence in the MESS-type models is formulated through a matrix exponential term, and the estimation of these models may require the computation of the matrix exponential terms many times in an estimation procedure. In the literature, it is well documented that the computation of the matrix exponential terms can pose challenges in terms of reliability, stability, accuracy, and efficiency. We propose a matrix-vector products approach based on the truncation of Taylor series expansion of the matrix exponential terms for the fast estimation of MESS-type models. We show how to efficiently implement this approach for a first-order MESS model, and provide extensive simulation evidence for its computational advantage over the default method utilized by a popular statistical software.

JEL-Classification: C13, C21, C31.

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1 Introduction

Spatial econometric models account for potential (weak) cross-sectional dependence among observations located on some relevant space. The Cliff-Ord-type spatial models (i.e., the spatial autoregressive (SAR)-type) impose an autoregressive specification for the variable of interest using a weights matrix which quantifies the strength of a measure of nearness between spatial units. (Anselin, 1988; Cliff and Ord, 1969, 1973; Whittle, 1954). The spatial autoregressive specification ensures that the spatial dependence among spatial units decays at a geometric rate. The likelihood based estimation of these models requires calculation of matrix determinant terms (the Jacobian terms) in each pass of the numerical optimization scheme, and for large matrices that are not sparse this can create computational challenges. See for example LeSage and Pace (2009, Chapter 4) for a variety of approximation methods suggested in the literature to alleviate this computational problem.

An alternative to SAR-type models is the matrix exponential spatial specification (MESS) suggested by LeSage and Pace (2007). In the MESS-type models, the spatial dependence is formulated with a matrix exponential term of type, $e^{\alpha A} = \sum_{i=0}^{\infty} (\alpha A)^i/i!$, where α is a scalar spatial parameter, A is the $n \times n$ spatial weights matrix and n is the number of spatial units. Therefore, the MESS-type models impose an exponential rate of decay for the cross-sectional dependence and have several features that make them more convenient for estimation (Chiu et al., 1996; Leonard and Hsu, 1992). The likelihood based estimation is greatly simplified because the likelihood function does not involve any matrix determinant terms. Since matrix exponential terms are always invertible, there is no need to impose restrictions on the parameter space of the spatial parameters, i.e., the MESS-type models always have reduced forms. Moreover, when the model involves heteroskedasticity of an unknown form in the error terms, the maximum likelihood estimator (MLE) remains consistent provided that the weights matrices are commutative (Debarsy et al., 2015). 1

Despite the aforementioned advantages of the MESS-type models, the likelihood and generalized method of moments (GMM) based estimations of these models require the computation of matrix exponential terms by the optimization solvers in each iteration, which can be computationally costly. The literature has suggested several alternative ways to compute $e^{\alpha A}$ such as Taylor series approximation, Padé approximation, ordinary differential equation methods, polynomial methods, matrix decomposition methods, splitting methods and Krylov space methods. Moler and Van Loan (1978, 2003) assess the effectiveness of nineteen methods according to the following attributes: (i) generality, (ii) reliability, (iii) stability, (iv) accuracy, (v) efficiency, (vi) storage requirements, (vii) ease of use, and (viii) simplicity. Although Moler and Van Loan (1978, 2003) state that "none (of the methods in their paper) are completely satisfactory", they claim that a scaling and squaring method with either the rational Padé or Taylor approximants can be the most effective one to compute the matrix exponential terms.² Popular software such as Python, R, MATLAB and Mathematica

¹See Debarsy et al. (2015) for the formal results on on the maximum likelihood (ML) and generalized method of moments (GMM) estimation of the MESS models.

²On the scaling and squaring method with either the rational Padé or Taylor approximants, see also Higham (2005)

have modules, packages and functions to compute matrix exponential of a given matrix. For example, MATLAB (function expm), Mathematica (function MatrixExp) and Python (function scipy.linalg.expm) use a scaling and squaring method combined with a Padé approximation for the computation of matrix exponential terms (Al-Mohy and Higham, 2010; Higham, 2005).

As pointed out by Moler and Van Loan (1978, 2003), all methods suggested in the literature for computing the matrix exponential terms are "dubious" in the sense that a sole method may not be entirely reliable for all applications.³ In other words, a method that is effective for a particular application may not be reliable for another application. In the context of MESS-type models, the scaling and squaring method combined with the Padé approximation as implemented in MATLAB through its expm function can be highly costly in terms of computation time. For example, a typical Monte Carlo simulation designed for a MESS model that has spatial dependence in the dependent variable and the disturbance term (for short MESS(1,1)) with 1000 resampling and a couple of different parameter combinations can take days or even weeks (see Section 4 on the details of the simulation setting). The expm function is also costly for estimating empirical applications with large sample sizes. For example, we show that the expm function for estimating the MESS (1,1) model for an empirical illustration involving 3107 observations takes 1072 seconds by the the quasi maximum likelihood (QML) estimator, 4805 seconds by the GMM estimator and 47742 seconds by the Bayesian estimator (see Section 5 for the details).

In this paper, we propose a matrix-vector products method based on the truncation of Taylor series expansion of the matrix exponential terms for the fast estimation of MESS-type models. Our analysis on the estimation of MESS-type models indicates that the estimation requires the computation of a matrix exponential term as a vector, rather than the matrix exponential term in isolation. For example, the estimation of the MESS (1,1) model requires the computation of terms such as $e^{\alpha A}e^{\tau B}v$ and $e^{\tau B}X$, where v is an $n \times 1$ vector, X is an $n \times k$ matrix, and α and τ are the spatial parameters. The matrix-vector products method provides approximations to $e^{\alpha A}e^{\tau B}v$ and $e^{\tau B}X$ in terms of matrix-vector products rather than providing approximations for $e^{\alpha A}$ and $e^{\tau B}$. In this paper, we show how this approach can be implemented for the QML, GMM and Bayesian estimation of the MESS(1,1) model. Using our suggested approach in the context of the MESS (1,1) model, we provide extensive simulation evidence on the computational time gains for three estimation methods (the QML, GMM and Bayesian methods).

In the literature, it is well known that the matrix-vector products approach can reduce the computational burden substantially. For example, Moler and Van Loan (2003) write "One of the most significant changes in numerical linear algebra in the past 25 years is the rise of iterative methods for sparse matrix problems, in which only matrix vector products are needed." In particular, LeSage and Pace (2009) extensively use sparse matrix-vector operations for the likelihood

and Bader et al. (2019). Sidje (1998) provides an extensive package named ExpoKIT (both Fortran and MATLAB versions are available) to compute matrix exponential terms with the Krylov subspace method using the Arnoldi process approximation.

³The computation of matrix exponential terms may be necessary for many applications from different fields. For example, MATLAB uses its expm function in its Control Toolbox, System Identification Toolbox, Neural Net Toolbox, Mu-Analysis and Synthesis Toolbox, Model Predictive Control toolbox, and Simulink.

and Bayesian estimation of spatial models to reduce the computational burden. The Krylov space methods (the twentieth method in Moler and Van Loan (2003)) suggested for the computation of matrix exponential terms also depends on the matrix-vector products approach. See, for example, Saad (1992), Gallopoulos and Saad (1992), Hochbruck and Lubich (1997), Sidje (1998), and related references. When A is sparse, in the first step of this approach, $e^{\alpha A}v$ is approximated by an element of Krylov subspace $K_m = \text{span}\{v, (\alpha A)v, \dots, (\alpha A)^m v\}$, where m, the dimension of the Krylov subspace, is small compare to n. The operations in the first step of this method involve only matrix-vector products (see Sidje (1998) for the implementation of this method).

The remainder of this paper is organized as follows. Section 2 presents the model under consideration and lays out briefly the details on the QML, GMM, and Bayesian MCMC estimation. This section also shows how the impact measures and their dispersion measures can be estimated. Section 3 provides the details on the computation of the matrix exponential terms using the matrix-vector products approach. We then show how the matrix-vector products approach can be applied to the QML, GMM and Bayesian estimation methods. Section 4 presents the setting for our Monte Carlo study and the simulation results. Section 5 illustrates the computational time advantage of the matrix-vector products method using a large dataset from the spatial econometric literature. We conclude in Section 6. Some simulation results are relegated to an appendix.

2 Model and Estimation Approaches

2.1 Model Specification

We consider the following first order matrix exponential spatial model (for short MESS(1,1))

$$e^{\alpha_0 W} y = X \beta_0 + u, \quad e^{\tau_0 M} u = \epsilon, \tag{2.1}$$

where $Y = (y_1, \ldots, y_n)'$ is the $n \times 1$ vector of observations on the dependent variable, X is the $n \times k$ matrix of non-stochastic exogenous variables with the associated parameter vector β_0 , W and M are the $n \times n$ spatial weights matrices of known constants with zero diagonal elements. The scalar parameters α_0 and τ_0 are called the spatial parameters. We call $U = (u_1, \ldots, u_n)'$ as the $n \times 1$ vector of regression disturbance terms and $\epsilon = (\epsilon_1, \ldots, \epsilon_n)'$ as the $n \times 1$ vector of disturbances (or innovations).

The matrix exponential terms $e^{\alpha W}$ and $e^{\tau M}$ in (2.1) are defined as $e^{\alpha W} = \sum_{i=0}^{\infty} (\alpha W)^i/i!$ and $e^{\tau M} = \sum_{i=0}^{\infty} (\tau M)^i/i!$, and are always invertable with the inverses $e^{-\alpha W}$ and $e^{-\tau M}$ (Chiu et al., 1996). Thus, the reduced form of the model always exists and is given by $y = e^{-\alpha W} X \beta_0 + e^{-\alpha W} e^{-\tau M} \epsilon$. On the other hand, in a SAR-type model, we need to restrict the parameter space of spatial parameters so that it has a reduced form. On the parameter space of spatial parameters in the SAR-type models, among others, see Elhorst (2014), Kelejian and Prucha (2010), Lee (2004), and LeSage and Pace (2009). In the SAR-type counterpart of the MESS(1,1) model, the matrix exponential terms $e^{\alpha_0 W} y$ and $e^{\tau_0 M} u$ in (2.1) are respectively replaced by $(I_n - \lambda_0 W) y$ and $(I_n - \lambda_0 W) y$

 $\rho_0 M)u$, where I_n is the $n \times n$ identity matrix, and λ_0 and ρ_0 are scalar spatial parameters. Under the assumption that $\|\lambda_0 W\| < 1$ for some matrix norm $\|\cdot\|$, we have $(I_n - \lambda_0 W)^{-1} = \sum_{i=0}^{\infty} (\lambda_0 W)^i$ (Horn and Johnson, 2012). Thus, the SAR model imposes a geometric decay pattern of spatial dependence among spatial units, while the MESS(1, 1) model exhibits an exponential decay.

2.2 Maximum Likelihood Approach

Under the assumption that ϵ_i 's are i.i.d normal with mean zero and variance σ_0^2 , the log-likelihood function of the model can be expressed as

$$\ln L(\theta) = -\frac{n}{2} \ln(2\pi\sigma^2) - \frac{1}{2\sigma^2} (e^{\alpha W}y - X\beta)' e^{\tau M'} e^{\tau M} (e^{\alpha W}y - X\beta), \tag{2.2}$$

where $\theta = (\alpha, \tau, \beta', \sigma^2)'$. Note that (2.2) does not include the Jacobian terms, because $\ln |e^{\alpha W}| = \ln (e^{\alpha \operatorname{tr}(W)}) = 0$ and $\ln |e^{\tau M}| = \ln (e^{\tau \operatorname{tr}(M)}) = 0$, where $|\cdot|$ denotes the determinant operator, and $\operatorname{tr}(\cdot)$ is the trace operator. Let $\psi = (\alpha, \tau)'$. Then, for a given value of ψ , the first-order conditions of (2.2) with respect to β and σ^2 yield

$$\widehat{\beta}(\psi) = (X' e^{\tau M'} e^{\tau M} X)^{-1} X' e^{\tau M'} e^{\tau M} e^{\alpha W} y, \tag{2.3}$$

$$\widehat{\sigma}^{2}(\psi) = \frac{1}{n} (e^{\alpha W} y - X \widehat{\beta})' e^{\tau M'} e^{\tau M} (e^{\alpha W} y - X \widehat{\beta}) = \frac{1}{n} y' e^{\alpha W'} e^{\tau M'} H(\tau) e^{\tau M} e^{\alpha W} y, \tag{2.4}$$

where $H(\tau) = I_n - e^{\tau M} X (X' e^{\tau M'} e^{\tau M} X)^{-1} X' e^{\tau M'}$. Then, ignoring constant terms, the concentrated likelihood function can be written as

$$\ln L^{c}(\theta) = -\frac{n}{2} \ln(\widehat{\sigma}^{2}(\psi)) = -\frac{n}{2} \ln\left(\frac{1}{n} y' e^{\alpha W'} e^{\tau M'} H(\tau) e^{\tau M} e^{\alpha W} y\right). \tag{2.5}$$

Thus, we can define the QMLE of ψ_0 as

$$\widehat{\psi} = \operatorname{argmin}_{\psi} \left(y' e^{\alpha W'} e^{\tau M'} H(\tau) e^{\tau M} e^{\alpha W} y \right). \tag{2.6}$$

Let $\gamma = (\alpha, \tau, \beta')'$ and $\gamma_0 = (\alpha_0, \tau_0, \beta'_0)'$ be the true parameter vector. Then, QMLE $\hat{\gamma}$ has the following asymptotic normal distribution (Debarsy et al., 2015),⁵

$$\sqrt{n}\left(\widehat{\gamma} - \gamma_0\right) \stackrel{d}{\to} N\left(0, \lim_{n \to \infty} C^{-1}\Omega C^{-1}\right). \tag{2.7}$$

⁴On the properties of MESS and SAR type models, see Debarsy et al. (2015), Kelejian and Prucha (2010), Lee (2004), and LeSage and Pace (2007, 2009).

⁵Note that under heteroskedasticity of an unknown form, the QMLE is still consistent and has an asymptotically normal distribution when W and M commute, i.e., WM = MW (Debarsy et al., 2015).

Here, $\Omega = 2\sigma_0^2 C + \Omega_1$ with

$$C = \frac{1}{n} \begin{pmatrix} \sigma_0^2 \operatorname{tr} \left(\mathbb{W}^s \mathbb{W}^s \right) + 2 \left(\mathbb{W} e^{\tau_0 M} X \beta_0 \right)' \left(\mathbb{W} e^{\tau_0 M} X \beta_0 \right) & * & * \\ \sigma_0^2 \operatorname{tr} \left(\mathbb{W}^s M^s \right) & \sigma_0^2 \operatorname{tr} \left(M^s M^s \right) & * \\ -2 \left(e^{\tau_0 M} X \right)' \mathbb{W} e^{\tau_0 M} X \beta_0 & 0 & 2 \left(e^{\tau_0 M} X \right)' \left(e^{\tau_0 M} X \right) \end{pmatrix},$$

$$(2.8)$$

$$\Omega_{1} = \frac{1}{n} \begin{pmatrix} (\mu_{4} - 3\sigma_{0}^{4}) \operatorname{vec}'_{D}(\mathbb{W}^{s}) \operatorname{vec}_{D}(\mathbb{W}^{s}) + 4\mu_{3} (\mathbb{W}e^{\tau_{0}M}X\beta_{0})' \operatorname{vec}_{D}(\mathbb{W}^{s}) & * & * \\ 0 & 0 & * \\ -2\mu_{3} (e^{\tau_{0}M}X)' \operatorname{vec}_{D}(\mathbb{W}^{s}) & 0 & 0 \end{pmatrix},$$
(2.9)

where $\mathbb{W} = e^{\tau_0 M} W e^{-\tau_0 M}$, $\mu_3 = \mathbb{E}\epsilon_i^3$, $\mu_4 = \mathbb{E}\epsilon_i^4$, $\operatorname{vec}_D(A)$ denotes a vector containing the diagonal elements of any square matrix A, and $B^s = B + B'$ for any $n \times n$ matrix B.

2.3 GMM Approach

When the error terms are homoskedastic, a set of moment functions that consists of linear and quadratic moment functions can be arranged such that the resulting GMME is as efficient as the QMLE under the normal case, and asymptotically more efficient than the QMLE under the nonnormal case. In the case of heteroskedasticity of an unknown form, an optimal GMME (OGMME) can be defined such that it is also more efficient than the QMLE (Debarsy et al., 2015).

To define the best set of moment functions when the error terms are simply i.i.d, we introduce the following notations. Let $\operatorname{Diag}(a)$ be the $n \times n$ matrix whose diagonal entries are the elements of the $n \times 1$ vector a, and $\operatorname{Diag}(A)$ be the $n \times n$ diagonal matrix whose diagonal entries are those of the $n \times n$ matrix A. Let $B^{(t)} = B - I_n \operatorname{tr}(B)/n$ for any $n \times n$ matrix B, and X_n^* be the submatrix of X with the intercept term removed. Define $P_1^* = \mathbb{W}, P_2^* = \operatorname{Diag}(\mathbb{W}), P_3^* = \operatorname{Diag}\left(e^{\tau_0 M}WX\beta_0\right)^{(t)}, P_4^* = M,$ $P_{l+4}^* = \operatorname{Diag}\left(e^{\tau_0 M}X_l^*\right)^{(t)}$ for $l = 1, \ldots, k^*$ and $F^* = (F_1^*, F_2^*, F_3^*, F_4^*)$ with $F_1^* = e^{\tau_0 M}X^*, F_2^* = e^{\tau_0 M}WX\beta_0, F_3^* = l_n$, and $F_4^* = \operatorname{vec}_D(\mathbb{W})$. Then, under the assumption that the error terms are i.i.d, the best set of moment functions suggested in Debarsy et al. (2015) takes the following form:

$$g^{*}(\gamma) = \frac{1}{n} \left(\epsilon'(\gamma) P_{1}^{*} \epsilon(\gamma), \dots, \epsilon'(\gamma) P_{k^{*}+4}^{*} \epsilon(\gamma), \epsilon'(\gamma) F^{*} \right)', \tag{2.10}$$

where $\epsilon(\gamma) = e^{\tau M} \left(e^{\alpha W} y - X \beta \right)$. Then, the best GMME (BGMME) is defined as

$$\widehat{\gamma}_B = \operatorname{argmin}_{\gamma} g^{*'}(\gamma) V^{*-1} g^*(\gamma), \tag{2.11}$$

where $V^* = nE\left(g^*\left(\gamma_0\right)g^{*'}\left(\gamma_0\right)\right)$. It can be shown that

$$V^* = \frac{1}{n} \begin{pmatrix} \frac{\sigma_0^4}{2} \omega' \omega + \frac{1}{4} \left(\mu_4 - 3\sigma_0^4 \right) \omega_d' \omega_d & \frac{1}{2} \mu_3 \omega_d' F^* \\ \frac{1}{2} \mu_3 F^{*\prime} \omega_d & \sigma_0^2 F^{*\prime} F^* \end{pmatrix}, \tag{2.12}$$

where $\omega_d = (\operatorname{vec}_D(P_1^{*s}), \dots, \operatorname{vec}_D(P_{k^*+4}^{*s}))$, $\omega = (\operatorname{vec}(P_1^{*s}), \dots, \operatorname{vec}(P_{k^*+4}^{*s}))$ and $\operatorname{vec}(A)$ denotes the vectorization of matrix A. Under some regularity conditions, it follows that (Debarsy et al., 2015)

$$\sqrt{n}\left(\widehat{\gamma}_B - \gamma_0\right) \stackrel{d}{\to} N\left(0, \lim_{n \to \infty} (G^{*'}V^{*-1}G^*)^{-1}\right),\tag{2.13}$$

where

$$G^* = \mathbb{E}\left(\frac{\partial g^*(\gamma_0)}{\partial \gamma'}\right) = \frac{1}{n} \begin{pmatrix} \frac{\sigma_0^2}{2} \omega' \operatorname{vec}\left(\mathbb{W}^s\right) & \frac{\sigma_0^2}{2} \omega' \operatorname{vec}\left(M^s\right) & 0\\ F^{*\prime} \mathbb{W} e^{\tau_0 M} X \beta_0 & 0 & -F^{*\prime} e^{\tau_0 M} X \end{pmatrix}. \tag{2.14}$$

Note that the BGMME defined (2.11) is not feasible, since V^* , P_1^* ,..., $P_{k^*+4}^*$ and F^* are functions of the unknown parameters. In practice, an initial consistent estimator of γ_0 can be used to replace the unknown parameters in these terms.⁶ A feasible estimator formulated in this way can be shown to have the same asymptotic distribution as that of $\widehat{\gamma}_B$ (Debarsy et al., 2015).

When the disturbance terms have heteroskedasticity of an uknown form, the matrices P_i 's used in the quadratic moment functions need to have zero diagonal elements. In the heteroskedastic case, an OGMME can be defined by using the following vector of moment functions,

$$\widehat{g}_d(\gamma) = \frac{1}{n} \left(\epsilon'(\gamma) \left(\widehat{\mathbb{W}} - \operatorname{Diag}(\widehat{\mathbb{W}}) \right) \epsilon(\gamma), \epsilon'(\gamma) M \epsilon(\gamma), \epsilon'(\gamma) \left(\widehat{\mathbb{W}} e^{\widehat{\tau} M} X \widehat{\beta}, e^{\widehat{\tau} M} X \right) \right)', \tag{2.15}$$

where $\widehat{\mathbb{W}} = e^{\widehat{\tau}M} W e^{-\widehat{\tau}M}$.

2.4 Bayesian Approach

To complete the model specification, we need to specify the prior distributions for α, τ, β and σ^2 . We assume the following prior distributions: $\alpha \sim N(\mu_{\alpha}, V_{\alpha}), \tau \sim N(\mu_{\tau}, V_{\tau}), \beta \sim N(\mu_{\beta}, V_{\beta}),$ and $\sigma^2 \sim IG(a_0, b_0)$, where IG denotes the inverse-gamma distribution. The likelihood is given by

$$y|\alpha, \tau, \beta, \sigma^2 \sim N(e^{-\alpha W} X \beta, \sigma^2 e^{-\alpha W} e^{-\tau M} e^{-\tau M'} e^{-\alpha W'}). \tag{2.16}$$

The standard Bayesian analysis for a linear regression model can be used to obtain the conditional posterior distributions of β and σ^2 . On the other hand, combining the likelihood function with the prior densities of spatial parameters indicates that the conditional posterior distributions of α and τ are non-standard. In the following algorithm, we suggest a Gibbs sampler that shows how to generate random draws from the joint posterior distribution $p(\beta, \sigma^2, \alpha, \tau|y)$.

Algorithm 1.

⁶Among other consistent estimators, the following initial GMME can be used: $\widehat{\gamma} = \operatorname{argmin}_{\gamma} g'(\gamma) g(\gamma)$, where $g(\gamma) = (W \epsilon(\gamma), M \epsilon(\gamma), W X, X)' \epsilon(\gamma)$ and $\epsilon(\gamma) = e^{\tau M} (e^{\alpha W} y - X \beta)$.

⁷We use $p(\cdot)$ to denote the relevant density function, and we omit X in the conditional sets for the simplicity of exposition.

1. Sampling step for β :

$$\beta | y, \alpha, \tau, \sigma^2 \sim N(\widehat{\beta}, K_\beta),$$
 (2.17)

where $K_{\beta} = (V_{\beta}^{-1} + X'e^{\tau M'}e^{\tau M}X/\sigma^2)^{-1}$ and $\widehat{\beta} = K_{\beta}(X'e^{\tau M'}e^{\tau M}e^{\alpha W}y/\sigma^2 + V_{\beta}^{-1}\mu_{\beta})$. 2. Sampling step for σ^2 :

$$\sigma^2|y,\alpha,\tau,\beta \sim IG(\widehat{\sigma}^2,K_{\sigma^2}),\tag{2.18}$$

where $\hat{\sigma}^2 = a_0 + \frac{n}{2}$ and $K_{\sigma^2} = b_0 + \frac{1}{2} (e^{\alpha W} y - X \beta)' e^{\tau M'} e^{\tau M} (e^{\alpha W} y - X \beta)$.

3. Sampling step for α :

$$p(\alpha|y,\beta,\tau,\sigma^2) \propto \exp\left(-\frac{1}{2}\left(\sigma^{-2}(e^{\alpha W}y - X\beta)'e^{\tau M'}e^{\tau M}(e^{\alpha W}y - X\beta) + V_{\alpha}^{-1}(\alpha^2 - 2\mu_{\alpha}\alpha)\right)\right),$$
(2.19)

which is a non-standard distribution. We can use a random-walk Metropolis-Hastings algorithm to sample from this distribution (LeSage and Pace, 2009). A candidate value α^{new} is generated according to

$$\alpha^{new} = \alpha^{old} + c_{\alpha} \times N(0, 1), \tag{2.20}$$

where c_{α} is the tuning parameter.⁸ The candidate value α^{new} is accepted with probability

$$\mathbb{P}(\alpha^{new}, \alpha^{old}) = min\left(1, \frac{p(\alpha^{new}|y, \beta, \sigma^2, \tau)}{p(\alpha^{old}|y, \beta, \sigma^2, \tau)}\right). \tag{2.21}$$

4. Sampling step for τ :

$$p(\tau|y,\beta,\alpha,\sigma^{2}) \propto \exp\left(-\frac{1}{2}\left(\sigma^{-2}(e^{\alpha W}y - X\beta)'e^{\tau M'}e^{\tau M}(e^{\alpha W}y - X\beta) + V_{\tau}^{-1}(\tau^{2} - 2\mu_{\tau}\tau)\right)\right). \tag{2.22}$$

We use the random-walk Metropolis-Hastings algorithm described in Step 3 to generate random draws from $p(\tau|y, \beta, \alpha, \sigma^2)$.

Remark 1. LeSage and Pace (2007) develop an efficient Bayesian estimation method for the MESS(1,0) model by assuming a normal-gamma prior for β and σ^2 , and a normal distribution prior for α . By using some properties of the multivariate normal distribution and the inverse gamma distribution, LeSage and Pace (2007) analytically derive the marginal posterior distribution of α . Since the marginal distribution of α is not in a known form, they suggest to use a univariate numerical integration method for the computation of posterior moments. Also, since the marginal posterior distributions of β and σ^2 depend on α , they suggest to compute the posterior moments of

 $^{^8}$ The tuning parameter is determined during the estimation such that the acceptance rate falls between 40% and 60%.

these parameters by numerical integration over α . Our approach presented in Algorithm 1 differs in two important ways. First, we use the random-walk Metropolis-Hastings algorithm suggested by LeSage and Pace (2009) to generate posterior draws for α . Second, we suggest independent prior distributions for β and σ^2 , and use the Gibbs sampler in Algorithm 1 to generate posterior draws for β and σ^2 .

2.5 Impact Measures

The dispersion of parameter estimators can be estimated in different ways (Arbia et al., 2020; Debarsy et al., 2015; Elhorst, 2014; LeSage and Pace, 2009; Taşpınar et al., 2018). In the case of the QMLE and GMME defined in Sections 2.2 and 2.3, the closed-forms of variance-covariance matrices are available. Thus, we can use the plug-in method for these estimators. That is, the unknown parameters in these variance-covariance matrices can be replaced by the corresponding estimates obtained from consistent estimators. In the case of Bayesian approach, we can use the empirical standard deviations of the random draws generated through Algorithm 1 as the estimate for the standard errors of parameters.

According to the model in (2.1), the derivative of y with respect to the kth explanatory variable x_k gives the marginal effect $e^{-\alpha_0 W}\beta_{0k}$, where β_{0k} is the kth element of the true coefficient vector β_0 . To ease the interpretation and presentation of this marginal effect, LeSage and Pace (2009) define three scalar measures for the marginal effect: the average direct impact, the average indirect impact, and the average total impact. The average direct impact is the average of the main diagonal elements of $e^{-\alpha_0 W}\beta_{0k}$, the average indirect impact is the average of the off-diagonal elements of $e^{-\alpha_0 W}\beta_{0k}$, and the total impact is the average of all elements of $e^{-\alpha_0 W}\beta_{0k}$. For statistical inference, one needs to determine the dispersions of these scalar impact measures. In the Bayesian approach, a sequence of random draws obtained through Algorithm 1 can be used to generate a sequence of random draws for each impact measure. Then, the mean and the standard deviation calculated from each sequence of impact measures can be used for inference.

In the QML and GMM cases, the classical delta method can be used to determine the dispersions of the impact measures. The estimator of the average direct effect is given by $\frac{1}{n} \operatorname{tr}(e^{-\widehat{\alpha}W_n}\widehat{\beta}_k)$. Then, by the mean value theorem, we obtain

$$\frac{1}{\sqrt{n}} \left(\operatorname{tr}(e^{-\widehat{\alpha}W}\widehat{\beta}_k) - \operatorname{tr}(e^{-\alpha_0 W}\beta_{0k}) \right)
= A_1 \times \sqrt{n} (\widehat{\alpha} - \alpha_0, \widehat{\beta}_k - \beta_{0k})' + o_p(1) \xrightarrow{d} N(0, \lim_{n \to \infty} A_1 B A_1'),$$
(2.23)

where $A_1 = \left(-\frac{1}{n} \operatorname{tr}(e^{-\alpha_0 W} W \beta_{0k}), \frac{1}{n} \operatorname{tr}(e^{-\alpha_0 W})\right)$, B is the asymptotic covariance of $\sqrt{n}(\widehat{\alpha} - \alpha_0, \widehat{\beta}_k - \beta_{0k})$. So the asymptotic variance of direct effects can be estimated by $\frac{1}{n} \widehat{A}_1 \widehat{B} \widehat{A}_1'$, where $\widehat{A}_1 = \left(-\frac{1}{n} \operatorname{tr}(e^{-\widehat{\alpha} W} W \widehat{\beta}_k), \frac{1}{n} \operatorname{tr}(e^{-\widehat{\alpha} W})\right)$, and \widehat{B} is the estimated asymptotic covariance of $\sqrt{n}(\widehat{\alpha} - \alpha_0, \widehat{\beta}_k - \beta_{0k})$. Applying the mean value theorem to the estimator of total effect $\frac{1}{n} \widehat{\beta}_k l_n' e^{-\widehat{\alpha} W} l_n$, where l_n is

the $n \times 1$ vector of ones, we obtain

$$\frac{1}{\sqrt{n}} \left(\widehat{\beta}_k l'_n e^{-\widehat{\alpha}W} l_n - \beta_{0k} l'_n e^{-\alpha_0 W} l_n \right)
= A_2 \times \sqrt{n} (\widehat{\alpha} - \alpha_0, \widehat{\beta}_k - \beta_{0k})' + o_p(1) \xrightarrow{d} N(0, \lim_{n \to \infty} A_2 B A'_2), \tag{2.24}$$

where $A_2 = \left(-\frac{1}{n}\beta_k l_n' e^{-\alpha_0 W}W l_n, \frac{1}{n}l_n' e^{-\alpha_0 W}l_n\right)$. Thus, $\operatorname{Var}(\frac{1}{n}\widehat{\beta}_k l_n' e^{-\widehat{\alpha}W}l_n)$ can be estimated by $\frac{1}{n}\widehat{A}_2\widehat{B}\widehat{A}_2'$, where $\widehat{A}_2 = \left(-\frac{1}{n}\widehat{\beta}_k l_n' e^{-\widehat{\alpha}W}W l_n, \frac{1}{n}l_n' e^{-\widehat{\alpha}W}l_n\right)$. Finally, applying the mean value theorem to the estimator of average indirect effects $\frac{1}{n}\left(\widehat{\beta}_k l_n' e^{-\widehat{\alpha}W}l_n - \operatorname{tr}(e^{-\widehat{\alpha}W}\widehat{\beta}_k)\right)$, we obtain

$$\frac{1}{\sqrt{n}} \left(\left(\widehat{\beta}_k l_n' e^{-\widehat{\alpha}W} l_n - \operatorname{tr}(e^{-\widehat{\alpha}W} \widehat{\beta}_k) \right) - \left(\beta_{0k} l_n' e^{-\widehat{\alpha}_0 W} l_n - \operatorname{tr}(e^{-\widehat{\alpha}_0 W} \beta_{0k}) \right) \right) =$$

$$(A_2 - A_1) \times \sqrt{n} (\widehat{\alpha} - \alpha_0, \widehat{\beta}_k - \beta_{0k})' + o_p(1) \xrightarrow{d} N(0, \lim_{n \to \infty} (A_2 - A_1) B(A_2 - A_1)'). \tag{2.25}$$

Then, an estimate of $\operatorname{Var}\left(\frac{1}{n}\left(\widehat{\beta}_k l_n' e^{-\widehat{\alpha}W} l_n - \operatorname{tr}(e^{-\widehat{\alpha}W}\widehat{\beta}_k)\right)\right)$ is given by $\frac{1}{n}(\widehat{A}_2 - \widehat{A}_1)\widehat{B}(\widehat{A}_2 - \widehat{A}_1)'$.

3 The matrix-vector products method

Our analysis in Section 2 indicates that the estimation of MESS(1,1) specifically requires the evaluation of $e^{\tau M}e^{\alpha W}y$ and $e^{\tau M}X$. In the case of the QMLE, the objective function in (2.6) is comprised of $e^{\tau M}e^{\alpha W}y$ and the term $H(\tau)$, which is a function of $e^{\tau M}X$. For the BGMME in (2.11), the vector of best moment functions $g^*(\gamma)$ contains the disturbances $\epsilon(\gamma)$, which is a function of $e^{\tau M}e^{\alpha W}y$ and $e^{\tau M}X$. Our Algorithm 1 indicates that the Bayesian estimator also requires the evaluation of $e^{\tau M}e^{\alpha W}y$ and $e^{\tau M}X$ in each pass through the Gibbs sampler. In this section, we show how the matrix-vector products approach can be used to compute $e^{\tau M}e^{\alpha W}y$ and $e^{\tau M}X$ based on the truncation of Taylor series expansion of matrix exponential terms.

We start with $e^{\tau M}e^{\alpha W}y$. By definition, we have $e^{\tau M}e^{\alpha W}y=\sum_{i=0}^{\infty}\frac{\tau^iM^i}{i!}\sum_{j=0}^{\infty}\frac{\alpha^jW^j}{j!}y$. Truncating the Taylor series at the (q+1)th order yields $e^{\tau M}e^{\alpha W}y\approx\sum_{i=0}^{q}\frac{\tau^iM^i}{i!}\sum_{j=0}^{q}\frac{\alpha^jW^j}{j!}y$. Note that we can express $\sum_{i=0}^{q}\frac{\tau^iM^i}{i!}\sum_{j=0}^{q}\frac{\alpha^jW^j}{j!}$ as

$$\sum_{i=0}^{q} \frac{\tau^{i} M^{i}}{i!} \sum_{j=0}^{q} \frac{\alpha^{j} W^{j}}{j!} = \sum_{i=1}^{q} \sum_{j=0}^{i-1} \left(\frac{\tau^{i} \alpha^{j} M^{i} W^{j}}{i! j!} + \frac{\tau^{j} \alpha^{i} M^{j} W^{i}}{i! j!} \right) + \sum_{i=0}^{q} \frac{\tau^{i} \alpha^{i} M^{i} W^{i}}{(i!)^{2}}.$$
 (3.1)

Let Diag (a_1, a_2, \ldots, a_n) be the $n \times n$ diagonal matrix with diagonal elements $\{a_1, a_2, \ldots, a_n\}$. Then, using (3.1), we can express $e^{\tau M} e^{\alpha W} y$ in the following way,

$$e^{\tau M} e^{\alpha W} y \approx \sum_{i=1}^{q} \sum_{j=0}^{i-1} \frac{\tau^{i} \alpha^{j} M^{i} W^{j}}{i! j!} y + \sum_{i=1}^{q} \sum_{j=0}^{i-1} \frac{\tau^{j} \alpha^{i} M^{j} W^{i}}{i! j!} y + \sum_{i=0}^{q} \frac{\tau^{i} \alpha^{i} M^{i} W^{i}}{(i!)^{2}} y$$

$$= Y_{1} D_{1} \nu_{1}(\alpha, \tau) + Y_{2} D_{2} \nu_{2}(\alpha, \tau) + Y_{3} D_{3} \nu_{3}(\alpha, \tau), \tag{3.2}$$

where Y_1 and Y_2 are $n \times \frac{q(q+1)}{2}$ matrices, Y_3 is an $n \times (q+1)$ matrix, D_1 and D_2 are $\frac{q(q+1)}{2} \times \frac{q(q+1)}{2}$ matrices, D_3 is an $(q+1) \times (q+1)$ matrix, $\nu_1(\alpha,\tau)$ and $\nu_2(\alpha,\tau)$ are $\frac{q(q+1)}{2} \times 1$ vectors, and $\nu_3(\alpha,\tau)$ is an $(q+1) \times 1$ vector. The terms in (3.2) are

$$Y_1 = [My, M^2y, M^2Wy, M^3y, M^3Wy, M^3W^2y, \dots, M^qy, M^qWy, \dots, M^qW^{q-1}y], \qquad (3.3)$$

$$Y_2 = [Wy, W^2y, MW^2y, W^3y, MW^3y, M^2W^3y, \dots, W^qy, MW^qy, \dots, M^{q-1}W^qy],$$
(3.4)

$$Y_3 = \left[y, MWy, M^2W^2y, \dots, M^qW^qy \right], \tag{3.5}$$

$$D_1 = D_2 = \operatorname{Diag}\left(\frac{1}{0!1!}, \frac{1}{0!2!}, \frac{1}{1!2!}, \dots, \frac{1}{0!q!}, \dots, \frac{1}{(q-1)!q!}\right), \tag{3.6}$$

$$D_3 = \operatorname{Diag}\left(\frac{1}{(0!)^2}, \frac{1}{(1!)^2}, \frac{1}{(2!)^2}, \dots, \frac{1}{(q!)^2}\right),\tag{3.7}$$

$$\nu_1(\alpha, \tau) = \left[\tau, \tau^2, \tau^2 \alpha, \tau^3, \tau^3 \alpha, \tau^3 \alpha^2, \dots, \tau^q, \tau^q \alpha, \dots, \tau^q \alpha^{q-1}\right]',\tag{3.8}$$

$$\nu_2(\alpha, \tau) = \left[\alpha, \alpha^2, \alpha^2 \tau, \alpha^3, \alpha^3 \tau, \alpha^3 \tau^2, \dots, \alpha^q, \alpha^q \tau, \dots, \alpha^q \tau^{q-1}\right]',\tag{3.9}$$

$$\nu_3(\alpha, \tau) = \left[1, \tau\alpha, \tau^2\alpha^2, \tau^3\alpha^3, \dots, \tau^q\alpha^q\right]'. \tag{3.10}$$

Next, we show how the matrix-vector products approach can be used to get an approximation of $e^{\tau M}X$. Let $X = [X_1, X_2, \dots, X_k]$, where X_i is the *i*th column. Then, we can write $e^{\tau M}X$ as

$$e^{\tau M} X = \left[e^{\tau M} X_1, e^{\tau M} X_2, \dots, e^{\tau M} X_k \right] \approx \mathbb{X} D_4 \mu(\tau), \tag{3.11}$$

where \mathbb{X} is an $n \times k(q+1)$ matrix, D_4 is an $k(q+1) \times k(q+1)$ matrix and $\mu(\tau)$ is an $k(q+1) \times k$ matrix. It can be shown that

$$\mathbb{X} = [X_1, MX_1, M^2X_1, \dots, M^qX_1, X_2, MX_2, M^2X_2, \dots, M^qX_2, \dots, X_k, MX_k, M^2X_k, \dots, M^qX_k],$$
(3.12)

$$D_4 = I_k \otimes \text{Diag}\left(\frac{1}{0!}, \frac{1}{1!}, \frac{1}{2!}, \dots, \frac{1}{q!}\right),$$
 (3.13)

$$\mu(\tau) = I_k \otimes \left[1, \tau, \tau^2, \dots, \tau^q\right]'. \tag{3.14}$$

Next, we discuss how the matrix-vector products approach can be applied to the estimators given in Section 2. Recall that the QMLE is defined by $\hat{\psi} = \operatorname{argmin}_{\psi} \left(y^{'} e^{\alpha W^{'}} e^{\tau M^{'}} H(\tau) e^{\tau M} e^{\alpha W} y \right)$, where the objective function requires the evaluation of $e^{\tau M} e^{\alpha W} y$ and $e^{\tau M} X$ in each iteration. Using the

matrix-vector products approach, we can avoid the evaluation of these terms in each iteration of the optimization routine. We can define \mathbb{X} , Y_i , and D_j for $i \in \{1, 2, 3\}$ and $j \in \{1, 2, 3, 4\}$, and then supply these terms as the inputs of the objective function in the optimization solver. In this way, we can avoid the computation of these terms in each iteration.

In the case of GMM approach, the vector of moment functions in (2.10) contains the disturbance terms $\epsilon(\gamma) = e^{\tau M} e^{\alpha W} y - e^{\tau M} X \beta$, which can also be expressed in the matrix-vector products approach by using (3.2) and (3.11). When implementing (2.11), similar to the case of QMLE, we can define \mathbb{X} , Y_i , and D_j for $i \in \{1, 2, 3\}$ and $j \in \{1, 2, 3, 4\}$, and then supply these terms as inputs for the objective function to the optimization solver.

Finally, in the case of Bayesian approach, we need to work with expressions in Algorithm 1. Before implementing the Gibbs sampler described in Algorithm 1, we can compute the required terms for the matrix exponential terms and then pass these terms to the sampler. That is, we can define \mathbb{X} , Y_i , and D_j for $i \in \{1, 2, 3\}$ and $j \in \{1, 2, 3, 4\}$, and then supply these terms to the Gibbs sampler. Thus, we can avoid computation of these terms in each pass of the Gibbs sampler.

Remark 2. The matrix-vector products approach is general enough and can be easily adjusted for some other MESS models. For example, the expressions for two special cases, MESS(1,0) and MESS(0,1), can be simply obtained by setting M and W to the zero matrix, respectively. Similarly, the Durbin versions of MESS(1,1), MESS(1,0) and MESS(0,1) can also be estimated by defining independent variables appropriately. The matrix-vector products approach also applies to the MESS models with the panel data by arranging the terms involving the matrix exponential terms appropriately. For higher order MESS models, or MESS(p,q), similar equations to (3.1)-(3.14) can be derived. For example, for the MESS(2,2), we have

$$e^{\tau_{1}M_{1}+\tau_{2}M_{2}}e^{\alpha_{1}W_{1}+\alpha_{2}W_{2}}y \approx \sum_{i=0}^{q} \frac{1}{i!} (\tau_{1}M_{1}+\tau_{2}M_{2})^{i} \sum_{j=0}^{q} \frac{1}{j!} (\alpha_{1}W_{1}+\alpha_{2}W_{2})^{j} y$$

$$= \sum_{i=1}^{q} \sum_{j=0}^{i-1} \left(\frac{1}{i!j!} \sum_{k_{1}=0}^{i} \sum_{k_{2}=0}^{j} {i \choose k_{1}} {j \choose k_{2}} \tau_{1}^{i-k_{1}} \tau_{2}^{k_{1}} \alpha_{1}^{j-k_{2}} \alpha_{2}^{k_{2}} M_{1}^{i-k_{1}} M_{2}^{k_{1}} W_{1}^{j-k_{2}} W_{2}^{k_{2}} \right)$$

$$+ \frac{1}{i!j!} \sum_{k_{1}=0}^{j} \sum_{k_{2}=0}^{i} {j \choose k_{1}} {i \choose k_{2}} \tau_{1}^{j-k_{1}} \tau_{2}^{k_{1}} \alpha_{1}^{i-k_{2}} \alpha_{2}^{k_{2}} M_{1}^{j-k_{1}} M_{2}^{k_{1}} W_{1}^{i-k_{2}} W_{2}^{k_{2}}$$

$$+ \sum_{i=0}^{q} \sum_{k_{1}=0}^{i} \sum_{k_{2}=0}^{i} {i \choose k_{1}} {i \choose k_{2}} \tau_{1}^{i-k_{1}} \tau_{2}^{k_{1}} \alpha_{1}^{i-k_{2}} \alpha_{2}^{k_{2}} M_{1}^{i-k_{1}} M_{2}^{k_{1}} W_{1}^{i-k_{2}} W_{2}^{k_{2}}$$

$$+ \sum_{i=0}^{q} \sum_{k_{1}=0}^{i} \sum_{k_{2}=0}^{i} {i \choose k_{1}} {i \choose k_{2}} \tau_{1}^{i-k_{1}} \tau_{2}^{k_{1}} \alpha_{1}^{i-k_{2}} \alpha_{2}^{k_{2}} M_{1}^{i-k_{1}} M_{2}^{k_{1}} W_{1}^{i-k_{2}} W_{2}^{k_{2}}$$

$$= \mathbb{Y}_{1} \mathbb{D}_{1} \omega_{1}(\alpha, \tau) + \mathbb{Y}_{2} \mathbb{D}_{2} \omega_{2}(\alpha, \tau) + \mathbb{Y}_{3} \mathbb{D}_{3} \omega_{3}(\alpha, \tau),$$

$$(3.15)$$

where \mathbb{Y}_l , \mathbb{D}_l and ω_l for $l \in \{1, 2, 3\}$ can be defined accordingly. In the case of $e^{\tau_1 M_1 + \tau_2 M_2} X$, we

have

$$e^{\tau_{1}M_{1}+\tau_{2}M_{2}}X \approx \left[\sum_{i=0}^{q} \frac{1}{i!} (\tau_{1}M_{1}+\tau_{2}M_{2})^{i}X_{1}, \dots, \sum_{i=0}^{q} \frac{1}{i!} (\tau_{1}M_{1}+\tau_{2}M_{2})^{i}X_{k}\right]$$

$$= \left[\sum_{i=0}^{q} \frac{1}{i!} \sum_{j=0}^{i} {i \choose j} (\tau_{1}M_{1})^{i-j} (\tau_{2}M_{2})^{j}X_{1}, \dots, \sum_{i=0}^{q} \frac{1}{i!} \sum_{j=0}^{i} {i \choose j} (\tau_{1}M_{1})^{i-j} (\tau_{2}M_{2})^{j}X_{k}\right]$$

$$= \mathbb{X} \, \mathbb{D}_{4}\kappa(\tau), \tag{3.16}$$

where

$$\mathbb{X} = \left[X_1, M_1 X_1, M_2 X_1, M_1^2 X_1, M_1 M_2 X_1, M_2^2 X_1, \dots, M_1^q X_1, M_1^{q-1} M_2 X_1, \dots M_1 M_2^{q-1} X_1, M_2^q X_1, \dots, M_1^q X_1, M_2^q X_1, \dots, M_1^q X$$

$$X_k, M_1 X_k, M_2 X_k, M_1^2 X_k, M_1 M_2 X_k, M_2^2 X_k, \dots, M_1^q X_k, M_1^{q-1} M_2 X_k, \dots, M_1 M_2^{q-1} X_k, M_2^q X_k$$
,

$$\mathbb{D}_4 = I_k \otimes Diag\left(\frac{\binom{0}{0}}{0!}, \frac{\binom{1}{0}}{1!}, \frac{\binom{1}{1}}{1!}, \dots, \frac{\binom{q}{0}}{q!}, \frac{\binom{q}{1}}{q!}, \dots, \frac{\binom{q}{q}}{q!}\right),$$

$$\kappa(\tau) = I_k \otimes \left[1, \tau_1, \tau_2, \tau_1^2, \tau_1 \tau_2, \tau_2^2, \dots, \tau_1^q, \dots, \tau_1^{q-1} \tau_2, \dots, \tau_1 \tau_2^{q-1}, \tau_2^q\right]'.$$

Remark 3. The matrix-vector products approach based on a large q value can provide a better accuracy, but it can increase the computation time. We may determine a satisfactory q value from the inverse error analysis for a scaling and squaring method with Taylor series approximation (the third method in Moler and Van Loan (1978, 2003)). Let $T_q(\alpha W) = \sum_{j=0}^q \frac{\alpha^j W^j}{j!}$. Since $e^{\alpha W} = \left(e^{2^{-m}\alpha W}\right)^{2^m}$, we may consider the approximation $\left(T_q(2^{-m}\alpha W)\right)^{2^m}$. By Corollary 1 in Moler and Van Loan (1978, 2003), if $\|2^{-m}\alpha W\|_{\infty} \leq 0.5$, then $\left(T_q(2^{-m}\alpha W)\right)^{2^m} = e^{\alpha W + E}$, where

$$\frac{\|E\|_{\infty}}{\|\alpha W\|_{\infty}} \le \left(\frac{1}{2}\right)^{q-3} \frac{1}{q+1} \approx \begin{cases} 1.5 \times 10^{-5} & if \quad q = 15\\ 3.6 \times 10^{-7} & if \quad q = 20\\ 2.4 \times 10^{-10} & if \quad q = 30 \end{cases}.$$

This inverse error analysis indicates that a value of q = 15 can be satisfactory.

Remark 4. In terms of reliability, stability and accuracy, researchers should bear in mind the choice of q in the Taylor series expansion of the matrix exponential terms. Remark 3 indicates that q=15 can be satisfactory for the estimation of MESS-type models in economics, which is also confirmed by the results from our simulations and the empirical illustration for the MESS(1,1) model. In terms of efficiency and storage requirements, the representation in the form of equation

(3.2) requires calculating and storing terms in the form of equations (3.3)–(3.10). Here the largest of these terms are (3.3) and (3.4), which are of size nq(q+1)/2. Note also that in (3.3)–(3.5) terms are computed sequentially such that in each pass a matrix-vector product is calculated. Therefore, (3.3) and (3.4) are computed in $O(n^2q(q+1)/2)$ operations for dense W and M. In terms of simplicity, note again that the representation in the form of equation (3.2) needs to be derived and the terms in this representation need to be defined before the estimation. This can be tedious for higher order MESS-type models, but we think that the computational advantage of the matrix-vector products method outweighs this cost.

4 Monte Carlo Simulations

In this section, we will investigate the implications of using the matrix-vector products method with the truncated Taylor series approximation versus the scaling and squaring algorithm with the Padé approximation (the expm function in MATLAB R2020b). To this end, we will explore the properties of the two competing methods in terms of computation time as well as their effects on the finite sample properties of the estimators described in Section 2.

We consider the following data generating process,

$$e^{\alpha W}y = \beta_1 X_1 + \beta_2 X_2 + u, \quad e^{\tau W}u = \epsilon,$$
 (4.1)

where the elements of X_1 and X_2 are independently drawn from $U(0, \sqrt{12})$ and N(0, 1), respectively. For the spatial weights matrix $W = (w_{ij})$, we consider two cases, the rook contiguity and queen contiguity. To this end, n spatial units are randomly allocated into $\sqrt{n} \times \sqrt{n}$ square lattice graph. In the rook contiguity case, $w_{ij} = 1$ if the j'th observation is adjacent (left/right/above or below) to the i'th observation on the graph. In the queen contiguity case, $w_{ij} = 1$ if the j'th observation is adjacent to, or shares a border with the i'th observation. The weights matrices are then row normalized. We set $(\beta_1, \beta_2)' = (2, 1)'$, and let α and τ take values from $\{-2, -0.2, 0.2, 2\}$. The disturbance terms are generated according to $\epsilon_i \sim \text{i.i.d.} N(0, 1)$. We consider two sample sizes, n = 169 and n = 361. For the QMLE and GMME, we set the number of repetitions to 1000. In the case of Bayesian estimator, we choose the following the priors: $\alpha \sim N(0, 10)$, $\tau \sim N(0, 10)$, $\beta \sim N(0_{2\times 1}, I_2)$, and $\sigma^2 \sim IG(6/2, 4/2)$. We set the number of repetitions to 100, the number of draws to 1500 and burn-ins to 500. We use the matrix-vector products approach with q = 15 in all cases.

We use the matrix-vector products method (denoted as mvp) and the scaling and squaring algorithm with the Padé approximation (denoted as expm) to obtain the parameter estimates, and the corresponding bias, root mean squared error (RMSE) and the coverage rate. We also compute the impact measures, which include the average direct effect, the average total effect and the average indirect effect, and their respective bias, RMSE and empirical coverage. We report the total computation time in seconds over 1000 resamples.⁹ In the case of the QMLE, the computation

 $^{^9\}mathrm{We}$ use a MacBook Pro 2016 with 2.4GHz intel core i7 processor and 8 GB 1867 MHz LPDDR3 memory to run

time for $\widehat{\psi} = (\widehat{\alpha}, \widehat{\tau})'$ includes the time to compute the estimates using the concentrated log-likelihood function in (2.5). The computation time for $\widehat{\beta}$ includes the time for $\widehat{\psi}$ and for $\widehat{\beta}$, which is computed using (2.3). The time for impact measures are consequently the sum of the computation time for $\widehat{\psi}$, $\widehat{\beta}$ and respective measures. In the case of BGMME, an initial GMM estimation is carried out to construct V^* , F^* and P^* 's. We use the following set of moments for the initial stage, $(W\epsilon(\gamma), M\epsilon(\gamma), WX, X)'\epsilon(\gamma)$. Thus, the computation time for the BGMME includes both stages for the estimation. The time for the impact measures are then computed by adding on corresponding computation time for impact measures. In the case of Bayesian estimation, the computation time includes the time for collecting 2000 draws including the burn-ins. Finally, the computation time for the impact measures are computed similarly to those in the cases of the QMLE and GMME.

We focus on the simulation results provided in Tables 1–5. Tables 1–3 report the simulation results for the QMLE case. The mvp method reduces the computation time by approximately 98% to 99% compared to the expm method for different sample sizes, while providing the same estimates for the parameters and the impact measures. In all cases, we obtain the same values for bias, RMSE and coverage rates under both methods. When n = 169, the computation time for the myp approach is about 2% of the computation time for the expm method. For example, when $\alpha = -2$ and $\tau = -2$, the bias, RMSE and coverage of $\hat{\alpha}$ using both methods are 0.003, 0.037 and 0.923, respectively. However, the computation time is 606.5 seconds for the expm method, and 11.8 seconds for the myp method. This means that on average, each computation takes 0.6065 seconds using the expm method, and 0.0118 seconds using the mvp method. For n = 361, the bias, RMSE and coverage of $\hat{\alpha}$ are again the same for both methods, but the computation time is 4168.4 seconds for the expm method, and 30.7 seconds for the myp method, leading to an average running time of 4.1684 and 0.0307 seconds, respectively. These results show the matrix-vector products method reduces the computational burden significantly, while maintaining the finite sample properties of the estimators. Table 2 presents the results for $\hat{\beta} = (\hat{\beta}_1, \hat{\beta}_2)'$. The findings are similar to those from Table 1. While providing a similar performance in term of the finite sample properties of the QMLE, the mvp method reduces the computation time by about 98% to 99% compared to the scaling and squaring algorithm. Table 3 presents the results for the average direct effect estimates for X_1 and X_2 . There is no difference between the two competing methods in terms of bias, RMSE and coverage, but we again observe that the mvp method is computationally more efficient than the expm method.

Table 4 presents the results for $\hat{\alpha}$ and $\hat{\tau}$ in the GMM estimation case. The findings are very similar to the findings in the QML estimation case. The mvp method reduces the computation time by about 95% for n=169 and 97% for n=361 compared to the expm method. For example, when n=169, the computation time for $\hat{\gamma}$ is 1407.9 seconds using the expm method, and 60.6 seconds using the mvp method. It means that on average, one resample takes 1.4079 seconds to run for the expm method, and 0.0606 seconds to run for the mvp method. When n=361, they are

our simulations.

¹⁰Some additional results are given in Tables A.1–A.10 of Appendix A. The simulation results in these additional tables also attest that our suggested mvp method is computationally more efficient than the expm method.

respectively 8663.3 seconds and 219.6 seconds, leading to an average time of 8.6633 seconds and 0.2196 seconds to run each resample.

Table 5 presents the results for $\hat{\alpha}$ and $\hat{\tau}$ in the Bayesian estimation case. The mvp method again provides the same computational performance over the expm method, while maintaining the same finite sample properties. It takes only about 1% when n = 169 and less than 1% when n = 361 of the total computation time for the expm method. For example, for $\alpha = -2$, $\tau = -2$ and n = 169, it takes 2779.4 seconds for the expm method, and 20.4 seconds for the mvp method to compute $\hat{\alpha}$. Thus, on average it takes respectively 27.794 seconds and 0.204 seconds to collect draws for one resample. When n = 361, the computation time for $\hat{\alpha}$ using the expm method increases by approximately sevenfold to 19741.3 seconds. But, the computation time for the mvp method only increases by about 20% to 25.9 seconds. Thus, it takes an average of 197.413 seconds to collect draws for each resample using the expm method, and 0.259 seconds using the mvp method.

Table 1: The QMLE results for $\widehat{\alpha}$ and $\widehat{\tau}$

					Ta	ble 1: L	The QIV	ILE res	sults for	lpha and $ au$							
		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\alpha}$																	
$\tau = -2$	Bias RMSE Coverage	.003 .037 .923	.003 .034 .934	.000 .032 .939	.002 .035 .945	.003 .037 .923	.003 .034 .934	.000 .032 .939	.002 .035 .945	001 .025 .950	.000 .023 .954	.000 .023 .957	.000 .024 .938	001 .025 .950	.000 .023 .954	.000 .023 .957	.000 .024 .938
$\tau = -0.2$	Time Bias RMSE Coverage Time	.000 .058 .927 637.1	.004 .053 .945 553.0	.001 .055 .943 572.2	.004 .060 .936 661.2	.000 .058 .927 14.0	.004 .053 .945 12.6	.001 .055 .943 12.9	.004 .060 .936 14.3	.000 .041 .937 4149.6	.001 .039 .945 3547.5	.000 .037 .944 3670.0	.000 .038 .956 4225.0	.000 .041 .937 37.2	.001 .039 .945 38.3	.000 .037 .944 39.8	.000 .038 .956 37.8
τ =0.2	Bias RMSE Coverage Time	.002 .055 .931 591.1	.002 .052 .934 534.6	.002 .053 .947 549.6	.002 .063 .931 624.3	.002 .055 .931 12.9	.002 .052 .934 12.2	.002 .053 .947 12.4	.002 .063 .931 13.4	.001 .040 .946 3838.9	001 .040 .956 3454.7	.001 .041 .925 3498.7	001 .038 .951 4008.7	.001 .040 .946 34.1	001 .040 .956 37.2	.001 .041 .925 37.7	001 .038 .951 35.6
τ=2	Bias RMSE Coverage Time	001 .024 .905 632.0	.000 .022 .935 651.5	.001 .027 .936 612.5	001 .029 .924 643.7	001 .024 .905 11.8	.000 .022 .935 12.6	.001 .027 .936 11.7	001 .029 .924 12.0	.000 .017 .934 4355.0	001 .018 .942 4512.8	.000 .021 .942 4270.1	001 .016 .942 4512.5	.000 .017 .934 31.3	001 .018 .942 40.3	.000 .021 .942 38.3	001 .016 .942 32.3
Results for $\hat{\tau}$																	
$\tau=-2$	Bias RMSE Coverage Time	012 .146 .948 606.5	003 .145 .940 607.9	008 .147 .941 591.1	011 .147 .945 619.5	012 .146 .948 11.8	003 .145 .940 11.9	008 .147 .941 11.6	011 .147 .945 11.6	006 .104 .940 4168.4	007 .102 .953 4264.6	007 .102 .945 4146.2	008 .101 .944 4228.9	006 .104 .940 30.7	007 .102 .953 38.2	007 .102 .945 37.2	008 .101 .944 30.8
τ =-0.2	Bias RMSE Coverage Time	002 .152 .944 637.1	.014 .149 .938 553.0	.004 .144 .947 572.2	.015 .151 .927 661.2	002 .152 .944 14.0	.014 .149 .938 12.6	.004 .144 .947 12.9	.015 .151 .927 14.3	.003 .102 .938 4149.6	.001 .103 .937 3547.5	.005 .102 .948 3670.0	.001 .101 .948 4225.0	.003 .102 .938 37.2	.001 .103 .937 38.3	.005 .102 .948 39.8	.001 .101 .948 37.8
τ =0.2	Bias RMSE Coverage Time	.010 .146 .937 591.1	.015 .149 .948 534.6	.014 .157 .929 549.6	.011 .154 .938 624.3	.010 .146 .937 12.9	.015 .149 .948 12.2	.014 .157 .929 12.4	.011 .154 .938 13.4	.011 .101 .951 3838.9	.007 .104 .939 3454.7	.015 .101 .940 3498.7	.004 .105 .948 4008.7	.011 .101 .951 34.1	.007 .104 .939 37.2	.015 .101 .940 37.7	.004 .105 .948 35.6
τ =2	Bias RMSE Coverage Time	.066 .167 .904 632.0	.059 .165 .908 651.5	.057 .170 .900 612.5	.055 .161 .916 643.7	.066 .167 .904 11.8	.059 .165 .908 12.6	.057 .170 .900 11.7	.055 .161 .916 12.0	.028 .107 .934 4355.0	.032 .106 .936 4512.8	.024 .106 .936 4270.1	.026 .105 .941 4512.5	.028 .107 .934 31.3	.032 .106 .936 40.3	.024 .106 .936 38.3	.026 .105 .941 32.3

					Table	C 4. 11.	ie Qivit	in resu.	ρ								
		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\boldsymbol{\beta}}_1$																	
$\tau = -2$	Bias	.004	.003	004	002	.004	.003	004	002	.001	.003	.001	.000	.001	.003	.001	.000
	RMSE	.059	.060	.056	.057	.059	.060	.056	.057	.039	.038	.036	.035	.039	.038	.036	.035
	Coverage	.942	.937	.949	.943	.942	.937	.949	.943	.952	.944	.946	.952	.952	.944	.946	.952
	Time	610.3	611.2	594.5	623.2	12.0	12.1	11.7	11.8	4193.3	4287.2	4168.7	4253.5	31.1	38.7	37.7	31.2
$\tau = -0.2$	Bias	.002	.002	002	001	.002	.002	002	001	002	.000	004	.001	002	.000	004	.001
	RMSE	.079	.076	.074	.083	.079	.076	.074	.083	.051	.052	.054	.053	.051	.052	.054	.053
	Coverage	.937	.939	.955	.936	.937	.939	.955	.936	.947	.935	.931	.941	.947	.935	.931	.941
	Time	640.5	556.1	575.4	664.7	14.1	12.7	13.0	14.5	4171.9	3567.4	3689.9	4247.2	37.6	38.7	40.2	38.2
τ =0.2	Bias	001	001	002	003	001	001	002	003	.002	.000	.000	.000	.002	.000	.000	.000
	RMSE	.084	.075	.079	.080	.084	.075	.079	.080	.053	.053	.054	.055	.053	.053	.054	.055
	Coverage	.938	.934	.942	.946	.938	.934	.942	.946	.951	.948	.936	.946	.951	.948	.936	.946
	Time	594.5	537.7	552.7	627.8	13.1	12.4	12.6	13.6	3861.0	3474.6	3518.6	4030.8	34.5	37.7	38.1	36.0
$\tau=2$	Bias	.002	.000	002	.000	.002	.000	002	.000	.001	002	002	.000	.001	002	002	.000
	RMSE	.036	.039	.036	.045	.036	.039	.036	.045	.028	.025	.029	.026	.028	.025	.029	.026
	Coverage	.921	.926	.940	.946	.921	.926	.940	.946	.943	.943	.944	.940	.943	.943	.944	.940
	Time	635.6	654.9	615.9	647.4	12.0	12.8	11.9	12.1	4379.6	4535.5	4292.8	4537.3	31.7	40.8	38.7	32.7
Results for $\widehat{\boldsymbol{\beta}}_2$																	
$\tau = -2$	Bias	.001	.000	001	.002	.001	.000	001	.002	.001	001	002	.000	.001	001	002	.000
	RMSE	.056	.051	.048	.060	.056	.051	.048	.060	.039	.037	.037	.041	.039	.037	.037	.041
	Coverage	.946	.955	.934	.946	.946	.955	.934	.946	.948	.957	.945	.937	.948	.957	.945	.937
	Time	610.3	611.2	594.5	623.2	12.0	12.1	11.7	11.8	4193.3	4287.2	4168.7	4253.5	31.1	38.7	37.7	31.2
$\tau = -0.2$	Bias	.000	.000	002	001	.000	.000	002	001	.002	002	004	.000	.002	002	004	.000
	RMSE	.080	.075	.076	.075	.080	.075	.076	.075	.054	.052	.051	.051	.054	.052	.051	.051
	Coverage	.928	.938	.955	.953	.928	.938	.955	.953	.942	.949	.941	.945	.942	.949	.941	.945
	Time	640.5	556.1	575.4	664.7	14.1	12.7	13.0	14.5	4171.9	3567.4	3689.9	4247.2	37.6	38.7	40.2	38.2
$\tau = 0.2$	Bias	.001	.002	005	003	.001	.002	005	003	.001	.002	.001	.002	.001	.002	.001	.002
	RMSE	.077	.086	.079	.090	.077	.086	.079	.090	.053	.053	.052	.052	.053	.053	.052	.052
	Coverage	.937	.929	.937	.944	.937	.929	.937	.944	.955	.941	.943	.946	.955	.941	.943	.946
	Time	594.5	537.7	552.7	627.8	13.1	12.4	12.6	13.6	3861.0	3474.6	3518.6	4030.8	34.5	37.7	38.1	36.0
$\tau=2$	Bias	.000	.001	001	001	.000	.001	001	001	001	.002	.000	.000	001	.002	.000	.000
	RMSE	.034	.034	.042	.036	.034	.034	.042	.036	.022	.025	.024	.027	.022	.025	.024	.027
	Coverage	.933	.941	.941	.926	.933	.941	.941	.926	.946	.941	.941	.931	.946	.941	.941	.931
	Time	635.6	654.9	615.9	647.4	12.0	12.8	11.9	12.1	4379.6	4535.5	4292.8	4537.3	31.7	40.8	38.7	32.7

Table 3: The QMLE results of average direct effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau = -2$	Bias	.003	.003	004	.000	.003	.003	004	.000	.003	.003	.001	001	.003	.003	.001	001
	RMSE	.106	.060	.056	.109	.106	.060	.056	.109	.076	.039	.037	.070	.076	.039	.037	.070
	Coverage	.947	.937	.952	.945	.947	.937	.952	.945	.946	.941	.944	.946	.946	.941	.944	.946
	Time	612.0	612.6	595.9	625.0	13.8	13.5	13.2	13.6	4205.5	4297.0	4178.6	4265.6	43.4	48.5	47.5	43.3
$\tau = -0.2$	Bias	.005	.003	001	.007	.005	.003	001	.007	001	.000	004	.003	001	.000	004	.003
	RMSE	.159	.077	.075	.153	.159	.077	.075	.153	.105	.052	.054	.101	.105	.052	.054	.101
	Coverage	.948	.941	.961	.937	.948	.941	.961	.937	.954	.939	.932	.940	.954	.939	.932	.940
	Time	642.2	557.5	576.8	666.5	15.9	14.2	14.5	16.3	4184.3	3577.2	3699.8	4259.6	50.0	48.6	50.1	50.6
τ =0.2	Bias	003	.000	001	.001	003	.000	001	.001	.003	.000	.000	.000	.003	.000	.000	.000
	RMSE	.163	.074	.080	.163	.163	.074	.080	.163	.106	.053	.055	.104	.106	.053	.055	.104
	Coverage	.940	.937	.942	.944	.940	.937	.942	.944	.949	.944	.941	.940	.949	.944	.941	.940
	Time	596.2	539.1	554.1	629.5	14.8	13.8	14.1	15.4	3873.4	3484.5	3528.4	4043.1	46.8	47.5	48.0	48.3
$\tau=2$	Bias	.005	.000	001	.000	.005	.000	001	.000	.003	002	002	.000	.003	002	002	.000
	RMSE	.079	.039	.037	.084	.079	.039	.037	.084	.055	.025	.030	.058	.055	.025	.030	.058
	Coverage	.922	.922	.946	.947	.922	.922	.946	.947	.930	.945	.939	.941	.930	.945	.939	.941
	Time	637.3	656.3	617.3	649.2	13.7	14.3	13.3	13.9	4391.6	4545.3	4302.7	4549.4	43.6	50.6	48.6	44.8
Results for X_2																	
$\tau = -2$	Bias	.000	.000	001	.006	.000	.000	001	.006	.002	001	002	.000	.002	001	002	.000
	RMSE	.096	.051	.048	.107	.096	.051	.048	.107	.068	.037	.038	.072	.068	.037	.038	.072
	Coverage	.942	.955	.936	.953	.942	.955	.936	.953	.943	.956	.945	.935	.943	.956	.945	.935
	Time	612.0	612.7	595.9	625.0	13.8	13.5	13.1	13.6	4205.5	4297.0	4178.6	4265.6	43.4	48.5	47.5	43.3
τ =-0.2	Bias	.001	.000	002	.003	.001	.000	002	.003	.004	002	004	.000	.004	002	004	.000
	RMSE	.144	.076	.076	.126	.144	.076	.076	.126	.095	.053	.051	.088	.095	.053	.051	.088
	Coverage	.934	.938	.952	.950	.934	.938	.952	.950	.948	.948	.939	.933	.948	.948	.939	.933
	Time	642.2	557.5	576.8	666.5	15.9	14.2	14.5	16.2	4184.3	3577.2	3699.8	4259.6	50.0	48.6	50.1	50.6
τ =0.2	Bias	.000	.002	005	002	.000	.002	005	002	.001	.002	.001	.003	.001	.002	.001	.003
	RMSE	.132	.086	.080	.156	.132	.086	.080	.156	.092	.053	.052	.089	.092	.053	.052	.089
	Coverage	.938	.930	.939	.940	.938	.930	.939	.940	.956	.942	.943	.952	.956	.942	.943	.952
	Time	596.2	539.1	554.1	629.5	14.8	13.8	14.1	15.4	3873.4	3484.5	3528.4	4043.1	46.9	47.5	48.0	48.3
$\tau=2$	Bias	.001	.001	001	003	.001	.001	001	003	001	.002	.000	001	001	.002	.000	001
	RMSE	.068	.034	.042	.061	.068	.034	.042	.061	.039	.025	.024	.045	.039	.025	.024	.045
	Coverage	.928	.937	.940	.939	.928	.937	.940	.939	.946	.941	.943	.936	.946	.941	.943	.936
	Time	637.4	656.3	617.3	649.2	13.7	14.3	13.3	13.9	4391.6	4545.3	4302.7	4549.4	43.7	50.6	48.6	44.7

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					Tabl	<u>e 4: Tl</u>	ne GMI	ME resu	ults for $\hat{\epsilon}$								
		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\alpha}$																	
$\tau = -2$	Bias	.002	.002	001	.001	.002	.002	001	.001	001	.000	.000	001	001	.000	.000	001
	RMSE	.039	.035	.034	.037	.039	.035	.034	.037	.026	.023	.024	.024	.026	.023	.024	.024
	Coverage	.919	.930	.916	.920	.919	.930	.916	.920	.939	.947	.946	.930	.939	.947	.946	.930
	Time	1407.9	1298.5	1257.3	1270.5	60.6	57.0	56.2	54.9	8663.3	8788.2	8576.1	8797.2	219.6	220.6	219.6	219.7
τ =-0.2	Bias	001	.002	.000	.002	001	.002	.000	.002	002	.001	001	001	002	.001	001	001
	RMSE	.058	.056	.057	.062	.058	.056	.057	.062	.042	.040	.038	.039	.042	.040	.038	.039
	Coverage	.924	.931	.925	.914	.924	.931	.925	.914	.934	.945	.936	.942	.934	.945	.936	.942
	Time	1512.6	1413.2	1378.5	1504.4	60.8	59.7	58.8	60.5	9794.5	8872.9	8787.4	9936.7	207.6	203.1	202.2	208.4
τ =0.2	Bias	.001	.001	.001	.000	.001	.001	.001	.000	.000	002	.000	002	.000	002	.000	002
	RMSE	.057	.055	.054	.065	.057	.055	.054	.065	.040	.041	.042	.038	.040	.041	.042	.038
	Coverage	.921	.920	.935	.911	.921	.920	.935	.911	.945	.943	.925	.942	.945	.943	.925	.942
	Time	1409.1	1312.1	1300.7	1409.1	58.5	57.5	57.2	58.6	9341.3	8504.4	8409.4	9463.1	205.9	201.4	200.1	205.8
τ =2	Bias	002	001	.000	002	002	001	.000	002	.000	001	001	001	.000	001	001	001
	RMSE	.025	.023	.029	.030	.025	.023	.029	.030	.017	.018	.021	.017	.017	.018	.021	.017
	Coverage	.902	.932	.923	.913	.902	.932	.923	.913	.933	.937	.940	.935	.933	.937	.940	.935
	Time	1420.2	1335.1	1340.9	1343.2	58.4	58.2	58.4	56.4	9321.8	9157.7	8956.1	9167.3	223.0	224.4	222.9	221.0
Results for $\hat{\tau}$																	
τ=-2	Bias	024	014	022	025	024	014	022	025	012	014	014	016	012	014	014	016
	RMSE	.151	.150	.152	.151	.151	.150	.152	.151	.105	.103	.105	.102	.105	.103	.105	.102
	Coverage	.941	.939	.930	.936	.941	.939	.930	.936	.937	.954	.934	.944	.937	.954	.934	.944
	Time	1407.9	1298.5	1257.3	1270.5	60.6	57.0	56.2	54.9	8663.3	8788.2	8576.1	8797.2	219.6	220.6	219.6	219.7
$\tau = -0.2$	Bias	015	.001	011	.001	015	.001	011	.001	003	007	002	008	003	007	002	008
	RMSE	.154	.152	.148	.154	.154	.152	.148	.154	.104	.105	.102	.103	.104	.105	.102	.103
	Coverage	.936	.937	.936	.926	.936	.937	.936	.926	.935	.932	.938	.947	.935	.932	.938	.947
	Time	1512.6	1413.2	1378.5	1504.4	60.8	59.7	58.8	60.5	9794.5	8872.9	8787.4	9936.7	207.6	203.1	202.2	208.4
τ=0.2	Bias	006	003	005	006	006	003	005	006	.003	.000	.007	004	.003	.000	.007	004
	RMSE	.151	.151	.161	.158	.151	.151	.161	.158	.104	.105	.103	.106	.104	.105	.103	.106
	Coverage	.929	.942	.912	.921	.929	.942	.912	.921	.946	.926	.935	.938	.946	.926	.935	.938
	Time	1409.1	1312.1	1300.7	1409.1	58.5	57.5	57.2	58.6	9341.3	8504.4	8409.4	9463.1	205.9	201.4	200.1	205.8
$\tau=2$	Bias	.039	.034	.030	.029	.039	.034	.030	.029	.016	.020	.011	.015	.016	.020	.011	.015
	RMSE	.164	.162	.167	.160	.164	.162	.167	.160	.107	.104	.105	.104	.107	.104	.105	.104
	Coverage	.907	.909	.898	.924	.907	.909	.898	.924	.935	.949	.940	.948	.935	.949	.940	.948
	Time	1420.2	1335.1	1340.9	1343.2	58.4	58.2	58.4	56.4	9321.8	9157.7	8956.1	9167.3	223.0	224.4	222.9	221.0

Table 5: The Bayesian results for $\widehat{\alpha}$ and $\widehat{\tau}$

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\alpha}$																	
$\tau = -2$	Bias	.006	004	.005	002	.006	004	.005	002	002	003	.001	.006	002	003	.001	.006
	RMSE	.036	.034	.038	.031	.036	.034	.038	.031	.028	.024	.022	.026	.028	.024	.022	.026
	Coverage	.960	.950	.940	.940	.960	.950	.940	.940	.950	.990	.980	.920	.950	.990	.980	.920
	Time	2779.4	2519.4	2522.4	2768.3	20.4	20.5	20.4	19.8	19741.3	17773.8	17742.6	19743.5	25.9	25.8	26.1	25.4
τ =-0.2	Bias	.001	.002	.002	004	.001	.002	.002	004	.012	.003	001	008	.012	.003	001	008
	RMSE	.059	.056	.050	.053	.059	.056	.050	.053	.039	.039	.035	.037	.039	.039	.035	.037
	Coverage	.950	.960	.950	.930	.950	.960	.950	.930	.900	.970	.960	.960	.900	.970	.960	.960
	Time	2569.1	2319.2	2319.2	2584.2	20.2	20.3	20.3	19.7	17645.8	15784.0	15748.3	17733.8	25.9	25.7	26.1	25.4
τ =0.2	Bias	.014	005	.003	008	.014	005	.003	008	.008	005	.001	002	.008	005	.001	002
	RMSE	.062	.051	.056	.065	.062	.051	.056	.065	.040	.041	.036	.042	.040	.041	.036	.042
	Coverage	.920	.970	.960	.940	.920	.970	.960	.940	.910	.950	.950	.970	.910	.950	.950	.970
	Time	2567.2	2333.8	2318.0	2570.9	20.3	20.0	20.1	20.1	17658.5	15786.3	15735.4	17797.9	26.1	25.4	25.8	25.2
τ =2	Bias	.000	.004	.002	002	.000	.004	.002	002	.000	.003	.000	002	.000	.003	.000	002
	RMSE	.025	.029	.035	.032	.025	.029	.035	.032	.023	.020	.020	.017	.023	.020	.020	.017
	Coverage	.960	.970	.960	.940	.960	.970	.960	.940	.930	.960	.940	.970	.930	.960	.940	.970
	Time	2771.2	2511.7	2526.1	2778.6	20.2	20.1	19.8	19.6	19692.2	17740.2	17745.2	19558.5	26.0	25.6	25.5	25.3
Results for $\hat{\tau}$																	
τ=-2	Bias	.055	.050	.038	.073	.055	.050	.038	.073	.010	.029	.030	.022	.010	.029	.030	.022
	RMSE	.168	.157	.142	.155	.168	.157	.142	.155	.106	.111	.107	.105	.106	.111	.107	.105
	Coverage	.930	.920	.960	.920	.930	.920	.960	.920	.930	.890	.920	.920	.930	.890	.920	.920
	Time	2768.4	2512.7	2515.5	2761.3	37.5	37.7	37.5	36.6	19666.9	17777.3	17735.6	19677.4	49.0	48.9	49.1	48.0
$\tau = -0.2$	Bias	.004	.022	.018	.002	.004	.022	.018	.002	.005	.015	.002	.004	.005	.015	.002	.004
	RMSE	.158	.134	.143	.154	.158	.134	.143	.154	.099	.110	.095	.108	.099	.110	.095	.108
	Coverage	.940	.970	.950	.930	.940	.970	.950	.930	.950	.930	.950	.910	.950	.930	.950	.910
	Time	2562.3	2314.1	2315.7	2576.4	37.2	37.5	37.4	36.6	17579.9	15779.5	15744.3	17671.9	48.9	48.6	49.0	48.0
τ =0.2	Bias	015	.012	003	.019	015	.012	003	.019	006	.000	.000	.006	006	.000	.000	.006
	RMSE	.165	.133	.141	.145	.165	.133	.141	.145	.094	.106	.104	.114	.094	.106	.104	.114
	Coverage	.870	.960	.950	.970	.870	.960	.950	.970	.960	.940	.920	.920	.960	.940	.920	.920
	Time	2560.7	2328.0	2313.5	2561.8	37.4	36.8	37.1	36.9	17602.7	15790.3	15729.3	17749.5	49.7	48.1	48.8	47.8
τ =2	Bias	042	047	035	055	042	047	035	055	016	013	016	014	016	013	016	014
	RMSE	.141	.155	.163	.164	.141	.155	.163	.164	.101	.103	.106	.102	.101	.103	.106	.102
	Coverage	.980	.960	.940	.960	.980	.960	.940	.960	.940	.980	.940	.970	.940	.980	.940	.970
	Time	2758.7	2504.3	2518.9	2769.1	37.5	36.9	36.7	36.3	19622.7	17732.9	17730.0	19463.7	49.0	48.4	48.2	47.8

5 An Empirical Illustration

In this section, we illustrate the computational time advantage of the matrix-vector products method using an empirical application. To this end, we use an example from Pace and Barry (1997) on the US presidential election in 1980. The dataset contains variables on the election results and county characteristics for 3107 US counties. In our model, the outcome variable is the (logged) proportion of voting age population that voted in the election (Y = Vote). The explanatory variables include the log percentage of population with a twelfth grade or higher education ($X_1 = \text{Educ}$), the log percentage of population with homeownership ($X_2 = \text{Homeowners}$), and the log per capita income ($X_3 = \text{Income}$). We consider the following MESS(1,1) specification

$$e^{\alpha W}Y = \beta_0 l_n + \beta_1 X_1 + \beta_2 X_2 + \beta_3 X_3 + u, \quad e^{\tau W}u = \epsilon.$$
 (5.1)

The spatial weights matrix W is the contiguity based weights matrix constructed using the latitude and longitude of the counties. We estimate this specification using the QML, GMM and Bayesian methods. For the Bayesian method, the same priors as those in the MC simulations are used. We use the expm and mvp methods to compute the estimation results and record the corresponding computation times. In the case of mvp method, the truncation order q is set to 15.

The parameter estimation results are summarized in Table 6. While the mvp method provides the same parameter estimates and standard errors (in parenthesis) as those obtained from the expm method, the computation time is significantly smaller for the mvp method. In the QML case, the computation time is 1072.4 seconds for the expm method, but only 6.0 seconds for the mvp method. In the GMM case, the computation time takes 4805.8 seconds for the expm method, and 21.1 seconds for the mvp method. For the Bayesian estimation, the number of draws is set to 1500 and the first 500 draws are discarded as burn-ins. The results show that the difference in terms of computation time is the biggest for the Bayesian estimation case, costing 47742.3 seconds (13.26 hours) for the expm method, and 11.4 seconds for the mvp method. Overall, these results show that the matrix-vector products method is not only useful in Monte Carlo simulations, where resamples are drawn for hundreds or thousands of times, but also useful in empirical applications with large sample sizes.

Table 6: The parameter estimates of the presidential election voting example

	Ql	ML	GN	ИМ	Baye	esian
	expm	mvp	expm	mvp	expm	mvp
α	-0.350***	-0.350***	-0.423***	-0.423***	-0.337***	-0.337***
	(0.045)	(0.045)	(0.045)	(0.045)	(0.041)	(0.041)
au	-0.443^{***}	-0.443^{***}	-0.374***	-0.374***	-0.458***	-0.458^{***}
	(0.055)	(0.055)	(0.055)	(0.055)	(0.050)	(0.050)
Educ	0.316***	0.316***	0.300***	0.300***	0.317^{***}	0.317^{***}
	(0.021)	(0.021)	(0.020)	(0.020)	(0.020)	(0.020)
Homeowners	0.572^{***}	0.572***	0.571^{***}	0.571^{***}	0.572***	0.572^{***}
	(0.016)	(0.016)	(0.016)	(0.016)	(0.016)	(0.016)
Income	-0.154***	-0.154***	-0.144***	-0.144***	-0.155***	-0.155^{***}
	(0.021)	(0.021)	(0.020)	(0.020)	(0.021)	(0.021)
Cons.	0.738***	0.738***	0.732^{***}	0.732^{***}	0.734^{***}	0.734^{***}
	(0.052)	(0.052)	(0.051)	(0.051)	(0.054)	(0.054)
Time	1072.4	6.0	4805.8	21.1	47742.3	11.4

Note: * p < 0.1; ** p < 0.05; *** p < 0.01.

To investigate the impact measures, we also compute the average direct effects, average total effects and average indirect effects discussed in Section 2.5, the measures of dispersion, and computation times. The results are summarized in Table 7. The reported dispersion measures are calculated by the delta method for the QML and GMM methods. In the case of Bayesian estimators, the standard deviation of posterior draws is used. The results show that the corresponding impact measures are very similar across the QML, GMM and Bayesian methods. For example, the average direct effect estimate for Educ is 0.320 using the QML method, 0.305 using the GMM method and 0.320 for the Bayesian method. However, in terms of computation time, the mvp method takes much less time than the expm method. For example, the computation times for the expm method are 1088.2, 4822.3 and 71584.0 seconds for the QML, GMM and Bayesian methods, respectively. On the other hand, the computation times for the mvp method are 22.0, 37.6 and 24491.0 seconds for the mvp method for the QML, GMM and Bayesian methods, respectively. ¹¹

¹¹Note that the impact measures consist of traces of $n \times n$ matrices. For example, the average direct effect equals $\frac{1}{n} \operatorname{tr}(e^{-\widehat{\alpha}W_n}\widehat{\beta}_k)$ for k=1,2 and 3, so we need to compute the trace of an $n \times n$ matrix for each draw through Algorithm 1 in the Bayesian method. This is why the time to compute the impact measures is relatively longer in the case of Bayesian method. However the time to sample the draws for the parameters is much shorter as shown in Table 6.

Table 7: The impact measures of the presidential election voting example

	Ql	ML	GN	ИΜ	Baye	esian
	expm	mvp	expm	mvp	expm	mvp
			Average di	rect effects		
Educ	0.320***	0.320***	0.305***	0.305***	0.320***	0.320***
	(0.020)	(0.020)	(0.020)	(0.020)	(0.020)	(0.020)
Homeowners	0.578***	0.578***	0.580***	0.580***	0.578***	0.578***
	(0.016)	(0.016)	(0.016)	(0.016)	(0.016)	(0.016)
Income	-0.156***	-0.156***	-0.147^{***}	-0.147^{***}	-0.156***	-0.156***
	(0.021)	(0.021)	(0.020)	(0.020)	(0.021)	(0.021)
			Average to	otal effects		
Educ	0.449***	0.449***	0.458***	0.458***	0.444***	0.444***
	(0.027)	(0.027)	(0.027)	(0.027)	(0.029)	(0.029)
Homeowners	0.812***	0.812***	0.872^{***}	0.872^{***}	0.802^{***}	0.802^{***}
	(0.043)	(0.043)	(0.044)	(0.044)	(0.039)	(0.039)
Income	-0.219***	-0.219***	-0.220***	-0.220***	-0.217^{***}	-0.217^{***}
	(0.028)	(0.028)	(0.030)	(0.030)	(0.029)	(0.029)
			Average inc	lirect effects		
Educ	0.129***	0.129***	0.153***	0.153***	0.124***	0.124***
	(0.017)	(0.017)	(0.017)	(0.017)	(0.018)	(0.018)
Homeowners	0.234***	0.234***	0.292^{***}	0.292***	0.224***	0.224***
	(0.036)	(0.036)	(0.038)	(0.038)	(0.032)	(0.032)
Income	-0.063^{***}	-0.063***	-0.074***	-0.074***	-0.060***	-0.060***
	(0.011)	(0.011)	(0.012)	(0.012)	(0.011)	(0.011)
Time	1088.2	22.0	4822.3	37.6	71584.0	24491.0

Note: * p<0.1; ** p<0.05; *** p<0.01.

6 Conclusion

The MESS-type models provide a class of alternatives to SAR-type models with attractive properties. However, the estimation of these models requires the computation of the matrix exponential terms in each iteration of a numerical optimization scheme, which can be computationally costly. In this paper, for the calculation of the matrix exponential terms, we propose a matrix-vector products method based on the truncation of Taylor series expansion of matrix exponential terms. Because the estimation of MESS-type models requires the computation of the matrix exponential terms as vectors, rather than the matrix exponential terms in isolation, our approach provides an efficient alternative to the default method available in several popular statistical software. The results from our extensive simulation study and empirical illustration confirm the computational time gains for three estimation methods for MESS-type models (i.e., the QML, GMM and Bayesian methods) using the matrix-vector products method.

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Appendix

A Additional Simulation Results

In this appendix, we provide some additional simulation results on the performance of both methods. Tables A.1 and A.2 include the simulation results for the indirect and total effects of X_1 and X_2 for the QMLE case. Tables A.3–A.6 provide additional results for the GMME. The remaining tables, Tables A.7–A.10, include additional simulation results for the Bayesian estimator. Overall, the simulation results in these tables attest that the mvp method is computationally more efficient than the expm method.

Table A.1: The QMLE results of average indirect effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau = -2$	Bias	006	005	.002	001	006	005	.002	001	.018	.000	.000	.001	.018	.000	.000	.001
	RMSE	.547	.079	.056	.107	.547	.079	.056	.107	.406	.054	.042	.071	.406	.054	.042	.071
	Coverage	.931	.943	.935	.945	.931	.943	.935	.945	.942	.951	.955	.943	.942	.951	.955	.943
	Time	613.8	614.1	597.3	626.7	15.5	15.0	14.6	15.3	4217.8	4306.7	4188.5	4277.7	55.7	58.2	57.3	55.3
$\tau = -0.2$	Bias	.029	005	.000	008	.029	005	.000	008	.007	002	.002	002	.007	002	.002	002
	RMSE	.875	.126	.098	.152	.875	.126	.098	.152	.615	.090	.065	.101	.615	.090	.065	.101
	Coverage	.932	.939	.941	.940	.932	.939	.941	.940	.949	.945	.945	.938	.949	.945	.945	.938
	Time	643.9	559.0	578.3	668.2	17.6	15.6	16.0	18.0	4196.7	3587.0	3709.6	4271.9	62.5	58.4	59.9	62.9
τ =0.2	Bias	012	002	002	002	012	002	002	002	.007	.004	.000	.000	.007	.004	.000	.000
	RMSE	.853	.120	.092	.166	.853	.120	.092	.166	.603	.092	.073	.104	.603	.092	.073	.104
	Coverage	.946	.942	.945	.939	.946	.942	.945	.939	.947	.953	.929	.938	.947	.953	.929	.938
	Time	598.0	540.6	555.6	631.3	16.5	15.2	15.5	17.2	3885.7	3494.3	3538.3	4055.5	59.1	57.4	57.9	60.6
$\tau=2$	Bias	.031	.001	001	.001	.031	.001	001	.001	.012	.001	.001	.001	.012	.001	.001	.001
	RMSE	.420	.049	.048	.083	.420	.049	.048	.083	.287	.041	.037	.058	.287	.041	.037	.058
	Coverage	.918	.936	.943	.942	.918	.936	.943	.942	.925	.943	.944	.942	.925	.943	.944	.942
	Time	639.1	657.8	618.8	650.9	15.4	15.7	14.8	15.6	4403.5	4555.0	4312.5	4561.4	55.6	60.3	58.4	56.8
Results for X_2																	
$\tau=-2$	Bias	008	002	.001	006	008	002	.001	006	.012	.000	.000	.000	.012	.000	.000	.000
	RMSE	.393	.042	.029	.101	.393	.042	.029	.101	.284	.029	.021	.068	.284	.029	.021	.068
	Coverage	.937	.931	.943	.928	.937	.931	.943	.928	.942	.954	.968	.896	.942	.954	.968	.896
	Time	613.8	614.1	597.3	626.7	15.5	15.0	14.6	15.3	4217.8	4306.7	4188.5	4277.7	55.6	58.2	57.3	55.4
$\tau = -0.2$	Bias	.009	003	.001	003	.009	003	.001	003	.020	001	.001	.000	.020	001	.001	.000
	RMSE	.615	.063	.048	.118	.615	.063	.048	.118	.415	.047	.033	.083	.415	.047	.033	.083
	Coverage	.922	.946	.954	.967	.922	.946	.954	.967	.959	.938	.949	.951	.959	.938	.949	.951
	Time	643.9	559.0	578.3	668.2	17.6	15.6	16.0	18.0	4196.7	3587.0	3709.6	4271.9	62.4	58.3	59.9	62.9
τ =0.2	Bias	003	.000	.000	.002	003	.000	.000	.002	.002	.002	.000	003	.002	.002	.000	003
	RMSE	.543	.063	.048	.147	.543	.063	.048	.147	.401	.047	.037	.084	.401	.047	.037	.084
	Coverage	.964	.933	.947	.917	.964	.933	.947	.917	.953	.958	.935	.952	.953	.958	.935	.952
	Time	598.0	540.6	555.6	631.3	16.5	15.2	15.5	17.1	3885.7	3494.2	3538.4	4055.5	59.1	57.4	57.9	60.6
$\tau=2$	Bias	.012	.001	.000	.003	.012	.001	.000	.003	001	.001	.000	.001	001	.001	.000	.001
	RMSE	.301	.024	.021	.058	.301	.024	.021	.058	.171	.021	.018	.042	.171	.021	.018	.042
	Coverage	.907	.944	.966	.981	.907	.944	.966	.981	.967	.917	.947	.958	.967	.917	.947	.958
	Time	639.1	657.7	618.8	650.9	15.4	15.7	14.8	15.6	4403.5	4555.1	4312.5	4561.4	55.5	60.4	58.4	56.8

Table A.2: The QMLE results of average total effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau=-2$	Bias	003	002	002	001	003	002	002	001	.021	.004	.000	.000	.021	.004	.000	.000
	RMSE	.643	.103	.072	.013	.643	.103	.072	.013	.477	.073	.047	.008	.477	.073	.047	.008
	Coverage	.938	.948	.936	.940	.938	.948	.936	.940	.941	.954	.945	.944	.941	.954	.945	.944
	Time	612.0	612.7	595.9	625.0	13.8	13.5	13.2	13.6	4205.5	4297.0	4178.6	4265.6	43.4	48.5	47.5	43.3
$\tau = -0.2$	Bias	.034	002	001	001	.034	002	001	001	.006	002	002	.000	.006	002	002	.000
	RMSE	1.021	.166	.107	.021	1.021	.166	.107	.021	.712	.108	.075	.013	.712	.108	.075	.013
	Coverage	.937	.934	.940	.934	.937	.934	.940	.934	.953	.950	.931	.952	.953	.950	.931	.952
	Time	642.2	557.5	576.8	666.5	15.9	14.2	14.5	16.2	4184.3	3577.2	3699.7	4259.5	50.0	48.5	50.0	50.6
τ =0.2	Bias	015	002	003	.000	015	002	003	.000	.010	.004	.000	.001	.010	.004	.000	.001
	RMSE	1.003	.142	.108	.020	1.003	.142	.108	.020	.700	.109	.079	.013	.700	.109	.079	.013
	Coverage	.945	.938	.936	.926	.945	.938	.936	.926	.944	.953	.926	.937	.944	.953	.926	.937
	Time	596.2	539.1	554.2	629.6	14.8	13.8	14.1	15.4	3873.4	3484.4	3528.5	4043.1	46.8	47.5	48.0	48.3
τ =2	Bias	.036	.001	002	.000	.036	.001	002	.000	.015	001	001	.000	.015	001	001	.000
	RMSE	.496	.061	.053	.010	.496	.061	.053	.010	.338	.046	.042	.004	.338	.046	.042	.004
	Coverage	.921	.942	.928	.931	.921	.942	.928	.931	.924	.945	.930	.940	.924	.945	.930	.940
	Time	637.3	656.3	617.3	649.2	13.7	14.3	13.3	13.9	4391.6	4545.3	4302.6	4549.4	43.6	50.6	48.6	44.8
Results for X_2																	
$\tau=-2$	Bias	009	003	.000	.000	009	003	.000	.000	.014	001	002	.000	.014	001	002	.000
	RMSE	.484	.076	.049	.009	.484	.076	.049	.009	.349	.055	.037	.006	.349	.055	.037	.006
	Coverage	.951	.947	.934	.938	.951	.947	.934	.938	.942	.948	.942	.942	.942	.948	.942	.942
	Time	612.0	612.7	595.9	624.9	13.7	13.5	13.1	13.6	4205.5	4297.0	4178.6	4265.6	43.4	48.5	47.5	43.2
$\tau = -0.2$	Bias	.010	003	001	.000	.010	003	001	.000	.024	003	002	.000	.024	003	002	.000
	RMSE	.752	.110	.081	.014	.752	.110	.081	.014	.504	.079	.051	.009	.504	.079	.051	.009
	Coverage	.930	.936	.944	.922	.930	.936	.944	.922	.945	.946	.935	.948	.945	.946	.935	.948
	Time	642.2	557.5	576.8	666.5	15.9	14.2	14.5	16.2	4184.3	3577.1	3699.7	4259.5	50.1	48.5	50.0	50.6
τ =0.2	Bias	003	.002	005	.000	003	.002	005	.000	.003	.004	.001	.000	.003	.004	.001	.000
	RMSE	.667	.118	.077	.015	.667	.118	.077	.015	.487	.076	.056	.009	.487	.076	.056	.009
	Coverage	.941	.935	.937	.936	.941	.935	.937	.936	.959	.951	.937	.946	.959	.951	.937	.946
	Time	596.2	539.1	554.1	629.6	14.8	13.8	14.0	15.4	3873.4	3484.4	3528.4	4043.1	46.8	47.5	48.0	48.3
τ =2	Bias	.014	.002	001	.000	.014	.002	001	.000	002	.003	.001	.000	002	.003	.001	.000
	RMSE	.367	.043	.049	.007	.367	.043	.049	.007	.208	.037	.027	.004	.208	.037	.027	.004
	Coverage	.922	.936	.940	.930	.922	.936	.940	.930	.946	.948	.949	.939	.946	.948	.949	.939
	Time	637.4	656.3	617.3	649.2	13.7	14.3	13.3	13.9	4391.6	4545.3	4302.6	4549.4	43.6	50.6	48.6	44.8

					Table A	1.3: Th	e GMN	1E resu	Its for β	$_1$ and β_2							
		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\boldsymbol{\beta}}_1$																	
$\tau = -2$	Bias	.002	.003	005	003	.002	.003	005	003	.000	.003	.000	001	.000	.003	.000	001
	RMSE	.060	.061	.057	.059	.060	.061	.057	.059	.040	.039	.037	.035	.040	.039	.037	.035
	Coverage	.930	.931	.935	.919	.930	.931	.935	.919	.950	.937	.939	.945	.950	.937	.939	.945
	Time	1407.9	1298.5	1257.3	1270.5	60.6	57.0	56.2	54.9	8663.3	8788.2	8576.1	8797.2	219.6	220.6	219.6	219.7
$\tau = -0.2$	Bias	001	001	004	004	001	001	004	004	003	002	006	.000	003	002	006	.000
	RMSE	.081	.078	.076	.086	.081	.078	.076	.086	.052	.053	.055	.053	.052	.053	.055	.053
	Coverage	.934	.931	.942	.916	.934	.931	.942	.916	.949	.936	.931	.938	.949	.936	.931	.938
	Time	1512.6	1413.2	1378.5	1504.4	60.8	59.7	58.8	60.5	9794.5	8872.9	8787.4	9936.7	207.6	203.1	202.2	208.4
τ =0.2	Bias	004	004	005	006	004	004	005	006	.000	002	002	001	.000	002	002	001
	RMSE	.086	.077	.082	.082	.086	.077	.082	.082	.054	.054	.054	.057	.054	.054	.054	.057
	Coverage	.925	.926	.924	.934	.925	.926	.924	.934	.950	.946	.934	.935	.950	.946	.934	.935
	Time	1409.1	1312.1	1300.7	1409.1	58.5	57.5	57.2	58.6	9341.3	8504.4	8409.4	9463.1	205.9	201.4	200.1	205.8
$\tau=2$	Bias	.001	.000	002	001	.001	.000	002	001	.001	002	002	.000	.001	002	002	.000
	RMSE	.037	.042	.038	.047	.037	.042	.038	.047	.029	.026	.030	.027	.029	.026	.030	.027
	Coverage	.909	.910	.932	.930	.909	.910	.932	.930	.940	.930	.938	.935	.940	.930	.938	.935
	Time	1420.2	1335.1	1340.9	1343.2	58.4	58.2	58.4	56.4	9321.8	9157.7	8956.1	9167.3	223.0	224.4	222.9	221.0
Results for $\widehat{\boldsymbol{\beta}}_2$																	
$\tau=-2$	Bias	.000	.000	001	.002	.000	.000	001	.002	.001	001	003	001	.001	001	003	001
	RMSE	.058	.052	.050	.062	.058	.052	.050	.062	.039	.038	.038	.042	.039	.038	.038	.042
	Coverage	.924	.947	.904	.926	.924	.947	.904	.926	.941	.950	.940	.930	.941	.950	.940	.930
	Time	1407.9	1298.5	1257.3	1270.5	60.6	57.0	56.2	54.9	8663.3	8788.2	8576.1	8797.2	219.6	220.6	219.6	219.7
$\tau = -0.2$	Bias	002	002	003	001	002	002	003	001	.001	003	004	001	.001	003	004	001
	RMSE	.083	.078	.078	.078	.083	.078	.078	.078	.055	.054	.051	.052	.055	.054	.051	.052
	Coverage	.915	.921	.940	.932	.915	.921	.940	.932	.935	.943	.937	.932	.935	.943	.937	.932
	Time	1512.6	1413.2	1378.5	1504.4	60.8	59.7	58.8	60.5	9794.5	8872.9	8787.4	9936.7	207.6	203.1	202.2	208.4
τ =0.2	Bias	.000	.001	007	006	.000	.001	007	006	001	.001	.001	.001	001	.001	.001	.001
	RMSE	.079	.089	.083	.094	.079	.089	.083	.094	.053	.054	.053	.052	.053	.054	.053	.052
	Coverage	.929	.914	.920	.919	.929	.914	.920	.919	.948	.930	.940	.938	.948	.930	.940	.938
	Time	1409.1	1312.1	1300.7	1409.1	58.5	57.5	57.2	58.6	9341.3	8504.4	8409.4	9463.1	205.9	201.4	200.1	205.8
τ =2	Bias	.000	.000	001	002	.000	.000	001	002	001	.002	.000	.000	001	.002	.000	.000
	RMSE	.035	.036	.044	.038	.035	.036	.044	.038	.022	.026	.024	.028	.022	.026	.024	.028
	Coverage	.916	.925	.924	.915	.916	.925	.924	.915	.947	.938	.938	.926	.947	.938	.938	.926
	Time	1420.2	1335.1	1340.9	1343.2	58.4	58.2	58.4	56.4	9321.8	9157.7	8956.1	9167.3	223.0	224.4	222.9	221.0

Table A.4: The GMME results of average direct effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau = -2$	Bias	.002	.003	005	004	.002	.003	005	004	.003	.003	.000	003	.003	.003	.000	003
	RMSE	.109	.061	.057	.113	.109	.061	.057	.113	.078	.039	.037	.071	.078	.039	.037	.071
	Coverage	.928	.935	.934	.918	.928	.935	.934	.918	.940	.934	.935	.944	.940	.934	.935	.944
	Time	1409.8	1299.9	1258.7	1272.2	62.5	58.4	57.7	56.7	8675.5	8798.0	8585.9	8809.3	231.8	230.4	229.4	231.8
$\tau = -0.2$	Bias	.003	.000	003	001	.003	.000	003	001	002	001	005	.000	002	001	005	.000
	RMSE	.162	.079	.077	.157	.162	.079	.077	.157	.107	.053	.055	.103	.107	.053	.055	.103
	Coverage	.932	.928	.941	.934	.932	.928	.941	.934	.949	.935	.930	.928	.949	.935	.930	.928
	Time	1514.4	1414.7	1380.0	1506.1	62.6	61.1	60.2	62.3	9807.0	8882.8	8797.4	9949.1	220.1	213.0	212.1	220.8
τ =0.2	Bias	005	003	004	008	005	003	004	008	.001	002	001	003	.001	002	001	003
	RMSE	.168	.076	.083	.168	.168	.076	.083	.168	.108	.053	.055	.107	.108	.053	.055	.107
	Coverage	.933	.925	.926	.925	.933	.925	.926	.925	.946	.943	.933	.932	.946	.943	.933	.932
	Time	1410.8	1313.5	1302.2	1410.8	60.2	58.9	58.7	60.3	9353.7	8514.4	8419.4	9475.4	218.2	211.3	210.1	218.1
$\tau=2$	Bias	.004	.000	001	004	.004	.000	001	004	.003	002	002	002	.003	002	002	002
	RMSE	.083	.041	.039	.087	.083	.041	.039	.087	.056	.026	.030	.059	.056	.026	.030	.059
	Coverage	.906	.910	.932	.931	.906	.910	.932	.931	.927	.933	.938	.932	.927	.933	.938	.932
	Time	1421.9	1336.5	1342.3	1344.9	60.1	59.6	59.9	58.1	9333.7	9167.5	8966.0	9179.2	235.0	234.2	232.7	232.8
Results for X_2																	
$\tau = -2$	Bias	001	.000	001	.005	001	.000	001	.005	.003	001	003	001	.003	001	003	001
	RMSE	.100	.053	.050	.111	.100	.053	.050	.111	.069	.038	.038	.073	.069	.038	.038	.073
	Coverage	.923	.946	.908	.930	.923	.946	.908	.930	.932	.947	.943	.930	.932	.947	.943	.930
	Time	1409.9	1299.9	1258.7	1272.2	62.5	58.4	57.6	56.7	8675.5	8798.1	8586.0	8809.2	231.9	230.4	229.4	231.8
τ =-0.2	Bias	001	002	003	.000	001	002	003	.000	.003	003	004	002	.003	003	004	002
	RMSE	.149	.079	.078	.130	.149	.079	.078	.130	.097	.054	.052	.090	.097	.054	.052	.090
	Coverage	.917	.922	.941	.934	.917	.922	.941	.934	.945	.943	.936	.927	.945	.943	.936	.927
	Time	1514.4	1414.7	1380.0	1506.1	62.6	61.1	60.2	62.3	9807.0	8882.8	8797.4	9949.1	220.1	213.1	212.1	220.8
τ =0.2	Bias	.000	.002	006	009	.000	.002	006	009	.000	.001	.001	.001	.000	.001	.001	.001
	RMSE	.136	.089	.083	.163	.136	.089	.083	.163	.094	.054	.053	.090	.094	.054	.053	.090
	Coverage	.931	.913	.923	.924	.931	.913	.923	.924	.952	.929	.940	.942	.952	.929	.940	.942
	Time	1410.8	1313.5	1302.2	1410.8	60.2	58.9	58.7	60.3	9353.7	8514.3	8419.3	9475.4	218.2	211.4	210.1	218.1
$\tau=2$	Bias	.002	.000	001	004	.002	.000	001	004	001	.002	.000	001	001	.002	.000	001
	RMSE	.070	.035	.043	.064	.070	.035	.043	.064	.039	.026	.024	.046	.039	.026	.024	.046
	Coverage	.919	.924	.925	.916	.919	.924	.925	.916	.943	.939	.940	.926	.943	.939	.940	.926
	Time	1421.9	1336.5	1342.3	1344.9	60.1	59.6	59.8	58.1	9333.7	9167.6	8966.0	9179.2	235.0	234.3	232.7	232.9

Table A.5: The GMME results of average indirect effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		au				au				au				au			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
tau=-2	Bias	.000	002	.004	.003	.000	002	.004	.003	.026	.002	.001	.003	.026	.002	.001	.003
	RMSE	.567	.081	.059	.112	.567	.081	.059	.112	.416	.056	.043	.071	.416	.056	.043	.071
	Coverage	.929	.931	.918	.916	.929	.931	.918	.916	.940	.949	.947	.941	.940	.949	.947	.941
	Time	1411.7	1301.4	1260.1	1273.9	64.6	59.8	59.1	58.4	8687.6	8807.8	8595.7	8821.3	244.0	240.1	239.2	243.8
tau=-0.2	Bias	.031	002	.002	.001	.031	002	.002	.001	.015	001	.004	.001	.015	001	.004	.001
	RMSE	.884	.131	.102	.157	.884	.131	.102	.157	.628	.091	.066	.103	.628	.091	.066	.103
	Coverage	.926	.924	.933	.937	.926	.924	.933	.937	.940	.943	.939	.930	.940	.943	.939	.930
	Time	1516.2	1416.1	1381.4	1507.8	64.3	62.6	61.7	64.0	9819.3	8892.7	8807.3	9961.4	232.5	222.9	222.0	233.0
tau=0.2	Bias	009	002	.000	.008	009	002	.000	.008	.010	.006	.001	.004	.010	.006	.001	.004
	RMSE	.883	.125	.094	.171	.883	.125	.094	.171	.610	.094	.074	.107	.610	.094	.074	.107
	Coverage	.930	.926	.936	.924	.930	.926	.936	.924	.941	.946	.924	.936	.941	.946	.924	.936
	Time	1412.5	1315.0	1303.6	1412.6	61.9	60.3	60.1	62.0	9365.9	8524.2	8429.3	9487.7	230.4	221.2	220.0	230.4
tau=2	Bias	.030	.002	.001	.005	.030	.002	.001	.005	.015	.002	.002	.002	.015	.002	.002	.002
	RMSE	.438	.052	.052	.086	.438	.052	.052	.086	.292	.041	.038	.059	.292	.041	.038	.059
	Coverage	.907	.930	.930	.929	.907	.930	.930	.929	.919	.936	.939	.932	.919	.936	.939	.932
	Time	1423.6	1337.9	1343.7	1346.6	61.7	61.1	61.3	59.8	9345.6	9177.3	8975.7	9191.0	246.9	244.0	242.4	244.7
Results for X_2																	
tau=-2	Bias	007	001	.002	005	007	001	.002	005	.017	.000	.001	.002	.017	.000	.001	.002
	RMSE	.411	.043	.031	.105	.411	.043	.031	.105	.290	.029	.022	.069	.290	.029	.022	.069
	Coverage	.911	.915	.923	.899	.911	.915	.923	.899	.932	.951	.961	.883	.932	.951	.961	.883
	Time	1411.7	1301.4	1260.1	1273.9	64.6	59.8	59.1	58.4	8687.5	8807.8	8595.7	8821.3	243.9	240.1	239.2	243.9
tau=-0.2	Bias	.010	002	.002	001	.010	002	.002	001	.022	001	.002	.003	.022	001	.002	.003
	RMSE	.632	.066	.050	.122	.632	.066	.050	.122	.425	.047	.034	.085	.425	.047	.034	.085
	Coverage	.907	.931	.939	.952	.907	.931	.939	.952	.945	.932	.936	.944	.945	.932	.936	.944
	Time	1516.1	1416.1	1381.4	1507.8	64.4	62.6	61.7	64.0	9819.4	8892.7	8807.3	9961.4	232.5	222.9	222.0	233.0
tau=0.2	Bias	.002	.000	.001	.008	.002	.000	.001	.008	.002	.003	.000	001	.002	.003	.000	001
	RMSE	.564	.066	.050	.154	.564	.066	.050	.154	.406	.048	.037	.085	.406	.048	.037	.085
	Coverage	.945	.915	.935	.887	.945	.915	.935	.887	.946	.953	.931	.945	.946	.953	.931	.945
	Time	1412.6	1315.0	1303.6	1412.6	61.9	60.3	60.1	62.0	9365.9	8524.3	8429.3	9487.7	230.4	221.3	220.0	230.4
tau=2	Bias	.017	.001	.001	.004	.017	.001	.001	.004	001	.002	.001	.001	001	.002	.001	.001
	RMSE	.310	.026	.023	.061	.310	.026	.023	.061	.174	.022	.019	.043	.174	.022	.019	.043
	Coverage	.899	.935	.961	.970	.899	.935	.961	.970	.961	.914	.941	.959	.961	.914	.941	.959
	Time	1423.6	1337.9	1343.7	1346.6	61.7	61.1	61.3	59.8	9345.6	9177.3	8975.7	9191.0	247.0	244.0	242.4	244.6

Table A.6: The GMME results of average total effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau = -2$	Bias	.002	.000	001	.000	.002	.000	001	.000	.030	.005	.001	.000	.030	.005	.001	.000
	RMSE	.665	.105	.075	.013	.665	.105	.075	.013	.488	.075	.048	.008	.488	.075	.048	.008
	Coverage	.931	.934	.923	.921	.931	.934	.923	.921	.939	.944	.938	.941	.939	.944	.938	.941
	Time	1409.9	1299.9	1258.7	1272.2	62.6	58.4	57.7	56.7	8675.5	8798.0	8585.9	8809.2	231.8	230.4	229.4	231.8
$\tau = -0.2$	Bias	.033	002	001	001	.033	002	001	001	.013	002	002	.001	.013	002	002	.001
	RMSE	1.033	.171	.109	.022	1.033	.171	.109	.022	.726	.109	.076	.013	.726	.109	.076	.013
	Coverage	.928	.922	.932	.909	.928	.922	.932	.909	.941	.944	.926	.940	.941	.944	.926	.940
	Time	1514.4	1414.7	1380.0	1506.1	62.6	61.1	60.2	62.3	9806.9	8882.8	8797.3	9949.1	220.1	213.0	212.1	220.7
τ =0.2	Bias	015	005	004	.000	015	005	004	.000	.012	.004	.000	.001	.012	.004	.000	.001
	RMSE	1.038	.147	.112	.021	1.038	.147	.112	.021	.710	.111	.080	.013	.710	.111	.080	.013
	Coverage	.930	.925	.926	.918	.930	.925	.926	.918	.942	.942	.924	.929	.942	.942	.924	.929
	Time	1410.8	1313.5	1302.2	1410.8	60.2	58.9	58.7	60.3	9353.6	8514.3	8419.3	9475.4	218.1	211.3	210.0	218.1
$\tau=2$	Bias	.034	.003	001	.000	.034	.003	001	.000	.018	.000	.000	.000	.018	.000	.000	.000
	RMSE	.517	.065	.056	.011	.517	.065	.056	.011	.344	.047	.043	.005	.344	.047	.043	.005
	Coverage	.904	.933	.922	.920	.904	.933	.922	.920	.922	.944	.929	.939	.922	.944	.929	.939
	Time	1421.9	1336.5	1342.3	1344.9	60.1	59.6	59.8	58.1	9333.7	9167.5	8966.0	9179.1	235.0	234.2	232.7	232.8
Results for X_2																	
$\tau = -2$	Bias	009	002	.001	.000	009	002	.001	.000	.020	.000	002	.000	.020	.000	002	.000
	RMSE	.506	.078	.051	.009	.506	.078	.051	.009	.356	.057	.038	.006	.356	.057	.038	.006
	Coverage	.930	.941	.917	.918	.930	.941	.917	.918	.934	.941	.934	.931	.934	.941	.934	.931
	Time	1409.8	1299.9	1258.7	1272.2	62.6	58.4	57.7	56.7	8675.4	8798.0	8585.9	8809.2	231.8	230.3	229.4	231.8
$\tau = -0.2$	Bias	.010	003	001	.000	.010	003	001	.000	.025	003	002	.000	.025	003	002	.000
	RMSE	.774	.115	.083	.015	.774	.115	.083	.015	.517	.081	.052	.009	.517	.081	.052	.009
	Coverage	.917	.926	.932	.917	.917	.926	.932	.917	.939	.933	.929	.941	.939	.933	.929	.941
	Time	1514.4	1414.7	1380.0	1506.1	62.6	61.1	60.2	62.3	9807.0	8882.8	8797.3	9949.1	220.1	213.0	212.1	220.7
τ =0.2	Bias	.001	.001	005	.000	.001	.001	005	.000	.001	.005	.001	.001	.001	.005	.001	.001
	RMSE	.691	.124	.079	.016	.691	.124	.079	.016	.494	.077	.056	.009	.494	.077	.056	.009
	Coverage	.920	.913	.917	.905	.920	.913	.917	.905	.954	.947	.930	.933	.954	.947	.930	.933
	Time	1410.8	1313.5	1302.2	1410.8	60.2	58.9	58.7	60.3	9353.6	8514.4	8419.3	9475.4	218.1	211.3	210.1	218.2
τ =2	Bias	.019	.001	.000	.000	.019	.001	.000	.000	002	.003	.001	.000	002	.003	.001	.000
	RMSE	.378	.045	.052	.007	.378	.045	.052	.007	.211	.038	.028	.005	.211	.038	.028	.005
	Coverage	.909	.928	.929	.910	.909	.928	.929	.910	.938	.942	.941	.929	.938	.942	.941	.929
	Time	1421.9	1336.5	1342.3	1344.9	60.0	59.6	59.8	58.1	9333.7	9167.5	8965.9	9179.1	235.0	234.1	232.6	232.8

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for $\widehat{\boldsymbol{\beta}}_1$																	
$\tau = -2$	Bias RMSE	002 .065	.001 .058	.001 .054	016 .057	002 .065	.001 .058	.001 .054	016 .057	007 .041	002 .034	004 .039	.010 .034	007 .041	002 .034	004 .039	.010 .034
	Coverage Time	.900 712.1	.960 657.7	.950 658.3	.970 710.1	.900 26.3	0.960 0.960	.950 27.2	.970 25.6	.960 4983.1	.970 4500.9	.950 4502.5	.960 4989.0	$.960 \\ 36.1$.970 35.1	$.950 \\ 36.2$	$.960 \\ 34.4$
$\tau = -0.2$	Bias RMSE Coverage	007 .071 .970	.000 .082 .940	.002 .082 .910	.002 .078 .920	007 .071 .970	.000 .082 .940	.002 .082 .910	.002 .078 .920	004 .060 .940	.004 .051 .950	002 .044 .980	.002 .056 .940	004 .060 .940	.004 .051 .950	002 .044	.002 .056 .940
	Time	660.6	604.6	605.1	664.7	26.0	27.1	26.9	25.6	4459.9	4004.1	4004.5	4487.9	36.1	34.3	004 .039 .950 36.2002 .044 .980 35.5 .008 .057 .950 34.8001 .030 .950 33.8008 .038 .970 36.2001 .054 .910 35.5 .004 .050 .960 34.8002 .023 .950	34.1
τ =0.2	Bias RMSE Coverage Time	002 .071 .990 659.9	005 .073 .960 608.4	.011 .076 .940 604.9	.005 .080 .950 661.6	002 .071 .990 26.2	005 .073 .960 26.3	.011 .076 .940 26.9	.005 .080 .950 26.6	004 .056 .940 4469.0	008 .056 .970 4010.0	.008 .057 .950 4004.6	007 .055 .950 4509.8	004 .056 .940 36.9	008 .056 .970 33.2	.057 .950	007 .055 .950 34.1
τ =2	Bias RMSE Coverage Time	.005 .045 .920 710.2	001 .039 .950 656.8	005 .049 .960 660.0	.005 .033 .930 713.0	.005 .045 .920 26.4	001 .039 .950 26.8	005 .049 .960 26.6	.005 .033 .930 25.4	.000 .024 .970 4972.3	.001 .032 .940 4496.2	001 .030 .950 4508.4	.002 .030 .940 4937.4	.000 .024 .970 36.2	.001 .032 .940 34.1	.030 .950	.002 .030 .940 34.2
Results for $\widehat{\beta}_2$																	
$\tau = -2$	Bias RMSE Coverage Time	009 .054 .940 712.1	.002 .058 .930 657.7	.011 .053 .940 658.3	014 .064 .910 710.1	009 .054 .940 26.3	.002 .058 .930 27.3	.011 .053 .940 27.2	014 .064 .910 25.6	.006 .040 .960 4983.1	001 .038 .950 4500.9	008 .038 .970 4502.5	003 .039 .920 4989.0	.006 .040 .960 36.1	001 .038 .950 35.1	.038 .970	003 .039 .920 34.4
τ =-0.2	Bias RMSE Coverage Time	010 .078 .940 660.6	.006 .088 .920 604.6	.005 .075 .930 605.1	002 .071 .960 664.7	010 .078 .940 26.0	.006 .088 .920 27.1	.005 .075 .930 26.9	002 .071 .960 25.6	.010 .048 .950 4459.9	.006 .045 .960 4004.1	001 .054 .910 4004.5	.010 .055 .950 4487.9	.010 .048 .950 36.1	.006 .045 .960 34.3	.054 .910	.010 .055 .950 34.1
τ =0.2	Bias RMSE Coverage Time	.007 .080 .970 659.9	.008 .074 .940 608.4	.006 .093 .900 604.9	014 .077 .940 661.6	.007 .080 .970 26.2	.008 .074 .940 26.3	.006 .093 .900 26.9	014 .077 .940 26.6	004 .052 .950 4469.0	.009 .054 .960 4010.0	.004 .050 .960 4004.6	007 .053 .950 4509.8	004 .052 .950 36.9	.009 .054 .960 33.2	.050 .960	007 .053 .950 34.1
τ =2	Bias RMSE Coverage Time	003 .040 .950 710.2	.010 .040 .920 656.8	.001 .032 .980 660.0	.003 .038 .950 713.0	003 .040 .950 26.4	.010 .040 .920 26.8	.001 .032 .980 26.6	.003 .038 .950 25.4	.004 .029 .960 4972.3	001 .030 .950 4496.2	002 .023 .950 4508.4	.003 .023 .940 4937.4	.004 .029 .960 36.2	001 .030 .950 34.1	.023	.003 .023 .940 34.2

Table A.8: The Bayesian results of average direct effects for X_1 and X_2

		n=169								n=361								
		expm				mvp				expm				mvp				
		α				α				α				α				
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	
Results for X_1																		
$\tau=-2$	Bias	011	.002	.002	028	011	.002	.002	028	008	002	003	.028	008	002	003	.028	
	RMSE	.113	.058	.055	.099	.113	.058	.055	.099	.074	.035	.039	.080	.074	.035	.039	.080	
	Coverage	.940	.950	.940	.970	.940	.950	.940	.970	.940	.960	.950	.940	.940	.960	.950	.940	
	Time	6623.1	5988.2	5995.2	6603.0	446.7	394.2	395.1	441.1	46853.9	42049.8	41972.0	46883.2	2521.7	2046.1	2060.1	2511.4	
τ =-0.2	Bias	010	.002	.004	.000	010	.002	.004	.000	023	.004	001	006	023	.004	001	006	
	RMSE	.143	.082	.084	.148	.143	.082	.084	.148	.121	.052	.044	.107	.121	.052	.044	.107	
	Coverage	.970	.950	.910	.950	.970	.950	.910	.950	.920	.950	.980	.980	.920	.950	.980	.980	
	Time	6159.4	5540.2	5542.9	6194.0	451.9	392.3	393.6	445.3	42115.0	37546.8	37469.3	42343.4	2513.3	2007.1	2016.9	2508.7	
τ =0.2	Bias	022	003	.013	.001	022	003	.013	.001	017	007	.009	012	017	007	.009	012	
	RMSE	.132	.073	.078	.174	.132	.073	.078	.174	.108	.055	.058	.110	.108	.055	.058	.110	
	Coverage	.970	.970	.950	.960	.970	.970	.950	.960	.970	.970	.950	.940	.970	.970	.950	.940	
	Time	6153.8	5574.3	5539.1	6161.3	450.2	388.4	393.2	450.4	42169.3	37567.6	37441.1	42523.0	2534.6	1993.2	2012.0	2520.8	
$\tau=2$	Bias	.009	001	004	.005	.009	001	004	.005	.000	.001	.000	001	.000	.001	.000	001	
	RMSE	.074	.039	.049	.083	.074	.039	.049	.083	.057	.032	.031	.051	.057	.032	.031	.051	
	Coverage	.950	.940	.950	.950	.950	.940	.950	.950	.970	.950	.970	.960	.970	.950	.970	.960	
	Time	6600.2	5968.9	6003.7	6623.9	446.6	386.7	386.6	440.4	46740.4	41953.4	41968.3	46378.7	2507.7	2005.1	2014.0	2459.6	
Results for X_2																		
$\tau=-2$	Bias	019	.002	.012	025	019	.002	.012	025	.012	001	008	.000	.012	001	008	.000	
	RMSE	.092	.059	.053	.106	.092	.059	.053	.106	.068	.039	.038	.065	.068	.039	.038	.065	
	Coverage	.950	.930	.930	.930	.950	.930	.930	.930	.960	.950	.970	.940	.960	.950	.970	.940	
	Time	6622.9	5988.2	5995.2	6602.9	441.1	388.9	389.8	435.5	46852.5	42047.8	41973.9	46884.1	2520.9	2044.7	2058.7	2509.7	
$\tau = -0.2$	Bias	016	.006	.006	005	016	.006	.006	005	.009	.006	001	.011	.009	.006	001	.011	
	RMSE	.137	.088	.076	.124	.137	.088	.076	.124	.087	.045	.054	.098	.087	.045	.054	.098	
	Coverage	.950	.930	.930	.940	.950	.930	.930	.940	.940	.970	.910	.950	.940	.970	.910	.950	
	Time	6158.9	5540.2	5542.7	6193.4	446.3	386.5	388.8	440.0	42113.2	37544.2	37471.8	42345.5	2512.1	2003.5	2012.7	2507.8	
τ =0.2	Bias	.001	.009	.007	027	.001	.009	.007	027	012	.010	.004	012	012	.010	.004	012	
	RMSE	.136	.075	.094	.143	.136	.075	.094	.143	.092	.054	.050	.093	.092	.054	.050	.093	
	Coverage	.950	.940	.900	.960	.950	.940	.900	.960	.960	.970	.960	.940	.960	.970	.960	.940	
	Time	6153.1	5574.5	5539.2	6160.9	444.8	382.6	387.8	443.6	42167.6	37565.5	37444.5	42525.4	2533.2	1989.7	2008.4	2519.7	
$\tau=2$	Bias	006	.010	.001	.004	006	.010	.001	.004	.007	001	002	.004	.007	001	002	.004	
	RMSE	.069	.040	.033	.073	.069	.040	.033	.073	.054	.030	.023	.042	.054	.030	.023	.042	
	Coverage	.940	.940	.970	.930	.940	.940	.970	.930	.950	.960	.970	.960	.950	.960	.970	.960	
	Time	6599.9	5969.2	6004.0	6623.8	441.5	381.5	381.1	435.7	46739.1	41951.6	41971.0	46380.8	2506.3	2003.2	2011.9	2458.5	

Table A.9: The Bayesian results of average indirect effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau = -2$	Bias	071	.010	008	.027	071	.010	008	.027	009	.007	001	028	009	.007	001	028
	RMSE	.546	.078	.067	.095	.546	.078	.067	.095	.401	.059	.040	.082	.401	.059	.040	.082
	Coverage	.940	.920	.950	.960	.940	.920	.950	.960	.970	.980	.960	.930	.970	.980	.960	.930
	Time	6968.0	6273.2	6280.8	6948.5	785.4	679.4	681.6	776.6	49273.4	43999.5	43930.4	49319.1	4908.2	3958.2	3985.9	4891.4
$\tau = -0.2$	Bias	018	.001	002	.002	018	.001	002	.002	161	004	.003	.009	161	004	.003	.009
	RMSE	.836	.132	.092	.147	.836	.132	.092	.147	.640	.092	.060	.107	.640	.092	.060	.107
	Coverage	.940	.960	.950	.940	.940	.960	.950	.940	.930	.950	.960	1.000	.930	.950	.960	1.000
	Time	6508.1	5827.4	5830.7	6544.0	796.4	675.1	680.5	785.2	44502.3	39476.8	39410.1	44758.0	4892.8	3883.1	3899.8	4887.4
τ =0.2	Bias	157	.017	004	.003	157	.017	004	.003	112	.012	002	.012	112	.012	002	.012
	RMSE	.798	.120	.101	.177	.798	.120	.101	.177	.595	.094	.066	.112	.595	.094	.066	.112
	Coverage	.960	.970	.960	.940	.960	.970	.960	.940	.930	.970	.970	.940	.930	.970	.970	.940
	Time	6500.7	5863.8	5827.2	6510.1	793.1	669.1	678.7	790.7	44566.5	39503.8	39383.6	44952.3	4933.2	3857.0	3890.5	4913.1
$\tau=2$	Bias	.030	008	001	003	.030	008	001	003	.004	007	.000	.002	.004	007	.000	.002
	RMSE	.352	.068	.060	.086	.352	.068	.060	.086	.355	.046	.036	.050	.355	.046	.036	.050
	Coverage	.950	.980	.940	.940	.950	.980	.940	.940	.930	.980	.950	.960	.930	.980	.950	.960
	Time	6942.3	6252.7	6289.8	6969.0	786.1	665.9	666.2	777.4	49150.4	43891.8	43925.0	48762.6	4880.2	3879.9	3897.7	4789.5
Results for X_2																	
$\tau = -2$	Bias	079	.006	005	.023	079	.006	005	.023	.052	.004	.001	001	.052	.004	.001	001
	RMSE	.370	.043	.033	.098	.370	.043	.033	.098	.284	.031	.021	.061	.284	.031	.021	.061
	Coverage	.960	.950	.950	.930	.960	.950	.950	.930	.960	.990	.950	.940	.960	.990	.950	.940
	Time	6968.5	6275.2	6282.5	6949.3	785.6	679.6	681.8	776.6	49275.7	44000.9	43933.2	49321.9	4911.6	3960.8	3987.6	4893.6
$\tau = -0.2$	Bias	047	.000	002	.006	047	.000	002	.006	009	002	.001	009	009	002	.001	009
	RMSE	.587	.063	.046	.118	.587	.063	.046	.118	.390	.046	.032	.092	.390	.046	.032	.092
	Coverage	.940	.970	.950	.950	.940	.970	.950	.950	.930	.960	.930	.960	.930	.960	.930	.960
	Time	6508.4	5828.7	5832.1	6544.5	796.7	675.4	680.9	785.2	44504.6	39479.7	39412.0	44761.5	4896.4	3882.8	3900.2	4889.8
τ =0.2	Bias	037	.012	001	.027	037	.012	001	.027	068	.009	.000	.011	068	.009	.000	.011
	RMSE	.572	.067	.051	.137	.572	.067	.051	.137	.404	.050	.033	.089	.404	.050	.033	.089
	Coverage	.910	.970	.940	.960	.910	.970	.940	.960	.960	.960	.950	.930	.960	.960	.950	.930
	Time	6501.3	5865.0	5828.6	6510.6	793.2	669.3	678.6	792.4	44568.6	39506.3	39385.6	44956.1	4935.0	3857.2	3890.8	4914.4
$\tau=2$	Bias	019	002	002	003	019	002	002	003	.024	004	.000	003	.024	004	.000	003
	RMSE	.273	.033	.033	.071	.273	.033	.033	.071	.243	.024	.018	.041	.243	.024	.018	.041
	Coverage	.910	.960	.950	.930	.910	.960	.950	.930	.940	.930	.970	.960	.940	.930	.970	.960
	Time	6942.9	6254.7	6291.6	6970.0	786.4	666.2	666.2	777.5	49152.6	43894.2	43927.7	48764.1	4882.3	3881.5	3899.1	4792.2

Table A.10: The Bayesian results of average total effects for X_1 and X_2

		n=169								n=361							
		expm				mvp				expm				mvp			
		α				α				α				α			
		-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2	-2	-0.2	0.2	2
Results for X_1																	
$\tau=-2$	Bias	082	.011	006	001	082	.011	006	001	018	.005	004	.000	018	.005	004	.000
	RMSE	.647	.103	.073	.012	.647	.103	.073	.012	.467	.081	.047	.007	.467	.081	.047	.007
	Coverage	.940	.920	.970	.930	.940	.920	.970	.930	.970	.940	.980	.950	.970	.940	.980	.950
	Time	6622.8	5988.8	5996.0	6602.9	440.6	388.3	389.4	435.0	46847.3	42042.2	41970.5	46880.4	2516.1	2040.8	2054.6	2505.6
$\tau = -0.2$	Bias	029	.003	.002	.002	029	.003	.002	.002	185	.000	.002	.003	185	.000	.002	.003
	RMSE	.965	.168	.096	.019	.965	.168	.096	.019	.753	.117	.071	.013	.753	.117	.071	.013
	Coverage	.960	.950	.980	.940	.960	.950	.980	.940	.930	.930	.980	.960	.930	.930	.980	.960
	Time	6158.9	5540.6	5543.4	6193.5	445.9	386.0	388.3	439.3	42109.2	37538.8	37467.6	42341.2	2507.6	2000.0	2009.6	2503.9
τ =0.2	Bias	180	.014	.010	.004	180	.014	.010	.004	129	.005	.008	.000	129	.005	.008	.000
	RMSE	.914	.147	.107	.020	.914	.147	.107	.020	.694	.111	.070	.013	.694	.111	.070	.013
	Coverage	.960	.980	.960	.950	.960	.980	.960	.950	.930	.970	.970	.960	.930	.970	.970	.960
	Time	6153.1	5575.0	5539.8	6160.9	444.3	382.1	387.2	443.1	42164.0	37560.7	37441.2	42521.2	2529.1	1986.4	2004.5	2515.6
$\tau=2$	Bias	.038	009	005	.002	.038	009	005	.002	.004	006	001	.001	.004	006	001	.001
	RMSE	.416	.082	.074	.009	.416	.082	.074	.009	.408	.058	.039	.007	.408	.058	.039	.007
	Coverage	.940	.920	.950	.920	.940	.920	.950	.920	.940	.940	.960	.960	.940	.940	.960	.960
	Time	6599.8	5969.7	6004.6	6623.9	441.0	380.7	380.5	435.2	46734.2	41946.5	41968.2	46376.3	2502.4	1999.5	2008.0	2454.7
Results for X_2																	
$\tau = -2$	Bias	097	.008	.006	001	097	.008	.006	001	.065	.003	007	001	.065	.003	007	001
	RMSE	.457	.087	.057	.010	.457	.087	.057	.010	.348	.059	.036	.007	.348	.059	.036	.007
	Coverage	.960	.920	.930	.890	.960	.920	.930	.890	.960	.940	.970	.960	.960	.940	.970	.960
	Time	6622.1	5989.0	5995.9	6602.3	439.8	387.5	388.5	434.2	46845.4	42038.7	41967.7	46877.6	2513.3	2037.6	2051.9	2502.8
$\tau = -0.2$	Bias	063	.006	.004	.001	063	.006	.004	.001	.000	.004	.000	.003	.000	.004	.000	.003
	RMSE	.717	.115	.075	.012	.717	.115	.075	.012	.472	.070	.050	.009	.472	.070	.050	.009
	Coverage	.940	.940	.950	.950	.940	.940	.950	.950	.950	.950	.970	.950	.950	.950	.970	.950
	Time	6158.1	5540.3	5543.2	6192.8	444.9	385.2	387.6	438.5	42107.0	37535.4	37466.3	42339.5	2505.1	1999.1	2009.4	2500.6
τ =0.2	Bias	036	.021	.006	.000	036	.021	.006	.000	080	.019	.004	.000	080	.019	.004	.000
	RMSE	.698	.120	.092	.013	.698	.120	.092	.013	.492	.082	.051	.009	.492	.082	.051	.009
	Coverage	.920	.950	.940	.940	.920	.950	.940	.940	.980	.960	.970	.950	.980	.960	.970	.950
	Time	6152.3	5574.7	5539.6	6160.3	443.4	381.0	386.3	442.1	42161.3	37558.1	37439.3	42518.5	2526.6	1985.5	2003.5	2512.8
$\tau=2$	Bias	025	.008	.000	.001	025	.008	.000	.001	.031	005	001	.001	.031	005	001	.001
	RMSE	.338	.053	.032	.006	.338	.053	.032	.006	.294	.043	.026	.004	.294	.043	.026	.004
	Coverage	.920	.930	.950	.970	.920	.930	.950	.970	.930	.940	.930	.950	.930	.940	.930	.950
	Time	6599.4	5970.1	6004.9	6623.3	440.1	380.1	379.6	434.2	46731.9	41943.0	41965.9	46373.9	2498.8	1997.6	2005.7	2451.6