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# The Final Report of Spectrometer Experiment (Group 4)

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## Abstract

In this experiment, the spectrometer was made by using the optical disc as the diffraction grating. The Planck constant was measured by using the LED lamp, and the absorption rate of copper sulfate solution was explored. At the same time, the problems related to the grating properties and diffraction phenomena are studied.

## 1 Introduction

### 1.1 Literature Review

In this experiment, we will mainly study the influence of measuring the solution on the light absorptivity with a self-made spectrometer, as well as some scientific facts in the DIY process and their influence on spectrometer. Therefore before we start our experiment, we first do some research and literature review in several topics: Diffraction Grating of Slits, CD and DVD discs, Planck's constant, LED.

#### 1.1.1 Diffraction Grating by slits

Diffraction is a phenomenon caused by the wave principle of light.<sup>1</sup> When the light enters the slit or the edge of the object, according to the explanation of Huygens wavelet principle, the light forms countless wavelets, and each wavelet forms a spherical wavefront origin at the slit and spreading in the whole space. On the projection screen, countless wavefronts form superposition at a certain angle to form a brighter spot and form diffraction fringes. In multiple-slit diffraction, more clear and uniform fringes can be formed than single slit diffraction. Because the position of the diffraction spot of light is related to the optical path difference (see formula), it is often related to the angle between the optical axes in the experiment. At the same time, because different colors of light have different wavelengths, it leads to fringes in different optical path differences. In the experiment, fringes of different colors are produced at different angles. We can separate light of different wavelengths.<sup>2</sup>

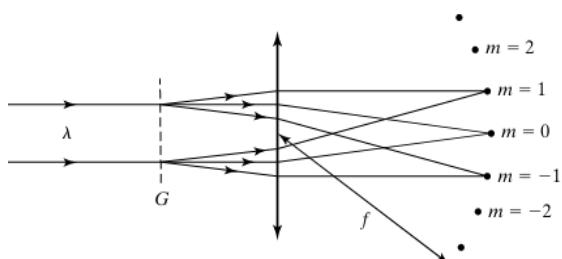


Figure 2: Distribution of the diffraction

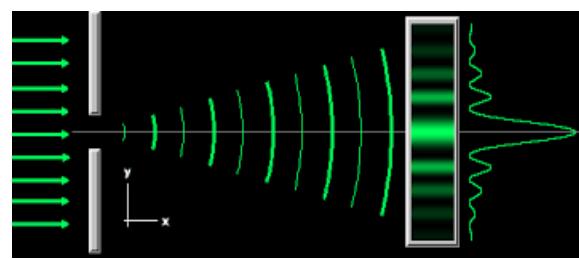


Figure 1: Phenomenon of Diffraction(Simulation)

$$a(\sin \theta_m) = m\lambda, m = 0, \pm 1, \pm 2, \dots$$

<sup>1</sup> Image from <http://electron9.phys.utk.edu/optics421/modules/m5/Diffraction.htm>

<sup>2</sup> Image from Introduction to optics, Pedrotti

Where  $m$  represent the index numbers of the diffraction fringe, that the center corresponding to  $m = 0$ . And  $\theta_m$  represent to the angle of the diffraction fringe corresponding to each index number of diffraction fringe.

In this process, we use angular dispersion to describe the ability of the medium to separate light of different wavelengths, which is often calculated by the following formula<sup>2</sup>.

$$\mathfrak{D} \equiv \frac{d\theta_m}{d\lambda} = \frac{m}{a \cos \theta_m}$$

<sup>3</sup>Where,  $a$  is grating constant or groove separation, the distance between each wavelet. For multiple slits diffraction it is the reciprocal of the density of the slits.



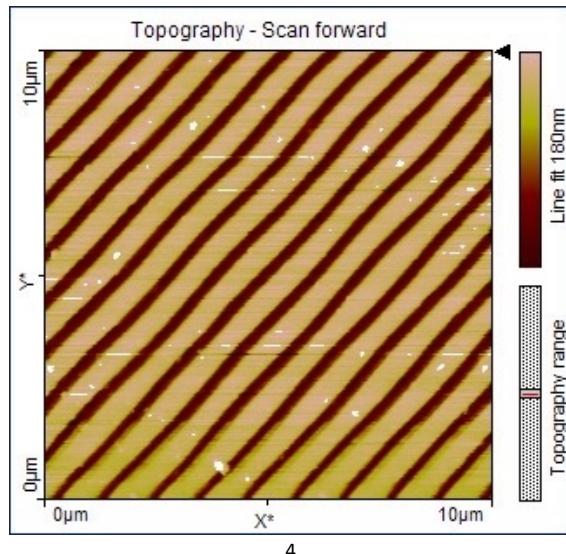
Figure 3 Optical Principle of diffraction Grating Spectrometer

### 1.1.2 CD and DVD discs

Next, we found that the plastic after removing the reflective film on CD / DVD was a very typical multi slit diffraction grating. They are composed of countless dents with equal spacing, which are highly similar to professional diffraction gratings and are easy to obtain. Therefore, in this experiment, we use the fragments of optical disc as the diffraction grating in the experiment.

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<sup>3</sup> Image from <https://www.nnn.org/>



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*Figure 4 AFM image of unburned DVD*

### 1.1.3 Planck's Constant

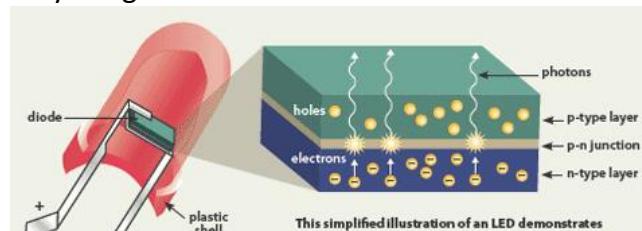
As one of the most important constants in modern physics, the Planck's Constant 'h' becomes the boundary between macroscopic and microscopic objects. It used to describe quantum size and shows for the first time the discontinuity of radiated energy.<sup>3,4</sup>

Planck believed that matter could radiate only an integral multiple of some minimum energy. Such a minimum energy is called a quantum of energy, each of which is  $E = h\nu$ .

### 1.1.4 LED light-emitting principle

Since the core parts of LEDs are wafers made up of P-type and N-type semiconductors. And there exists a transition layer between P-N junction and P-N junction.<sup>5</sup>

<sup>5</sup>As the p-N junction of the semiconductor material is applied to a forward voltage, the combination of injected minority and majority carriers releases excess energy as light, converting electrical energy directly into light energy. As the current flows from LED anode flows to the cathode, the semiconductor body emits different colors lights from ultraviolet to infrared. And the intensity of light in different colors is related to current.



*Figure 5 Principle of LED emitting*

<sup>4</sup> Image from <https://www.nnn.org/>

<sup>5</sup> Image from <https://www.xedlight.com/blog/post/what-are-the-disadvanges-of-led>

### 1.1.5 Absorption spectrum of copper sulfate

In an experiment carried out by the Los Angeles City College, aimed at finding out the relationship between concentration of the solution and the absorbance of solution, which draw out a colored solution's distribution of absorbance line is highly related to its color and the absorbance and concentration is of linear relationship<sup>6</sup>.

In this experiment, two wavelengths of light are used, one is close to the actual light the color CuSO<sub>4</sub> shows (500nm, shallow blue), another is far from its original color (650nm, red). For the blue-colour CuSO<sub>4</sub> solution, theoretically, the absorbance of 500nm light will be significantly lower than the 650nm light.

### 1.1.6 Beer–Lambert law

The Beer–Lambert law, also known as Beer's law, the Lambert–Beer law, or the Beer–Lambert–Bouguer law relates the light to the properties of the material through which the light is travelling.

$$A = \varepsilon \ell c$$

- $A$  is the absorbance
- $\varepsilon$  is the molar attenuation coefficient or absorptivity of the attenuating species
- $L$  is the optical path length in cm
- $c$  is the concentration of the attenuating species

$$c = \frac{\mu_{10}(\lambda)}{\varepsilon(\lambda)}$$

## 1.2 Objective

### 1.2.1 Diffraction Grating by slits

- To investigate the relation between grating constant and grating angle.
- To determine the grating constant of CD and DVD discs.
- To determine the disc that is more suitable using in Spectrometer.
- To verify the diffraction equation.

### 1.2.2 Planck's Constant with LEDs

- To investigate the relation between LED forward voltage and LED wavelength.
- To measure the value of the Planck's Constant with LEDs bulbs series and Spectrometer.

### 1.2.3 Investigate the absorption spectrum of copper sulfate solution

- To determine the absorption spectrum of copper sulfate solution.
- To compare the absorptivity of different wavelengths
- To investigate the changes of absorbance of light with specified wavelength in different concentrations of solution
- To determine the wavelength (color) of maximum absorbance for a copper (II) sulfate solution.



## 2 Experiment Setup

### 2.1 Experiment 1: Diffraction grating of CD/DVD discs

In this experiment we have some simple instrument to conduct the experiment:

#### 2.1.1 Projection Screen

In the experiment, we use a piece of cardboard as the projection screen of the experiment, use a bracket to fix it vertically on the desktop, and apply a small amount of glue on the place facing the experimental equipment to fix the coordinate paper. We use weights to fix

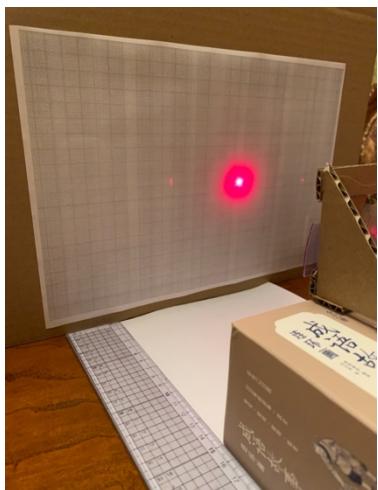


Figure 6 The Projection Screen with Red Laser emitting



Figure 8 The distance ruler

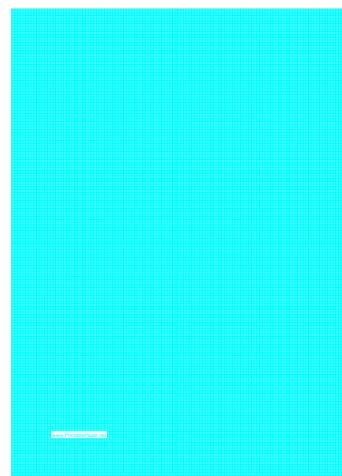


Figure 7 The Blank coordinate paper

the screen position to prevent displacement when marking later.

#### 2.1.2 Coordinate Paper

In the initial experiment, we found that because the slits of the disc are not straight lines, but rings, it is difficult to ensure that the grating is perpendicular to the ground and the light spot is located on a straight line parallel to the ground. Therefore, here we use two position coordinate paper to replace the scale commonly used in traditional experiments as the recording medium of experimental results.

In the experiment, we use coordinate paper with an interval of 1mm to record the abscissa and ordinate position of the light spot in the experiment. Before the experimental measurement, we record the test parameters on the coordinate paper for statistics later. We recorded the results of the same experiment under different conditions.

#### 2.1.3 Distance Ruler

In the experiment, the distance between the grating and the screen is one of the control variables. Therefore, in this experiment, we use a 30cm scale pair to adjust the distance between the grating platform and the screen. The 0-scale line of the scale is close to the projection screen. And keep it perpendicular to the screen.

#### 2.1.4 Grating platform

In the experiment, we need to fix the grating (disc fragments) vertically on the desktop and control its height from the desktop and the distance from the screen. Therefore, in the experiment, we made a simple grating platform with cardboard. The platform can install two (30mmx30mm) square disc fragments at the same time. At the same time, the platform fixes the grating at the height of 10cm, about half of the screen height, to ensure that all the light spots can fall on the screen.



Figure 9 Grating Platform

#### 2.1.5 Grating / disc fragment

In the experiment, we cut the disc with an art knife and remove the reflective coating on it, leaving only the plastic information layer. The fragments are cut into a square size of 30mmx30mm, in which a pair of edges are parallel to the radius of the ring, which will help to ensure the horizontal distribution of the light spot as much as possible in the later design.

### 2.2 Design of Spectrometer

In this section, we will discuss the design of the spectrometer.

According to experiment 1, we determined that DVD has stronger diffraction dispersion, that is, for light with different wavelengths, DVD can disperse them more widely, which will help to improve the accuracy of spectrometer. Therefore, in spectrometer, we use DVD as the grating. In spectrometer, we use a 720p camera module as the data acquisition terminal. We remove the infrared filter on the camera to reduce the transmittance of long wavelength visible light and improve the accuracy of the experiment. At the same time, the camera uses manual focusing to avoid the blur caused by automatic focusing on the dark spectrometer.

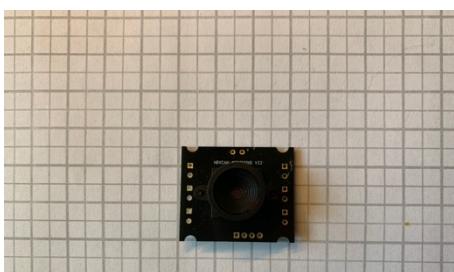


Figure 12 Camera Module

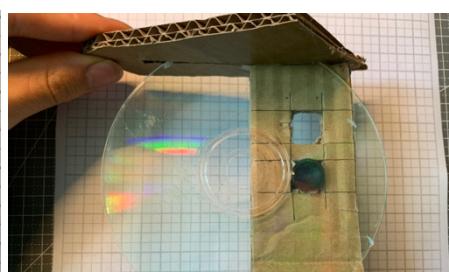


Figure 11 Setup of Disc and Camera



Figure 10 Structure of prototype

Since the use distance generally does not change, manual focusing has higher stability and accuracy.

The camera module and the DVD grating are placed close together to ensure that the camera can capture clear diffraction fringes. Both are placed vertically in the chamber with a cardboard platform that can be rotated to allow us to adjust the pointing of the camera so that clear and accurate diffraction fringes can be captured in the camera.

Firstly, according to Theremino's DIY experimental guide,<sup>7</sup> we made a prototype of spectrometer in a carton, as shown in the figure. Here we use cardboard to make a slit so that the light can keep the vertical incidence as much as possible.

The light will pass through the inclined grating through the slit, and the light of different colors will be diffracted to different angles and finally recorded by the sensor of the camera module. The following stripes are obtained.



Figure 14 The view of camera directly toward the slit

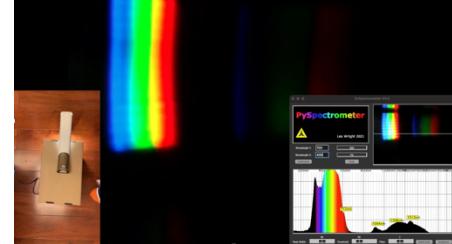


Figure 13 Testing of prototype

In the software, we use the python program created by Les Wright.[7] He obtains a luminous intensity spectrum by measuring the light brightness of each position on the central horizontal line of the camera, as shown in the figure. After calibration with red laser and green laser, each point corresponds to different wavelengths of incident light one by one. Finally, we will get a data set of light intensity spectrum of incident light.

It is worth noting that the obtained light intensity spectrum is a relative value calculated based on the brightest point as 100%, so we need to change the process of data analysis during the subsequent data analysis, which will be shown later in the presentation.

After testing the prototype, we found that it is feasible to experiment with this principle, and we can get some key data we need:

- Length and width of slit
- Angle between camera and filter group and incident light
- Relative position between camera and slit

Therefore, we can use these data to draw the 3D drawing of our formal spectrometer. The spectrometer will be divided into several different parts. Through assembly, we can flexibly adjust the instrument.

- Main body
- Spectrometer cover
- Camera dock
- Disc dock

### 2.2.1 Main body

The main body is the main part of the spectrometer, which constitutes the cavity of the spectrometer and has the connection parts with various components on it. The head of spectrometer is a protruding circular ring, which is used to connect with the test module. The protruding part of the shear ensures that there will be no interference of ambient light at the connection. In the middle of the circular ring is a slit, which filters out the obliquely incident light.

There is a capsule shaped mounting slot at the back end of the main body, which will be fixed with the base of the camera dock. At the same time, it allows the camera to move and rotate left and right, and adjust the imaging position of the incident spectrum in the camera. A set of magnetic ports for connecting with other parts are reserved at the upper end and front end of the main body. In this design, we use magnets to connect each part. It helps to adjust the connection strength and ensure the stability of the connection.

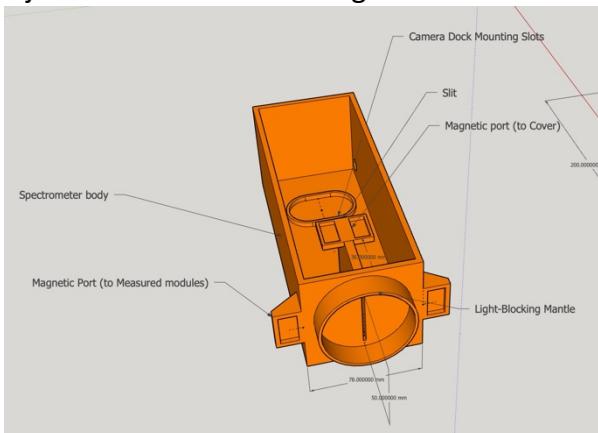


Figure 16 Main Body

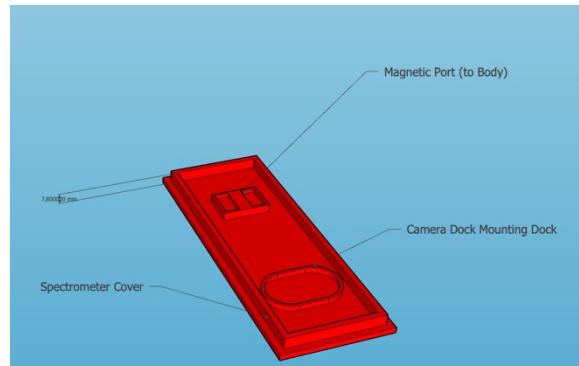


Figure 15 Spectrometer Cover

### 2.2.2 Spectrometer cover

The spectrometer cover is connected with the main body to form a closed cavity of the spectrometer to prevent the influence of ambient light on the internal sensor. Like the main body, the cover has a mounting slot for fixing the camera dock and a magnetic port for fixing with the main body. The four sides of the cover extend outward, so that the ambient light cannot enter through the gap.

### 2.2.3 Camera Dock

The camera dock consists of two circular bases and a middle fixing plate. The space and screw hole for inserting the camera module are reserved on the middle fixing plate, as well

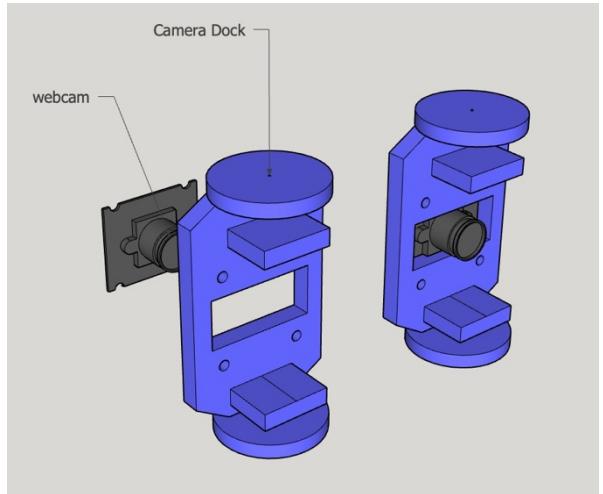


Figure 18 Camera Dock

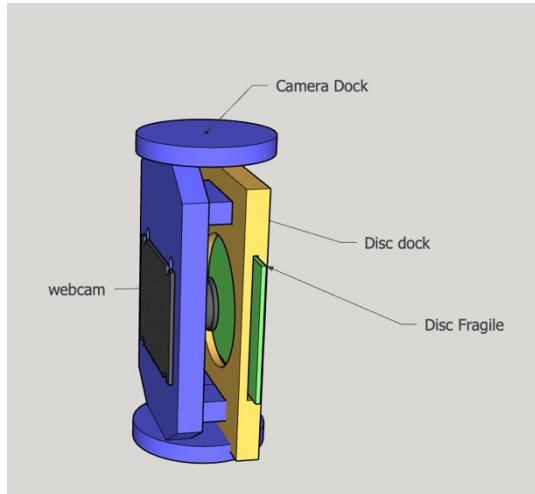


Figure 17 Disc Dock mounting on the Camera Dock

as the fixed square column for connecting with disc dock.

### 2.2.4 Disc Dock

The disc dock has three openings of different shapes, with square holes connected with camera dock on both sides and a slit for fixing the grating in the center. After the disc dock is installed on the camera dock, it still has the activity of moving back and forth. To adapt to the expansion and contraction of the lens during manual zoom.

### 2.2.5 Others

All experimental parts are printed in 3D with white resin, which has certain light transmittance. In order to avoid the influence of ambient light. All the other parts are covered with light blocking tape, so that the light cannot pass through.



Figure 19 The shading tape is covered on the inside of the equipment

## 2.3 Parameters of the spectrometer

### 2.3.1 Inclined angle between grating and input light

In the application process we need to adjust the Camera Dock to an appropriate angle to the slit so that the dispersed spectrum can enter the camera lens vertically, so we can analyze it by the above schematic.



*Figure 20 The angle of the grating and camera could adjust in the spectrometer*

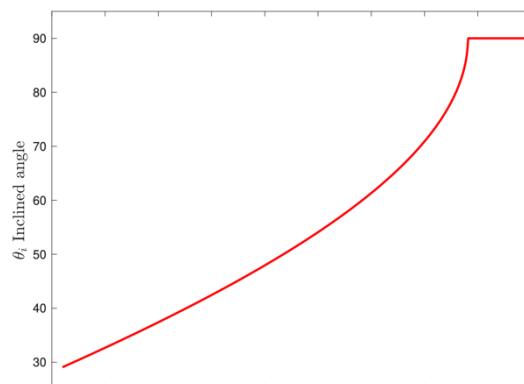
Here we assume we using the laser light with wavelength  $\lambda = 650\text{nm}$  emitting into the spectrometer and we require the one of the diffraction light ( $m=1$ ) of it perpendicularly shooting into the camera. Next we can calculate the incline angle between the incident light and the normal line of the grating. The equation could be list that:

$$a(\sin \theta_i + \sin \theta_m) = m\lambda$$

Where  $\lambda = 600\text{mm}$ ,  $m = 1$ ,  $\theta_m = 0$ . Therefore we can turn out the incident angle as:

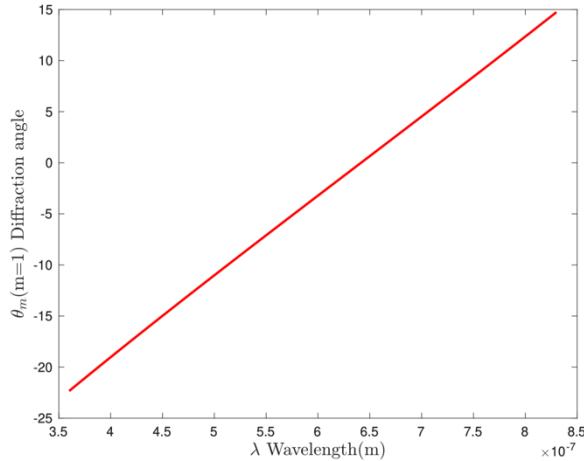
$$\theta_i = \arcsin \frac{m\lambda}{a} = 54.0959^\circ$$

Since there is a range in the spectrum, so that not all wavelengths of light can be directed perpendicularly into the center of the camera, but necessarily along different directions, we can plot the relationship between the images that can make different wavelengths of light emitted and the tilt angle.



*Figure 21 Where  $\lambda=600\text{mm}$ ,  $m=1$ ,  $\theta_m=0$ . Therefore we can turn out the incident angle as:*

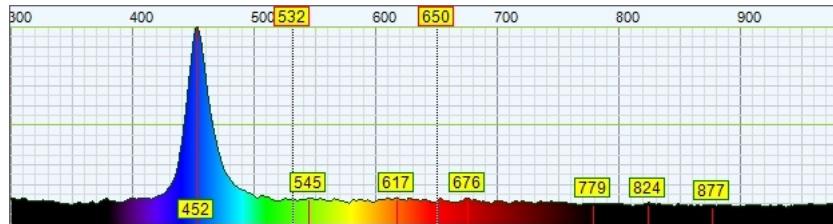
The visible spectrum is between about 400nm and 800nm, so here we can use the angle of 54 degrees of vertical intake at 600nm as the tilt angle, and here we can plot the corresponding angle change when the corresponding light is diffracted out. It can be seen that the two show a linear variation relationship, this is helpful for software to analyze the data.



### 2.3.2 Resolution

The resolution of a Spectrometer is influenced by many factors, such as the slit width of the mainbody, the density of the optical disc grating, the resolution of the sensor, etc. Therefore, we make a simple estimation here by using the results obtained from actual measurements.

After calibration, we placed a blue LED with a wavelength of about 450 nm at the front slit of the spectrometer and obtained the following image.



We can calculate the resolution from the obtained data as follows

$$\Delta\lambda_{\min} = 2\Delta\lambda_{1/2} = 20.90 \text{ nm}$$

And the resolution could be obtained that:

$$\mathfrak{R} \equiv \frac{\lambda}{(\Delta\lambda)_{\min}} = 21.64$$

### 2.3.3 Free Spectral Range

Firstly, due to the limitation of the sensor, the spectrometer can only measure in the visible region, and we can determine whether the spectrometer can cover the whole visible region by calculation.

m	800nm	400nm
0	-54	-54
1	15.7228	-15.6058
2	exceed(maximum at 670nm in 90)	15.7228

Through the calculation we can find that the spectrometer can cover the entire visible interval, but at the same time there may be overlap at 400nm at m=2. However, since the intensity of diffraction decays sharply with angle, there is no need to worry about overlap here, which affects the experimental results.

#### 2.4 Experiment 2: Measure the Planck constant by LED

The LED module is relatively simple. There are two guide rails on it. These two guide rails correspond to the screw holes on the back of the mini bread plate. The LED can be fixed in the same straight line of the slit, and the light released by the LED can enter the spectrometer vertically.

In the experiment, a series connection method is used to connect multiple LED lamp beads. This method not only improves the brightness of the experiment, but also widens the gap between the forward voltages of LED lamps with different colors, which is conducive to the error caused by Voltmeter measurement. There is a space on the side wall of the module to connect internal and external circuits. The voltmeter can measure the forward voltage of

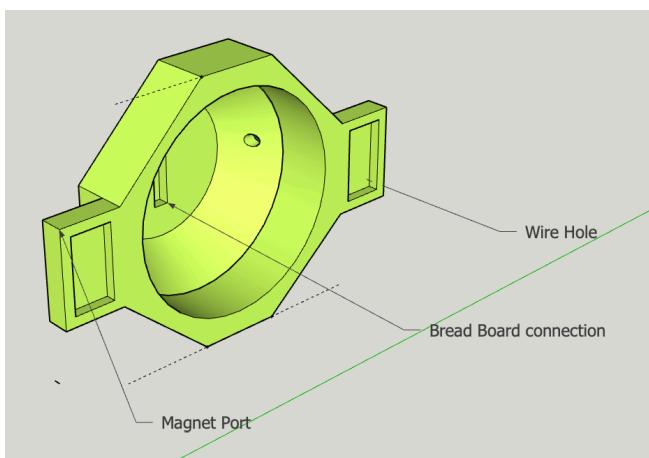


Figure 23 LED Module

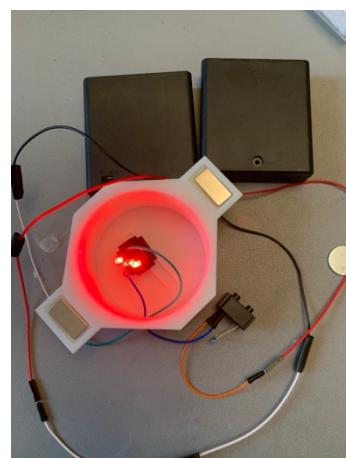


Figure 22 Circuit mounted on the LED module

the internal LED by measuring the port left outside.

#### 2.5 Experiment 3: Investigate the absorption spectrum of copper sulfate solution

The cuvette module is composed of a front cover, a rear plate and a cuvette shelf in the middle. The front cover has a port connected with the spectrometer, and the rear cover is equipped with a lighting circuit and a white LED light board to provide light. The cuvette shelf in the middle can fix five cuvettes for measurement at the same time. The cuvettes measured by spectrometer can be adjusted by sliding and translating the rods on both sides to realize the measurement of multiple experimental sides without disassembly, which can greatly improve the efficiency.

Here, the white LED light board is used to provide light, and the white LED light board can evenly and stably blue and red light, which will help to explore the absorption rate in Experiment 3.

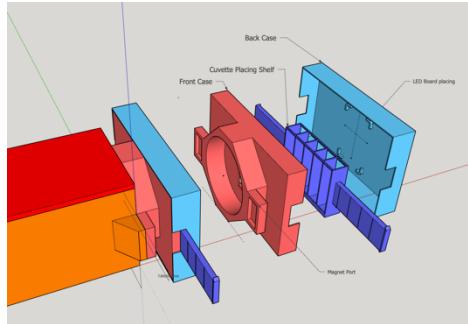


Figure 25 Cuvette Module

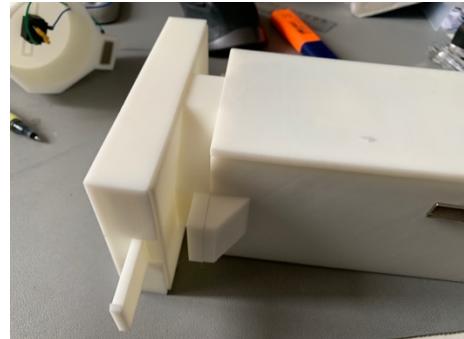


Figure 24 Cuvette Module mounting with spectrometer

### 3 Diffraction grating of CD/DVD discs

#### 3.1 Hypothesis

- Other conditions remain unchanged, the shorter the wavelength of the laser, the closer the spot position.
- When other conditions remain unchanged, CD has longer distance of spots compare with the DVD.

#### 3.2 Methodology

1. Firstly, conduct preliminary experiment to confirm the test parameters that can conduct the test normally and accurately.
  - In the experiment, the distance before the spot ( $m = 0, \pm 1$ ) is too far, so that the coordinate paper cannot completely record the positions of the three points, making the data incomplete.
  - In the experiment, if the distance between the light spots is too close, the data is inaccurate due to the size of the light spot, and the error of the ruler.
  - Therefore, after the preliminary test. We have confirmed several groups of test parameters that can be tested, as shown in the table below (Experimental group)

Wavelength(nm)	Grating	5	10	15
532	CD			
	DVD			
650	CD			
	DVD			
		Spots too close		
		Spots out of range		
		Able to conduct the experiment		

2. Then we use these data for grouping experiments.

3. Firstly, record the experimental control variables on the coordinate paper, including grating type (CD / DVD), grating screen distance (50 / 100 / 150mm) and laser type (650 / 550nm). It shall be pasted and fixed in parallel along the edge of the projection screen to ensure that the coordinate paper is pasted flat enough.
4. Fix the light screen with heavy objects. Press the scale vertically against the screen.
5. Fix the grating platform on the point marked by the scale. Make the grating a given distance from the screen.
6. Use the specified laser vertical incidence grating and keep it fixed.
7. Use a fluorescent marker to mark the position of the laser spot on the screen, which is generally 3 spots, and there is the possibility of 5 spots at the same time. Repeat the experiment with markers of different colors. During this period, the grating platform can be moved horizontally parallel to the screen plane, or the grating can be rotated to make the light spot fall at different positions of the coordinate paper and reduce the error in the experiment. 3-5 groups of experiments were carried out for each experimental parameter.

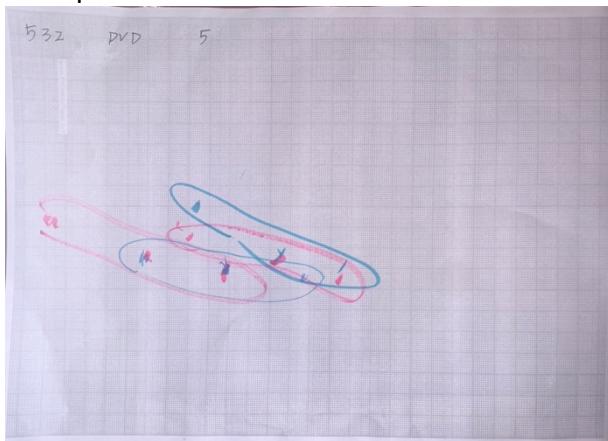


Figure 26 One coordinate paper with result

8. After the experiment with the same experimental parameters. Take off the coordinate paper with the experimental results, take photos and archive, and carry out the next group of experiments.
9. All experimental parameters. After the experiment, start the transcription and analysis of the test results. See the data analysis section.

### 3.3 Data Analysis method

In this part, the recording and analysis of test results will be introduced. Matlab script will be used for automatic calculation in the process of data calculation.

1. The coordinate paper recording the experimental results will have several groups of marks of different colors. First, we confirm the center point of each group of marks, which corresponds to the center spot ( $m = 0$ ) in the diffraction spot. The next two marks correspond to the spot of  $m = \pm 1$  respectively. Therefore, to calculate the distance between the two, we can first take the center point as the origin, record the relative positions, horizontal and vertical coordinates of the surrounding two points,

and record these distance coordinates in the same column of the CSV document with the experimental parameters.

Each column represents one experiment, and the recorded results are as follows:

2. Then we can calculate the diffraction angle  $\theta_m$  through the average value of the distance (d) between the light spots and the distance (L) between the grating and the screen,

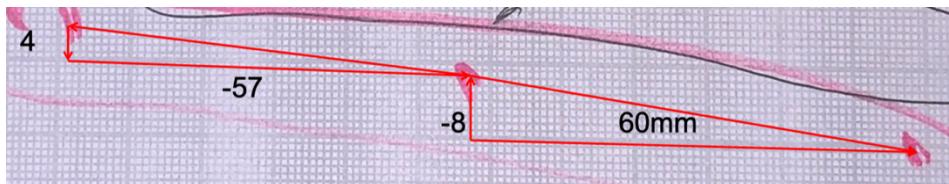
$$\theta_m = \arctan \frac{d}{L}$$

3. Next, we can calculate the grating constant of the disk, where we use the diffraction formula in the literature review.

4. So, we can calculate the required results.

For example

Here we list out one experiment and its result to prove the feasibility of experiment



$$d_1 = 60.310, d_2 = 57.1402, L = 150\text{mm}, \lambda = 532\text{nm}$$

Therefore

$$\theta_1 = 0.373158$$

And we can turn out the diffraction

$$a = 1.45930 \times 10^{-6}$$

The reciprocal is

$$a^{-1} = 685259/\text{m}$$

The standard CD has 625 rows per 1mm. Therefore we can find that the result is very close to the exact value.

### 3.4 Result Analysis

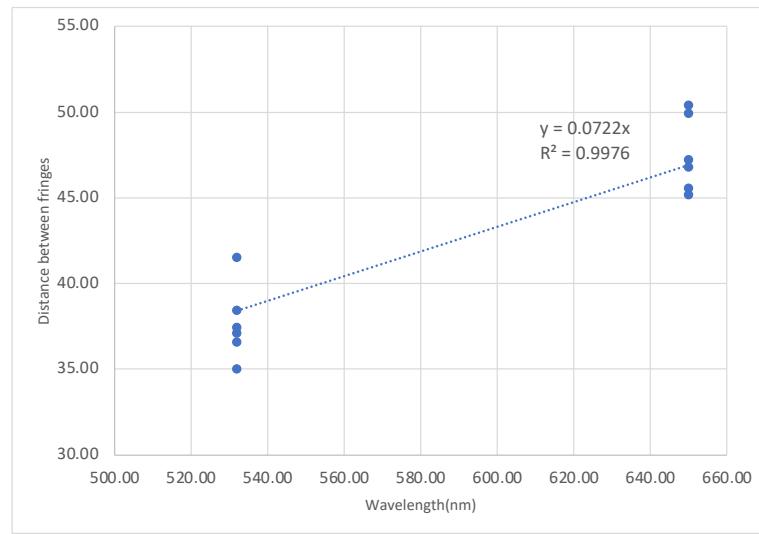
We organize the data to get the result below, the detailed graph could refer to the appendix or Github. And we can find the measured average slit density below, as can be seen, that the average DVD is about 1250/mm compared to the actual result of 1350. The error is small, within the acceptable range. As for CD, the measured value is about 660/mm compared to the actual result of 625. Also in the acceptable range.

	DVD	real	CD	real(/mm)
av	1249.81246	1350	667.314071	625
stdev	69.90020042		-29.261238	

Table 1 The Final result of Experiment 1

Next, we can analyze the hypothesis, first, we compare the result with CD and distance at the same wavelength laser light. The result is below. Clearly that the disc-screen distance and the distance between spots is linear, and when the disc screen distance increase, the spots

distance will increase linearly. The Hypothesis was proved.



## 4 Measure the Planck constant by LED

### 4.1 Hypothesis

- The shorter the wavelength of light, the greater the voltage across the LEDs (forward voltage).

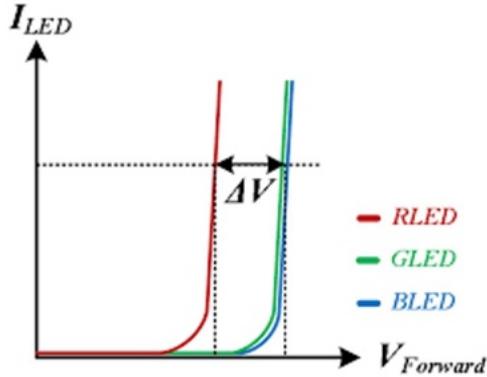


Figure 27 the forward voltage with different colors lights

### 4.2 Methodology

- In this experiment, we use multiple LEDs of the same type in series and measure their total potential to reduce the error due to with the multimeter as the measurement tool. And to save cost, here we use 8 households 1.5V batteries as the power supply, which constitute a 12V DC voltage.
- The figure below demonstrates the circuit simulation of this experiment in Multisim. We can see that when different switches are closed, different voltages can be obtained, which is the result of multiple LEDs forward voltage.
- Here we mainly measuring the forward potential, which is the amount of voltage needed to get current to flow across a diode. After the forward potential get flow, the current will increase dramatically, but the voltage will increase slowly. Here we don't have varistor to adjust the potential between the LEDs, instead we measure the voltage which is a bit larger than the forward voltage, the error is around 0.05V, which is in acceptable range.



Figure 29 Circuit simulation of this experiment

#### 4.2.1 Specific Experiment Steps:

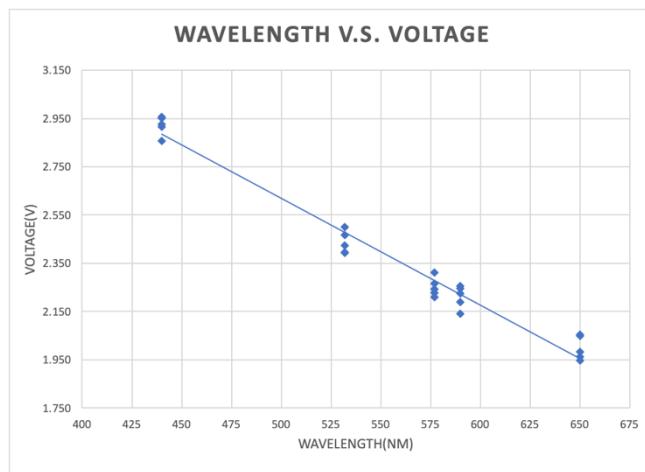
1. Calibrate the voltmeter
2. Calibrate the spectrometer with a red laser and a green laser
3. Install the specified number of LED lights of the same color in series on the breadboard and fix the breadboard on the testing module. Connect the power supply through the wire, make sure the circuit is connected correctly and the LED lights up normally.
4. Connect the Module to the spectrometer, switch on the power. Then record the experimental spectrum after the software displays the stable light intensity spectrum.
5. Record the voltage at both ends of the LED lights and record the data.
6. Turn off the power, turn it on again, repeat the experiment, repeat the above operation, record more than 3 groups of data to reduce the error.
7. Replace the LED lights with different colors and repeat the above operation.

#### 4.3 Data Analysis method

- After obtaining the measured voltage value, divide it by the number of LED lights  $U_0 = \frac{U}{n}$ , here  $U$  is the voltmeter reading,  $n$  is the number of LED lights, and the  $U_0$  is the final voltage we need.
- Then substitute the result into the formula:  $E = qU$ , here the value of the constant  $q$  is  $1.6 \cdot 10^{-19} C$ , then we get the energy.
- Since that  $E = hf$ , and we already know the wavelength of the LED lights, with the formula  $f = \frac{1}{\lambda}$  we can get the value of frequency.
- Finally using  $E = hf$  we can obtain the value of  $h$ , which is the Planck's Constant.

#### 4.4 Result Analysis

First, we noticed that the voltage had a linear relationship with the wavelength, which proved the relationship between the two presented in the formula.



We can through the observation of error in data evaluate the scientific rigor of the experiment, through calculation we can find in our final experiments, the standard deviation control within 5%, this shows that the experimental data is constant, the data is effective, we found that the relative theory of relative error is only 5% of the Planck constant, this proves that the experiment is scientific and accurate.

At the same time, we can notice that in the experiment, the Planck constant result of the blue-purple lamp has a larger relative error than other lamps, which may be due to the internal defects of this type of bulb during production, such as resistance, leading to the measured voltage is greater than our theoretical calculation value. Finally, the Planck constant is larger.

## 5 Investigate the absorption spectrum of copper sulfate solution

### 5.1 Hypothesis

- Under certain wavelength of light (500nm and 650nm), the absorbance of CuSO<sub>4</sub> solution is positively related to the concentration of the solution, i.e., the higher the concentration of the solution is, the more significant the absorbance will be.
- Under certain concentration of CuSO<sub>4</sub> solution. The absorbance of light at 650nm is higher than the absorbance of light at 500nm

### 5.2 Methodology

#### 5.2.1 Material

Spectrometer, LED white light source, cuvette, deionized water, volumetric flask, 4 set of CuSO<sub>4</sub> solution separately with concentration of 0.1 mol/L, 0.5 mol/L , 1 mol/L and 2mol/L



Figure 30 Cuvette(10mm)



Figure 31 LED panel



Figure 32 Copper sulfate crystal

#### 5.2.2 Solution Preparation

- i. Calculate the amount of CuSO<sub>4</sub>, and deionized water needed.
- ii. Weigh out the amount of material determined in step i.
- iii. Prepare the volumetric flask, add in the solute (CuSO<sub>4</sub> Crystal) first.
- iv. Add in the deionized water about 2/3 full and mix the solution.
- v. Add in the deionized water until the water level reaches the mark on the flask accurately.
- vi. Mix the current solution again and obtain the solution with accurate molar concentration.

#### 5.2.3 Experiment Procedure

- i. Turn on the spectrometer, and turn on the LED panel in the testing module.
- ii. Insert the cuvette with deionized water into the spectrometer, assure there's no dirt or fingerprint on the cuvette, if there is, wipe them off with light duty wipe.
- iii. Record the spectrum of the deionized water.
- iv. Repeat the experiment step ii and iii with another cuvette filled with each concentration of solution.

vi. Each solution or solvent must be measured at least 3 times to assure the accuracy.

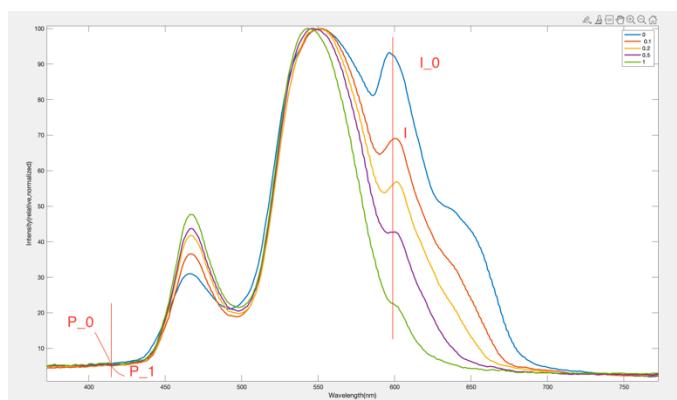
### 5.3 Data Analysis method

- i. Mark certain wavelength (424nm and 650nm) on the spectrum recorded in the experiment.
- ii. Read the absorbance in such wavelength, i.e., the intensity of certain wavelength of light.
- iii. Fill the data into such analysis form below:
- iv. Calculate the relative intensity loss caused by the increase of concentration, compare the data, and draw the solution.

Wavelength/	424nm	650nm
Intensity		
Concentration		
Deionized water		
0.1 mol/L Solution		
0.2 mol/L Solution		
0.5 mol/L Solution		
1 mol/L Solution		

Table 2 The experiment Sheet

Here we use the algorithm to calculate the absorbance:



Where:

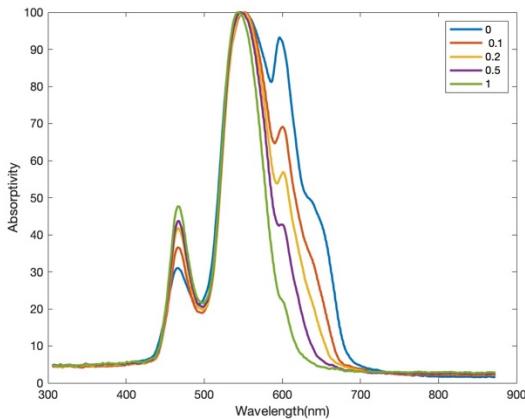
- P0: the intensity of background noise in pure water
- P1: the intensity of background noise in Xmol/L copper sulfate
- I0: the intensity of certain wavelength in pure water
- I1: the intensity of certain wavelength in Xmol/L copper sulfate

The absorbance could be written as

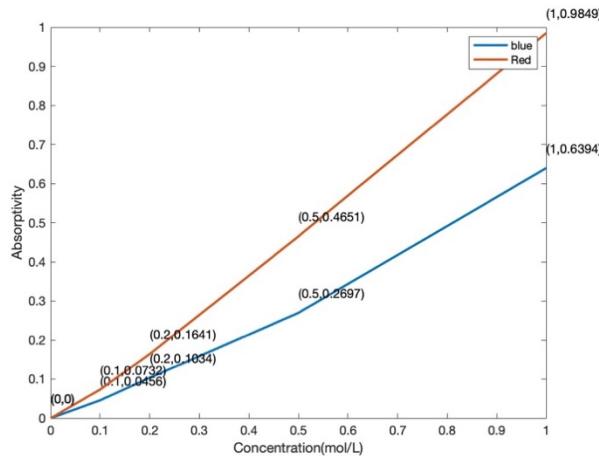
$$A = 1 - \frac{(I - P_1)P_0}{P_1(I_0 - P_0)}$$

#### 5.4 Result Analysis

Here give the original data without modified, which is normalized with the maximum density:



And then we perform some algorithms to determine the absorptance spectrum of each solution. As we can focus on 2 wavelengths, and plot the graph to show the relation between concentration and absorptance:



Since in the experiment, the thickness of the solution is fixed, which determined the optical path length of the light in the solution is fixed. According to Beer-Lambert law, the absorbance is given by:

$$A = \varepsilon lc$$

The  $\varepsilon$  is the molar absorptivity of the substance, in the experiment, this value is fixed since we only use copper sulfate as material,  $l$  is the optical path length, which is also fixed in the experiment. Thus, theoretically, the absorbance is proportional to the molar concentration of the solution (denoted as  $c$  in the equation).

From the concentration graph above, we can observe that both wavelengths of light, blue or red, has their absorptivity approximately proportional to their concentration.

Thus, we can claim that, the regular pattern of the absorptivity of Copper Sulfate solution is accord with the Beer-Lambert law.

Additionally, due to the blue color of the transparent solution, the absorptivity of blue light should be significantly lower than the red light, which we can also observe from the graph, the graph of the red light is significantly steeper than the blue light graph, which indicates a larger molar absorptivity.

Hence, generally, we can state that the absorptivity of the solution is of linear proportional to the concentration of the solution, which accords with the Beer-Lambert Law; and for copper sulfate solution, the molar absorptivity of blue light is significantly lower than the red light, which accords with its blue color and transparency.

## 6 Conclusion and Improvement

### 6.1 Conclusion

#### 6.1.1 Measuring Diffraction Grating of CD/DVD

In the experiment, we can conclude that DVD has larger angular dispersion than CD, and calculate that DVD has more dense slits (CD 600-700/cm, DVD 120-1300/cm), which makes DVD more suitable for light splitting. For spectrometer.

#### 6.1.2 Measuring the Planck Constant

In the experiment, we can conclude that the shorter the wavelength of light, the lower the voltage across the LEDs. With the average value of the experiment data 'h', we get 6.19005, and its relative error is 6.6%.

#### 6.1.3 Absorption Spectrum of Copper Sulfate

In the first hypothesis, we can conclude that, under same wavelength of light, the higher the concentration, the more significant the absorption, and therefore calculate under different concentration, the rate of absorption varies.

In the second hypothesis, we can conclude that, under different wavelength, the copper sulfate solution tends to absorb more light with higher wavelength, and with 2 pairs of solution with different concentration, and with molar absorptivity at 635nm ( $2.81 \text{ M}^{-1}\text{cm}^{-1}$ ), we can calculate the molar absorptivity of the solution, where we may find that comparing with lower wavelength (500nm), the absorption at higher wavelength is larger, which corresponding to the blue colour of the solution.

### 6.2 Further Improvement

There are some aspects still that can be improved in the experiment.

- First, because the DIY spectrometer is used in the experiment, it cannot accurately measure the spectral intensity, only the brightness of stripes photographed by the photographic module can be used to judge the results of normalization, which is difficult to be very precise and scientific.
- In the experiment, the fringe is too thick and overlaps with the surrounding fringe. This is because the incident light is too thick due to the large width of the slit, which forms a large fringe in the diffraction process. Therefore, the experimental results can be accurate by further reducing the width of the slit.

In Experiment 2,

- we can use more LED beads in series and parallel at the same time to improve the sum of voltage and brightness and install an astigmatism plate at the front end of the spectrometer to make the incident light more uniform, so as to reduce the possible problem of distortion of experimental results caused by the uneven incident light in the experiment.

- At the same time, we can choose more LED beads and a wider wavelength range to conduct experiments, which will help improve the accuracy and universality of the experiment and also reduce the errors generated in the experiment by taking the average.

In Experiment 3,

- we can try to use a more uniform light source to irradiate to get more accurate spectral information. Meanwhile, we can adjust the brightness of the LED beads to make the result more accurate and avoid the problem that the camera cannot accurately measure the brightness relationship of each point due to the high brightness. Here LED light will have some wavelength with zero intensity which made the spectrum we plot has huge error, therefore using more uniform light in each wavelength is required to improve the result.
- We can increase the solute concentration again to verify the accuracy of Beer-Lambert Law at high concentrations.
- The software we using in the measurement doesn't perform well because of lack of the function to output the absolute intensity but relative intensity, which cause problem in data processing.

## 7 Experiment Management

### 7.1 Budget

Budget List of Lab Project 2022/04(Group 4)								
Author	Guo Yiming	Update Date	1/7/2022					
Member	Guo Yiming, Yang Ziou, Wang Yiran							
exchange rate(RM-CNY)		1.5	Update Date	7/1/2022				
No	Item	Quanta	CNY Price	RM Price	Total (RM)	reimbursement(1/0)	Purchase Link	Shop Link
	Part 1: Experiment on Diffraction Grating							
1	CD	2	0.00	0.00	0.00	1		
2	DVD	5	2.68	1.79	8.93	1		
3	650nm red laser emitter	1	11.42	7.61	7.61	1		
					0.00			
	Part 2: Make a Spectrometer				0.00			
4	3D Model Printing Fee	1	230.00	153.33	153.33	1		
5	magnet	20	0.30	0.20	4.00	1		
6	blackout carpet	1	22.00	14.67	14.67	1		
7	Camera Module	1	20.80	13.87	13.87	1		
					0.00			
	Part 3: Application 1:Planck Constant				0.00			
8	LED light group(8 color)	1	13.80	9.20	9.20	0		
9	DuPont line	1	0.00	0.00	0.00	0		
10	Bread board	1	0.00	0.00	0.00	0		
11	Battery box	2	0.00	0.00	0.00	0		
12	Universal electric meter	1	0.00	0.00	0.00	0		
13	resistance	1	0.00	0.00	0.00	0		
14	switch	1	0.00	0.00	0.00	0		
					0.00			
	Part 4: Absorption spectrum of solution							
14.00	cuvette cover		5.00	4.10	2.73	13.67	1.00	
15.00	Cuvette(10mm)		10.00	2.90	1.93	19.33	1.00	
16.00	LED Light board		2.00	7.61	5.07	10.15	1.00	
					0.00			
	Others				0.00			
17	Equipment damage(Unexpected expenditure)	1	50.00	33.33	33.33	1		
18	Logistics price	1	20.00	13.33	13.33	1		
	Total				301.43			

Table 3 Budget of Project (update at 1/7/2022)

### 7.2 Distribution of Work

Work	Participants
Hardware Design	Guo Yiming
Software	Wang Yiran, Yang Ziou
Background Research/Literature Review	Wang Yiran, Yang Ziou
The Experiment of Diffraction Grating measurement	Guo Yiming
Plank's Constant with LEDs	Yang Ziou
Spectrum absorption of some solutions (for example, CuSO <sub>4</sub> )	Wang Yiran
Experiment plan design, setup and conduction	Guo Yiming, Wang Yiran, Yang Ziou
Planning project and Budget	Wang Yiran, Yang Ziou
Data analysis	Guo Yiming, Wang Yiran, Yang Ziou

Table 4 Distribution of work

### 7.3 Timeline & Milestone

Milestone and Gantt Chart of Group 4 (2022/04)																
No	Milestone	Week														
		1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
1	experimental plan, and background research															
2	prototype testing and software testing															
3	3D modeling, circuit design, assembly and performance testing															
4	Completed the experiment diffraction grating															
5	Completed the experiment Application 1															
6	Completed the experiment Application 2															
7	data collection and organization analyze															
8	Completed the experimental report															

Table 5 Timeline and Milestone

## 8 Reference

1. Born, M. & Wolf, E. *Principles of optics*. (Cambridge University Press, 2019).
2. Pedrotti, F. L., Pedrotti, L. M. & Pedrotti, L. S. *Introduction to optics*. (Pearson/Prentice Hall, 2007).
3. Blundell, S. & Blundell, K. M. *Concepts in thermal physics*. (Oxford University Press, 2010).
4. Harris, R. *Modern physics: Randy Harris*. (Pearson, 2014).
5. Boylestad, R., Nashelsky, L. & Boylestad, R. L. *Electronic devices and circuit theory*. (Pearson, 2013).
6. Formation of complex ions.
7. Theremino. Theremino Spectrometer Construction.
8. Wright, L. *PySpectrometer 2021-03-05*. (2022).
9. 用单色 LED 光源测量普朗克常数 - 百度文库.  
<https://wenku.baidu.com/view/994884787b3e0912a21614791711cc7931b7782c.html>.
10. Doe, J. & Smith, S. Using Absorbance to Determine the Concentration of CuSO<sub>4</sub>. 4.

## 9 Appendix

### 9.1 Experiment 1 Data

Index	Wavelength	Frequency	Voltage	Number	V_av	Planck constant				
1	650	4.61538E+14	7.85	4	1.963	6.81538E-34	Planck constant av	6.94143E-34	Absolute Error	3.15359E-35
2	650	4.61538E+14	8.20	4	2.049	7.11245E-34	Planck constant stdev	1.69397E-35	Relative	0.047593681
3	650	4.61538E+14	7.79	4	1.948	6.76356E-34				
4	650	4.61538E+14	7.94	4	1.984	6.8871E-34				
5	650	4.61538E+14	8.21	4	2.054	7.12866E-34				
6	590	5.08475E+14	9.02	4	2.255	7.10474E-34	Planck constant av	6.86417E-34	Absolute Error	2.38099E-35
7	590	5.08475E+14	8.57	4	2.142	6.74863E-34	Planck constant stdev	9.20313E-36	Relative	0.035933669
8	590	5.08475E+14	8.98	4	2.246	7.07727E-34				
9	590	5.08475E+14	8.76	4	2.189	6.89723E-34				
10	590	5.08475E+14	8.90	4	2.225	7.00967E-34				
11	577	5.19931E+14	8.97	4	2.243	6.91285E-34	Planck constant av	6.91954E-34	Absolute Error	2.9347E-35
12	577	5.19931E+14	8.84	4	2.210	6.81073E-34	Planck constant stdev	1.32358E-35	Relative	0.044290202
13	577	5.19931E+14	9.06	4	2.266	6.98334E-34				
14	577	5.19931E+14	9.24	4	2.311	7.11993E-34				
15	577	5.19931E+14	8.91	4	2.228	6.86533E-34				
16	532	5.6391E+14	9.86	4	2.466	7.00681E-34	Planck constant av	6.93844E-34	Absolute Error	3.12367E-35
17	532	5.6391E+14	9.69	4	2.423	6.88444E-34	Planck constant stdev	1.19651E-35	Relative	0.047142117
18	532	5.6391E+14	10.00	4	2.500	7.10267E-34				
19	532	5.6391E+14	9.58	4	2.395	6.80386E-34				
20	532	5.6391E+14	9.57	4	2.393	6.79993E-34				
21	440	6.81818E+14	11.66	4	2.914	6.84848E-34	Planck constant av	6.91E-34	Absolute Error	2.84435E-35
22	440	6.81818E+14	11.43	4	2.858	6.71594E-34	Planck constant stdev	1.20435E-35	Relative	0.042926701
23	440	6.81818E+14	11.80	4	2.950	6.93212E-34				
24	440	6.81818E+14	11.83	4	2.956	6.94708E-34				
25	440	6.81818E+14	11.71	4	2.927	6.87722E-34				
							Planck constant av	6.94912E-34	Absolute Error	3.23054E-35
							Planck constant stdev	1.36305E-35	Relative	0.048755027

## 9.2 Experiment 2 Data

	CD1/DVD2	wavelength(532/650)	distance(50/100/150)	d1av	d2av	sintheta1	sintheta2	m1	m2	a1	slit density//m	slit density(/m)
1	1	532	150	60.75615997		0.375415067		1	2	1.4171E-06	705667.4187	705.6674187
2	1	532	150	59.25789269		0.367420617		1	2	1.44793E-06	690640.2582	690.6402582
3	1	532	150	54.45906044		0.341264862		1	2	1.55891E-06	641475.3045	641.4753045
4	1	532	150	52.89060916		0.332537389		1	2	1.59982E-06	625070.281	625.070281
5	1	532	150	53.66918917		0.336880566		1	2	1.57919E-06	633234.1465	633.2341465
6	1	532	150	56.86368745		0.354475128		1	2	1.50081E-06	666306.6308	666.3066308
7	1	532	100	36.60208777	102.8840913	0.343720021	0.717086181	1	2	1.54777E-06	646090.2654	646.0902654
8	1	532	100	37.44886408	103.3282103	0.350703565	0.718584912	1	2	1.51695E-06	659217.2282	659.2172282
9	1	532	100	38.43046617	101.9040616	0.358726401	0.713743433	1	2	1.48302E-06	674297.7456	674.2977456
10	1	532	100	41.54817179		0.383682802		1	2	1.38656E-06	721208.2751	721.2082751
11	1	532	100	37.11454959		0.347953306		1	2	1.52894E-06	654047.5677	654.0475677
12	1	532	100	35.02940858		0.330597674		1	2	1.60921E-06	621424.1989	621.4241989
13	1	650	100	49.9398218		0.44678284		1	2	1.45485E-06	687358.2152	687.3582152
14	1	650	100	50.38942553		0.449993587		1	2	1.44447E-06	692297.8261	692.2978261
15	1	650	100	46.81222563		0.423967705		1	2	1.53314E-06	652258.0072	652.2580072
16	1	650	100	45.18840741		0.411791962		1	2	1.57847E-06	633526.0959	633.5260959
17	1	650	100	45.58900233		0.414816626		1	2	1.56696E-06	638179.4247	638.1794247
18	1	650	100	47.23563439		0.427105464		1	2	1.52187E-06	657085.3287	657.0853287
19	1	532	150	59.04674277		0.366287334		1	2	1.45241E-06	688510.0261	688.5100261
20	1	532	150	60.29958809		0.372987648		1	2	1.42632E-06	701104.6014	701.1046014
21	1	532	150	60.10287459		0.371939428		1	2	1.43034E-06	699134.263	699.134263
22	1	532	150	59.46993908		0.368557068		1	2	1.44347E-06	692776.4428	692.7764428
23	2	532	50	45.08066157		0.669626424		1	2	7.94473E-07	1258696.286	1258.696286
24	2	532	50	39.82332253		0.623008229		1	2	8.53921E-07	1171068.099	1171.068099
25	2	532	50	38.18514336		0.606946988		1	2	8.76518E-07	1140877.797	1140.877797
26	2	532	50	39.34556159		0.618402973		1	2	8.6028E-07	1162411.604	1162.411604
27	2	532	100	93.77770646		0.684048685		1	2	7.77722E-07	1285805.798	1285.805798
28	2	532	100	97.93631349		0.699696449		1	2	7.6033E-07	1315218.889	1315.218889
29	2	532	100	86.70063437		0.655077089		1	2	8.12118E-07	1231347.911	1231.347911
30	2	532	100	97.75858758		0.699047475		1	2	7.61036E-07	1313999.013	1313.999013
31	2	650	50	76.1987817		0.836075342		1	2	7.77442E-07	1286269.757	1286.269757
32	2	650	50	86.62404029		0.866079139		1	2	7.50509E-07	1332429.444	1332.429444

CD1/DVD2	If CD then input 1 or DVD input 2
wavelength(532/650)	532nm or 650nm laser using
d1av	The average distance from center to (m=1) points
d2av	The average distance from center to (m=2) points
sintheta1	The sin for diffraction angle from center to (m=1) points
sintheta2	The sin for diffraction angle from center to (m=2) points
m1	diffraction spots index(m=1)
m2	diffraction spots index(m=2)

