

## §6.5 Consistency Conditions

Useful to assume that all symmetry currents (even global symmetries) are coupled to gauge fields. At the end, we can always take  $g \rightarrow 0$  for those symmetries and return to a global symmetry

→ apart from anomalies, the effective action  $\Gamma[A]$  is invariant under  
 ↗  
 background gauge field

$$A_{\beta\mu}(y) \mapsto A_{\beta\mu}(y) + i \int d^4x \Sigma_\alpha(x) \mathcal{T}_\alpha(x) A_{\beta\mu}(y)$$

where we must take

$$-i \mathcal{T}_\alpha(x) = -\frac{\partial}{\partial x^\mu} \frac{\delta}{\delta A_{\alpha\mu}(x)} - C_{\alpha\beta\gamma} A_{\beta\mu}(x) \frac{\delta}{\delta A_{\gamma\mu}(x)}$$

in order to reproduce

$$\delta A_m^\beta = \partial_m \varepsilon^\beta + i \varepsilon^\beta (\gamma_\alpha^\lambda)^{\rho} \gamma_\lambda A_m^\alpha - i C_{\alpha\beta}^\rho \gamma_\alpha$$

Taking anomalies into account, have

$$\mathcal{T}_\alpha(x) \Gamma[A] = G_\alpha[x; A]$$

where

$$D_m \langle j_\alpha^\mu(x) \rangle = -i G_\alpha[x; A], \quad \text{recall } S_{\text{gauge}} = \int \lambda_\alpha D_m j_\alpha^\mu$$

$$\text{and} \quad \langle j_\alpha^\mu(x) \rangle = \frac{\delta}{\delta A_{\alpha m}(x)} T[A] \quad \begin{matrix} \uparrow \\ \text{gauge trf.} \\ \text{parameter} \end{matrix}$$

with  $D_m = \partial_m - i A_m^\beta(x) (t_\beta)_e^m$  the gauge-covariant derivative

The commutation relations

$$[\tilde{J}_\alpha(x), \tilde{J}_\beta(y)] = i C_{\alpha\beta\gamma} \delta^4(x-y) \tilde{J}_\gamma(x)$$

imply the "Wess-Zumino" consistency conditions:

$$\begin{aligned} & \tilde{J}_\alpha(x) G_\beta[y; A] - \tilde{J}_\beta(y) G_\alpha[x; A] \\ &= i C_{\alpha\beta\gamma} \delta^4(x-y) G_\gamma[y; A] \end{aligned} \quad (1)$$

Reformulate these in terms of BRST-trfs.:

introduce ghost field  $\omega_\alpha$

→ define nilpotent BRST operator  $s$ :

$$s A_{\alpha m} = \partial_m \omega_\alpha + C_{\alpha\beta\gamma} A_{\beta m} \omega_\gamma,$$

$$s \omega_\alpha = -\frac{1}{2} C_{\alpha\beta\gamma} \omega_\beta \omega_\gamma,$$

Work with functional

$$G[\omega, A] = \int \omega_\alpha(x) G_\alpha[x; A] d^4x = i S_{\text{gauge}} \quad \begin{matrix} \text{with gauge trf. parameter } \lambda_2 = \omega_2! \end{matrix}$$

$$\begin{aligned}
\rightarrow sG[\omega, A] &= -\frac{1}{2} C_{\alpha\beta\gamma} \int d^4x \omega_\beta(x) \omega_\gamma(x) G_\alpha[x; A] \\
&\quad - \int d^4x \omega_\alpha(x) \int d^4y \left[ \frac{\partial \omega_\beta(y)}{\partial y^\mu} + C_{\beta\gamma\mu} A_{\gamma\mu}(y) \omega_\gamma(y) \right] \frac{\delta G_\alpha[x; A]}{\delta A_{\beta\mu}(y)} \\
&= \int d^4x \int d^4y \omega_\alpha(x) \omega_\beta(y) \left[ -\frac{1}{2} C_{\alpha\beta\gamma} \delta^4(x-y) G_\gamma[x; A] \right. \\
&\quad \left. + i \tilde{J}_\beta(y) G_\alpha[x; A] \right]
\end{aligned}$$

As ghost fields anti-commute, write

$$\begin{aligned}
sG[\omega, A] &= -\frac{1}{2} i \int d^4x \int d^4y \omega_\alpha(x) \omega_\beta(y) \\
&\quad \times \left[ i C_{\alpha\beta\gamma} \delta^4(x-y) G_\gamma[x; A] + \tilde{J}_\beta(y) G_\alpha[x; A] - \tilde{J}_\alpha(x) G_\beta[x; A] \right]
\end{aligned}$$

$\rightarrow$  (1) holds iff

$$(2) \quad sG[\omega, A] = 0 \quad \text{for all ghost fields } \omega_\alpha(x)$$

Now consider possibility that

$$G[\omega; A] = \bar{s}F[A]$$

As  $s$  satisfies  $s^2=0$ , we automatically get

$$sG = 0. \quad \text{Since } sF[A] \text{ is } S_{\text{gauge}} F[A]$$

with  $\pi_\alpha = \omega_\alpha$  and  $G[\omega; A]$  is the anomalous trf. of the action  $S$ , the modified action

$$S - sF[A] \quad (*)$$

will be anomaly free!

- we are interested in anomalies which cannot be subtracted as in (\*)
- these are in the "cohomology" of the s-operator:  $\text{Ker}(s)/\text{Im}(s)$

Write the anomaly as  $G = \int d^4x \mathcal{G}(x)$

$$\text{Then } sG=0 \Leftrightarrow s\mathcal{G}(x) = 2_m \cancel{\mathcal{J}^\mu(x)} \quad (3)$$

$$\text{Also write } sF[A] = \int d^4x \underbrace{sF(x)}_{\substack{\text{ghost number } 1 \\ \downarrow \\ \text{ghost number } 0}} \quad$$

→ cohomology of s at ghost number 1:

$$H^1(s|d)$$

Now recall triangle anomaly for 3 flavor currents:

$$[\langle \partial_m \mathcal{J}_\alpha^\mu(x) \rangle]_{\text{anom}} = -\frac{i}{24\pi^2} D_{\alpha\beta\gamma} \epsilon^{\kappa\nu\rho} \partial_\kappa A_\nu^\beta(x) \partial_\rho A_\mu^\alpha(x)$$

→ of dimensionality 4 (in units of energy)

WZ-consistency condition (1) relates operators of same dimensionality

$$\begin{aligned} \rightarrow G_\alpha &= i[\langle \partial_m \mathcal{J}_\alpha^\mu \rangle]_{\text{anom}} \quad (4) \\ &= -\frac{i}{24\pi^2} \epsilon^{\kappa\nu\rho} \text{Tr} \left\{ T_\alpha [\partial_\kappa A_\nu \partial_\rho A_\mu + i c_1 \partial_\kappa A_\nu A_\mu A_\rho] \right\} \end{aligned}$$

$$+ i c_2 A_{\mu} \partial_{\nu} A_{\lambda} A_{\rho} + i c_3 A_{\mu} A_{\nu} \partial_{\lambda} A_{\rho} - c_4 A_{\mu} A_{\nu} A_{\lambda} A_{\rho} \Big] \Big\},$$

where  $A_{\mu} = A_{\alpha\mu} T_{\alpha}$ , and  $c_i$  are constants to be determined.

→ rewrite in language of differential forms:

- $dx^{\mu} \wedge dx^{\nu} \wedge dx^{\lambda} \wedge dx^{\rho} = \epsilon^{\mu\nu\lambda\rho} d^4 x$ ,  $d^4 x = dx^0 dx^1 dx^2 dx^3$

- $d \equiv dx^m \frac{\partial}{\partial x^m}$ ,  $d^2 = 0$ ,  $ds + sd = 0$

- anti-commuting quantities

$$A = i A_{\mu} dx^{\mu} = i A_{\alpha\mu} T_{\alpha} dx^{\mu}, \quad \omega = i \omega_{\alpha} T_{\alpha}$$

→ eq. (4) becomes:

$$G[\omega, A] = \frac{1}{24\pi^2} \int \text{Tr} \left\{ \omega \left[ (dA)^2 + g(dA) A^2 \right. \right. \\ \left. \left. + c_2 A(dA)A + c_3 A^2(dA) + c_4 A^4 \right] \right\}$$

rewrite BRST trfs. as :

$$sA = -d\omega + \{A, \omega\}$$

$$s\omega = \omega^2$$

$$\rightarrow s \text{Tr} [\omega A^4] = \text{Tr} \left[ \omega^2 A^4 - \omega \cancel{\{A, \omega\}} A^3 + \omega \cancel{A \{A, \omega\}} A^2 \right. \\ \left. - \omega \cancel{A^2 \{A, \omega\}} A + \omega \cancel{A^3 \{A, \omega\}} \right. \\ \left. + \omega d\omega A^3 \right] \text{ terms}$$

$$= \text{Tr} [\omega^2 A^4] + \omega d\omega A^3 \text{ terms}$$

Thus  $sG[\omega, A] = 0$  can only be satisfied  
iff  $c_4 = 0$ .

$$\begin{aligned} \rightarrow sG &= \frac{1}{24\pi^2} \int \text{Tr} \left\{ -(\text{d}A)^2 \omega^2 + \omega d\omega A dA - d\omega \omega dA A \right. \\ &\quad - A \omega dAd\omega - \omega Ad\omega dA \\ &\quad + c_1 [\omega dA d\omega A - \omega dA A d\omega] \\ &\quad + c_2 [\omega d\omega dAA - \omega AdAd\omega] \\ &\quad + c_3 [\omega d\omega AdA - \omega Ad\omega dA] \\ &\quad - c_1 [-\omega Ad\omega A^2 + \omega d\omega A^3 + \omega dAA^2 \omega] \\ &\quad - c_2 [\omega A^2 d\omega A - \omega Ad\omega A^2 + \omega AdAd\omega] \\ &\quad \left. - c_3 [-\omega A^3 d\omega + \omega A^2 d\omega A + \omega A^2 dA \omega] \right\} \end{aligned}$$

$\rightarrow$  can write the integrand as  $dF$  if

$$c_1 = -c_2 = +c_3 = -\frac{1}{2}$$

In this case, we get

$$\begin{aligned} G[\omega, A] &= \frac{1}{24\pi^2} \int \text{Tr} \left\{ \omega d \left[ \underbrace{AdA - \frac{1}{2} A^3}_{\text{Chern-Simons form}} \right] \right\} \\ &= \frac{1}{24\pi^2} \int \text{Tr} \left\{ \omega d \left[ A F + \frac{1}{2} A^3 \right] \right\}, \quad (5) \end{aligned}$$

$$\text{where } F = \frac{1}{2} i \epsilon_{\mu\nu} F_{\mu\nu} dx^\mu \wedge dx^\nu = dA - A^2 \quad (*)$$

Stora-Zumino descent equations:

take spacetime dimensionality to be  $2n$

Note that

$$dF \stackrel{(*)}{=} -d(A^2) = -(dA)A + A(dA) = [A, F]$$

$\rightarrow \text{Tr } F^{n+1}$  is closed:

$$d\text{Tr } F^{n+1} = (n+1)\text{Tr}\{(dF)F^n\} = \overline{\text{Tr}}\{[A, F^{n+1}]\} = 0$$

$\rightarrow$  if spacetime is simply connected,  
 $\text{Tr}\{F^{n+1}\}$  is exact:

$$\overline{\text{Tr}}\{F^{n+1}\} = d\Omega_{2n+1}$$

Since  $\overline{\text{Tr}}\{F^{n+1}\}$  is gauge invariant, and  
BRST trf. is a special gauge trf., we get

$$s\overline{\text{Tr}}\{F^{n+1}\} = 0$$

Using  $s^2 = 0$ , we get

$$d(s\Omega_{2n+1}) = -s\overline{\text{Tr}}\{F^{n+1}\} = 0$$

$$\rightarrow s\Omega_{2n+1} = d\Omega'_{2n}$$

$$\text{Furthermore, } d(s\Omega'_{2n}) = -s^2\Omega_{2n+1} = 0$$

$$\rightarrow s\Omega'_{2n} = d\Omega''_{2n-1} \text{ and we have}$$

$$s \int_{\text{spacetime}} \Omega'_{2n} = 0$$

Thus  $\int \Omega_{2n}^1$  is a candidate for  $G[\omega, A]$

by integrating the two 1st order diff. eqs.

$$d\Omega_{2n+1} = \text{Tr}\{F^{n+1}\} \quad \text{and} \quad d\Omega_{2n}^1 = S\Omega_{2n+1}$$

→ solutions are given by:

$$\Omega_{2n+1} = (n+1) \int_0^1 dt \text{Tr}\{AF_t^n\},$$

$$(*) \quad \Omega_{2n}^1 = -(n+1) \sum_{r=0}^{n-1} \int_0^1 dt (1-t) \text{Tr}\{\omega d(F_t^r AF_t^{n-1-r})\},$$

$$\text{where } F_t \equiv tF + (t-t^2)A^2.$$

Evaluation of (\*) shows that (5) gives a result for  $G[\omega, A]$  proportional to  $\int \Omega_4^1$  in case of 4 spacetime dimensions.