

CSL7320: Digital Image Analysis

Image Restoration and Reconstruction

Image Restoration and Reconstruction



http://fireoracleproductions.com/services__samp

Image degradation due to

- noise in transmission
- imperfect image acquisition
 - environmental condition
 - quality of sensor



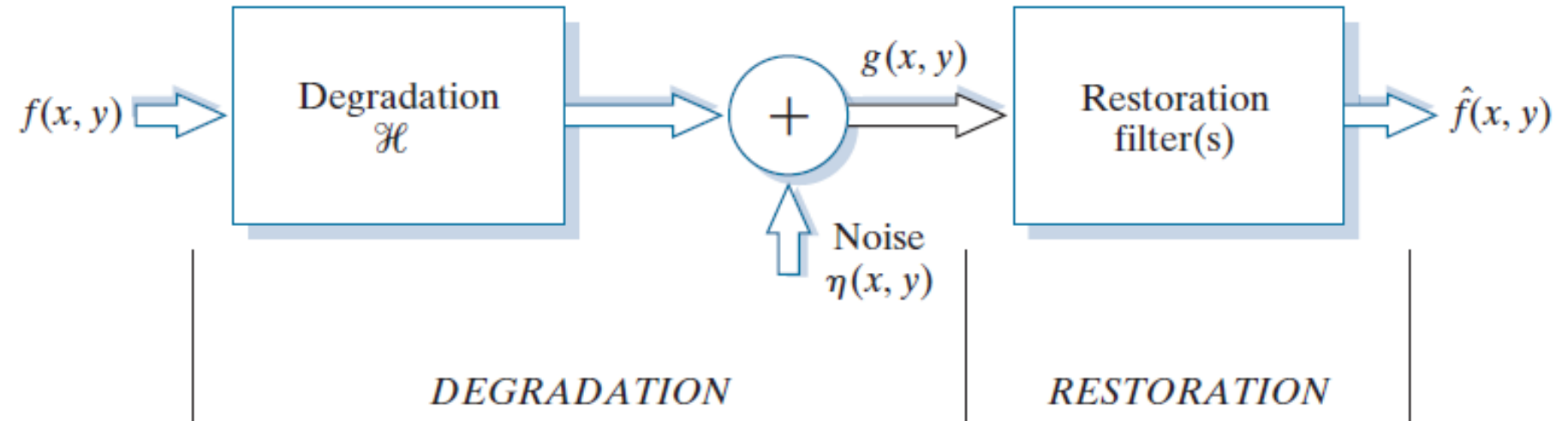
Google image

Slide credit: Yan Tong

Image Restoration and Reconstruction

FIGURE 5.1

A model of the image degradation/restoration process.



$$g(x, y) = h(x, y) \otimes f(x, y) + \eta(x, y)$$

$$G(u, v) = H(u, v)F(u, v) + N(u, v)$$

Identity $H(u, v) \rightarrow$ degradation only comes from additive noise

Image Restoration and Reconstruction

$$g(x, y) = f(x, y) + \eta(x, y)$$

$$G(u, v) = F(u, v) + N(u, v)$$

Noise models:

- Impulse noise: pepper and salt
- Continuous noise model:
 - Gaussian, Rayleigh, Gamma, Exponential, Uniform

Properties of Noise

- **Spatial properties**

- Spatially periodic noise
- Spatially independent noise

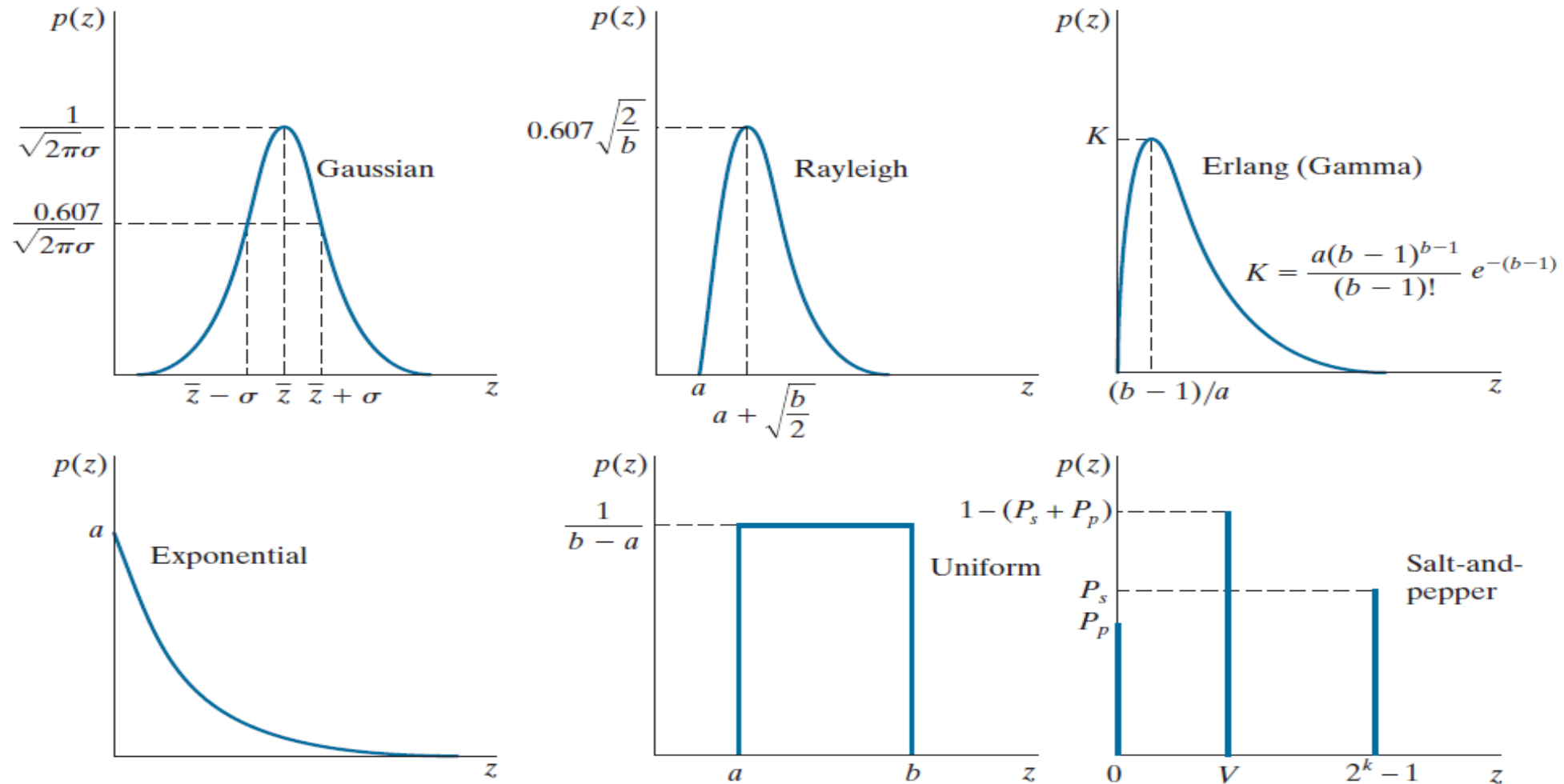
It is uncorrelated with respect to the image itself (that is, there is no correlation between pixel values and the values of noise components)

- **Frequency properties**

- White noise – noise containing all frequencies within a bandwidth

If Fourier spectrum of noise is constant, the noise is called *white noise*

Noise Probability Density Functions



a	b	c
d	e	f

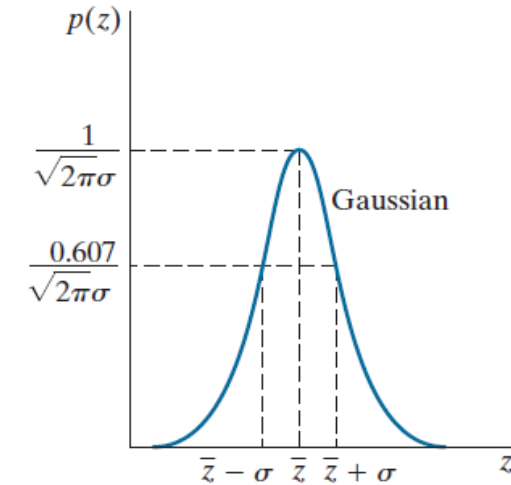
FIGURE 5.2 Some important probability density functions.

Noise Probability Density Functions

Gaussian Noise

The PDF of a *Gaussian* random variable, z , is defined by the following familiar expression:

$$p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{(z - \bar{z})^2}{2\sigma^2}} \quad -\infty < z < \infty \quad (5-3)$$



Rayleigh Noise

The PDF of *Rayleigh* noise is given by

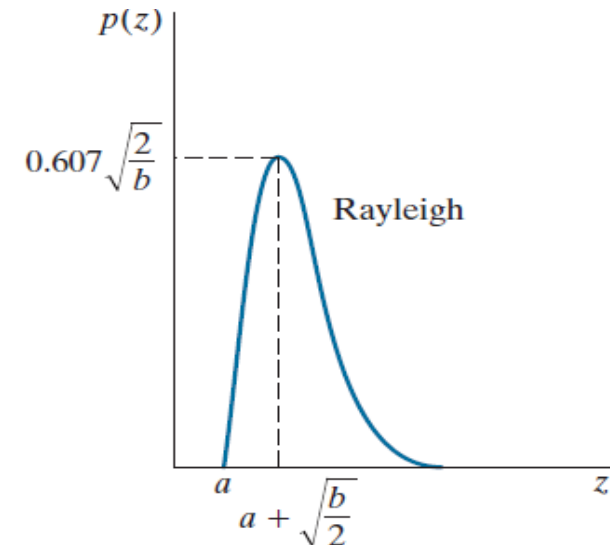
$$p(z) = \begin{cases} \frac{2}{b}(z - a)e^{-(z-a)^2/b} & z \geq a \\ 0 & z < a \end{cases} \quad (5-4)$$

The mean and variance of z when this random variable is characterized by a Rayleigh PDF are

$$\bar{z} = a + \sqrt{\pi b/4} \quad (5-5)$$

and

$$\sigma^2 = \frac{b(4 - \pi)}{4} \quad (5-6)$$



Noise Probability Density Functions

Exponential Noise

The PDF of *exponential* noise is given by

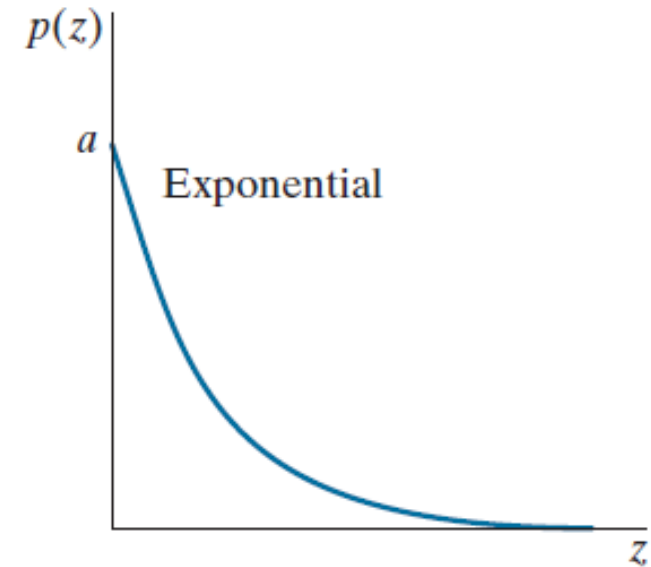
$$p(z) = \begin{cases} ae^{-az} & z \geq 0 \\ 0 & z < 0 \end{cases} \quad (5-10)$$

where $a > 0$. The mean and variance of z are

$$\bar{z} = \frac{1}{a} \quad (5-11)$$

and

$$\sigma^2 = \frac{1}{a^2} \quad (5-12)$$



Noise Probability Density Functions

Uniform Noise

The PDF of *uniform* noise is

$$p(z) = \begin{cases} \frac{1}{b-a} & a \leq z \leq b \\ 0 & \text{otherwise} \end{cases} \quad (5-13)$$

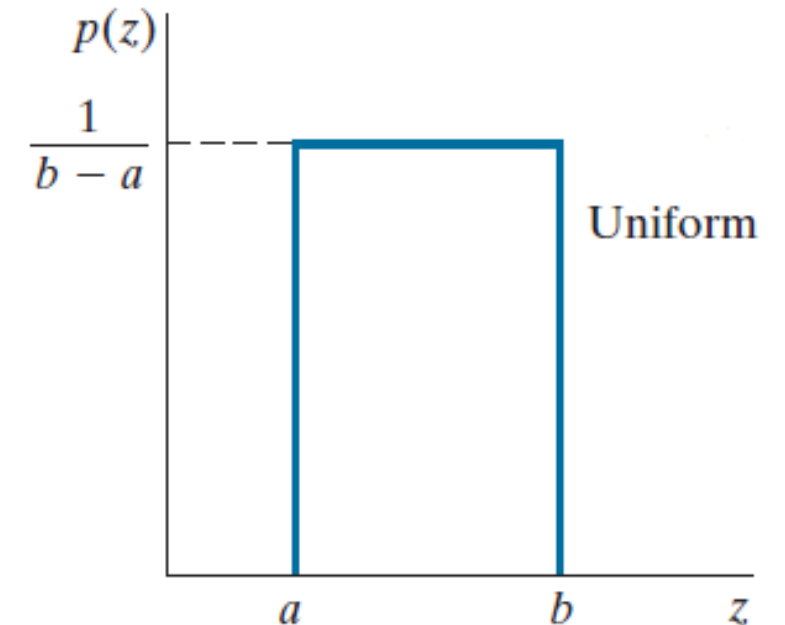
The mean and variance of z are

$$\bar{z} = \frac{a+b}{2} \quad (5-14)$$

and

$$\sigma^2 = \frac{(b-a)^2}{12} \quad (5-15)$$

Figure 5.2(e) shows a plot of the uniform density.



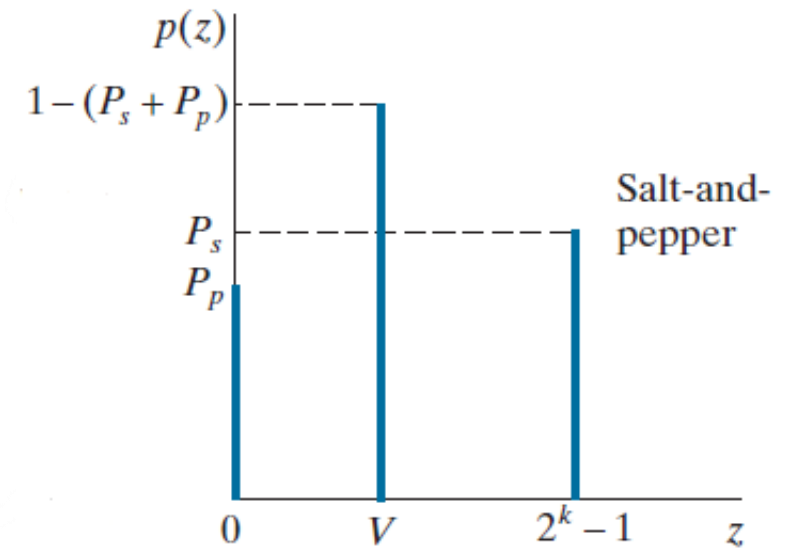
Noise Probability Density Functions

Salt-and-Pepper Noise

If k represents the number of bits used to represent the intensity values in a digital image, then the range of possible intensity values for that image is $[0, 2^k - 1]$ (e.g., $[0, 255]$ for an 8-bit image). The PDF of *salt-and-pepper* noise is given by

$$p(z) = \begin{cases} P_s & \text{for } z = 2^k - 1 \\ P_p & \text{for } z = 0 \\ 1 - (P_s + P_p) & \text{for } z = V \end{cases} \quad (5-16)$$

where V is any integer value in the range $0 < V < 2^k - 1$.



Noisy Images and Their Histograms

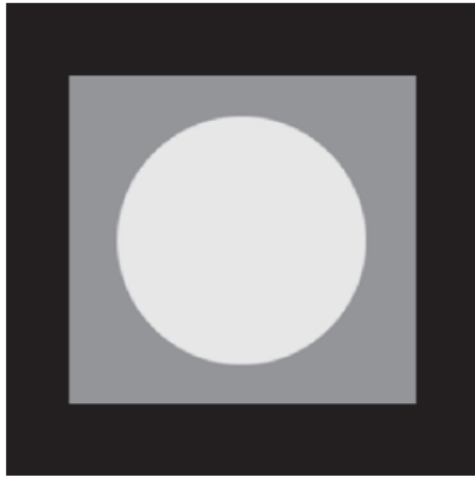
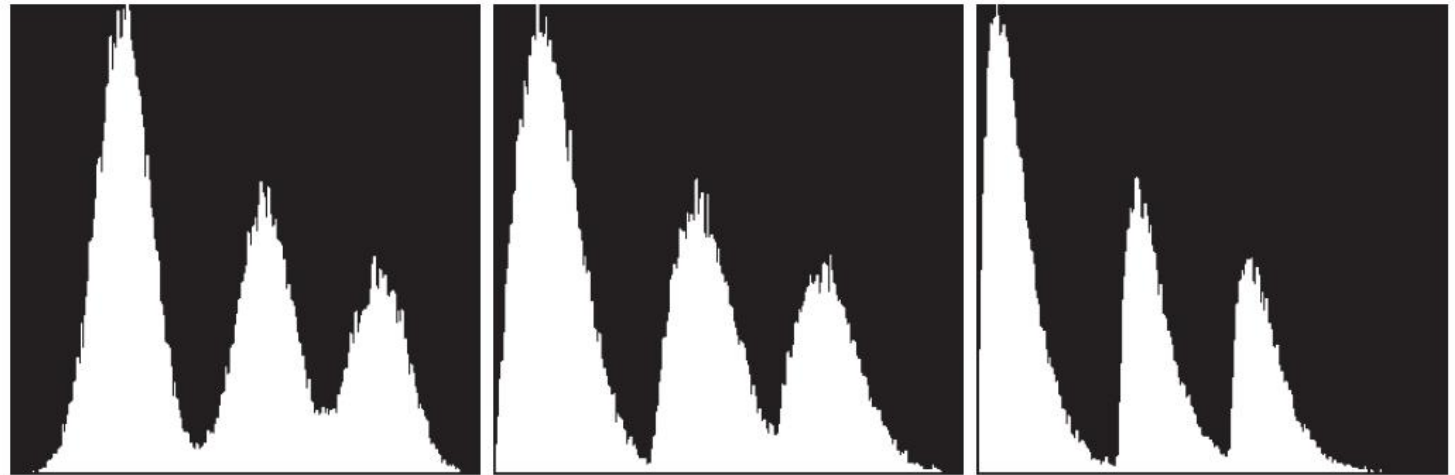
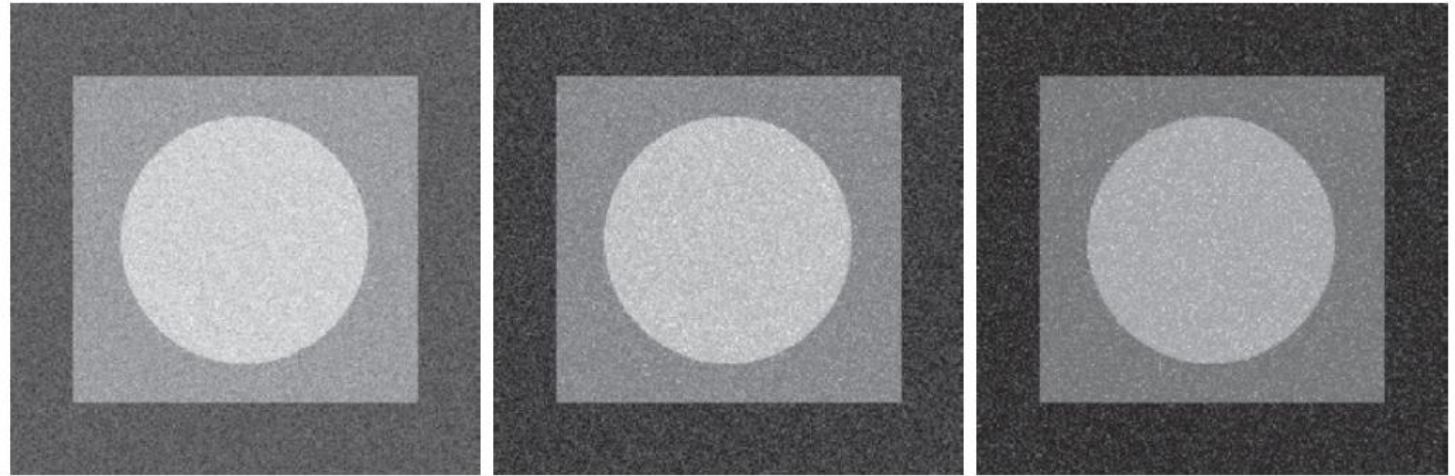


FIGURE 5.3

Test pattern used to illustrate the characteristics of the PDFs from Fig. 5.2.



a	b	c
d	e	f

FIGURE 5.4 Images and histograms resulting from adding Gaussian, Rayleigh, and Erlanga noise to the image in Fig. 5.3.

Noisy Images and Their Histograms

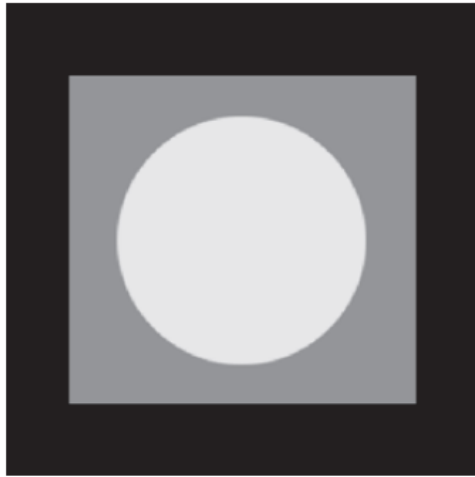
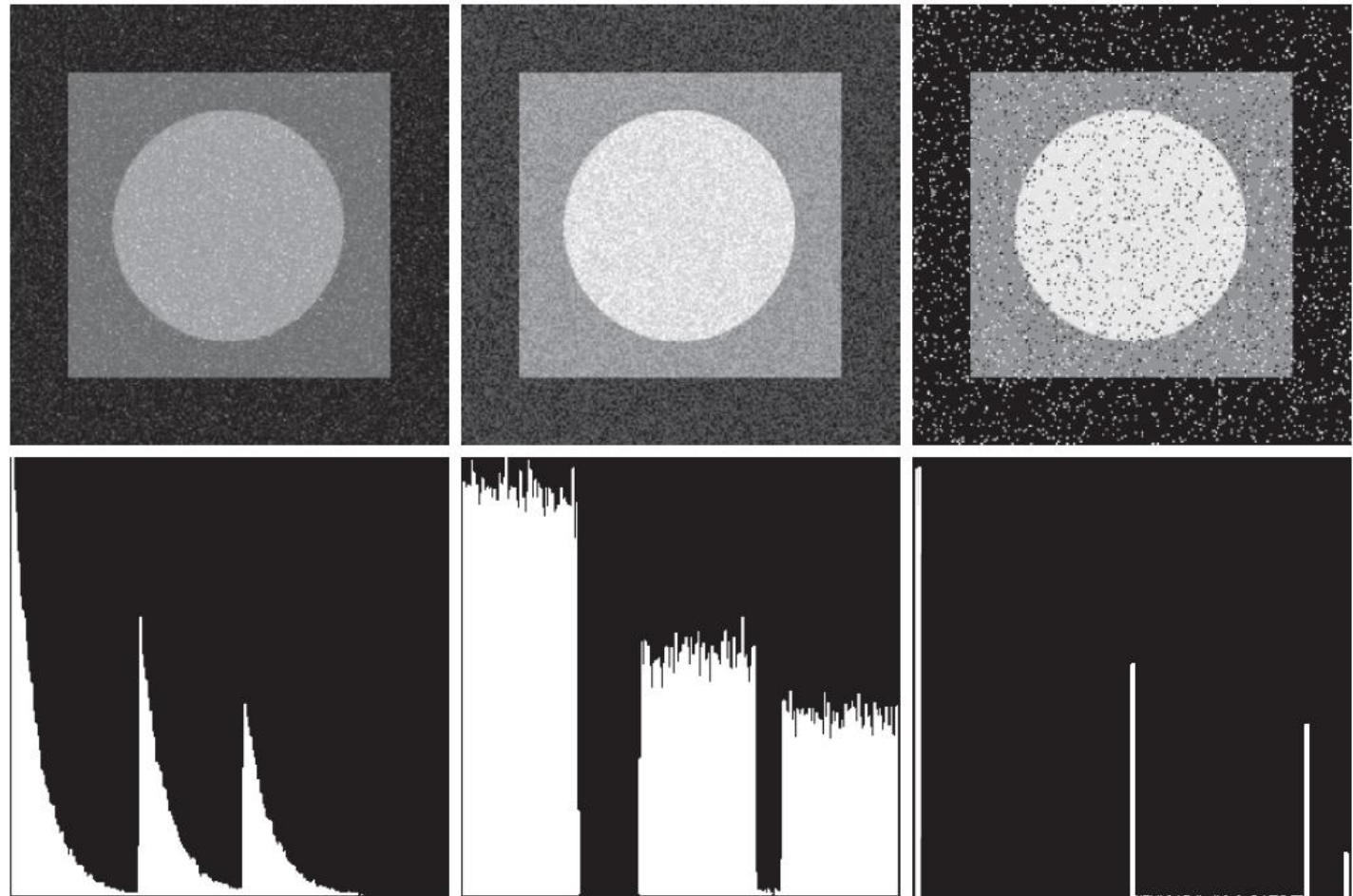


FIGURE 5.3

Test pattern used to illustrate the characteristics of the PDFs from Fig. 5.2.



g h i
j k l

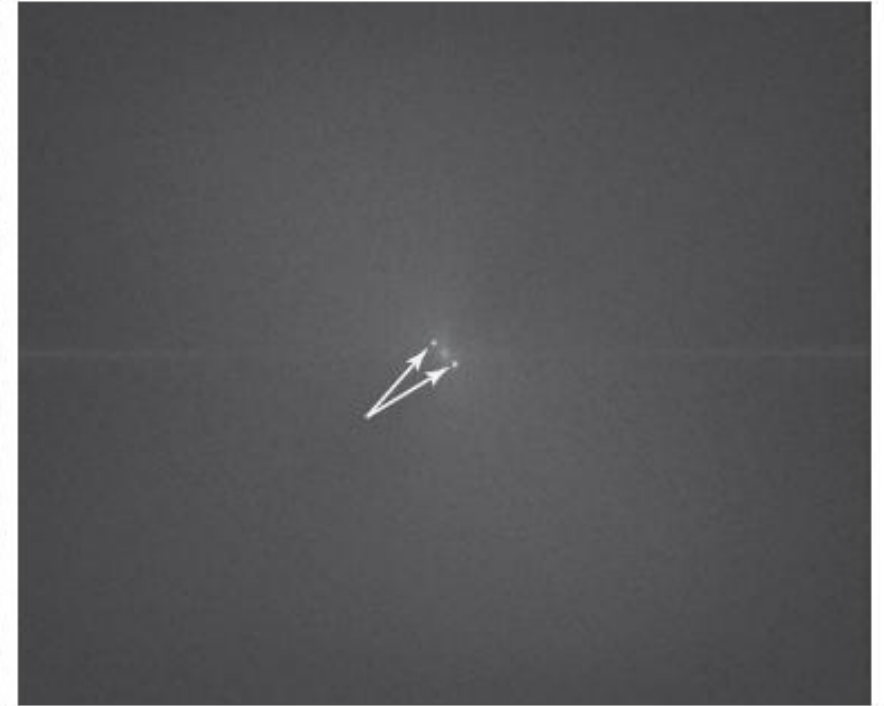
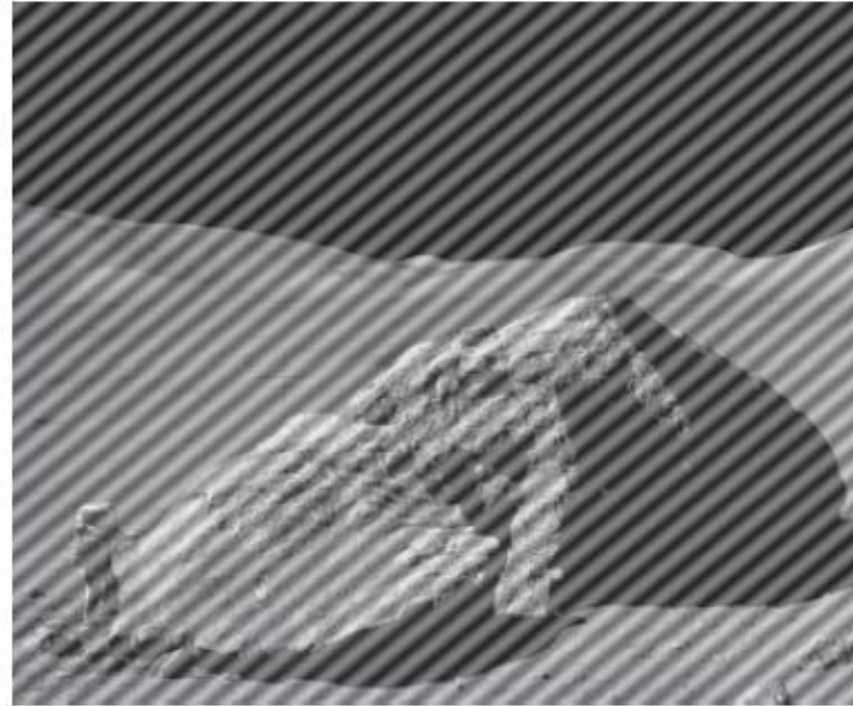
FIGURE 5.4 (*continued*) Images and histograms resulting from adding exponential, uniform, and salt-and-pepper noise to the image in Fig. 5.3. In the salt-and-pepper histogram, the peaks in the origin (zero intensity) and at the far end of the scale are shown displaced slightly so that they do not blend with the page background.

Noisy Images

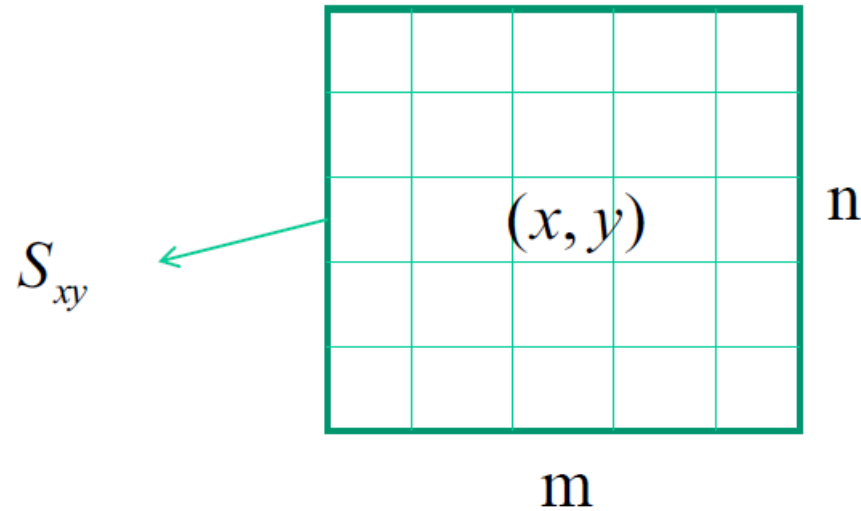
a b

FIGURE 5.5

(a) Image corrupted by additive sinusoidal noise.
(b) Spectrum showing two conjugate impulses caused by the sine wave.
(Original image courtesy of NASA.)



Mean Filters for Continuous Noise Models



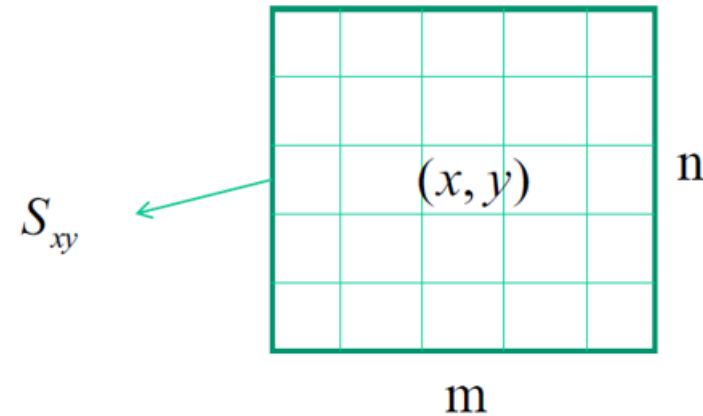
Arithmetic Mean Filter: a linear filter

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s,t) \in S_{xy}} g(s, t)$$

Non-linear Mean Filter: Geometric Mean Filter

$$\hat{f}(x,y) = \left[\prod_{(r,c) \in S_{xy}} g(r,c) \right]^{\frac{1}{mn}}$$

- Removing the salt noise
- Fail in the pepper noise

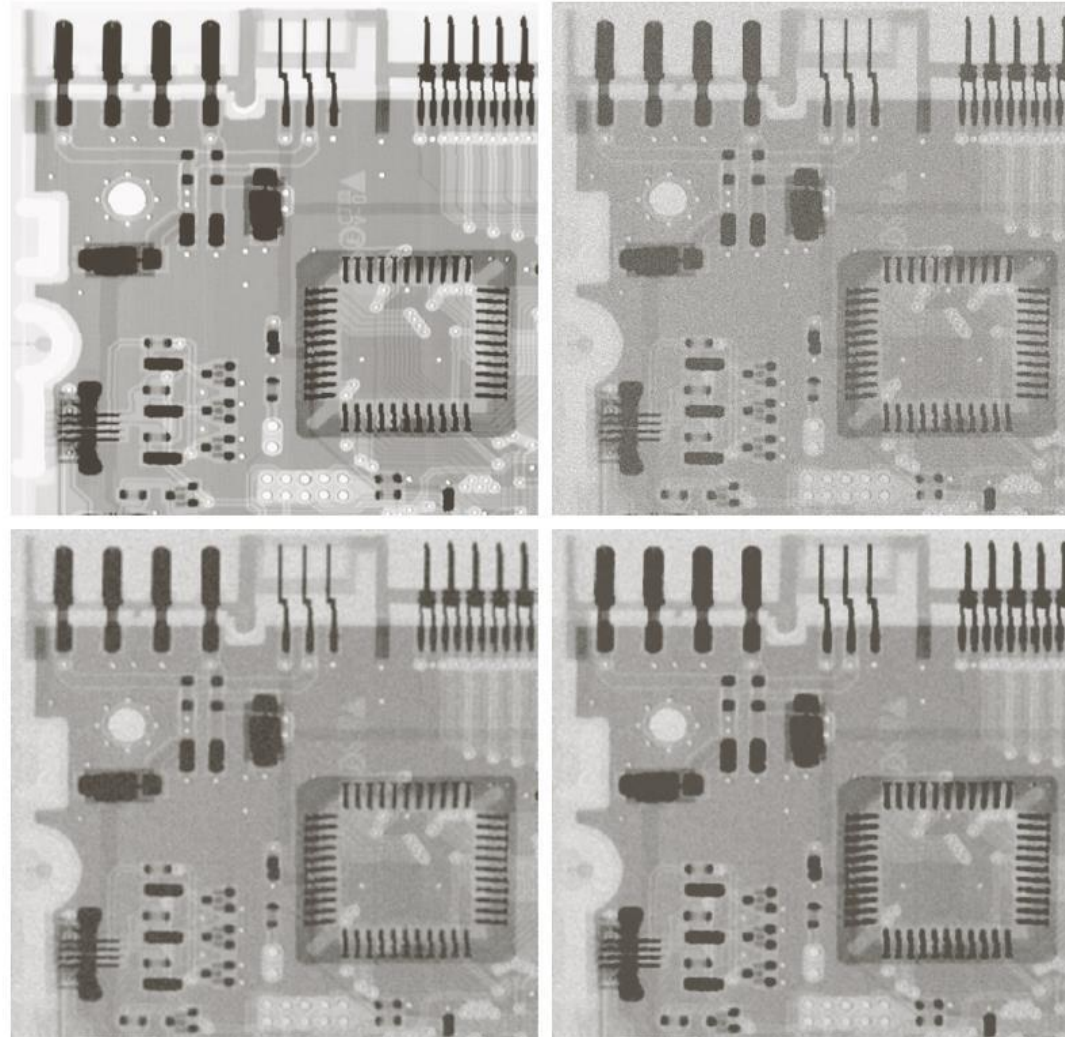


Example: Geometric Mean Filter

a	b
c	d

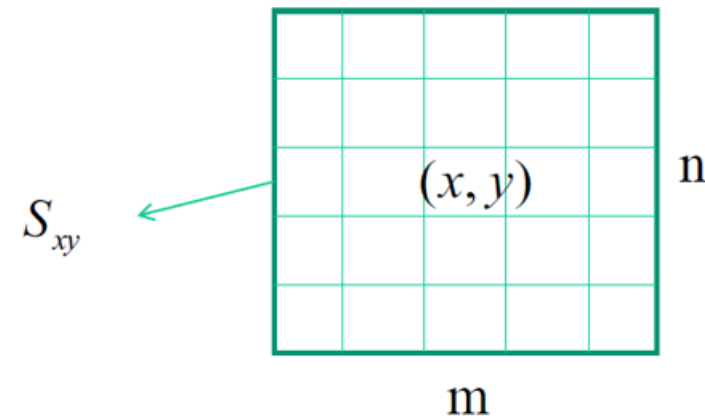
FIGURE 5.7

(a) X-ray image.
(b) Image corrupted by additive Gaussian noise.
(c) Result of filtering with an arithmetic mean filter of size 3×3 .
(d) Result of filtering with a geometric mean filter of the same size.
(Original image courtesy of Mr. Joseph E. Pascente, Lixi, Inc.)



Non-linear Mean Filter: Harmonic Mean Filter

$$\hat{f}(x, y) = \frac{1}{\frac{1}{mn} \sum_{(s,t) \in S_{xy}} \frac{1}{g(s,t)}}$$

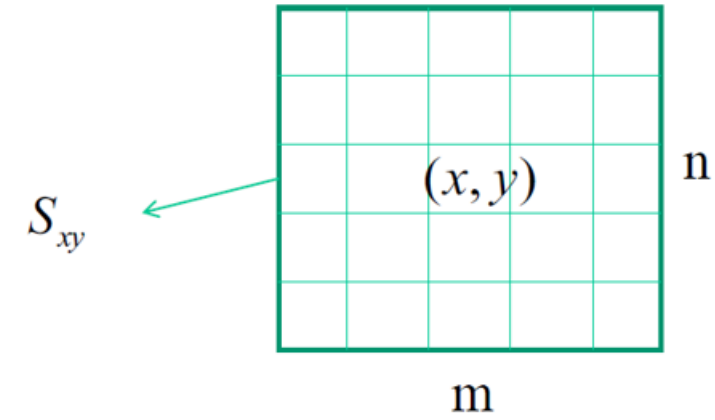


- Removing the salt noise
- Fail in the pepper noise

$$H(x_1, x_2, \dots, x_n) = \frac{n}{\frac{1}{x_1} + \frac{1}{x_2} + \dots + \frac{1}{x_n}} = \frac{n}{\sum_{i=1}^n \frac{1}{x_i}}$$

Non-linear Mean Filter: Contraharmonic Mean Filter

$$\hat{f}(x, y) = \frac{\sum_{(s,t) \in S_{xy}} g(s, t)^{Q+1}}{\sum_{(s,t) \in S_{xy}} g(s, t)^Q}$$



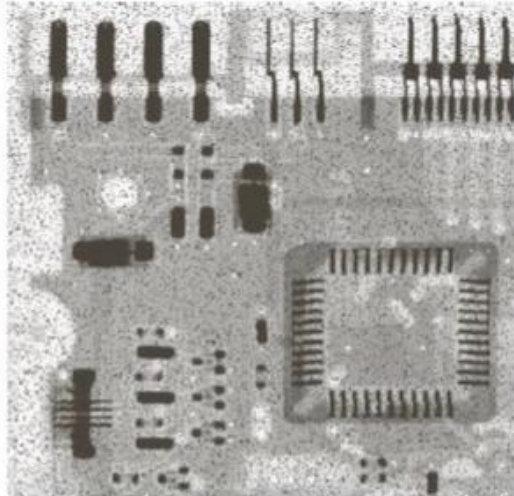
Q is the order of the filter

- positive Q removes pepper noise
- negative Q removes salt noise
- Special cases: Q=0, Q=-1

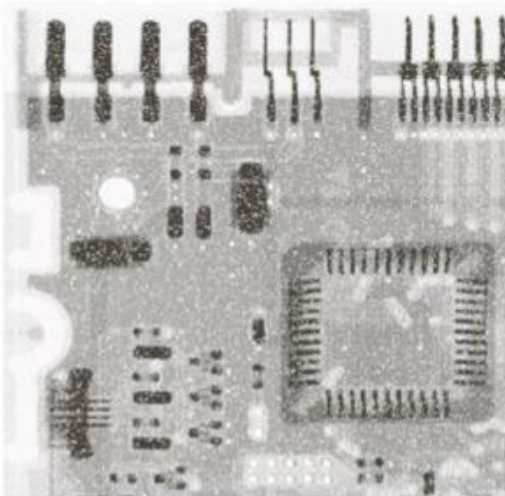
$$\begin{aligned} C(x_1, x_2, \dots, x_n) &= \frac{\frac{1}{n} (x_1^2 + x_2^2 + \dots + x_n^2)}{\frac{1}{n} (x_1 + x_2 + \dots + x_n)}, \\ &= \frac{x_1^2 + x_2^2 + \dots + x_n^2}{x_1 + x_2 + \dots + x_n}. \end{aligned}$$

Example: Contraharmonic Mean Filter

Pepper noise



Salt noise

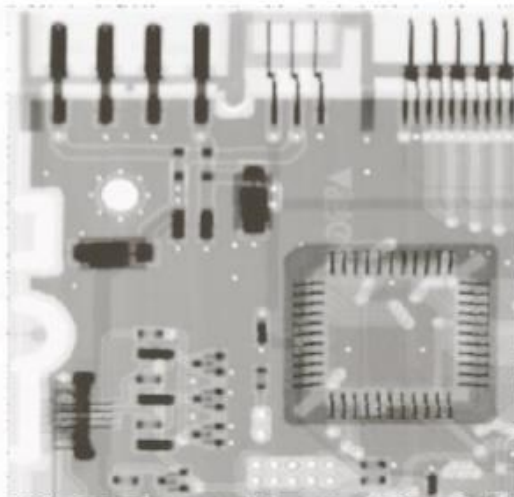


a b
c d

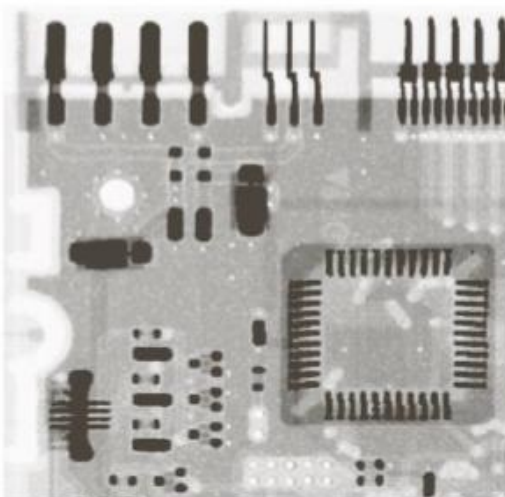
FIGURE 5.8

(a) Image corrupted by pepper noise with a probability of 0.1. (b) Image corrupted by salt noise with the same probability. (c) Result of filtering (a) with a 3×3 contra-harmonic filter of order 1.5. (d) Result of filtering (b) with $Q = -1.5$.

$Q=1.5$



$Q=-1.5$



A Failed Case with Wrong Sign of Contraharmonic Filter

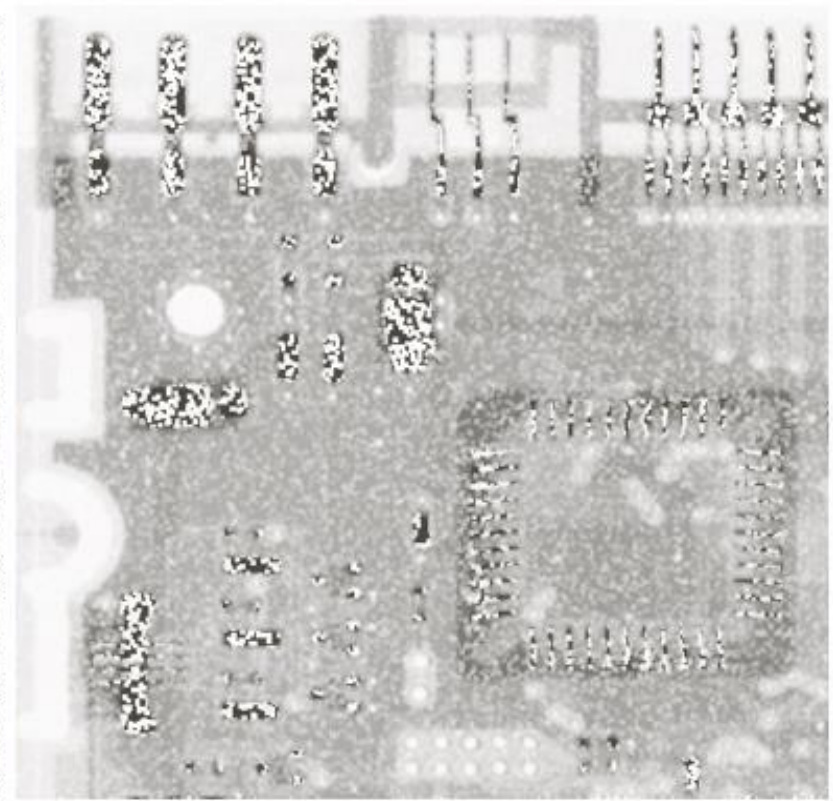
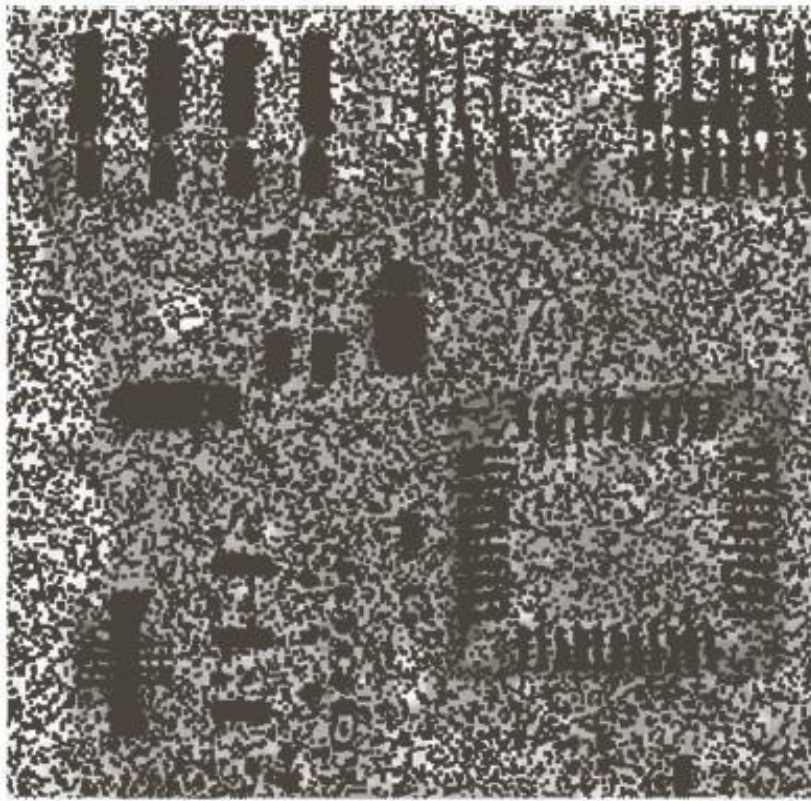
a b

FIGURE 5.9

Results of selecting the wrong sign in contraharmonic filtering.

(a) Result of filtering Fig. 5.8(a) with a contraharmonic filter of size 3×3 and $Q = -1.5$.

(b) Result of filtering 5.8(b) with $Q = 1.5$.



Summary: Mean Filters for Continuous Noise Models

Arithmetic Mean Filter: a linear filter

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s,t) \in S_{xy}} g(s, t)$$

- Work well for continuous noise

Summary: Non-linear Mean Filters

Geometric Mean Filter

$$\hat{f}(x, y) = \left[\prod_{(s,t) \in S_{xy}} g(s, t) \right]^{\frac{1}{mn}}$$

Work well for

- Continuous noise
- Salt noise

Harmonic Mean Filter

$$\hat{f}(x, y) = \frac{1}{\frac{1}{mn} \sum_{(s,t) \in S_{xy}} \frac{1}{g(s, t)}}$$

Fail for the pepper noise

Contraharmonic Mean Filter

$$\hat{f}(x, y) = \frac{\sum_{(s,t) \in S_{xy}} g(s, t)^{Q+1}}{\sum_{(s,t) \in S_{xy}} g(s, t)^Q}$$

Q is the order of the filter

- positive Q removes pepper noise
- negative Q removes salt noise
- Special cases: Q=0, Q=-1

Order-Statistic Filters

Median Filter

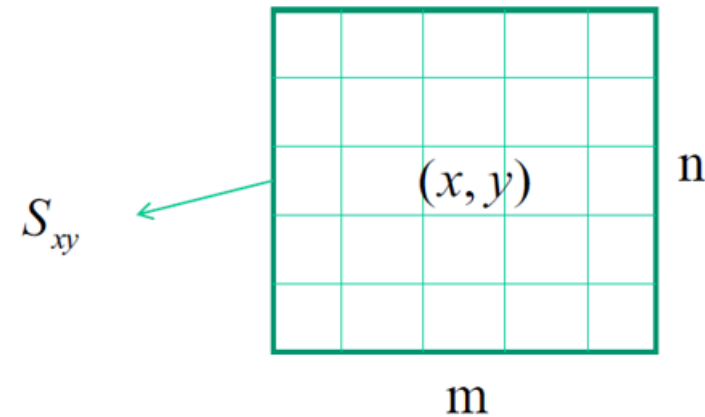
$$\hat{f}(x, y) = \operatorname{median}_{(r, c) \in S_{xy}} \{g(r, c)\}$$

Max and Min Filters

$$\hat{f}(x, y) = \max_{(r, c) \in S_{xy}} \{g(r, c)\}$$

Midpoint Filter

$$\hat{f}(x, y) = \frac{1}{2} \left[\max_{(r, c) \in S_{xy}} \{g(r, c)\} + \min_{(r, c) \in S_{xy}} \{g(r, c)\} \right]$$



Alpha-Trimmed Mean Filter

$$\hat{f}(x, y) = \frac{1}{mn - d} \sum_{(r, c) \in S_{xy}} g_R(r, c)$$

Order-Statistic Filters -- Median Filter

a	b
c	d

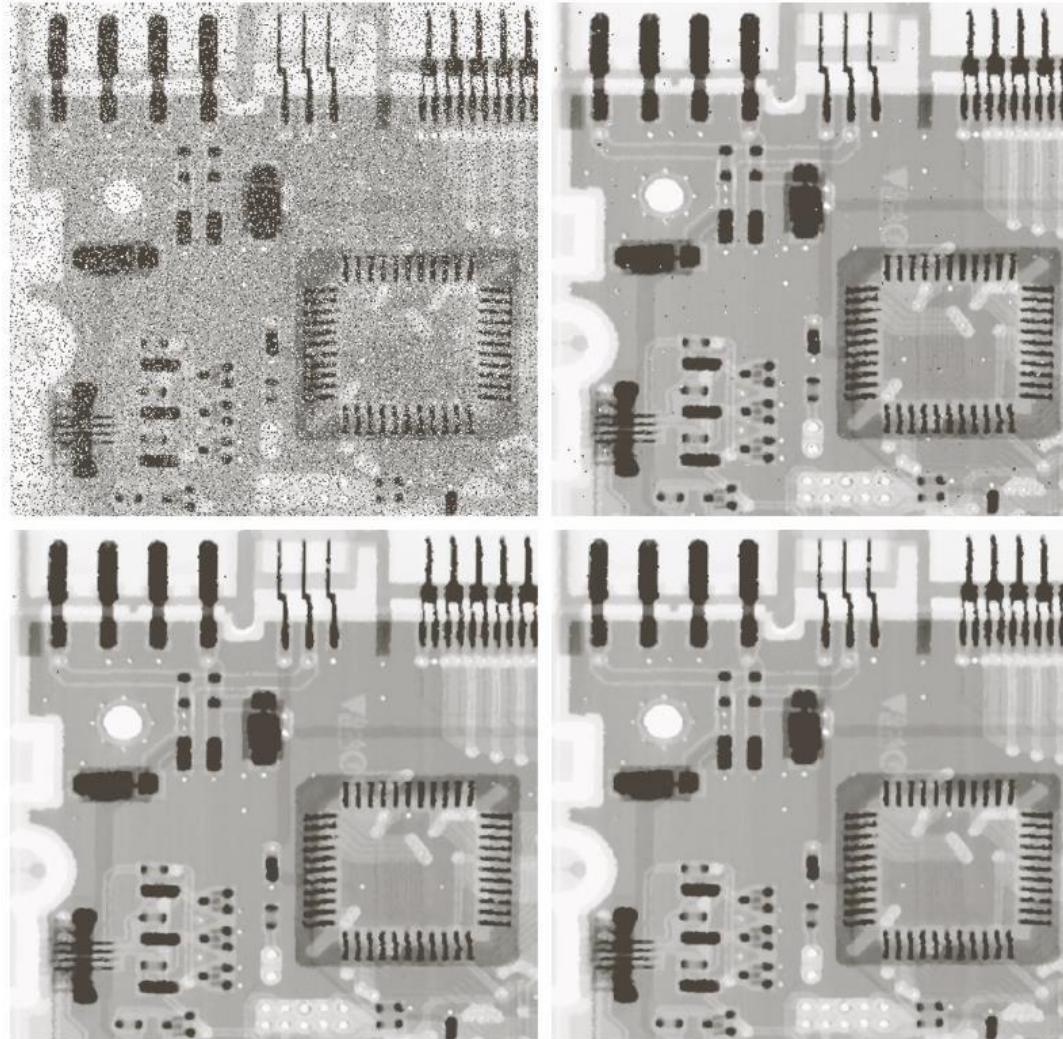
FIGURE 5.10

(a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.1$.

(b) Result of one pass with a median filter of size 3×3 .

(c) Result of processing (b) with this filter.

(d) Result of processing (c) with the same filter.



Repeating median filter will remove most of the noise while increase image blurring

Slide credit: Yan Tong

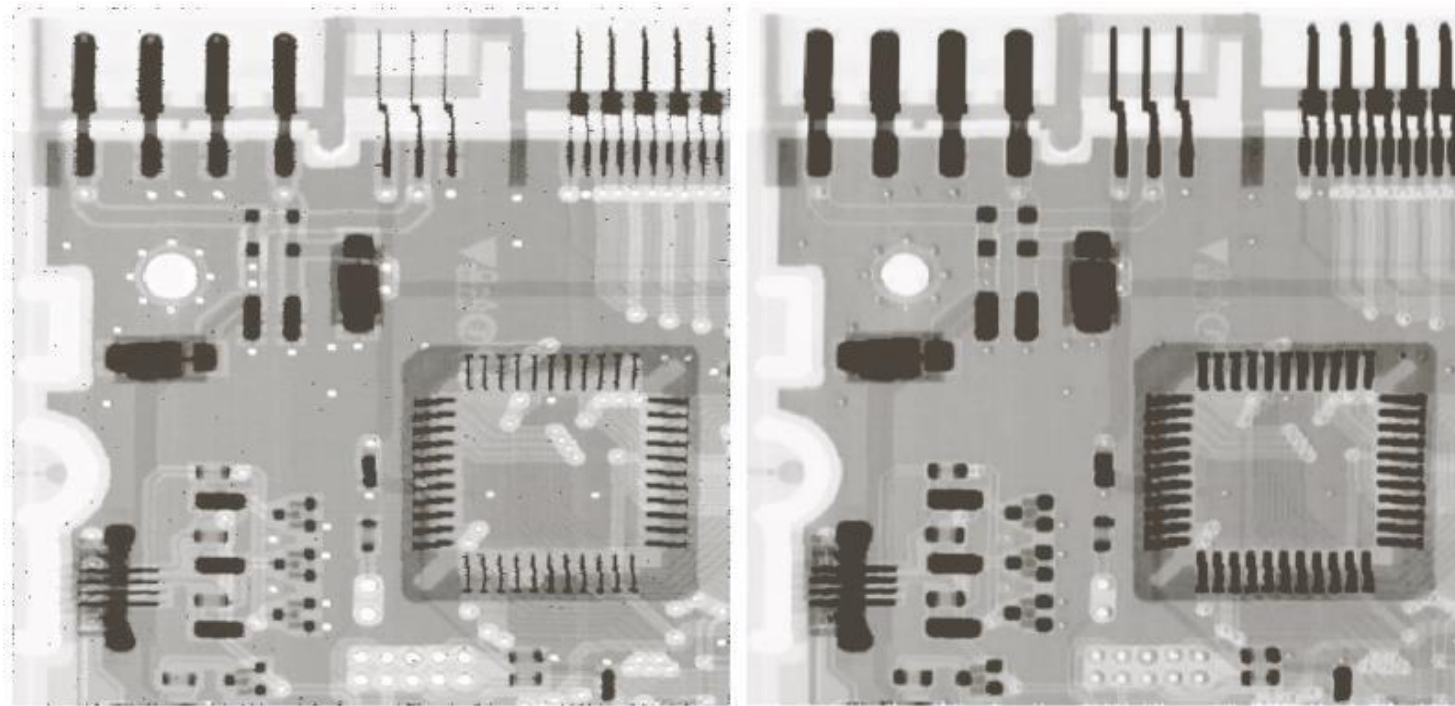
Order-Statistic Filters -- Max/Min Filters

a b

FIGURE 5.11

(a) Result of filtering Fig. 5.8(a) with a max filter of size 3×3 .

(b) Result of filtering 5.8(b) with a min filter of the same size.



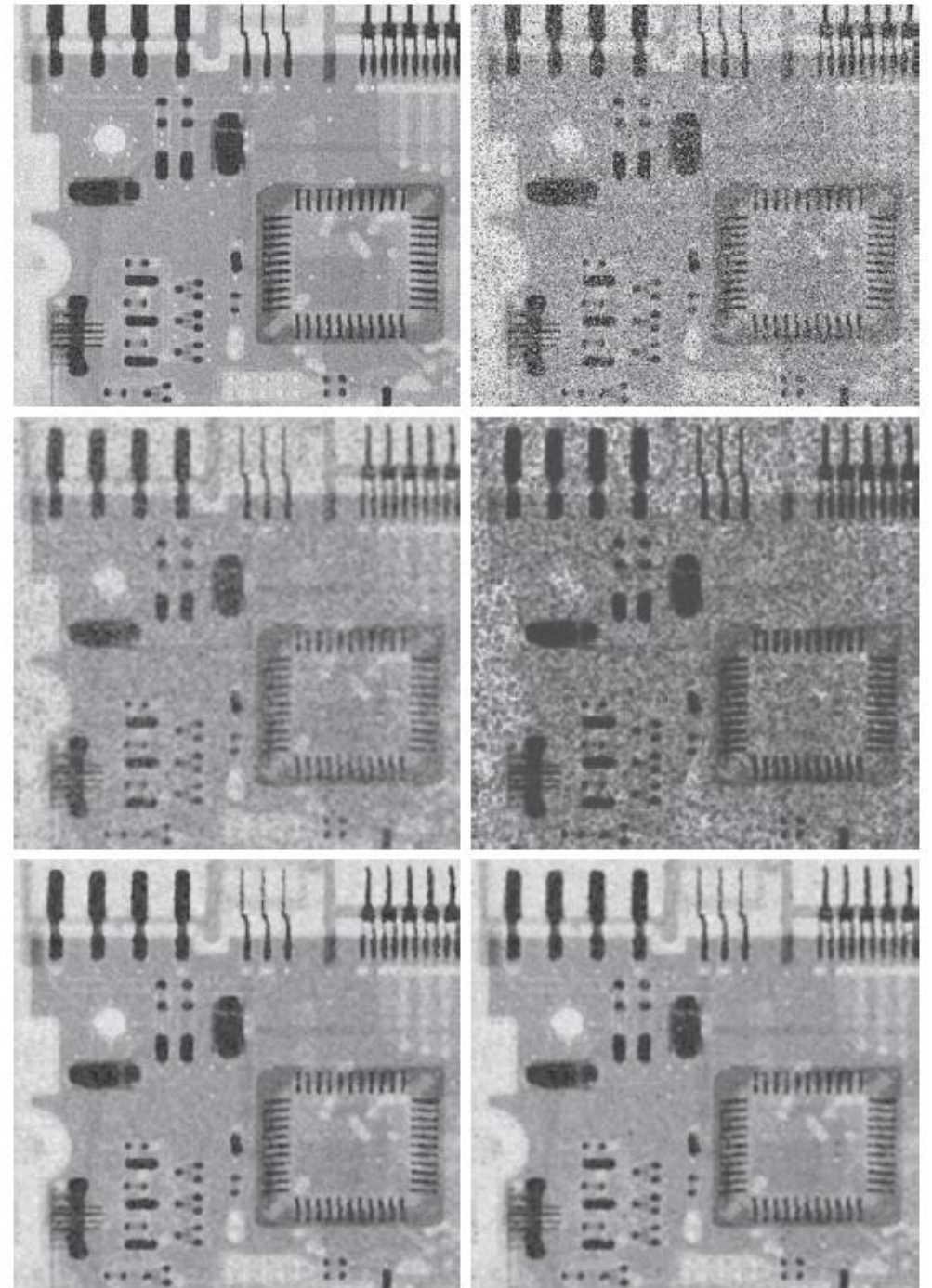
- Find the extreme points
- Remove the targeting impulse noise

Order-Statistic Filters

a b
c d
e f

FIGURE 5.12

(a) Image corrupted by additive uniform noise. (b) Image additionally corrupted by additive salt-and-pepper noise. (c)-(f) Image (b) filtered with a 5×5 :
(c) arithmetic mean filter;
(d) geometric mean filter;
(e) median filter;
(f) alpha-trimmed mean filter, with $d = 6$.



Adaptive Median Filter

Stage A: check if the median value is an extreme value

$$A1 = z_{\text{med}} - z_{\text{min}}$$

$$A2 = z_{\text{med}} - z_{\text{max}}$$

If $A1 > 0$ AND $A2 < 0$, go to stage B

Else increase the window size

If window size $\leq S_{\text{max}}$ repeat stage A

Else output z_{med}

Goal 1: remove salt-and-pepper noise with higher probability

Goal 2: smoothing the noise other than impulses

Goal 3: reduce distortion

Stage B: check if the center pixel is an extreme value

$$B1 = z_{xy} - z_{\text{min}}$$

$$B2 = z_{xy} - z_{\text{max}}$$

If $B1 > 0$ AND $B2 < 0$, output z_{xy}

Else output z_{med}

z_{min} = Minimum gray level value in S_{xy} .

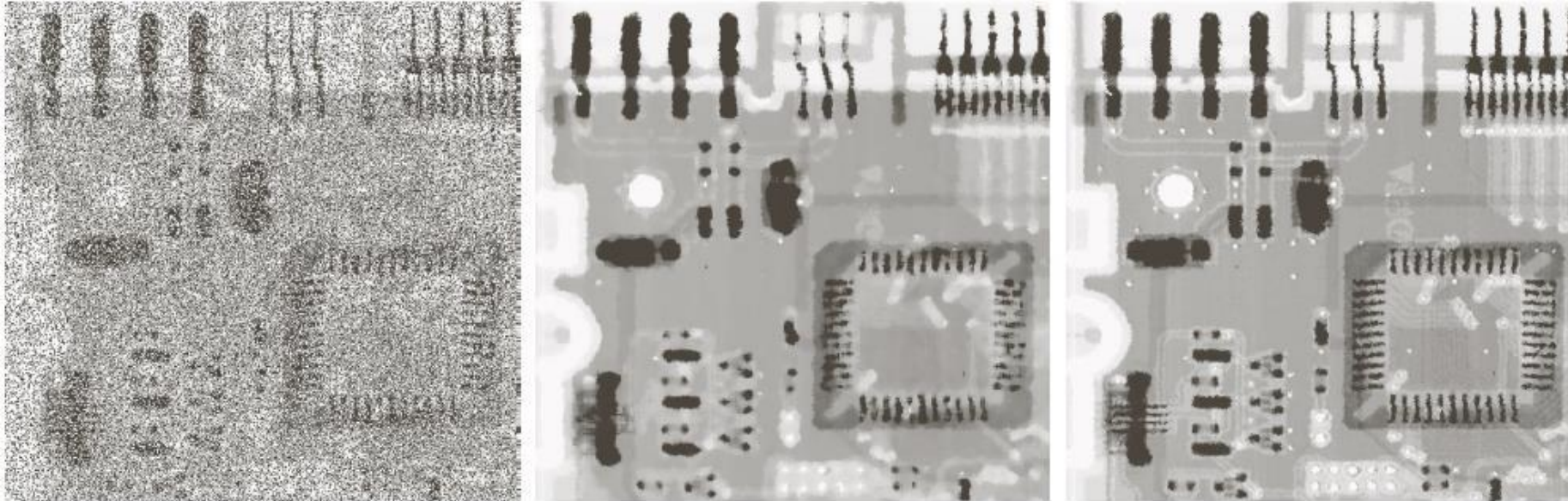
z_{max} = Maximum gray level value in S_{xy}

z_{med} = Median of gray levels in S_{xy}

z_{xy} = gray level at coordinates (x, y)

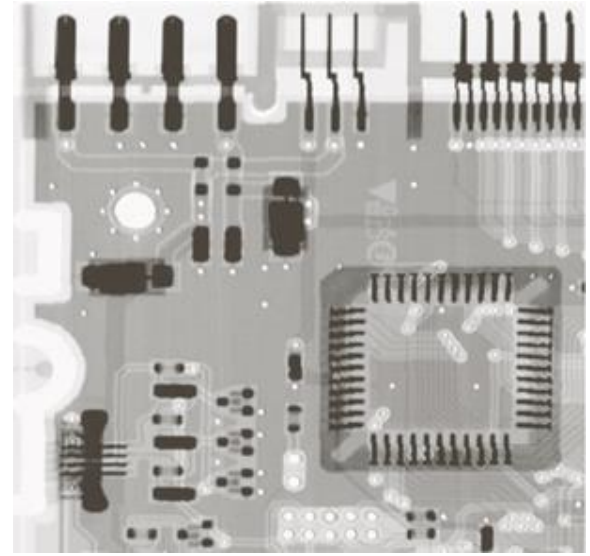
S_{max} = Maximum allowed size of S_{xy}

Example: Adaptive Median Filter



a b c

FIGURE 5.14 (a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.25$. (b) Result of filtering with a 7×7 median filter. (c) Result of adaptive median filtering with $S_{\max} = 7$.



Clean Image

Linear, Position-Invariant Degradations – Noise Free Case

$$g(x, y) = H[f(x, y)]$$

Linearity $H[af_1(x, y) + bf_2(x, y)] = aH[f_1(x, y)] + bH[f_2(x, y)]$

Position/space invariant $H[f(x - \alpha, y - \beta)] = g(x - \alpha, y - \beta)$



H does not depend on the location (x, y) ,
only represented by the input and output

Impulse Response for Linear H

$$f(x, y) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\alpha, \beta) \delta(x - \alpha, y - \beta) d\alpha d\beta \quad \leftarrow \text{Sifting property}$$

$$g(x, y) = H[f(x, y)]$$



$$g(x, y) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\alpha, \beta) \boxed{H[\delta(x - \alpha, y - \beta)]} d\alpha d\beta$$

Impulse response

Image Degradations

$$g(x, y) = H[f(x, y)] + \eta(x, y)$$



$$g(x, y) = \boxed{h(x, y) \otimes f(x, y)} + \eta(x, y)$$

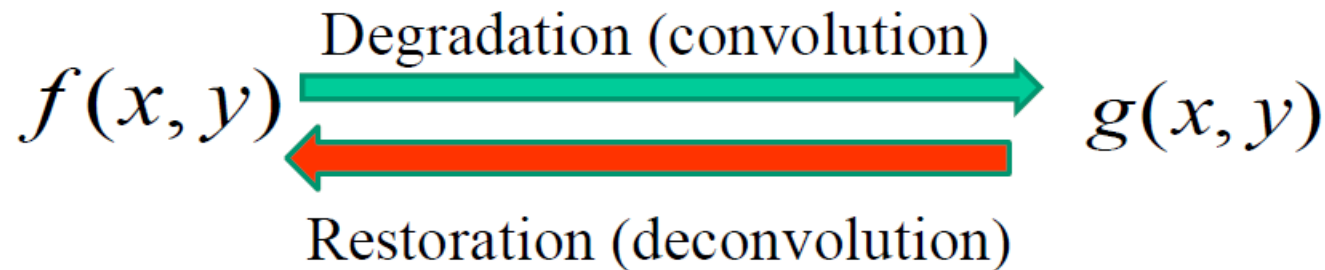
Degradation VS Restoration

$$g(x, y) = H[f(x, y)] + \eta(x, y)$$



Note: a linear, position invariant degradation system with additive noise can be modeled as the convolution of the degradation function with the image plus the additive noise.

$$g(x, y) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\alpha, \beta) h(x - \alpha, y - \beta) d\alpha d\beta + \eta(x, y)$$



Degradation VS Restoration

$$g(x, y) = (h \star f)(x, y) + \eta(x, y)$$



$$G(u, v) = H(u, v)F(u, v) + N(u, v)$$

Estimate the Degradation Function

- **Observation**
- **Experimentation**
- **Mathematical modeling**

Estimate Degradation Function - Observation

Assumptions:

- The degradation function is linear and position-invariant
- No other knowledge about the degradation function

Estimation by image observation:

- Extract a subimage with strong signal → Higher signal-to-noise ratio
- Perform restoration on the subimage

degradation function in the subimage ← $H_s(u, v) = \frac{G_s(u, v)}{\hat{F}_s(u, v)}$ → Observed subimage

↓ $H(u, v)$ → Restored subimage

Application: restoring old pictures

Estimate Degradation Function - Experimentation

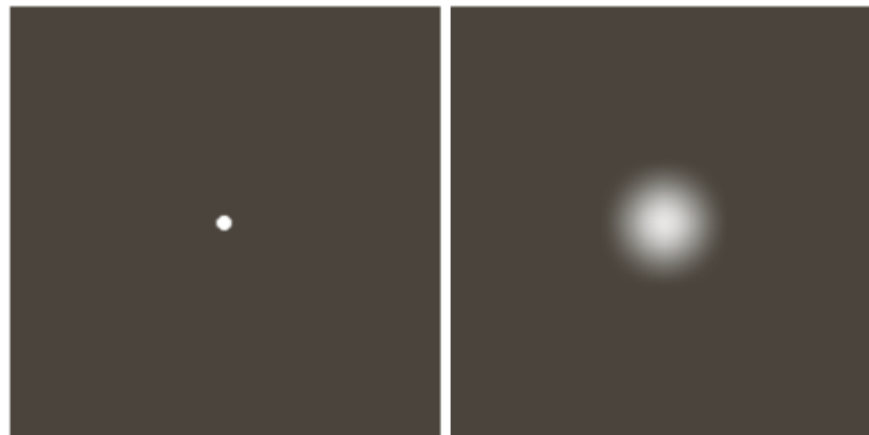
Assumptions:

- A similar equipment is available
- Change the system setting can achieve similar degraded images

$$H(u, v) = \frac{G(u, v)}{A}$$

→ Observed image
→ Impulse signal (A constant describing strength of the impulse)

a b
FIGURE 5.24
Degradation
estimation by
impulse
characterization.
(a) An impulse of
light (shown
magnified).
(b) Imaged
(degraded)
impulse.



Estimation by Modeling – Motion Blur

Constant velocity along x and y direction:

$$x_0(t) = \frac{at}{T} \quad y_0(t) = \frac{bt}{T}$$

Suppose that an image $f(x, y)$ undergoes planar motion. These x_0 and y_0 are the time-varying components of motion in the x - and y -directions, respectively.

T is the duration of the exposure

The image has been displaced by a total distance a along x and b along y




a b

FIGURE 5.26

(a) Original image.
(b) Result of blurring using the function in Eq. (5.6-11) with $a = b = 0.1$ and $T = 1$.

Estimation by Modeling – Motion Blur

An example of motion blur

$$g(x, y) = \int_0^T f[x - x_0(t), y - y_0(t)] dt$$


Motion in both x and y direction during acquisition

Estimation by Modeling – Motion Blur

An example of motion blur

$$G(u, v) = F(u, v) \int_0^T e^{-j2\pi[ux_0(t) + vy_0(t)]} dt$$

➡ $H(u, v) = \int_0^T e^{-j2\pi[ux_0(t) + vy_0(t)]} dt$

Estimation by Modeling – Motion Blur

Constant velocity along x and y direction:

$$x_0(t) = at / T \quad y_0(t) = bt / T$$

What is $H(u, v)$?

$$H(u, v) = T \frac{\sin[\pi(ua + vb)]}{\pi(ua + vb)} e^{-j\pi(ua + vb)}$$

Estimation by Modeling – Motion Blur

a b

FIGURE 5.26

(a) Original image. (b) Result of blurring using the function in Eq. (5-77) with $a = b = 0.1$ and $T = 1$.

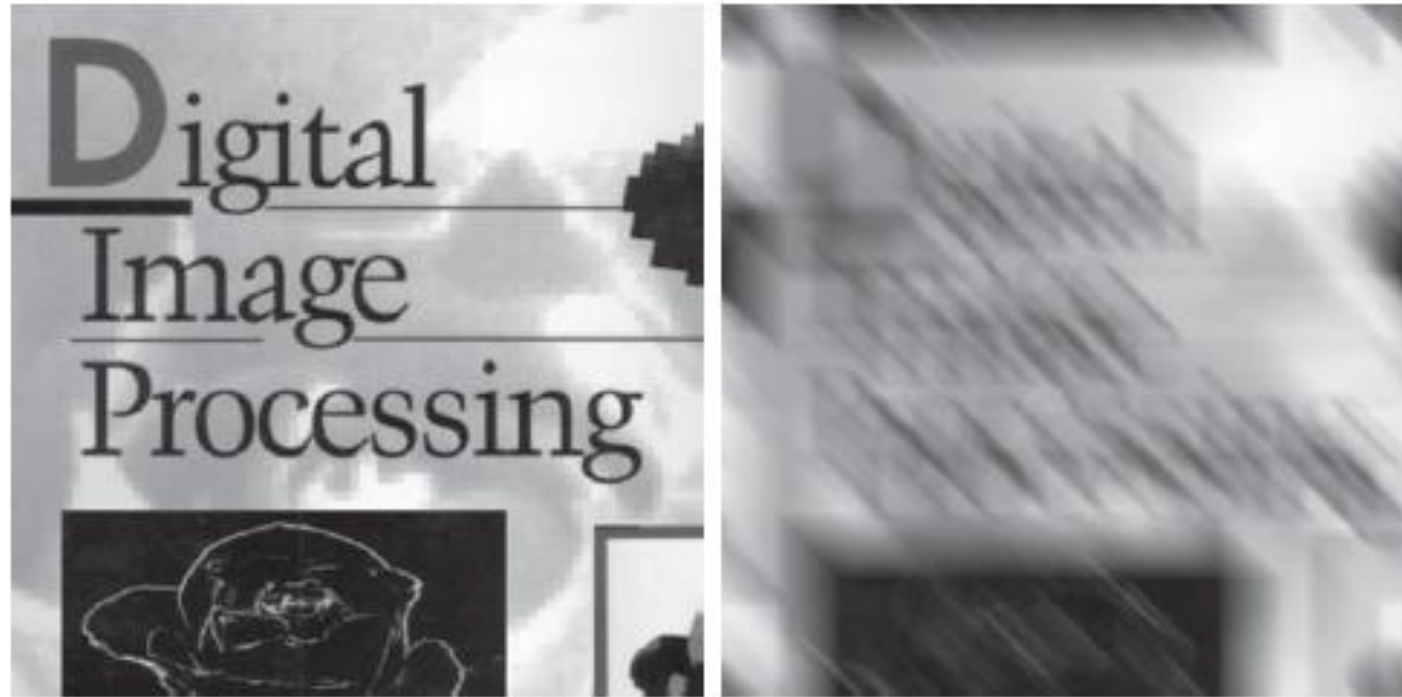


Figure 5.26(b) is an image blurred by computing the Fourier transform of the image in Fig. 5.26(a), multiplying the transform by $H(u,v)$ from Eq. (5-77), and taking the inverse transform. The images are of size 688 x 688 pixels, and we used $a = b = 0.1$ and $T = 1$ in Eq. (5-77).

Image Restoration

Given the degradation system \mathbf{H} and the input image \mathbf{G} ,
recover the original image \mathbf{F}

- Inverse filtering

Inverse Filtering

Ideally:

$$G(u, v) = H(u, v)F(u, v)$$



$$\hat{F}(u, v) = \frac{G(u, v)}{H(u, v)}$$

In practice:


$$G(u, v) = H(u, v)F(u, v) + N(u, v)$$



$$\hat{F}(u, v) = F(u, v) + \frac{N(u, v)}{H(u, v)}$$



Limiting the analysis to
frequencies near the origin
(0,0)

Low pass filtering 

An Example of Inverse Filtering

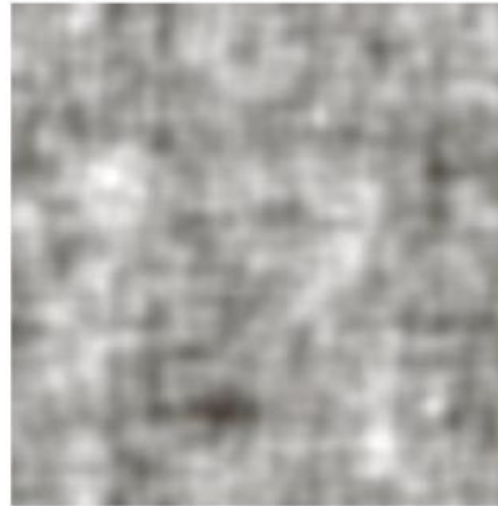
Original image



Degraded image



$$\frac{G(u, v)}{H(u, v)} \rightarrow$$



The reasons for this poor result are

$$\hat{F}(u, v) = F(u, v) + \frac{N(u, v)}{H(u, v)}$$

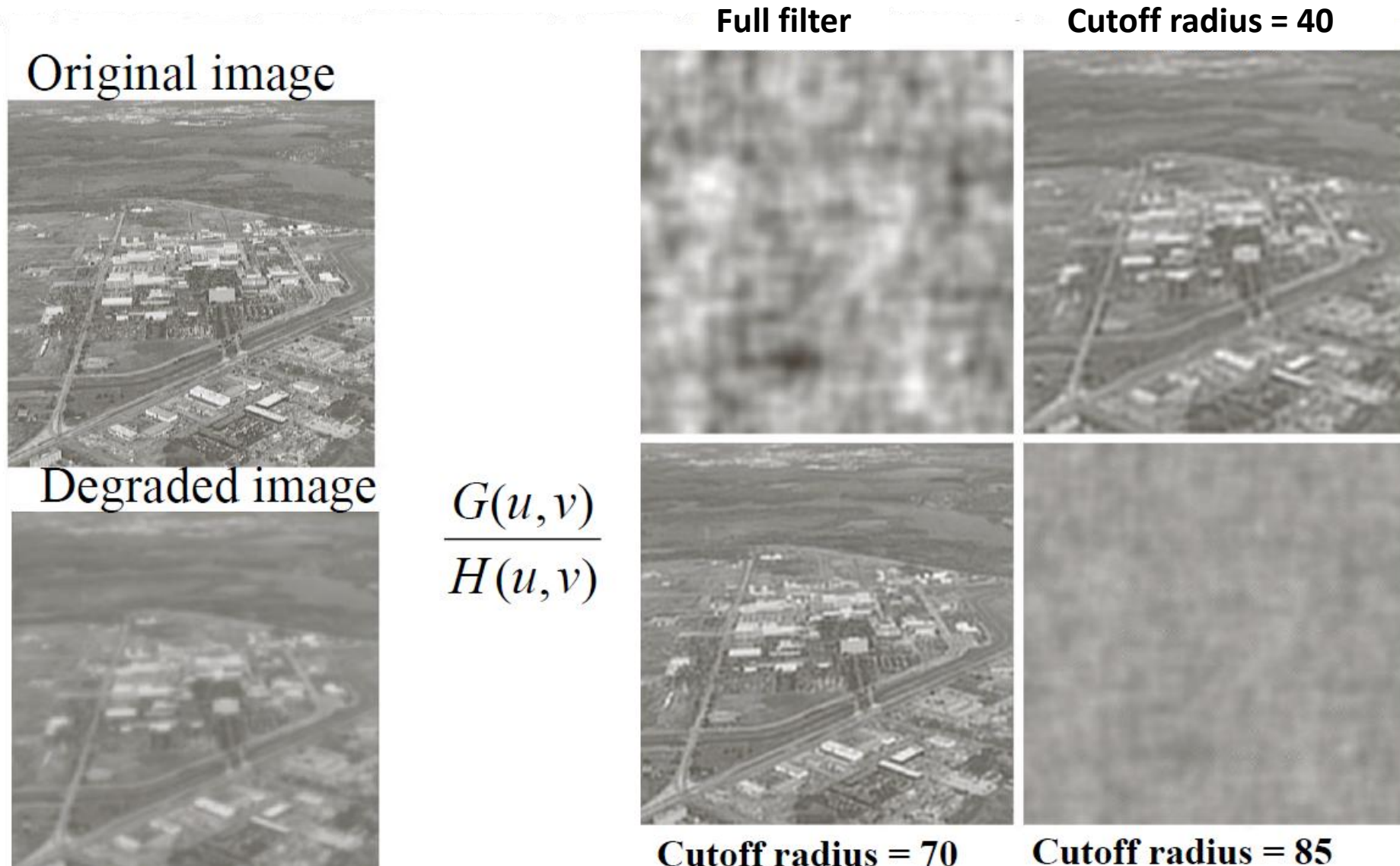
That is, the degradation function used was

$$H(u, v) = e^{-k[(u + M/2)^2 + (v - N/2)^2]^{5/6}}$$

with $k = 0.0025$. The $M/2$ and $N/2$ constants are offset values; they center the function so that it will correspond with the centered Fourier transform

$$M = N = 480$$

An Example of Inverse Filtering



Cutting off values of the ratio $G(u,v)/H(u,v)$ outside a radius of 40, 70, and 85, respectively

The cut off was implemented by applying to the ratio a Butterworth lowpass function of order 10.

$$H(u,v) = \frac{1}{1 + [D(u,v) / D_0]^{2n}}$$

n=10

$$D(u,v) = [(u - M/2)^2 + (v - N/2)^2]^{1/2}$$

D₀ is radius

Slide credit: Yan Tong