# Physics Notes

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#### Measurement 1

#### 1.1 Instrument Uncertainty

All instruments have uncertainty:

- 1. Analogue Instruments: Half the smallest measurement unit
- 2. Digital Instruments: The smallest significant figure
- 3. Human reaction time:  $\pm 0.10$ s

# Significant Figures

- 1. Adding or subtracting: Follow term with least decimal place
- 2. Multiplying or Dividing: Follow term with least significant figure

# Propagation of error

For any  $f(a, \dots)$  the general formula for  $\Delta f$  is:

$$\Delta f = \sqrt{\left(\frac{\partial f}{\partial a} \Delta a\right)^2 + \cdots}$$

Some specific examples:

1. 
$$f = a \pm b$$

$$\Delta f = \sqrt{(\Delta a)^2 + (\Delta b)^2}$$

2. 
$$f = ab$$
 or  $f = \frac{a}{b}$ 

$$\frac{\Delta f}{f} = \sqrt{\left(\frac{\Delta a}{a}\right)^2 + \left(\frac{\Delta b}{b}\right)^2}$$

#### 2 Mechanics

#### 2.1 Statics

When all objects are motionless (or have constant velocity),

$$\sum \mathbf{F}_{net} = 0$$
$$\sum \mathbf{\tau}_{net} = 0$$

Four basic forces to consider:

**Tension** Pulling force felt by a rope, string, etc. Every piece of rope feels a pulling force in both directions.

**Friction** Parallel to surface of contact, can be static or kinetic.

**Normal** Perpendicular to surface of contact, prevents object from falling through surface.

**Gravity** Force acting between two objects with mass. Always acts downwards for objects on surface of earth. Static friction does no work.

#### 2.2**Kinematics**

$$egin{aligned} oldsymbol{v} &= \lim_{\Delta t o 0} rac{\Delta oldsymbol{x}}{\Delta t} = rac{doldsymbol{x}}{dt} = \dot{oldsymbol{x}} \ oldsymbol{a} &= rac{doldsymbol{v}}{dt} = rac{d^2oldsymbol{x}}{dt^2} = \dot{oldsymbol{x}} = \ddot{oldsymbol{x}} \end{aligned}$$

# 2.2.1 Polar Coordinates

Differentiation of unit vectors:

$$\dot{\hat{m{r}}} = \dot{ heta}\hat{m{ heta}}$$
 $\dot{\hat{m{ heta}}} = -\dot{m{ heta}}\hat{m{r}}$ 

Velocity and acceleration in polar form:

$$egin{align} oldsymbol{r} &= r \hat{oldsymbol{r}} \ oldsymbol{v} &= \dot{oldsymbol{r}} = \dot{r} \hat{oldsymbol{r}} + r \dot{oldsymbol{ heta}} \hat{oldsymbol{ heta}} \ oldsymbol{a} &= \dot{oldsymbol{v}} = (\ddot{r} - \dot{oldsymbol{ heta}}^2 r) \hat{oldsymbol{r}} + (r \ddot{oldsymbol{ heta}} + 2 \dot{r} \dot{oldsymbol{ heta}}) \hat{oldsymbol{ heta}} \end{split}$$

#### **Dynamics** 2.3

$$m{F} = mm{\ddot{x}}$$
  $m{F}_{action} = -m{F}_{reaction}$ 

Free body diagram techniques:

- 1.  $\Sigma \mathbf{F}_{net} = 0$  for massless pulleys
- 2. Conservation of string

Solving differential equations in 1dimension:

1. 
$$F = f(t)$$

$$m \int_{v_0}^{v(t)} dv' = \int_{t_0}^{t} f(t')dt'$$
$$m \int_{x_0}^{x(t)} dx' = \int_{t_0}^{t} v(t')dt'$$

2. 
$$F = f(x)$$

$$a = \frac{dv}{dt} = \frac{dv}{dt}\frac{dx}{dx} = v\frac{dv}{dx}$$
$$m\int_{v_0}^{v(x)} v'dv' = \int_{x_0}^{x} f(x')dx'$$

3. 
$$F = f(v)$$

$$m \int_{v_0}^{v(t)} \frac{dv'}{f(v')} = \int_{t_0}^t dt'$$

#### 2.3.1 Friction

Kinetic and static friction:

$$egin{aligned} oldsymbol{f_k} &= \mu_k oldsymbol{N} \ oldsymbol{f_s} &< \mu_s oldsymbol{N} \end{aligned}$$

#### 2.3.2 **Constraining Forces**

For any rigid body, there are 6 degrees of freedom (DF). There can be constraining forces (C) acting on the body.

- Statics: C + DF = 6
- Dynamics C + DF > 6

There are 3 assumptions made for a body moving without any constraint:

- 1.  $f_{ij} \| r_{ij}$
- 2.  $r_{ij}$  is constant for any 2 points in a rigid body
- 3.  $\mathbf{f}_{12} + \mathbf{f}_{21} = 0$

# 2.3.3 Fictitious Forces

Force felt by an object in a non-inertial frame. Let r to be the position vector of the object in the accelerated frame and R be the position vector of the accelerated frame, then the possible forces that acts on r are:

$$\frac{\delta^2 \mathbf{r}}{\delta t^2} = \frac{\mathbf{F}}{m} - \frac{d^2 \mathbf{R}}{dt^2} - \boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r})$$
$$-2\boldsymbol{\omega} \times \boldsymbol{v} - \frac{d\boldsymbol{\omega}}{dt} \boldsymbol{r}$$

- 1. Translational force:  $-m\frac{d^2\mathbf{R}}{dt^2}$
- 2. Centrifugal force:  $-m\boldsymbol{\omega} \times (\boldsymbol{\omega} \times \boldsymbol{r})$
- 3. Coriolis force:  $-2m\boldsymbol{\omega} \times \boldsymbol{v}$
- 4. Azimuthal force:  $-m\frac{d\omega}{dt}\mathbf{r}$

# 2.4 Conservation Laws

Energy  $W_{NC}=0$ 

Momentum  $\Sigma \mathbf{F}_{net} = 0$ 

Angular Momentum  $\Sigma \tau_{net} = 0$ 

### Energy

For a force in one dimension:

$$mv\frac{dv}{dx} = F(x)$$
$$\frac{1}{2}mv^2 = E + \int_{x_0}^x F(x')dx'$$

We can then define potential energy:

$$U(x) = -\int_{x_0}^x F(x')dx'$$

Work-Energy theorem:

$$W_{AB} = \int_{x_1}^{x_2} F(x')dx'$$
$$W_{\text{total}} = \Delta KE$$

Conservative forces are forces that only depend on *position*. For conservative forces:

$$\oint \mathbf{F} \cdot d\mathbf{l} = 0$$

$$\nabla \times \mathbf{F} = 0$$

$$\mathbf{F} = -\nabla U$$

$$W_C = -\Delta U$$

For non-conservative forces:

$$W_{NC} = \Delta(K+U) = \Delta E$$

Where E is defined as the mechanical energy of the system.

# 2.5.1 Energy Analysis

The Lagrangian method is based on the *principle of stationary action*.

$$\mathcal{L}(\dot{x}, x, t) = T - V$$
$$\frac{d}{dt}(\frac{\partial \mathcal{L}}{\partial \dot{x}}) - \frac{\partial \mathcal{L}}{\partial x} = 0$$

The Hamiltonian  $\mathcal{H}$  can be used for the conservation of energy:

$$\mathcal{H}(\dot{x}, x, t) = T + V$$
$$\dot{\mathcal{H}} = 0$$

Where T is the kinetic energy, and V is the potential energy of the system.

### 2.5.2 Power

Power is the rate of work done per unit time:

$$P = \frac{dW}{dt}$$

Mechanical power:

$$P = \frac{d}{dt} \oint \mathbf{F} \cdot d\mathbf{x} = \frac{d}{dt} \oint \mathbf{F} \cdot \frac{d\mathbf{x}}{dt} dt$$
$$= \mathbf{F} \cdot \mathbf{v}$$

### 2.6 Momentum

Momentum is defined as:

$$p = mv$$

When there is no net force on the system,

$$\sum \mathbf{F}_{net} = 0 \Rightarrow \frac{d\mathbf{p}}{dt} = 0$$
$$\Rightarrow \mathbf{p} \text{ is conserved}$$

Impulse is defined as:

$$\mathcal{I} = \int_{t_1}^{t_2} \mathbf{F}(t)dt = \int_{t_1}^{t_2} \frac{d\mathbf{p}}{dt}dt$$
$$\mathcal{I} = \mathbf{p}(t_2) - \mathbf{p}(t_1) = \Delta \mathbf{p}$$

For perfectly elastic collisions of two objects in 1-D, relative velocity is constant.

$$v_1 - v_2 = -(v_1' - v_2')$$

For other collisions in 1-D, we have the coefficient of restitution e:

$$e = -\frac{v_2' - v_1'}{v_2 - v_1} \qquad 0 \le e \le 1$$

# 2.7 Central Forces

For any particle subjected to a central force,

$$F(r) = mr\dot{\theta}^2 - m\ddot{r}$$
$$L = mr^2\dot{\theta}$$

Because angular momentum L is constant, we can look at central forces systems in 1-dimension.

$$V_{\text{eff}}(r) = \frac{L^2}{2mr^2} + V(r)$$
$$E = V_{\text{eff}} + \frac{1}{2}m\dot{r}^2$$

# **2.7.1** Gravity

For any two point masses of  $m_1$  and  $m_2$  in empty space, the gravitational force between them is:

$$oldsymbol{F} = rac{Gm_1m_2}{|oldsymbol{r}|^2}\hat{oldsymbol{r}}$$

Where r is the position vector of one mass respect to the other, and G is the gravitational constant.

$$F = mg$$

For a mass m at the Earth's surface, where  $g=9.81m/s^2$  pointing downwards.

# 2.8 Uniform Circular Motion

For a point mass moving in uniform circular motion, we define:

$$\omega = \frac{v}{r}$$

The centripetal acceleration a and the force required to keep the object in its circular path:

$$a = \frac{v^2}{r} = \omega^2 r$$
$$F = \frac{mv^2}{r} = m\omega^2 r$$

# 2.9 Rotational Dynamics (Constant $\hat{L}$ )

# 2.9.1 Angular Momentum

The angular momentum of a point mass is defined as:

$$oldsymbol{L} = oldsymbol{r} imes oldsymbol{p}$$

For a flat object lying on a 2-D plane rotating with angular speed  $\omega$ :

$$oldsymbol{L} = \int oldsymbol{r} imes oldsymbol{p} = \int r^2 \omega oldsymbol{\hat{z}} dm$$

If we define the moment of intertia about the z-axis to be  $I_z = \int (x^2 + y^2) dm$ , we have:

$$L_z = I_z \omega$$

$$T = \int \frac{1}{2} m \mathbf{v}^2 = \int \frac{r^2 \omega^2}{2} dm$$

$$= \frac{1}{2} I_z \omega^2$$

For the z-component of  $\boldsymbol{L}$  and kinetic energy T.

### 2.9.2 General Motion

For an object with a moving center of mass, and rotating at  $\omega$  about it,

$$oldsymbol{L} = oldsymbol{r}_{\mathrm{CM}} imes oldsymbol{p}_{\mathrm{CM}} + I_{\mathrm{CM}} \omega \hat{oldsymbol{z}}$$
 $T = rac{1}{2} m v_{\mathrm{CM}}^2 + rac{1}{2} I_{\mathrm{CM}} \omega^2$ 

# **2.9.3** Torque

Torque is defined as:

$$oldsymbol{ au} = oldsymbol{r} imes oldsymbol{F}$$

Using an origin satisfying any of the following conditions to calculate L,

- 1. The origin is the center of mass
- 2. The origin is not accelerating
- 3.  $(\mathbf{R} \mathbf{r_0})$  is parallel to  $\mathbf{r_0}$ , the position of the origin in a fixed coordinate system

$$\frac{d\mathbf{L}}{dt} = \sum \boldsymbol{\tau}_{\mathrm{ext}}$$

When there is no external torque, we have the conservation of angular momentum.

$$\tau_{\rm ext} = I\alpha$$

Where  $\alpha = \frac{d\omega}{dt}$  is the angular acceleration.

#### 2.9.4 Angular Impulse

Angular impulse is defined as:

$$\mathcal{I}_{ heta} = \int_{t_1}^{t_2} oldsymbol{ au}(t) dt = \Delta oldsymbol{L}$$

If we have a force F(t) applied at a constant distance R from the origin.

$$au(t) = \mathbf{R} \times \mathbf{F}(t)$$

$$\mathcal{I}_{\theta} = \mathbf{R} \times \mathcal{I}$$

$$\Delta \mathbf{L} = \mathbf{R} \times (\Delta \mathbf{p})$$

#### 2.9.5Parallel-axis Theorem

Let an object of mass M rotate about its center of mass with the same frequency  $\omega$  as the center of mass rotates about the origin (with radius R):

$$L_z = (MR^2 + I_{\rm CM})\omega$$

Thus if the moment of inertia of an object is  $I_0$  about a particular axis, its moment of inertia about a parallel axis separated by R is:

$$I = MR^2 + I_0$$

# 2.9.6 Perpendicular-axis Theo- Loss of simultaneity

For flat 2-D objects in the x-y plane, and orthogonal axes x, y and z:

$$I_z = I_x + I_y$$

# Moments of Inertia

Center of mass for an object of mass M:

$$m{R}_{\mathrm{CM}} = rac{\int m{r} dm}{M}$$

Common moments of inertia (taken about center of mass unless stated):

- 1. Point mass at r from axis:  $mr^2$
- 2. Rod of length L:  $\frac{1}{3}mL^2$

- 3. Rod of length L about one end: Transverse velocity addition
- 4. Solid disk of radius r perpendicular to axis:  $\frac{1}{2}mr^2$
- 5. Hollow sphere with radius  $r: \frac{2}{3}mr^2$
- 6. Solid sphere with radius  $r: \frac{2}{5}mr^2$

# Special Relativity

#### 3.1 **Postulates**

- 1. The speed of light has the same value in all inertial frames
- 2. Physical laws remain the same in all inertial frames

#### 3.2 **Kinematics**

# 3.2.1 Lorentz Transform

$$x = \gamma(x' + \beta ct')$$

$$y = y'$$

$$z = z'$$

$$ct = \gamma(\beta x' + ct')$$

Where 
$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$
 and  $\beta = \frac{v}{c}$ .

# Fundamental Effects Length contraction

$$l' = \frac{l}{\gamma}$$

Where l is the proper length.

# Time dilation

$$t' = \gamma t$$

Where t is the proper time.

$$\Delta t = \frac{Lv}{c^2}$$

Two events separated by L and  $\Delta t$ in the rest frame will appear simultaneous to an observer moving at

# Longitudinal velocity addition

$$v_x' = \frac{u+v}{1+uv/c^2}$$

Where u is the velocity of an object in the frame traveling at v respect to the lab frame, and  $v'_x$  is the xvelocity of the object viewed by the lab frame. 4

$$v_y' = \frac{u_y}{\gamma_v(1 + u_x v/c^2)}$$

Where  $u_y$  and  $u_x$  are velocity components of an object in the frame traveling at v respect to the lab frame, and  $v'_{y}$  is the y-velocity of the object viewed by the lab frame.

# Longitudinal Doppler effect

$$f' = f\sqrt{\frac{1+\beta}{1-\beta}}$$

Where f' is the frequency observed of a moving source emitting at frequency f in its rest frame.

# 3.2.3 Minkowski Diagrams

Space-time diagrams with x and ctaxes. Some properties are:

- 1. Light travels at 45° to horizontal.
- 2. x' and ct' axes of another moving frame are  $\theta$  to the x and ct axes respectively, with

$$\tan(\theta) = \beta$$

3. Units on axes of the moving and stationary frames are related by:

$$\frac{x'}{x} = \frac{ct'}{ct} = \sqrt{\frac{1+\beta^2}{1-\beta^2}}$$

### **Dynamics** 3.3

#### 3.3.1 Momentum

$$oldsymbol{p} = \gamma_v m oldsymbol{v} = rac{m oldsymbol{v}}{\sqrt{1 - rac{v^2}{c^2}}}$$

#### 3.3.2 Energy

$$E^2 = p^2 c^2 + m^2 c^4$$

For massive particles:

$$E = \gamma mc^2 = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

For massless particles(such as photons):

$$E = pc = \frac{hc}{\lambda}$$

### 3.4 4-vectors

A 4-vector  $\vec{A} = (A_1, A_2, A_3, A_4)$  is a quantity that transforms as follows:

$$A'_{1} = \gamma(A_{1} + i\beta A_{4})$$
  
 $A'_{2} = A_{2}$   
 $A'_{3} = A_{3}$   
 $A'_{4} = \gamma(A_{4} - i\beta A_{1})$ 

The dot product of two 4-vectors is invariant under Lorentz transformations:

$$\vec{A} \cdot \vec{B} = \vec{A'} \cdot \vec{B'}$$

# 3.4.1 Different 4-vectors

# **4-position** (dx, dy, dz, icdt)

4-vectors originate from the invariant interval ds.

$$\vec{ds}^2 = (dx, dy, dz, icdt)^2$$
$$= dx^2 + dy^2 + dz^2 - c^2 dt^2$$

# **4-velocity** $\gamma_v(\boldsymbol{v},ic)$

To obtain other 4-vectors, we can multiply invariant quantities to the 4position vector, such as proper time:

$$d\tau = \frac{dt}{\gamma}$$

$$\vec{v} = \frac{ds}{d\tau}$$

$$= \gamma_v \left( \frac{dx}{dt}, \frac{dy}{dt}, \frac{dz}{dt}, ic \right)$$

$$= \gamma_v(\mathbf{v}, ic)$$

# 4-momentum $(\boldsymbol{p}, i\frac{E}{c})$

As mass is invariant,

$$\vec{p} = m\vec{v}$$

$$= (\gamma_v m \mathbf{v}, i\gamma_v mc)$$

$$= \left(\mathbf{p}, i\frac{E}{c}\right)$$

For photons in x-direction, the 4-momentum vector is:

$$\vec{p} = \left(\frac{h}{\lambda}, 0, 0, i\frac{h}{\lambda}\right)$$

4-wave  $(\mathbf{k}, i\frac{\omega}{c})$ 

For electromagnetic waves,

$$k = \frac{2\pi}{\lambda} = \frac{\omega}{c}$$

$$p = \frac{h}{\lambda} = \hbar k$$

$$E = hf = \hbar \omega$$

$$\vec{p} = \hbar \left(k, i\frac{\omega}{c}\right)$$

$$\vec{k} = \frac{\vec{p}}{\hbar}$$

**4-force**  $\gamma_v\left(\boldsymbol{f}, \frac{i}{c} \frac{dE}{dt}\right)$ 

$$ec{F} = rac{dec{p}}{d au}$$

$$= \gamma_v \left( f, rac{d}{dt} \left( i rac{E}{c} 
ight) 
ight)$$

# 4 Electricity and Magnetism

# 4.1 Electrostatics

Coulomb's law The force between a point charge q and test charge Q:

$$\mathbf{F} = \frac{1}{4\pi\epsilon_0} \frac{Qq}{\mathbf{r}^2} \mathbf{\hat{r}}$$

Where  $\mathbf{r} = \mathbf{r} - \mathbf{r'}$  is the displacement vector from Q at  $\mathbf{r}$  and q at  $\mathbf{r'}$ .

Superposition principle The interaction between any two charges is unaffected by any other charges

### 4.1.1 Electric Field

The electric field of a point charge is defined as:

$$\mathbf{E} = \frac{\mathbf{F}}{Q} = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \hat{\mathbf{r}}$$

For a continuous volume charge distribution  $\rho(\mathbf{r'})$ , we can use the superposition principle to get:

$$\boldsymbol{E}(\boldsymbol{r}) = \frac{1}{4\pi\epsilon_0} \int_{\mathcal{V}} \frac{\rho(\boldsymbol{r'})}{r^2} \hat{\boldsymbol{r}} d\tau'$$

Taking the divergence of  $\boldsymbol{E}$ , we get Gauss' law:

$$\nabla \cdot \boldsymbol{E} = \frac{\rho(\boldsymbol{r})}{\epsilon_0}$$

$$\oint_{\mathcal{S}} \boldsymbol{E} \cdot d\boldsymbol{a} = \frac{Q_{\text{enc}}}{\epsilon_0}$$

Taking the curl of E:

$$\nabla \times \mathbf{E} = 0$$

$$\oint \mathbf{E} \cdot d\mathbf{l} = 0$$

For any surface charge in an electric field E, the field felt by an area element on the surface is:

$$m{E}_{ ext{felt}} = rac{1}{2} \left( m{E}_{ ext{above}} + m{E}_{ ext{below}} 
ight)$$

### 4.1.2 Electric Potential

As the line integral of the electrostatic field is path independent, we can define the potential at a point r:

$$V(\boldsymbol{r}) = -\int_{\mathcal{O}}^{\boldsymbol{r}} \boldsymbol{E} \cdot d\boldsymbol{l}$$

Where  $\mathcal{O}$  is a standard reference point, usually set to infinity. The potential of a point charge can then be found, and with the superposition principle we can find the potential of any charge distribution:

$$V = \frac{1}{4\pi\epsilon_0} \frac{q}{r}$$

$$V(\mathbf{r}) = \frac{1}{4\pi\epsilon_0} \int_{\mathcal{V}} \frac{\rho(\mathbf{r'})}{r} d\tau'$$

Taking the gradient of the potential:

$$\mathbf{E} = -\nabla V$$
$$\nabla^2 V = -\frac{\rho}{\epsilon_0}$$

# 4.1.3 Work and Energy

The work needed to bring a charge Q from infinity to a point a is:

$$W = \int_{\infty}^{a} \mathbf{F} \cdot d\mathbf{l}$$
$$= -Q \int_{\infty}^{a} \mathbf{E} \cdot d\mathbf{l}$$
$$= QV(\mathbf{a})$$

The energy in a continuous charge distribution is:

$$W = \frac{1}{2} \int \rho V d\tau$$
$$= \frac{\epsilon}{2} \int E^2 d\tau$$

Where the integral is taken over all space.

### 4.1.4 Conductors

A perfect conductor has an unlimited supply of free charges.

- 1. E = 0 and  $\rho = 0$  inside a conductor
- 2. Any conductor is an equipotential
- 3. Just outside a conductor,  $\boldsymbol{E}$  is perpendicular to the surface.

If we charge up two conductors with +Q and -Q, the potential between them is proportional to the charge Q (because the electric field is proportional to Q), and we define the constant of proportionality capacitance:

$$C = \frac{Q}{V}$$

The work done by charging a capacitor is:

$$W = \int_0^Q \frac{q}{C} dq = \frac{Q^2}{2C}$$
$$= \frac{1}{2}CV^2$$

# 4.1.5 Image Charges

In certain special cases, a charge placed next to a grounded conductor has equivalents.

- 1. A point charge and a conducting sheet: An opposite charge in the mirror image position.
- 2. A point charge and a conducting sphere, or an infinite line charge and conducting cylinder: Opposite image charge and charge forms the Apollonius sphere/cylinder.

# 4.1.6 Uniqueness Theorems

First uniqueness theorem The solution to Laplace's equation  $(\nabla^2 V = 0)$  in some volume  $\mathcal{V}$  is uniquely determined if V is specified on the boundary surface  $\mathcal{S}$ .

Second uniqueness theorem In a volume  $\mathcal{V}$  surrounded by conductors and containing a specified charge density  $\rho$ , the electric field is uniquely determined if the total charge on each conductor is given.

### 4.2 Magnetostatics

# 4.2.1 Lorentz Force Law

The force felt by:

1. A point charge q moving at velocity v through a magnetic field B:

$$oldsymbol{F} = qoldsymbol{v} imes oldsymbol{B}$$

2. A line current I:

$${m F} = I \int (d{m l} \times {m B})$$

3. A general volume current J per unit area perpendicular to flow:

$$\mathbf{F} = \int (\mathbf{J} \times \mathbf{B}) d\tau$$

# 4.2.2 Biot-Savart Law

The magnetic field created by a steady line current:

$$\boldsymbol{B}(\boldsymbol{r}) = \frac{\mu_0}{4\pi} I \int \frac{d\boldsymbol{l} \times \hat{\boldsymbol{r}}}{r^2}$$

# 4.2.3 Magnetic Fields

The magnetic field is divergence-free:

$$\nabla \cdot \mathbf{B} = 0$$

$$\oint \mathbf{B} \cdot d\mathbf{a} = 0$$

Taking the curl of the magnetic field gives Ampere's Law:

$$\nabla \times \boldsymbol{B} = \mu_0 \boldsymbol{J}$$
$$\oint \boldsymbol{B} \cdot d\boldsymbol{l} = \mu_0 I_{\text{enc}}$$

# 4.3 Electrodynamics

### 4.3.1 Electromotive Force

For an electric field applied in a material:

$$oldsymbol{J} = rac{oldsymbol{E}}{
ho}$$

Where  $\rho$  is the resistivity constant depending on the material. This leads to Ohm's law:

$$V = IR$$
$$R = \rho \frac{l}{A}$$

The power delivered:

$$P = VI = I^2R$$

The electromotive force (emf)  $\mathcal{E}$  is the line integral of the force per unit charge driving the current:

$$\mathcal{E} = \oint \mathbf{E} \cdot d\mathbf{l}$$

 $\mathcal{E} = V$  for an ideal source.

# 4.3.2 Faraday's Law

Faraday's law states that a changing magnetic flux  $\Phi$  induces an electric field:

$$\mathcal{E} = \oint \mathbf{E} \cdot d\mathbf{l} = \frac{d\Phi}{dt}$$
$$\mathbf{\nabla} \times \mathbf{E} = \frac{d\mathbf{B}}{dt}$$

# 4.3.3 Inductance

If we have two current loops 1 and 2, the flux  $\Phi_2$  through loop 2 is proportional to the current through loop 1:

$$\Phi_2 = M_{21}I_1$$

Where  $M_{21} = M_{12}$  is the mutual inductance between these two loops. We can also define an self inductance L, for a single loop:

$$\Phi = LI$$

$$\mathcal{E} = -L\frac{dI}{dt}$$

### 4.4 Electric Circuits

# 6 Oscillations and Waves

Many questions involve solving linear differential equations. For such equations, linear combinations of solutions will also be a solution.

### 5.1 Oscillations

# 5.1.1 Simple Harmonic Motion

We have a spring force, F = -kx.

$$\ddot{x} + \omega^2 x = 0$$
, where  $\omega = \sqrt{\frac{k}{m}}$   
 $x(t) = A\cos(\omega t + \phi)$ 

# 5.1.2 Damped Oscillators

In addition to the spring force, we now have a drag force  $F_f = -bv$ , and the total force  $F = -kx - b\dot{x}$ .

$$\ddot{x} + 2\gamma \dot{x} + \omega^2 x = 0$$

Where  $2\gamma = b/m$  and  $\omega^2 = k/m$ . Let  $\Omega = \sqrt{\gamma^2 - \omega^2}$ .

$$x(t) = e^{-\gamma t} (Ae^{\Omega t} + Be^{-\Omega t})$$

# Underdamping $(\Omega^2 < 0)$

$$x(t) = e^{-\gamma t} (Ae^{i\tilde{\omega}t} + Be^{-i\tilde{\omega}t})$$
$$= e^{-\gamma t} C\cos(\tilde{\omega}t + \phi)$$

Where  $\tilde{\omega} = \sqrt{\omega^2 - \gamma^2}$ . The system will oscillate with its amplitude decreasing over time. The frequency of oscillations will be smaller than in the undamped case.

# Overdamping $(\Omega^2 > 0)$

$$x(t) = Ae^{-(\gamma - \Omega)t} + Be^{-(\gamma + \Omega)t}$$

The system will not oscillate, and the motion will go to zero for large t.

# Critical damping $(\Omega^2 = 0)$

We have  $\gamma = \omega$ , and:

$$\ddot{x} + 2\gamma \dot{x} + \gamma^2 x = 0$$

In this special case,  $x = te^{-\gamma t}$  is also a solution:

$$x(t) = e^{-\gamma t}(A + Bt)$$

Systems with critical damping go to zero the quickest.

### 5.1.3 Driven Oscillators

We have to solve differential equations of this form:

$$\ddot{x} + 2\gamma \dot{x} + ax = \sum_{n=1}^{N} C_n e^{i\omega_n t}$$

We first find particular solutions for each n, by guessing solutions of the form  $x_{p_n}(t) = Ae^{i\omega_n t}$ :

$$-A\omega_n^2 + 2iA\gamma\omega_n + Aa = C_n$$
$$x_{p_n}(t) = \frac{C_n}{-\omega_n^2 + 2i\gamma\omega_n + a}e^{i\omega_n t}$$

Using the superposition principle, the final solution is a linear combination of the general solution and the particular solutions, with the combination constants determined by initial conditions.

# 5.1.4 Coupled Oscillators

Normal modes are states of a system where all parts are moving with the same frequency. General strategy to find normal modes:

- 1. Write down the n equations of motions corresponding to the n degrees of freedom the system has.
- 2. Substitute  $x_i = A_i e^{i\omega t}$  into the differential equations to get a system of linear equations in  $A_i$ , with  $i = 1, 2, \dots, n$
- 3. Non-trivial solutions exist if and only if the determinant of the matrix is zero. Solve for  $\omega$ , and subsequently find  $A_i$

The motion of the system can then be decomposed into linear combinations of its normal modes.

# 5.1.5 Small Oscillations

For an object at a local minimum of a potential well, we can expand V(x) about the equilibrium point:

$$V(x) = V(x_0) + V'(x_0)(x - x_0) + \frac{1}{2!}V''(x_0)(x - x_0)^2 + \cdots$$

As  $V(x_0)$  is an additive constant, and  $V'(x_0) = 0$  by definition of equilibrium,

$$V(x) \approx \frac{1}{2}V''(x_0)(x - x_0)^2$$
$$F = -\frac{dV}{dx} = -V''(x_0)(x - x_0)$$
$$\omega = \sqrt{\frac{V''(x_0)}{m}}$$

# 5.2 Wave Equation

A wave is a disturbance of a continuous medium that propagates with a fixed shape at constant velocity. In one dimension:

$$f(z,t) = f(z - vt, 0) = g(z - vt)$$

All such functions f are the solutions to the wave equation:

$$\frac{\partial^2 f}{\partial x^2} = \frac{1}{v^2} \frac{\partial^2 f}{\partial t^2}$$

Where v is the speed of propagation.

# 5.2.1 Electromagnetic Waves

# 6 Optics

# 6.1 Geometric Optics

Results from Fermat's principle of least time:

$$\theta_{\text{incidence}} = \theta_{\text{reflection}}$$
 $n_1 \sin \theta_1 = n_2 \sin \theta_2$ 

Sign convention:

- Light rays travel from left to right
- f is positive if surface makes rays more convergent
- Distances are measured from the surface (left is negative)
- $s_o$  is negative for real objects
- $s_i$  is positive for real images
- y above optical axis is positive

$$\frac{1}{f} = \frac{1}{s_i} + \frac{1}{s_o}$$
$$M = \frac{y_i}{y_o} = -\frac{s_i}{s_o}$$

For thin lenses and mirrors:

$$\frac{1}{f} = \frac{2}{R}$$

For composite thin lenses:

$$\frac{1}{f} = (n-1)\left(\frac{1}{R_1} + \frac{1}{R_2}\right)$$

Lens formed by interface of two materials with different n:

$$\frac{n_2 - n_1}{R} = \frac{n_2}{s_i} + \frac{n_1}{s_o}$$

### 6.2 Polarization

For polarized light:

$$E = E_0 \cos \theta$$
$$I = I_0 \cos^2 \theta$$

For unpolarized light:

$$\langle I \rangle = I_0 \langle \cos^2 \theta \rangle = \frac{I_0}{2}$$

Brewster angle at which all reflected light at an interface is polarized:

$$\tan \theta_i = \frac{n_t}{n_i}$$

# 6.3 Physical Optics

Interference is the superposition of wave amplitudes when waves overlap.

### 6.3.1 Double Slit:

Occurs when slits are of negligible width, distance between slits comparable to wavelength, such that diffraction effects are insignificant. For bright fringes:

$$d\sin\theta = m\lambda$$
 
$$y_m = R\frac{m\lambda}{d} \qquad m \in \mathbb{Z}$$

For incident medium's refractive index  $n_i$ , reflection medium's refractive index  $n_r$ , if  $n_i < n_r$ , the reflected wave undergoes a  $\frac{\pi}{2}$  phase shift.

# 6.3.2 Single Slit:

Occurs when size of slit is comparable to wavelength. Location of dark fringes when wavelets at distance  $\frac{a}{2}$  destructively interfere:

$$\sin \theta = \frac{m\lambda}{d}$$

$$y_m = x \frac{m\lambda}{a} \qquad m \in \mathbb{Z}$$

# 6.3.3 Intensity in Diffraction Patterns

For double slit interference:

$$I = I_{\text{max}} \cos^2 \left( \frac{\pi d \sin \theta}{\lambda} \right)$$

For single slit diffraction:

$$I = I_{\text{max}} \left[ \frac{\sin(\pi a \sin \theta / \lambda)}{\pi a \sin \theta / \lambda} \right]^2$$

Double slit including effects of diffraction:

$$I = I_{\text{max}} \cos^2 \left( \frac{\pi d \sin \theta}{\lambda} \right)$$
$$\cdot \left[ \frac{\sin(\pi a \sin \theta / \lambda)}{\pi a \sin \theta / \lambda} \right]^2$$

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