

Navier-Stokes Equations

Application:

Incompressible Newtonian fluid simulation

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OUTLINE

- 1. Project Topic**
- 2. Background Knowledge**
- 3. Preliminary**
- 4. Schedule of the Project**

1. Project Topic

We all have know how to express the motion of the particle.

How does those expressions apply to macroscopic fluids?



1. Project Topic



This project will mainly go back to the **nature** of **fluid molecular motion**, which would be the basic of computer animation theory.

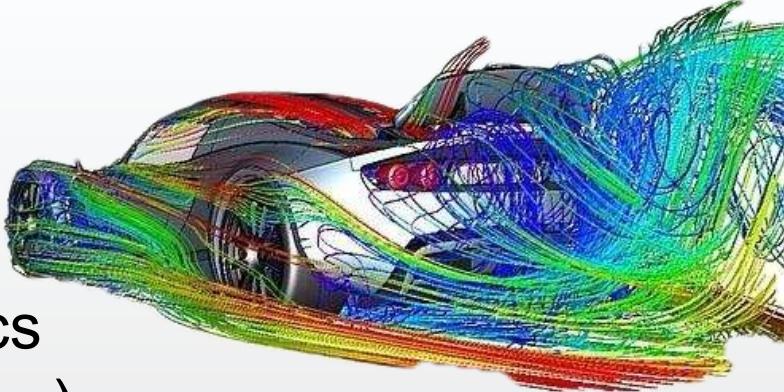
The motion **trajectory** of the fluid will be simulated by using the given boundary conditions and initial value conditions.

1. Project Topic



This project will focus on the **Incompressible Newtonian fluid simulation**. Common **laminar** and **turbulent** flow phenomena are the main research targets.

2. Background



Hydrodynamics
(gas, water, etc...)



Essence: Laws of Conservation



The Conservation Law of:

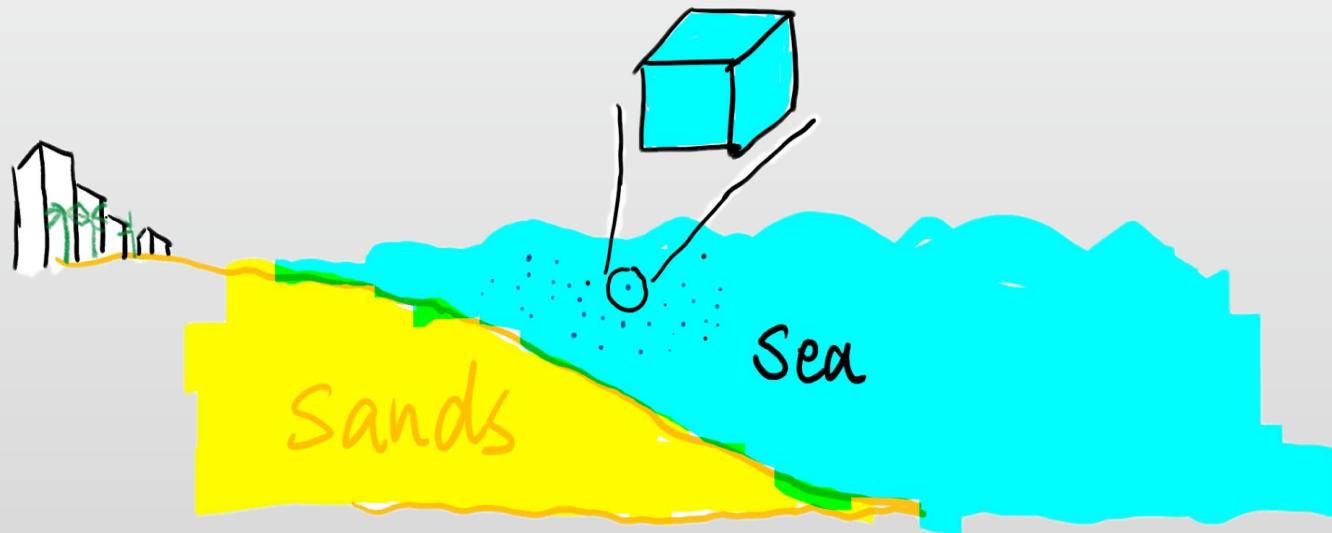
Mass: $F = m \cdot a$

Momentum: $I = F \cdot t = m \cdot v$

Energy: $\Delta E = Q + W$

2. Background

Try to imagine, there is a sea.
The sea is made up of
countless tiny drops of water...

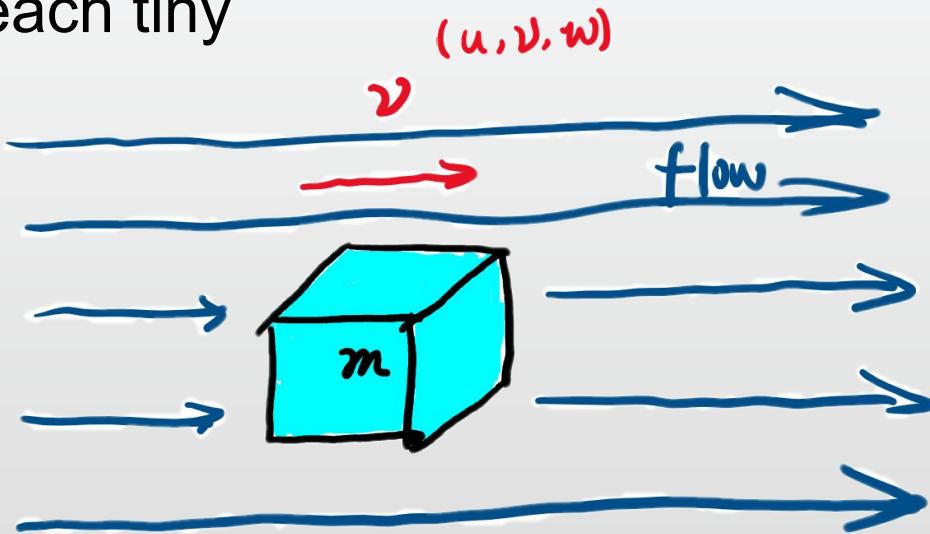


So how to describe the motion status of each tiny drop of water?

2. Background

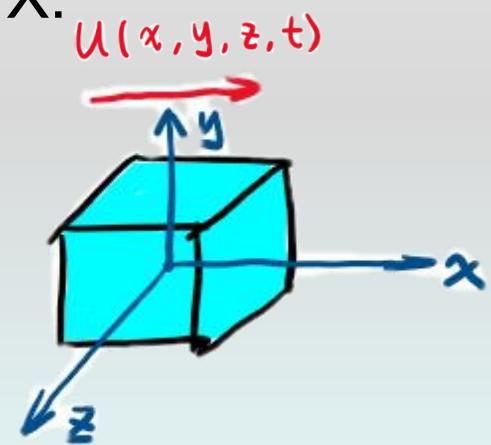
The motion status of each tiny drop of water?

$$\begin{aligned} F &= m^*a \\ \Rightarrow F &= m^* (dv/dt) \\ \Rightarrow F &= (\rho^*V)^*(dv/dt) \end{aligned}$$

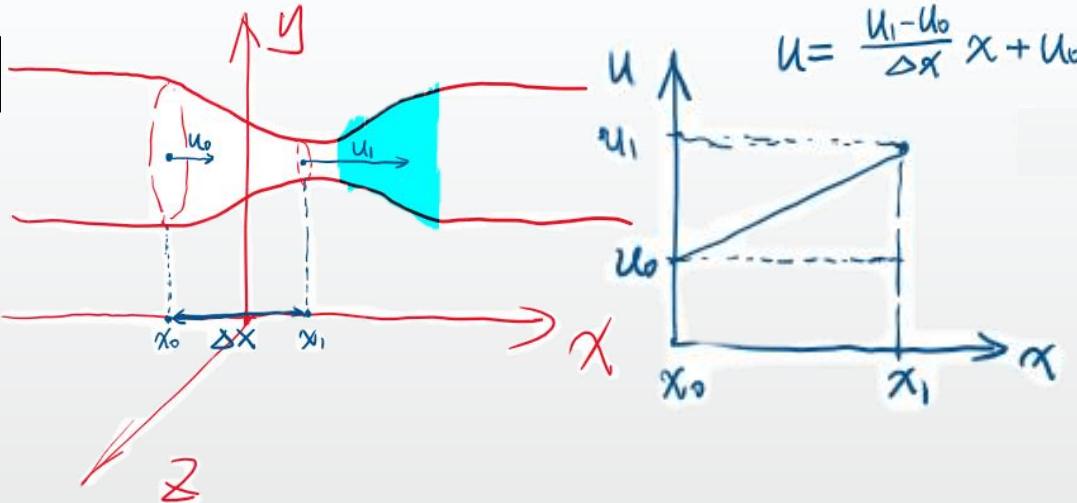


In 3D, If we want to analys the direction of X:

$$\begin{aligned} \Rightarrow a_x &= du/dt = (\partial u / \partial t) * (dt/dt) \\ &\quad + (\partial u / \partial x) * \underline{(dx/dt)} \rightarrow u \\ &\quad + (\partial u / \partial y) * \underline{(dy/dt)} \rightarrow v \\ &\quad + (\partial u / \partial z) * \underline{(dz/dt)} \rightarrow w \end{aligned}$$



2. Background



Thus, we get:

$$F_x = \rho^* [(\partial u / \partial t) + (\partial u / \partial x) * u + (\partial u / \partial y) * v + (\partial u / \partial z) * w]$$



Local Acceleration



Convective Acceleration

$$= \rho^* [(\partial u / \partial t) + \nabla \cdot \nabla u]$$

Now let's talk about the **force (F)** on the water cube :

$$F_x = f_{x\text{inside}} \rightarrow \tau \text{ (viscous shear stress)} + \sigma \text{ (viscous normal stress)}$$

$$+ F_{x\text{outside}} \rightarrow m^*g \text{ (can be ignored if we don't think gravity)}$$

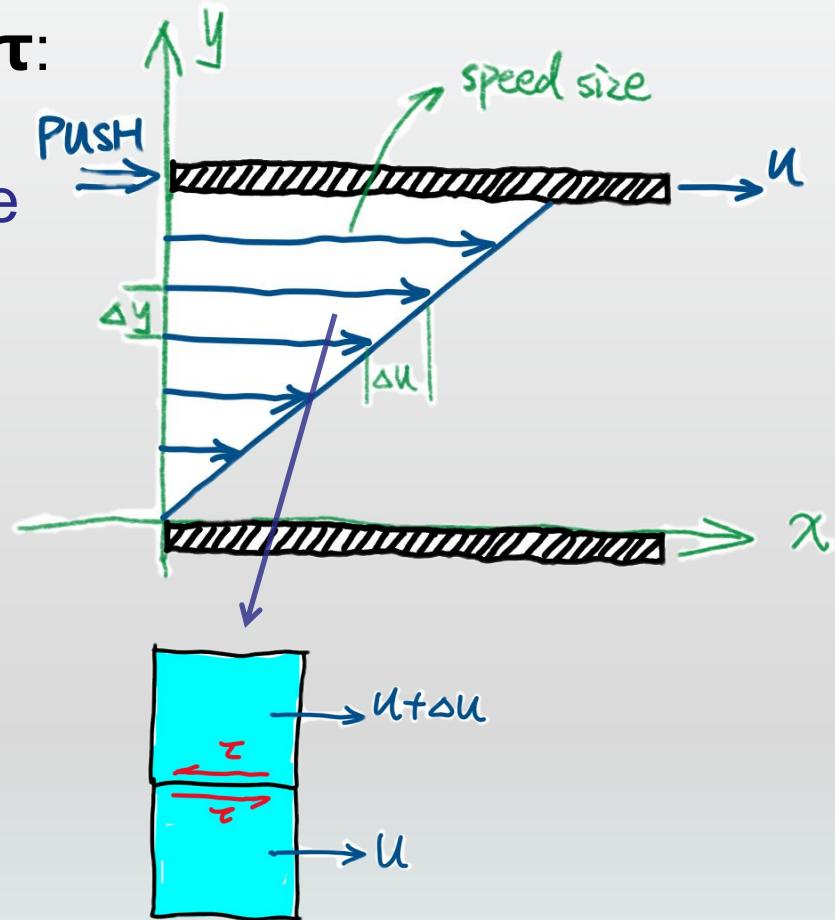
2. Background

Let's talk about the **shear stress τ** :

For the viscous Newton fluid, if we think about the 2D model, we will get the **viscous shear stress τ** between fluid molecules:

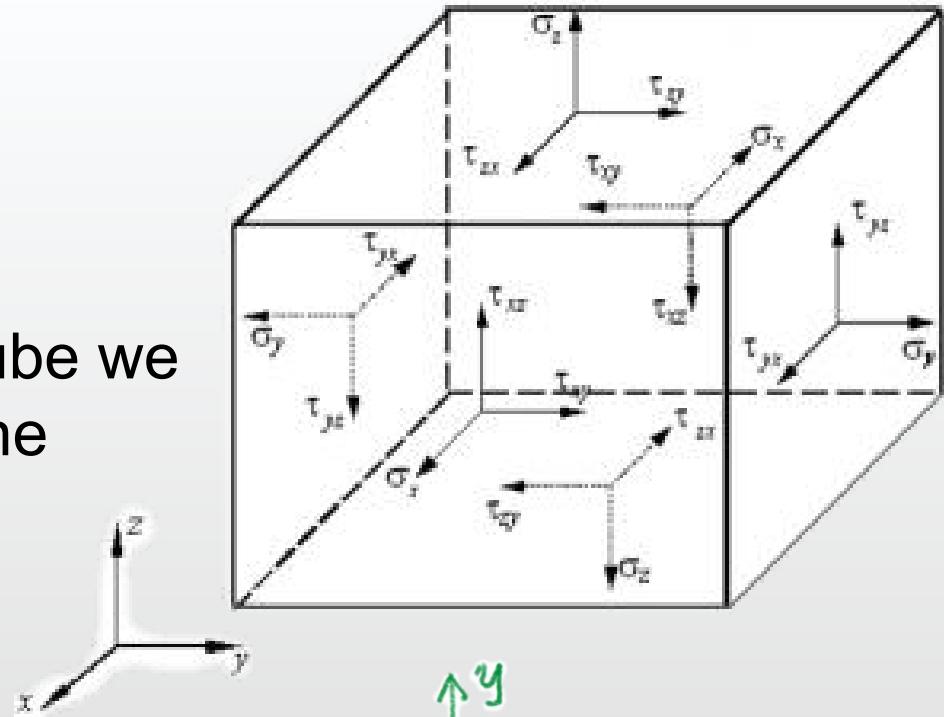
$$\tau = \mu^*(du/dy)$$

μ is the coefficient of viscosity



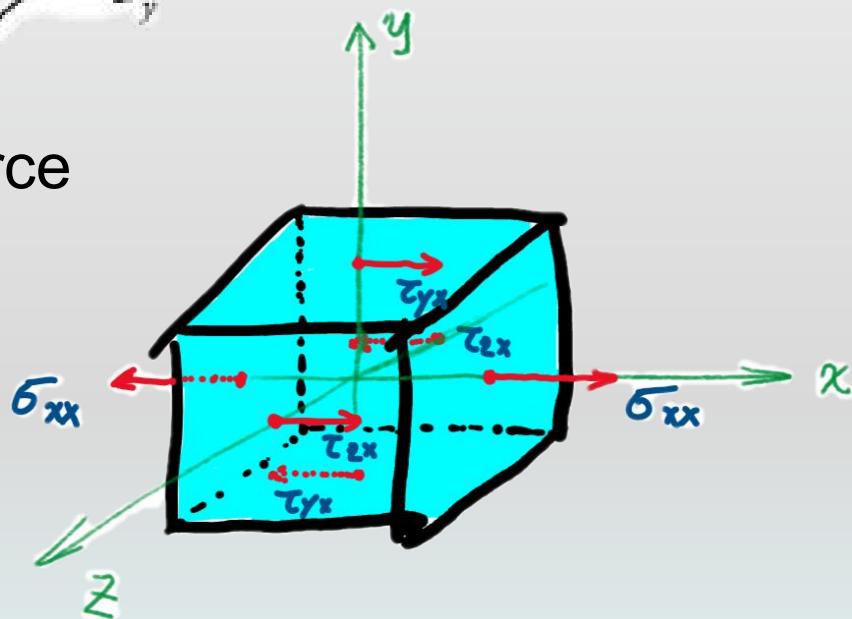
2. Background

So for the 3D model, for the cube we not only need to considerate the **shear stress τ** , but also the **normal stress σ** .



Let's think about those internal force in the direction of x:

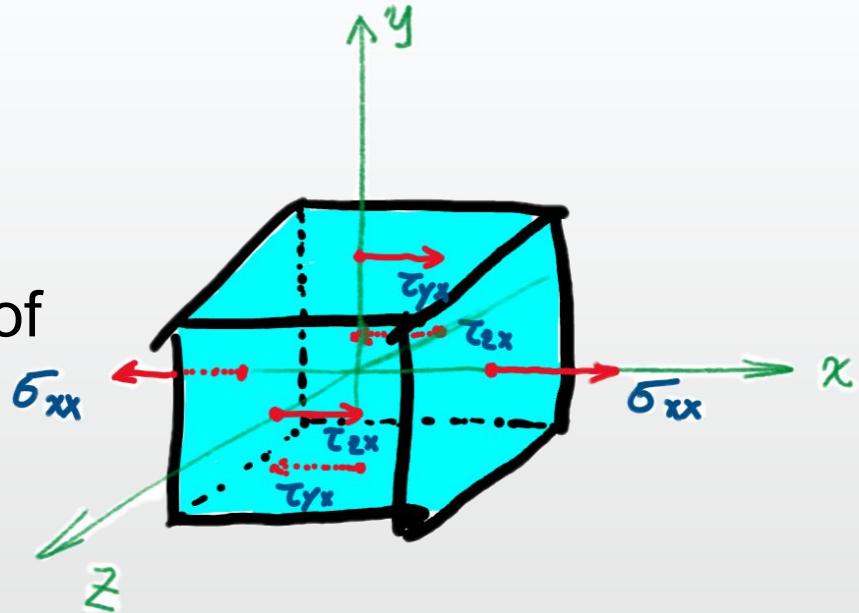
$$\begin{aligned} f_x &= (\partial\sigma_{xx}/\partial x)dxdydz \rightarrow (\partial\sigma_{xx}/\partial x)*1 \\ &+ (\partial\tau_{yx}/\partial y)dxdydz \rightarrow (\partial\tau_{yx}/\partial y)*1 \\ &+ (\partial\tau_{zx}/\partial z)dxdydz \rightarrow (\partial\tau_{zx}/\partial z)*1 \end{aligned}$$



2. Background

The internal force in the direction of x:

$$f_x = (\partial\sigma_{xx}/\partial x) + (\partial\tau_{yx}/\partial y) + (\partial\tau_{zx}/\partial z)$$



Based on the **Constitutive Equation** (for incompressible newton fluid) we've talked in 2D situation:

$$\sigma_{xx} = -p + \underline{\Delta\sigma_{xx}} = -p + 2\mu^*(\partial V_x/\partial x) - (2/3)\mu^*[(\partial V_x/\partial x) + (\partial V_y/\partial y) + (\partial V_z/\partial z)]$$

(Additional viscous stress) (Incompressible part)

(Compressible part)

$$= -p + 2\mu^*(\partial u/\partial x) - (2/3)\mu^*(\nabla V)$$

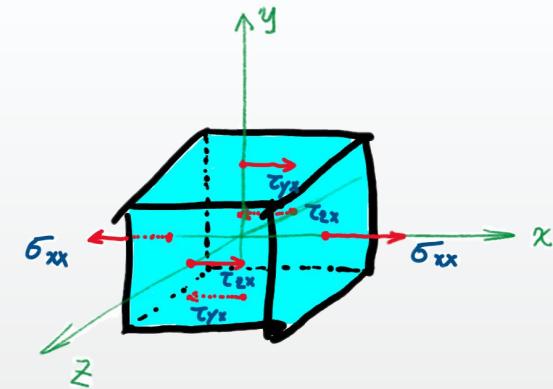
$$\tau_{yx} = \mu^*[(\partial V_y/\partial x) + (\partial V_x/\partial y)] = \mu^*[(\partial v/\partial x) + (\partial u/\partial y)]$$

$$\tau_{zx} = \mu^*[(\partial V_z/\partial x) + (\partial V_x/\partial z)] = \mu^*[(\partial w/\partial x) + (\partial u/\partial z)]$$

2. Background

Thus, for the **incompressible** newton fluid we have:

$$\begin{aligned} F_x &= f_x + F_{x\text{outside}} \\ &= [(\partial \sigma_{xx}/\partial x) + (\partial \tau_{yx}/\partial y) + (\partial \tau_{zx}/\partial z)] + [(\partial/\partial x)*\rho*g] \\ &= \{-(\partial/\partial x)*p + 2\mu*(\partial^2 u/\partial x^2) + \mu*[(\partial^2 v/\partial x\partial y) + (\partial^2 u/\partial y^2)] \\ &\quad + \mu*[(\partial^2 w/\partial x\partial z) + (\partial^2 u/\partial^2 z)]\} + \{(\partial/\partial x)*\rho*g\} \end{aligned}$$



If we ignore the gravity:

$$\begin{aligned} F_x &= -(\partial/\partial x)*p + \mu*[2*(\partial^2 u/\partial x^2) + (\partial^2 v/\partial x\partial y) + (\partial^2 u/\partial y^2) + (\partial^2 w/\partial x\partial z) + (\partial^2 u/\partial^2 z)] \\ &= -(\partial/\partial x)*p + \mu*[(\partial^2 u/\partial x^2) + (\partial^2 u/\partial y^2) + (\partial^2 u/\partial^2 z) \\ &\quad + (\partial^2 u/\partial x^2) + (\partial^2 v/\partial x\partial y) + (\partial^2 w/\partial x\partial z)] \\ &= -(\partial/\partial x)*p + \mu*[\nabla^2 u + (\partial^2 u/\partial x\partial x) + (\partial^2 v/\partial x\partial y) + (\partial^2 w/\partial x\partial z)] \\ &= -(\partial/\partial x)*p + \mu*[\nabla^2(\partial V/\partial x) + (\partial^2 \nabla V/\partial x^2)] \\ &= -(\partial/\partial x)*p + \nabla \mu*[\nabla u + (\partial u/\partial x)] \end{aligned}$$

2. Background

In conclusion, if we have x-y-z directions, combining motivation equations, we will have:

$$\left. \begin{aligned} F_x &= \rho^*[(\partial u / \partial t) + V \cdot \nabla u] \\ F_y &= \rho^*[(\partial v / \partial t) + V \cdot \nabla v] \\ F_z &= \rho^*[(\partial w / \partial t) + V \cdot \nabla w] \\ F_x &= -(\partial / \partial x)*p + \nabla \mu^*[\nabla u + (\partial u / \partial x)] \\ F_y &= -(\partial / \partial y)*p + \nabla \mu^*[\nabla v + (\partial v / \partial x)] \\ F_z &= -(\partial / \partial z)*p + \nabla \mu^*[\nabla w + (\partial w / \partial x)] \end{aligned} \right\}$$



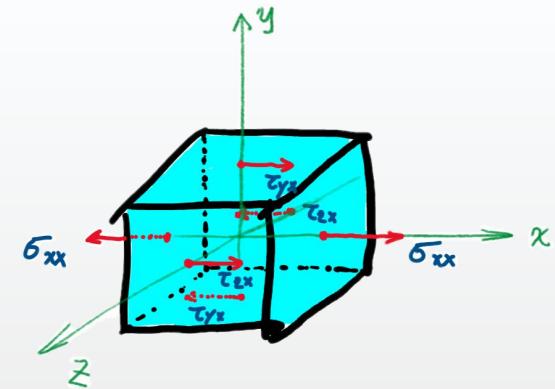
Let's put the system in vector form:

$$\left. \begin{aligned} F &= \rho^*[(\partial V / \partial t) + V \cdot \nabla V] \dots\dots \textcircled{1} \\ F &= -\nabla p + \nabla \mu^*(\nabla V + \nabla V^T) \dots\dots \textcircled{2} \end{aligned} \right\}$$



Of course if you wanna talk about the compressible fluid with gravity, ② can be expressed as:

$$F = -\nabla p + \nabla \mu^*(\nabla V + \nabla V^T) - \nabla(2/3)\mu^2(\nabla V)^*I + \rho^*g \dots\dots \textcircled{2}'$$





2. Background

Thus, combining ① and ②, the **Navier-Stokes Equations** can be expressed as below:

$$\rho^*[(\partial V/\partial t) + V \cdot \nabla V] = -\nabla p + \nabla \mu^*(\nabla V + \nabla V^T)$$

| | | | |
|------------------------------|---|------------------|---------------------|
| Coordinates: (x,y,z) | Time : t | Pressure: p | Heat Flux: q |
| Velocity Components: (u,v,w) | Density: D | Stress: τ | Reynolds Number: Re |
| | | Total Energy: Et | Prandtl Number: Pr |
| Continuity: | $\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u)}{\partial x} + \frac{\partial(\rho v)}{\partial y} + \frac{\partial(\rho w)}{\partial z} = 0$ | | |
| X - Momentum: | $\frac{\partial(\rho u)}{\partial t} + \frac{\partial(\rho u^2)}{\partial x} + \frac{\partial(\rho uv)}{\partial y} + \frac{\partial(\rho uw)}{\partial z} = -\frac{\partial p}{\partial x} + \frac{1}{Re_\tau} \left[\frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} \right]$ | | |
| Y - Momentum: | $\frac{\partial(\rho v)}{\partial t} + \frac{\partial(\rho uv)}{\partial x} + \frac{\partial(\rho v^2)}{\partial y} + \frac{\partial(\rho vw)}{\partial z} = -\frac{\partial p}{\partial y} + \frac{1}{Re_\tau} \left[\frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{yz}}{\partial z} \right]$ | | |
| Z - Momentum: | $\frac{\partial(\rho w)}{\partial t} + \frac{\partial(\rho uw)}{\partial x} + \frac{\partial(\rho vw)}{\partial y} + \frac{\partial(\rho w^2)}{\partial z} = -\frac{\partial p}{\partial z} + \frac{1}{Re_\tau} \left[\frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} \right]$ | | |
| Energy: | $\frac{\partial(E_I)}{\partial t} + \frac{\partial(uE_I)}{\partial x} + \frac{\partial(vE_I)}{\partial y} + \frac{\partial(wE_I)}{\partial z} = -\frac{\partial(ue)}{\partial x} - \frac{\partial(ve)}{\partial y} - \frac{\partial(we)}{\partial z} - \frac{1}{Re_\tau Pr} \left[\frac{\partial q_x}{\partial x} + \frac{\partial q_y}{\partial y} + \frac{\partial q_z}{\partial z} \right]$ | | |
| | $+ \frac{1}{Re_\tau} \left[\frac{\partial}{\partial x} (u \tau_{xx} + v \tau_{xy} + w \tau_{xz}) + \frac{\partial}{\partial y} (u \tau_{xy} + v \tau_{yy} + w \tau_{yz}) + \frac{\partial}{\partial z} (u \tau_{xz} + v \tau_{yz} + w \tau_{zz}) \right]$ | | |

$$\underbrace{\rho \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right)}_1 = \underbrace{-\nabla p + \nabla \cdot (\mu(\nabla \mathbf{u} + (\nabla \mathbf{u})^T)) - \frac{2}{3}\mu(\nabla \cdot \mathbf{u})\mathbf{I}}_2 + \underbrace{\mathbf{F}}_4$$

<https://www.comsol.com/multiphysics/navier-stokes-equations?parent=modeling-conservation-mass-energy-momentum-0402-432-302>

Don't forget! our aim is researching a **SEA** not only one tiny water cube. Thus, we need to have continuity equation to confine our NS equation:

$$\partial \rho / \partial t + \nabla \cdot (\rho \mathbf{V}) = \partial \rho / \partial t + \partial \rho u / \partial x + \partial \rho v / \partial y + \partial \rho w / \partial z = 0$$

2. Background

Continuity equation shows our fluid is always a whole with a balanced mass distribution, which will eliminate the compressible part of the NS equation.

$$\rho^*[(\partial \mathbf{V}/\partial t) + \mathbf{V} \cdot \nabla \mathbf{V}] = -\nabla p + \nabla \mu^*(\nabla \cdot \mathbf{V} + \nabla \mathbf{V}^T)$$

$$\underbrace{\rho \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right)}_1 = \underbrace{-\nabla p}_2 + \underbrace{\nabla \cdot (\mu(\nabla \mathbf{u} + (\nabla \mathbf{u})^T) - \frac{2}{3}\mu(\nabla \cdot \mathbf{u})\mathbf{I})}_3 + \underbrace{\mathbf{F}}_4$$

<https://www.comsol.com/multiphysics/navier-stokes-equations?parent=modeling-conservation-mass-energy-momentum-0402-432-302>

Compressibility can be measured by Match Number:

$$M = |\mathbf{V}|/C, \text{ (C is the sound speed in the fluid)}$$

High Match Number means compressible flow;
Low Match Number means incompressible flow;

2. Background

By the way, we can also have the restricted condition by adding the energy equation like:

| Coordinates: (x,y,z) | | Time : t | Pressure: p | Heat Flux: q |
|------------------------------|--|---|----------------|---------------------|
| Velocity Components: (u,v,w) | | Density: D | Stress: τ | Reynolds Number: Re |
| | | Total Energy: Et | | Prandtl Number: Pr |
| Continuity: | | $\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u)}{\partial x} + \frac{\partial(\rho v)}{\partial y} + \frac{\partial(\rho w)}{\partial z} = 0$ | | |
| X - Momentum: | | $\frac{\partial(\rho u)}{\partial t} + \frac{\partial(\rho u^2)}{\partial x} + \frac{\partial(\rho uv)}{\partial y} + \frac{\partial(\rho uw)}{\partial z} = -\frac{\partial p}{\partial x} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} \right]$ | | |
| Y - Momentum: | | $\frac{\partial(\rho v)}{\partial t} + \frac{\partial(\rho uv)}{\partial x} + \frac{\partial(\rho v^2)}{\partial y} + \frac{\partial(\rho vw)}{\partial z} = -\frac{\partial p}{\partial y} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{yz}}{\partial z} \right]$ | | |
| Z - Momentum: | | $\frac{\partial(\rho w)}{\partial t} + \frac{\partial(\rho uw)}{\partial x} + \frac{\partial(\rho vw)}{\partial y} + \frac{\partial(\rho w^2)}{\partial z} = -\frac{\partial p}{\partial z} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} \right]$ | | |
| Energy: | | $\begin{aligned} \frac{\partial(E_I)}{\partial t} + \frac{\partial(uE_I)}{\partial x} + \frac{\partial(vE_I)}{\partial y} + \frac{\partial(wE_I)}{\partial z} &= -\frac{\partial(ue_p)}{\partial x} - \frac{\partial(ve_p)}{\partial y} - \frac{\partial(we_p)}{\partial z} - \frac{1}{Re \cdot Pr_r} \left[\frac{\partial q_x}{\partial x} + \frac{\partial q_y}{\partial y} + \frac{\partial q_z}{\partial z} \right] \\ &+ \frac{1}{Re_r} \left[\frac{\partial}{\partial x} (u \tau_{xx} + v \tau_{xy} + w \tau_{xz}) + \frac{\partial}{\partial y} (u \tau_{xy} + v \tau_{yy} + w \tau_{yz}) + \frac{\partial}{\partial z} (u \tau_{xz} + v \tau_{yz} + w \tau_{zz}) \right] \end{aligned}$ | | |

$$\frac{\partial}{\partial t} \left[\rho \left(e + \frac{1}{2} u^2 \right) \right] + \nabla \cdot \left[\rho \mathbf{u} \left(e + \frac{1}{2} u^2 \right) \right] = \nabla \cdot (k \nabla T) + \nabla \cdot (-p \mathbf{u} + \boldsymbol{\tau} \cdot \mathbf{u}) + \mathbf{u} \cdot \mathbf{F} + \mathcal{Q}$$

<https://www.comsol.com/multiphysics/navier-stokes-equations?parent=modeling-conservation-mass-energy-momentum-0402-432-302>

However, for incompressible flows, the temperature equation is **completely decoupled** from the NS equation, so we won't talk about the equation here.



2. Background

Next for the NS equations, let's eliminate the **dimensions** of the equation:

$$\rho^*[(\partial V/\partial t) + V \cdot \nabla V] = -\nabla p + \nabla \mu^*(\nabla V + \nabla V^\top)$$

We can multiply the both side of the equation by a factor $L/(\rho^* \tilde{V}^2)$. L is the characteristic length, ρ is the fluid density, \tilde{V} is the mean velocity. Then we will get:

$$(L/\tilde{V}^2)^*[(\partial V/\partial t) + V \cdot \nabla V] = -\nabla p^* L/(\rho^* \tilde{V}^2) + \nabla \mu L/(\rho^* \tilde{V}^2)^*(\nabla V + \nabla V^\top)$$

Set $V' = V/\tilde{V}$, $p' = p^*(1/(\rho^* \tilde{V}^2))$, $(\partial/\partial t') = (L/\tilde{V})^*(\partial/\partial t)$, $\nabla' = L \nabla$:

$$\Rightarrow [(\partial V'/\partial t') + V' \cdot \nabla' V'] = -\nabla' p' + \mu/(\rho^* L^* \tilde{V})^* \nabla' (\nabla' V' + \nabla' V'^\top)$$

$$\Rightarrow [(\partial V/\partial t) + V \cdot \nabla V] = -\nabla p + \mu/(\rho^* L^* \tilde{V})^* \nabla (\nabla V + \nabla V^\top)$$



Inverse of Reynolds number

$$\Rightarrow [(\partial V/\partial t) + V \cdot \nabla V] = -\nabla p + 1/Re^* \nabla (\nabla V + \nabla V^\top)$$

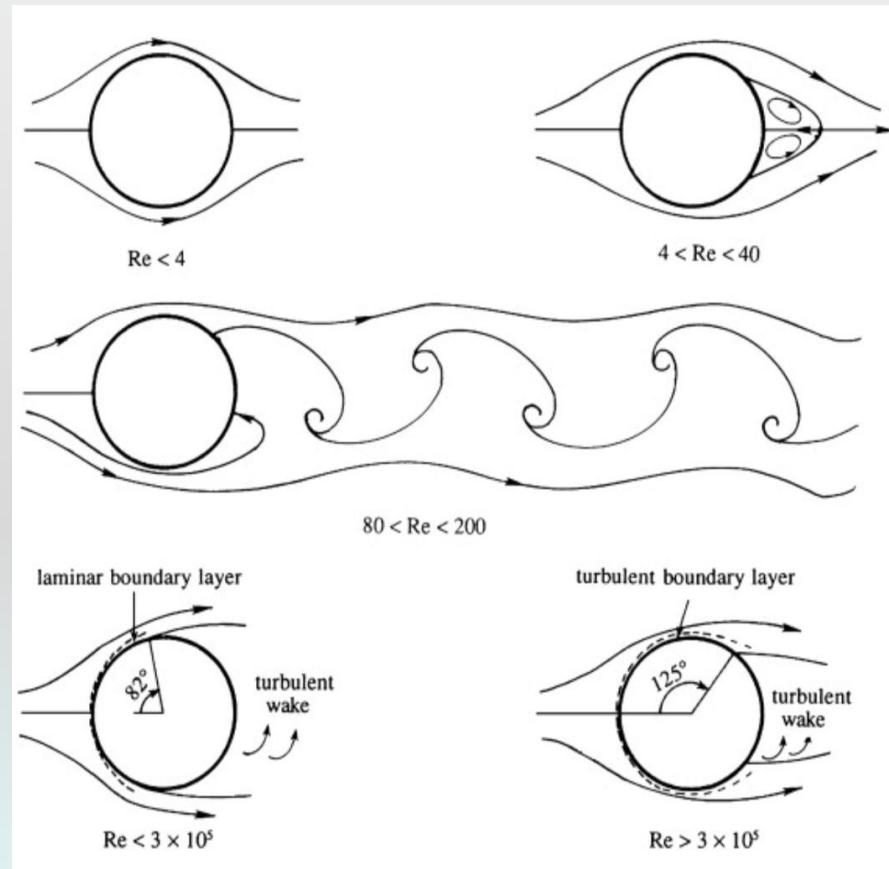
| Coordinates: (x,y,z) | Time: t | Pressure: p | Heat Flux: q |
|------------------------------|---|------------------|---------------------|
| Velocity Components: (u,v,w) | Density: D | Stress: τ | Reynolds Number: Re |
| | | Total Energy: Et | Prandtl Number: Pr |
| Continuity: | $\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u)}{\partial x} + \frac{\partial(\rho v)}{\partial y} + \frac{\partial(\rho w)}{\partial z} = 0$ | | |
| X - Momentum: | $\frac{\partial(\rho u)}{\partial t} + \frac{\partial(\rho u^2)}{\partial x} + \frac{\partial(\rho uv)}{\partial y} + \frac{\partial(\rho uw)}{\partial z} = -\frac{\partial p}{\partial x} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} \right]$ | | |
| Y - Momentum: | $\frac{\partial(\rho v)}{\partial t} + \frac{\partial(\rho uv)}{\partial x} + \frac{\partial(\rho v^2)}{\partial y} + \frac{\partial(\rho vw)}{\partial z} = -\frac{\partial p}{\partial y} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{yz}}{\partial z} \right]$ | | |
| Z - Momentum: | $\frac{\partial(\rho w)}{\partial t} + \frac{\partial(\rho uw)}{\partial x} + \frac{\partial(\rho vw)}{\partial y} + \frac{\partial(\rho w^2)}{\partial z} = -\frac{\partial p}{\partial z} + \frac{1}{Re \cdot \epsilon_r} \left[\frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} \right]$ | | |
| Energy: | $\begin{aligned} \frac{\partial(E_I)}{\partial t} + \frac{\partial(uE_I)}{\partial x} + \frac{\partial(vE_I)}{\partial y} + \frac{\partial(wE_I)}{\partial z} &= -\frac{\partial(up)}{\partial x} - \frac{\partial(vp)}{\partial y} - \frac{\partial(wp)}{\partial z} - \frac{1}{Re \cdot Pr_r} \left[\frac{\partial q_x}{\partial x} + \frac{\partial q_y}{\partial y} + \frac{\partial q_z}{\partial z} \right] \\ &+ \frac{1}{Re_r} \left[\frac{\partial}{\partial x} (u \tau_{xx} + v \tau_{xy} + w \tau_{xz}) + \frac{\partial}{\partial y} (u \tau_{xy} + v \tau_{yy} + w \tau_{yz}) + \frac{\partial}{\partial z} (u \tau_{xz} + v \tau_{yz} + w \tau_{zz}) \right] \end{aligned}$ | | |

2. Background

Reynolds number: A number shows the ratio between internal forces and viscous forces.

$$Re = (\rho * L * \tilde{V}) / \mu$$

1. Because there is no dimension, incompressible Newtonian flows with the **same Reynolds number** are comparable.
2. **High** Reynolds number will cause **turbulent**. At this time flow is **inviscid**.
3. **Low** Reynolds number will cause **laminar**. At this time flow is **viscous**.



2. Background

$$\left\{ \begin{array}{l} \frac{\partial p}{\partial t} + \nabla(pV) = 0 \\ \left(\frac{\partial V}{\partial t} + V \cdot \nabla V \right) = -\nabla p + 1/\text{Re}^* \nabla (\nabla V + \nabla V^T) \end{array} \right.$$



Navier-Stokes Equations 3-dimensional - unsteady

Glenn
Research
Center

Coordinates: (x, y, z)

Velocity Components: (u, v, w)

Time: t

Density: ρ

Total Energy: E_t

Pressure: p

Stress: τ

Energy: E_t

Heat Flux: q

Reynolds Number: Re

Prandtl Number: Pr

Continuity:
$$\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u)}{\partial x} + \frac{\partial(\rho v)}{\partial y} + \frac{\partial(\rho w)}{\partial z} = 0$$

X - Momentum:
$$\frac{\partial(\rho u)}{\partial t} + \frac{\partial(\rho u^2)}{\partial x} + \frac{\partial(\rho uv)}{\partial y} + \frac{\partial(\rho uw)}{\partial z} = -\frac{\partial p}{\partial x} + \frac{1}{Re_r} \left[\frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} \right]$$

Y - Momentum:
$$\frac{\partial(\rho v)}{\partial t} + \frac{\partial(\rho uv)}{\partial x} + \frac{\partial(\rho v^2)}{\partial y} + \frac{\partial(\rho vw)}{\partial z} = -\frac{\partial p}{\partial y} + \frac{1}{Re_r} \left[\frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{yz}}{\partial z} \right]$$

Z - Momentum
$$\frac{\partial(\rho w)}{\partial t} + \frac{\partial(\rho uw)}{\partial x} + \frac{\partial(\rho vw)}{\partial y} + \frac{\partial(\rho w^2)}{\partial z} = -\frac{\partial p}{\partial z} + \frac{1}{Re_r} \left[\frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} \right]$$

Energy:

$$\begin{aligned} \frac{\partial(E_t)}{\partial t} + \frac{\partial(uE_t)}{\partial x} + \frac{\partial(vE_t)}{\partial y} + \frac{\partial(wE_t)}{\partial z} &= -\frac{\partial(up)}{\partial x} - \frac{\partial(vp)}{\partial y} - \frac{\partial(wp)}{\partial z} - \frac{1}{Re_r Pr_r} \left[\frac{\partial q_x}{\partial x} + \frac{\partial q_y}{\partial y} + \frac{\partial q_z}{\partial z} \right] \\ &+ \frac{1}{Re_r} \left[\frac{\partial}{\partial x} (u \tau_{xx} + v \tau_{xy} + w \tau_{xz}) + \frac{\partial}{\partial y} (u \tau_{xy} + v \tau_{yy} + w \tau_{yz}) + \frac{\partial}{\partial z} (u \tau_{xz} + v \tau_{yz} + w \tau_{zz}) \right] \end{aligned}$$

3. Preliminary

Normally, the **NS equation** is classed as **Nonlinear Partial Differential Equation**, which means the equation is hard to find the analytical solution.

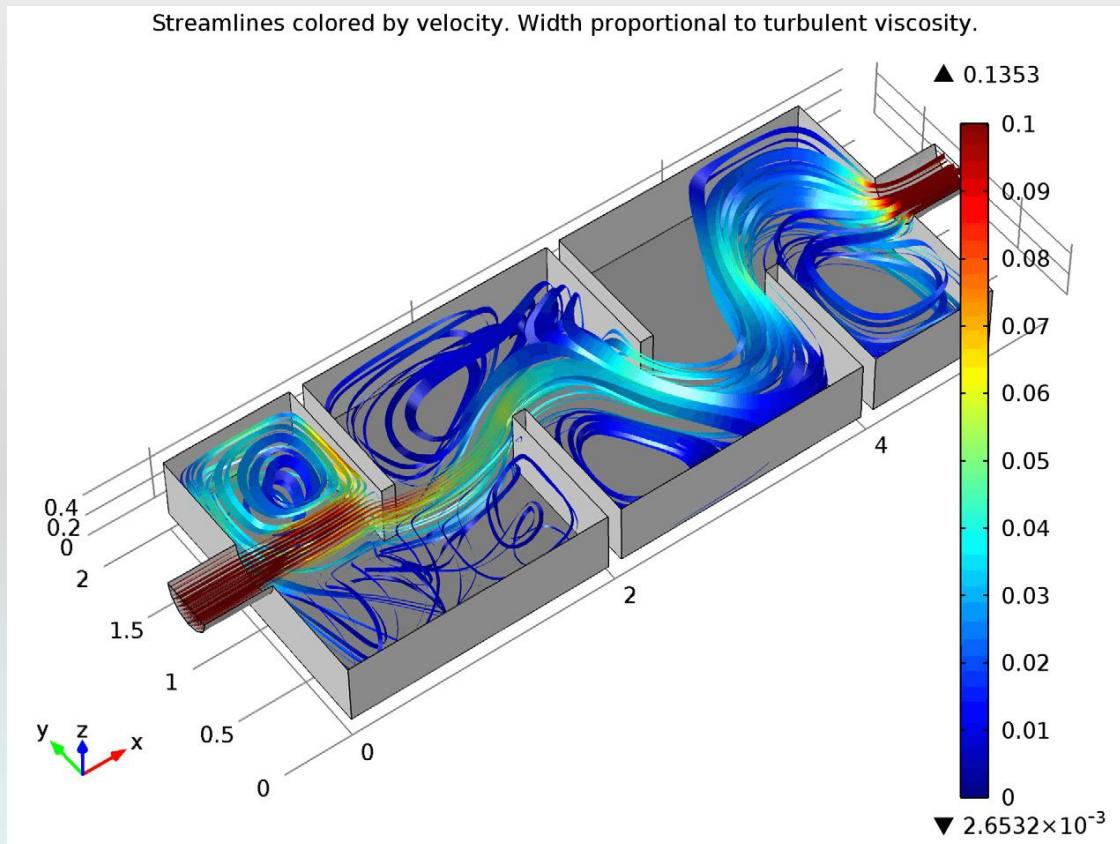
Furthermore, because of the complexity of the NS equation, fluid movement without any constraints takes up a huge amount of computing resources.

Thus, the simulation should be based on the approximated equation and the **spectral boundary conditions**.

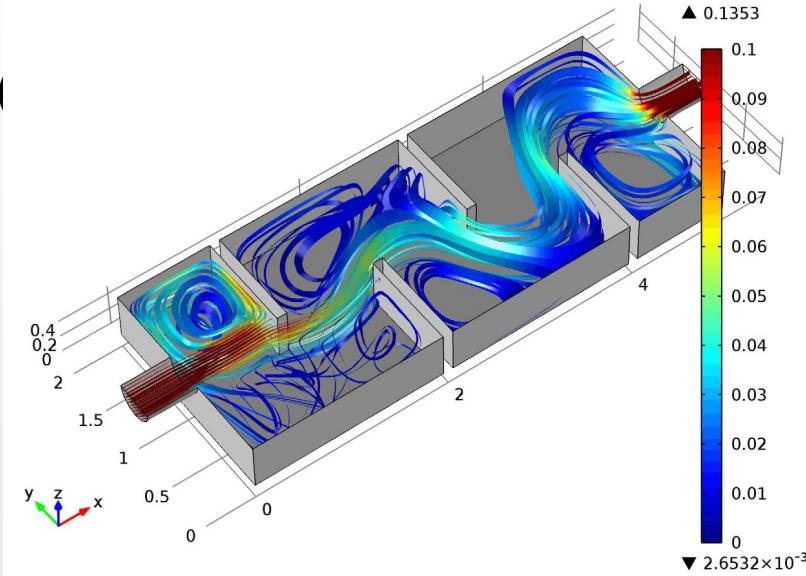
3. Cosmol Experiment

Cosmol would be a useful modeling tool.

The NS equation based on the approximated equation and the spectial boundary conditions (inlet, outlet, walls, etc..) can be simulated in Cosmol.



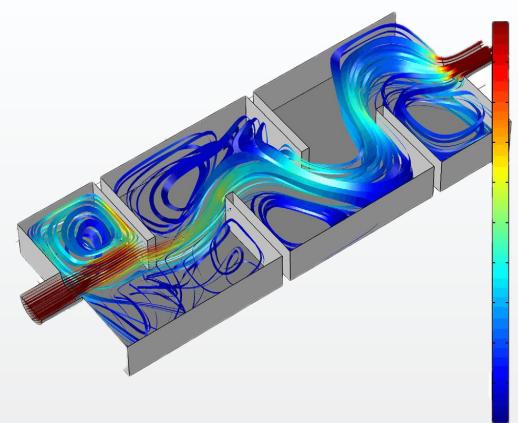
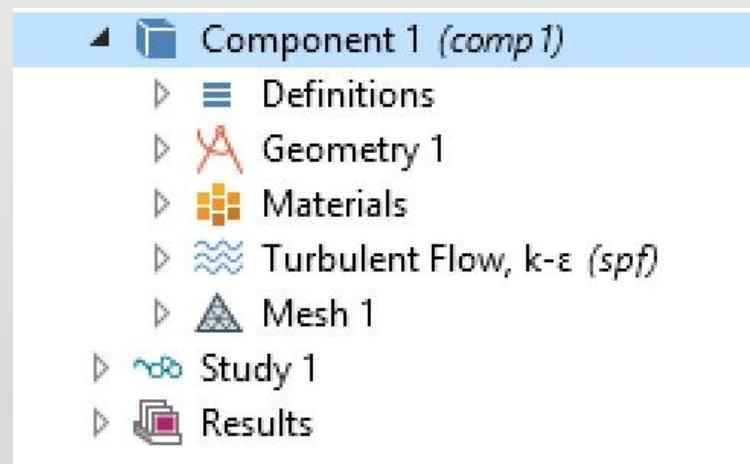
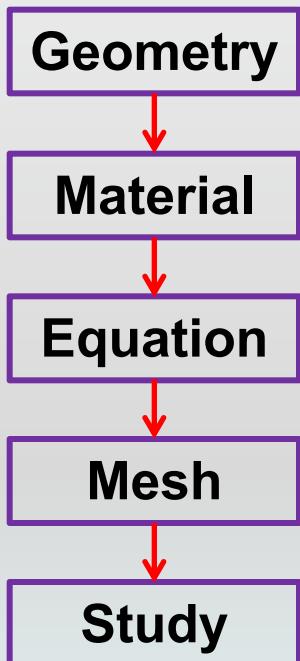
3. Cosmol Experiment



The [initial phase](#) of the project will be used to simulate and compare different NS family equation systems.

3. Comsol Experiment

Comsol has its simulating pipeline for single-phase flow. The pipeline can be simplified as below:



3. Comsol Experiment

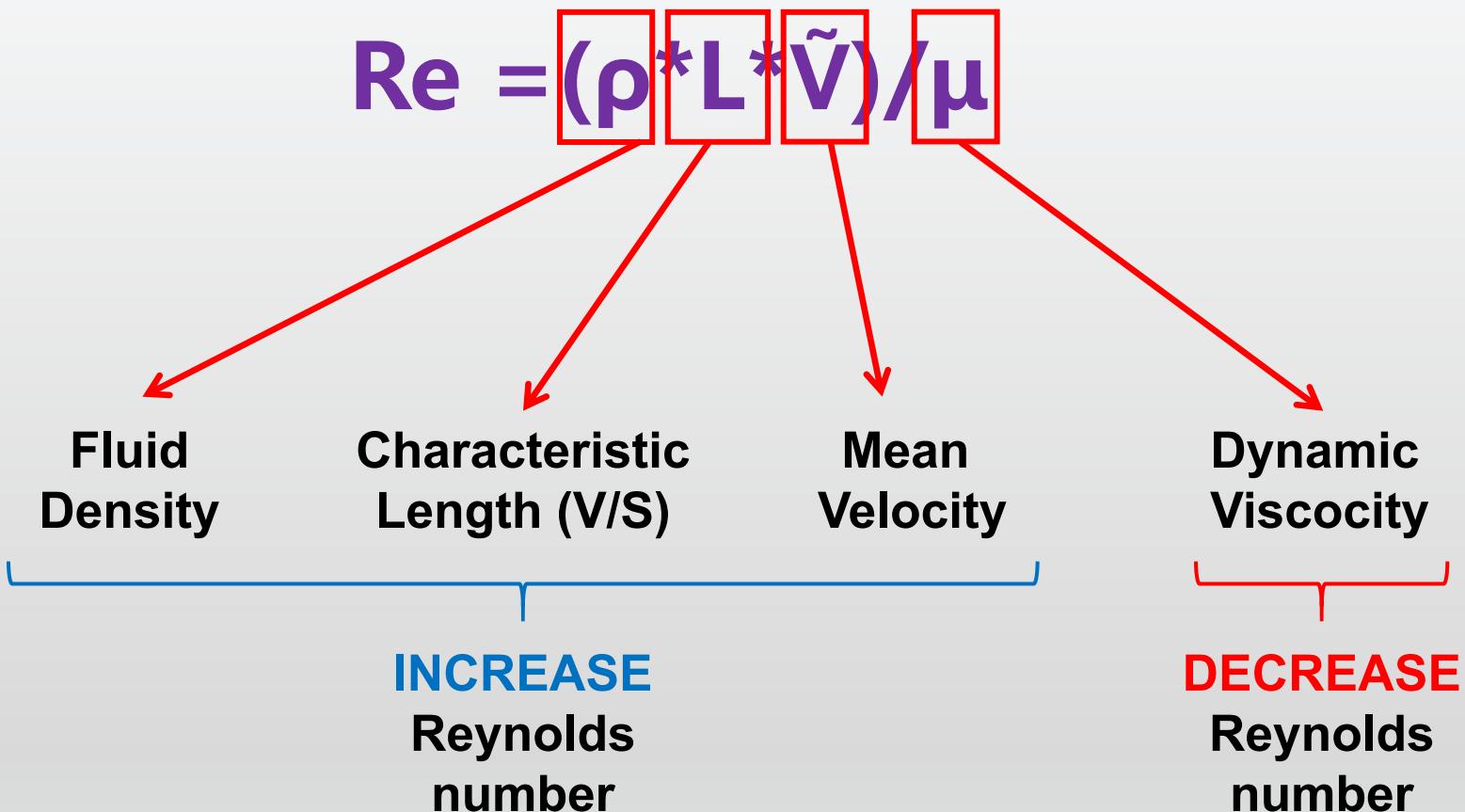
Equation

Comsol provides different kinds of equations for different conditions of flow:

- Laminar Flow
 - Navier-Stokes Equations
- Creeping Flow
 - Stokes Equations
- Non-Newtonian
 - Carreau, Power-Law, and user-defined models
- Turbulent Flow
 - k- ε , k- ω , Low Re k- ε , Shear Stress Transport, and Spalart-Allmaras turbulence models

-  Fluid Flow
-  Single-Phase Flow
 -  Creeping Flow (spf)
 -  Laminar Flow (spf)
-  Turbulent Flow
 -  Turbulent Flow, Algebraic yPlus (spf)
 -  Turbulent Flow, L-VEL (spf)
 -  Turbulent Flow, k- ε (spf) (highlighted)
 -  Turbulent Flow, Realizable k- ε (spf)
 -  Turbulent Flow, k- ω (spf)
 -  Turbulent Flow, SST (spf)
 -  Turbulent Flow, Low Re k- ε (spf)
 -  Turbulent Flow, Spalart-Allmaras (spf)
 -  Turbulent Flow, v²-f (spf)
-  Large Eddy Simulation
-  Rotating Machinery, Fluid Flow
-  Thin-Film Flow
-  Multiphase Flow
-  Porous Media and Subsurface Flow
-  Nonisothermal Flow
-  High Mach Number Flow
-  Particle Tracing

3. Parameter Experiments



3. Parameter Experiments

EXP1: Temperature

Experiment 1: Experimental Group

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

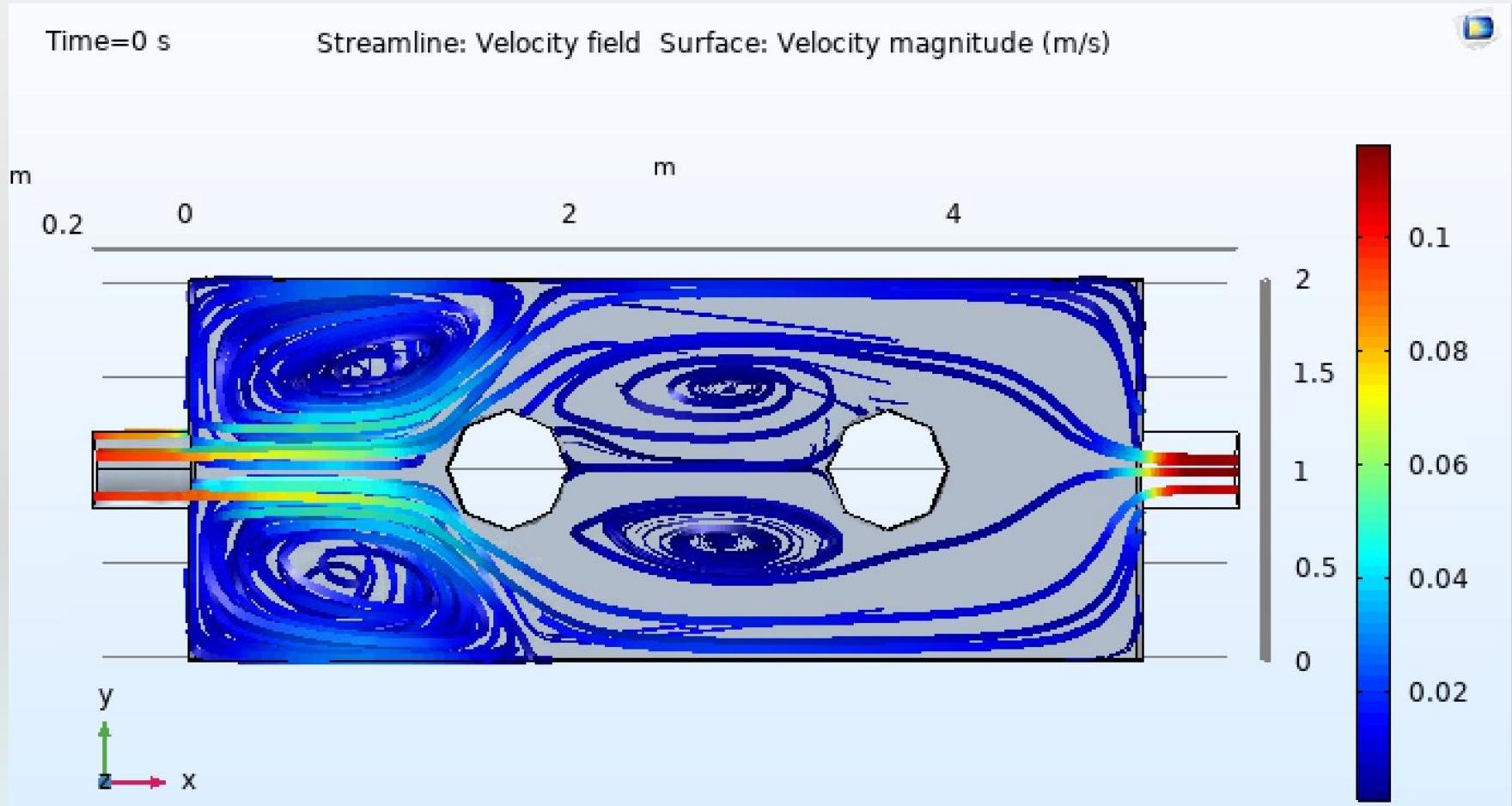
Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 1: Controlled Group 1 - Low Temp.

Material: Water

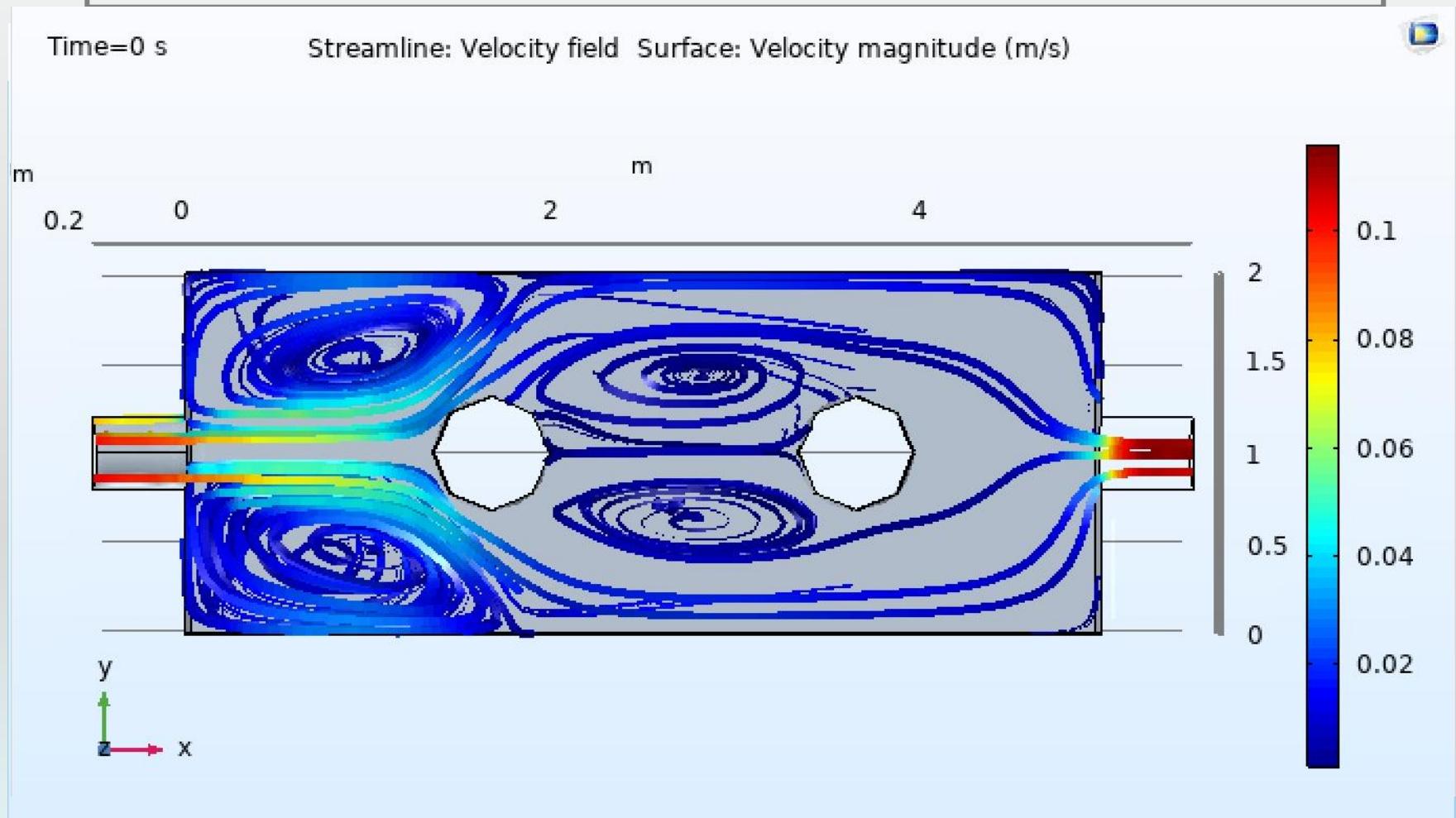
Temperature: 20°C

Density: 1000 kg/m³

Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 1.01×10^{-3} Pa·s



Experiment 1: Controlled Group 2 - High Temp.

Material: Water

Temperature: 100°C

Density: 958.4 kg/m³

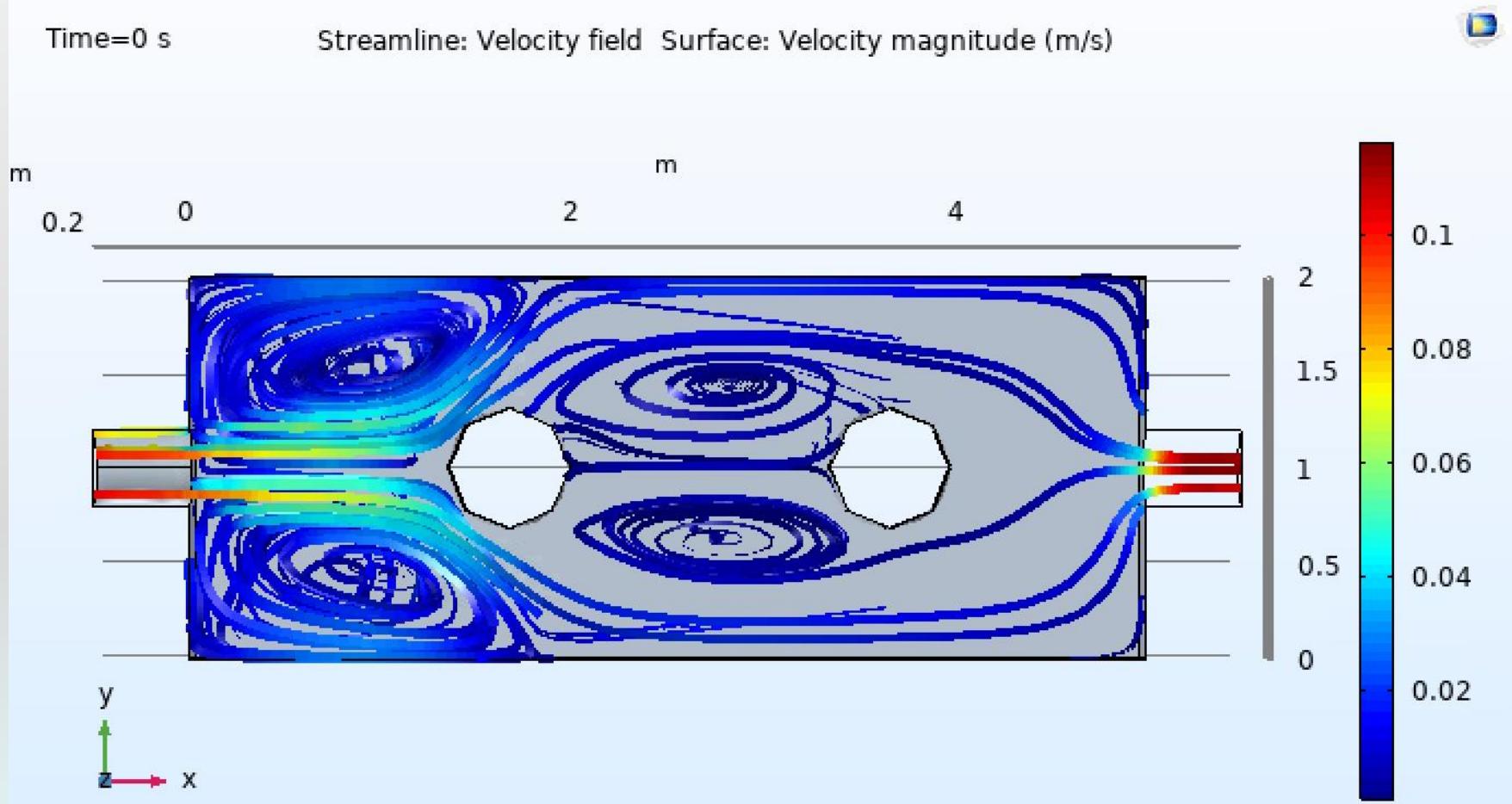
Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 2.8×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



3. Parameter Experiments

EXP2: Density

Experiment 2: Experimental Group

Material: Ethyl Alcohol

Temperature: 20°C

Density: 789 kg/m³

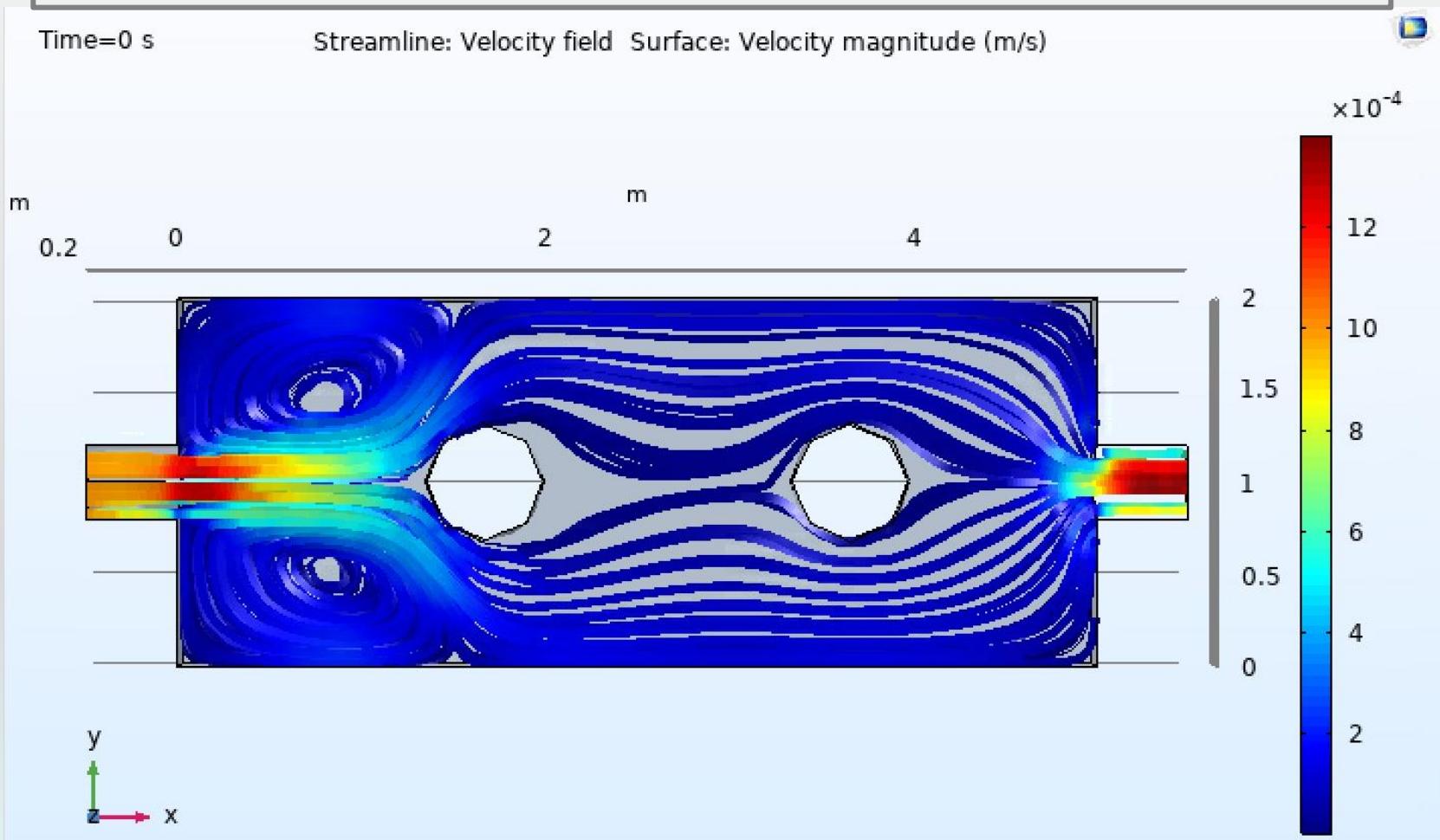
Characteristic Length: Normal

Mean Velocity: 0.001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 1.19×10^{-3} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 2: Controlled Group 1 - High Density

Material: Mercury

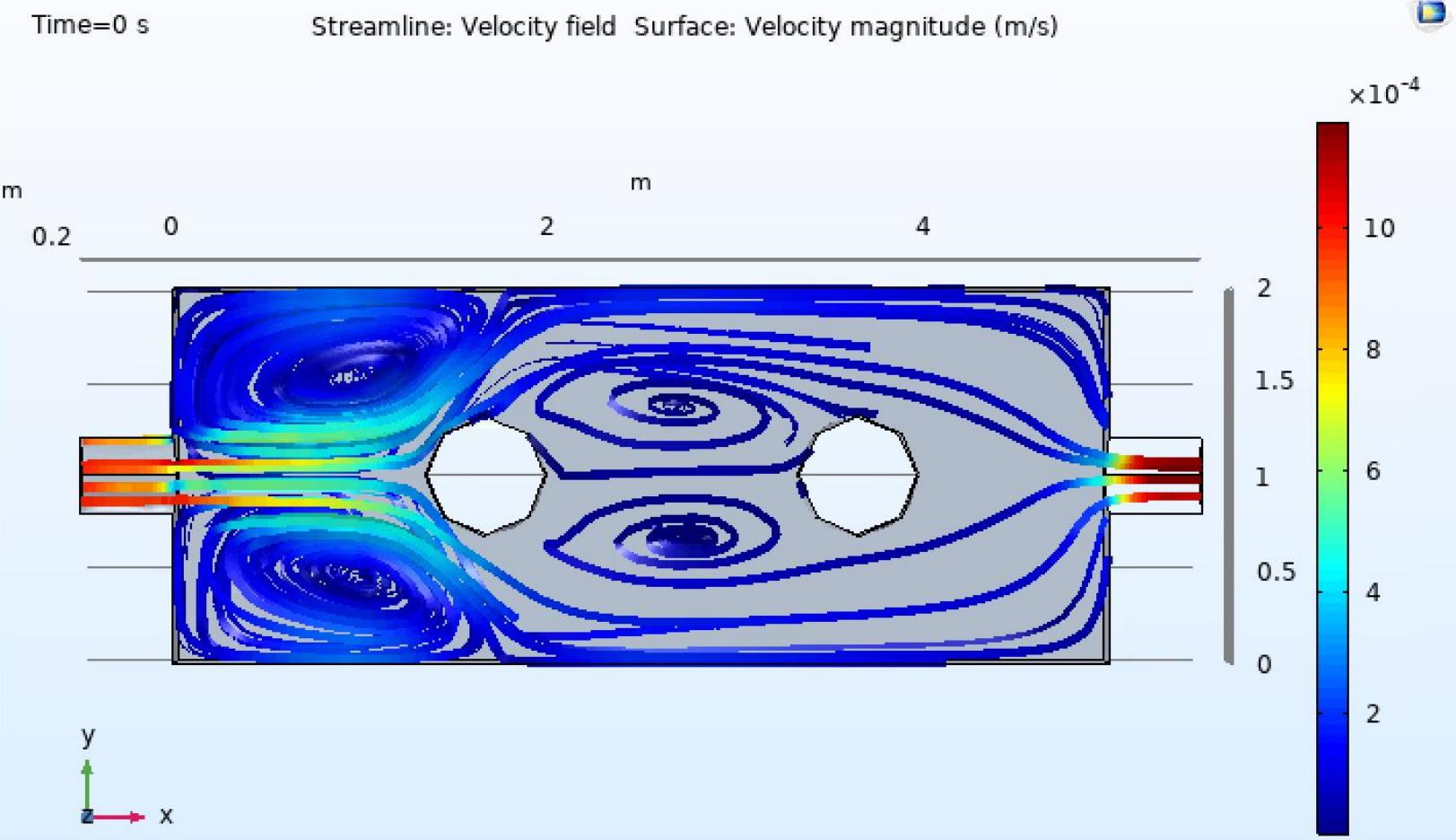
Temperature: 20°C

Density: 13600 kg/m³

Characteristic Length: Normal

Mean Velocity: 0.001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 1.57×10^{-3} Pa·s



3. Parameter Experiments

EXP3: Characteristic Length

Experiment 3: Experimental Group

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

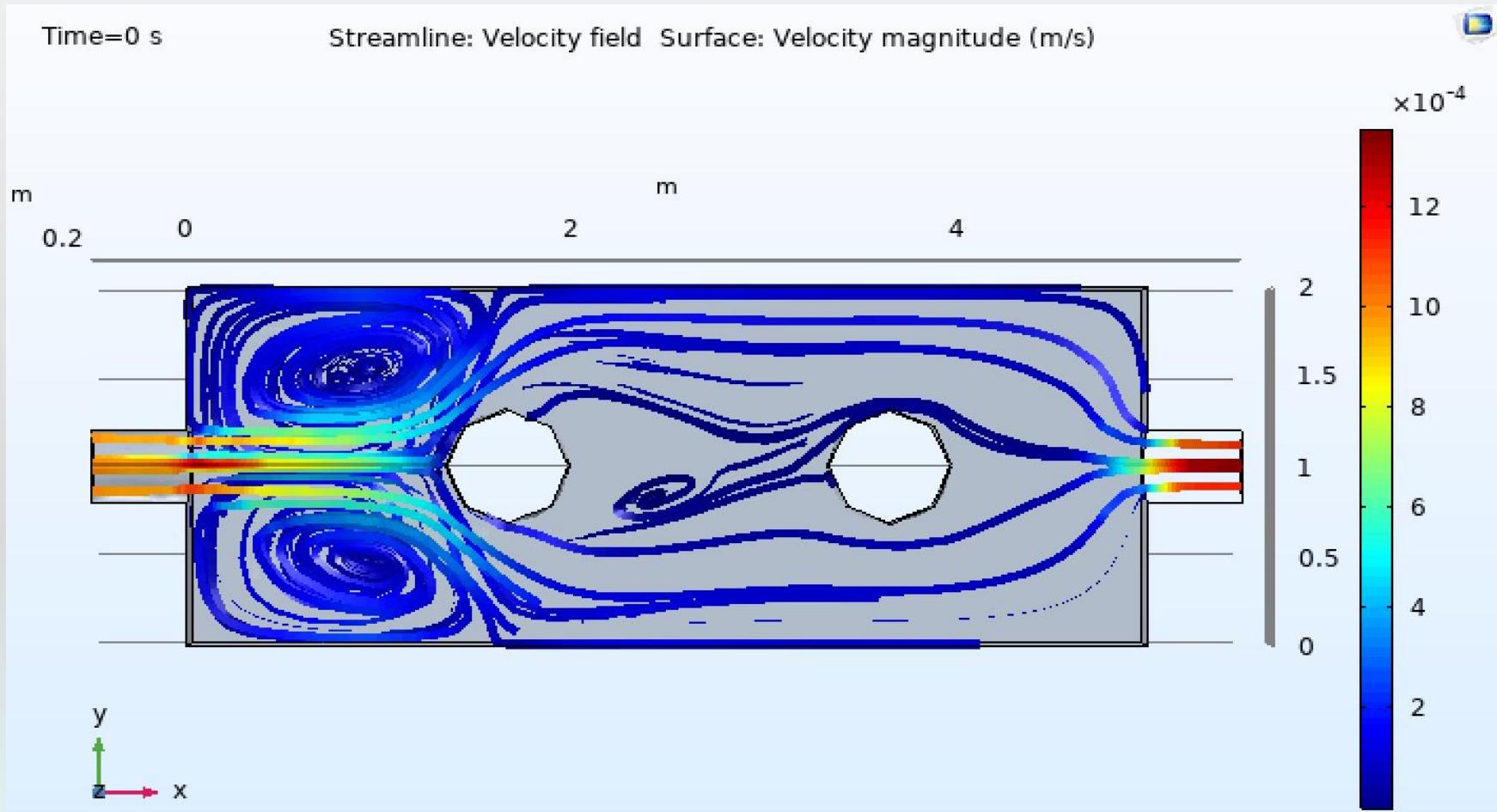
Characteristic Length: Normal

Mean Velocity: 0.001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 3: Controlled Group 1 - Square Size

Material: Water

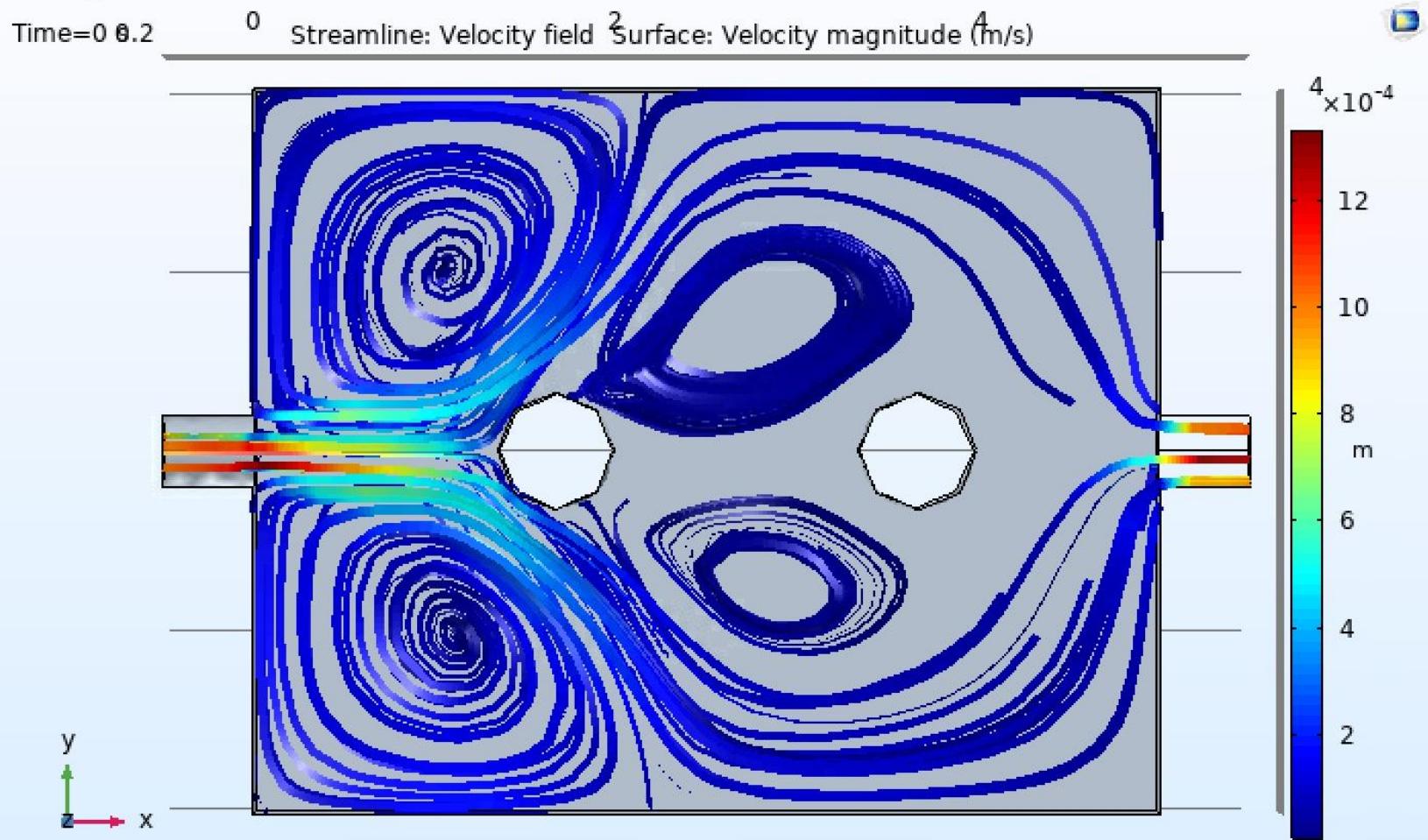
Temperature: 25°C

Density: 997.05 kg/m³

Characteristic Length: Square

Mean Velocity: 0.001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s



Experiment 3: Controlled Group 2 - Long Size

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

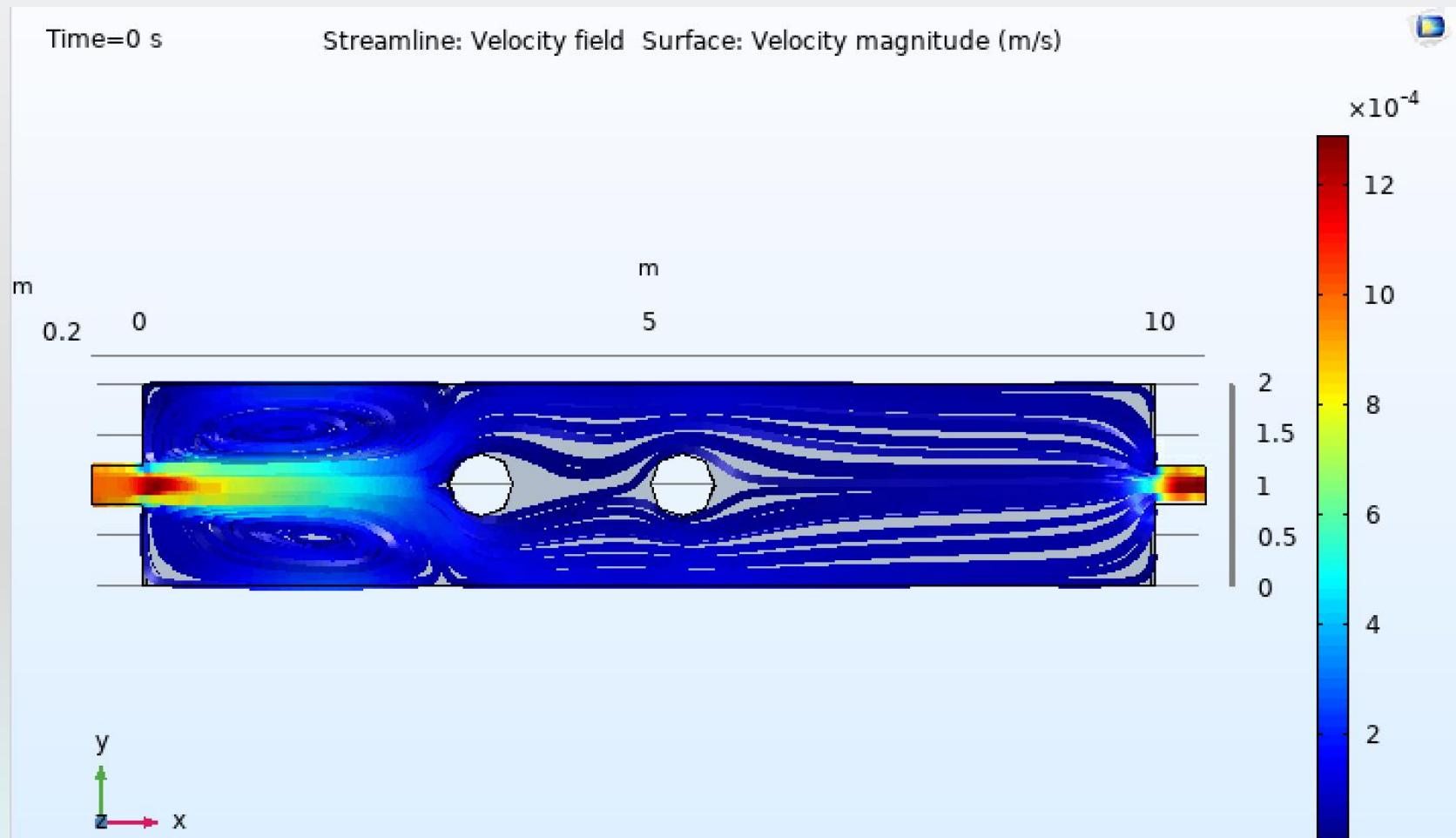
Characteristic Length: Long

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 3: Controlled Group 3 - Large Input

Material: Water

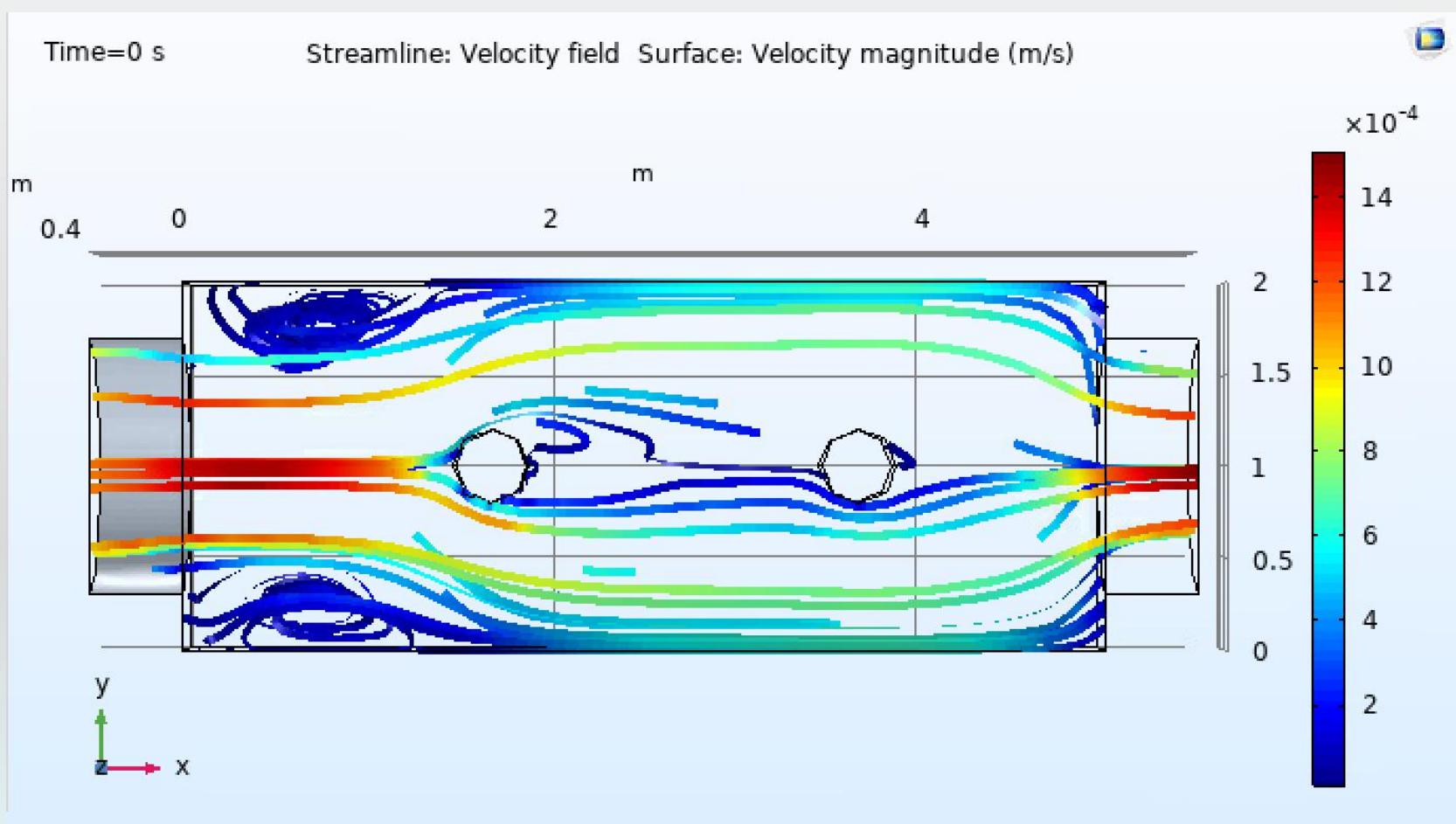
Temperature: 25°C

Density: 997.05 kg/m³

Characteristic Length: Large input

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} pa·s



3. Parameter Experiments

EXP4: Average Speed

Experiment4: Experimental Group

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

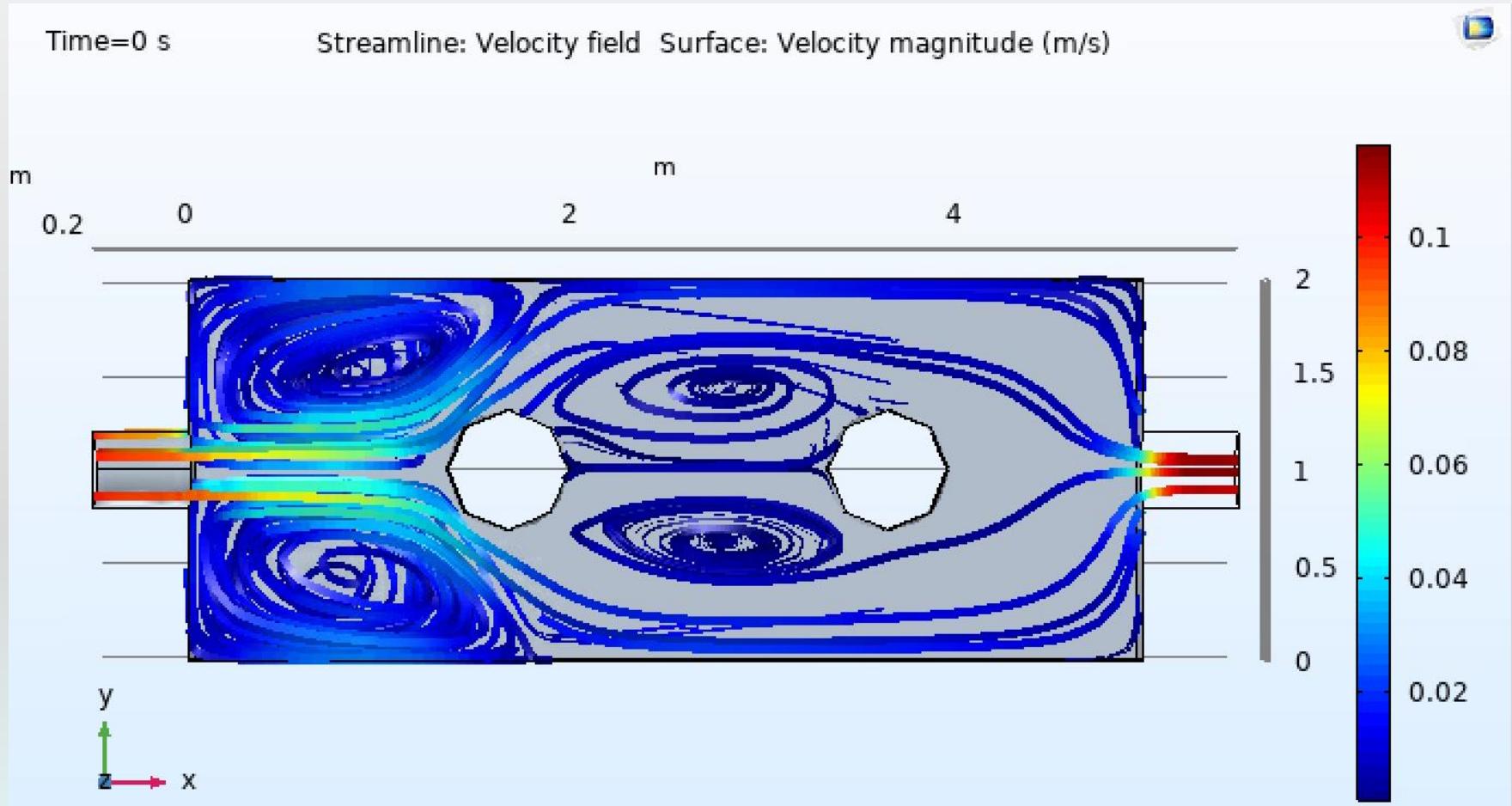
Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 4: Controlled Group 1 - Low Speed

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

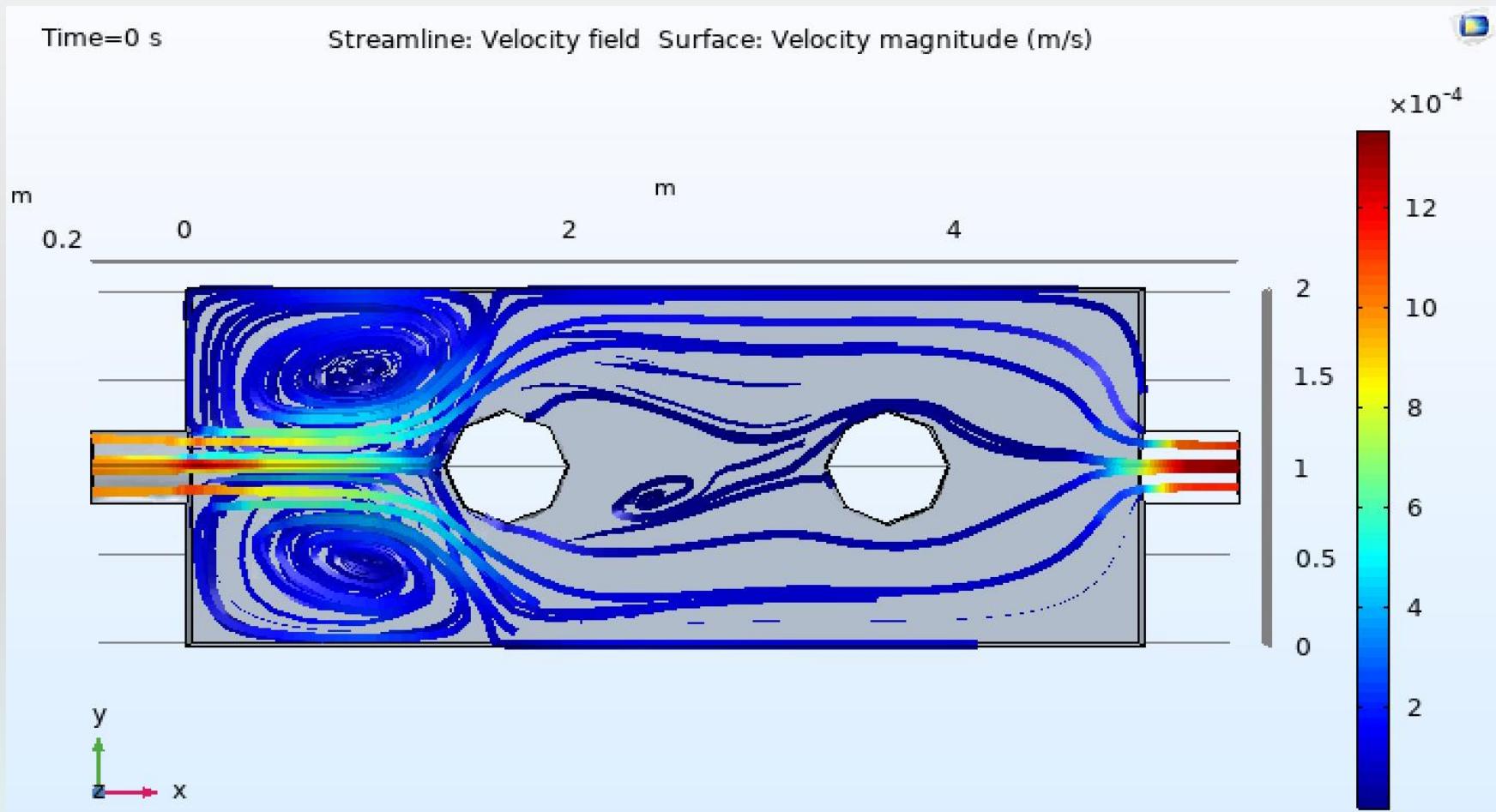
Characteristic Length: Normal

Mean Velocity: 0.001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 4: Controlled Group 2 - Lower Speed

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

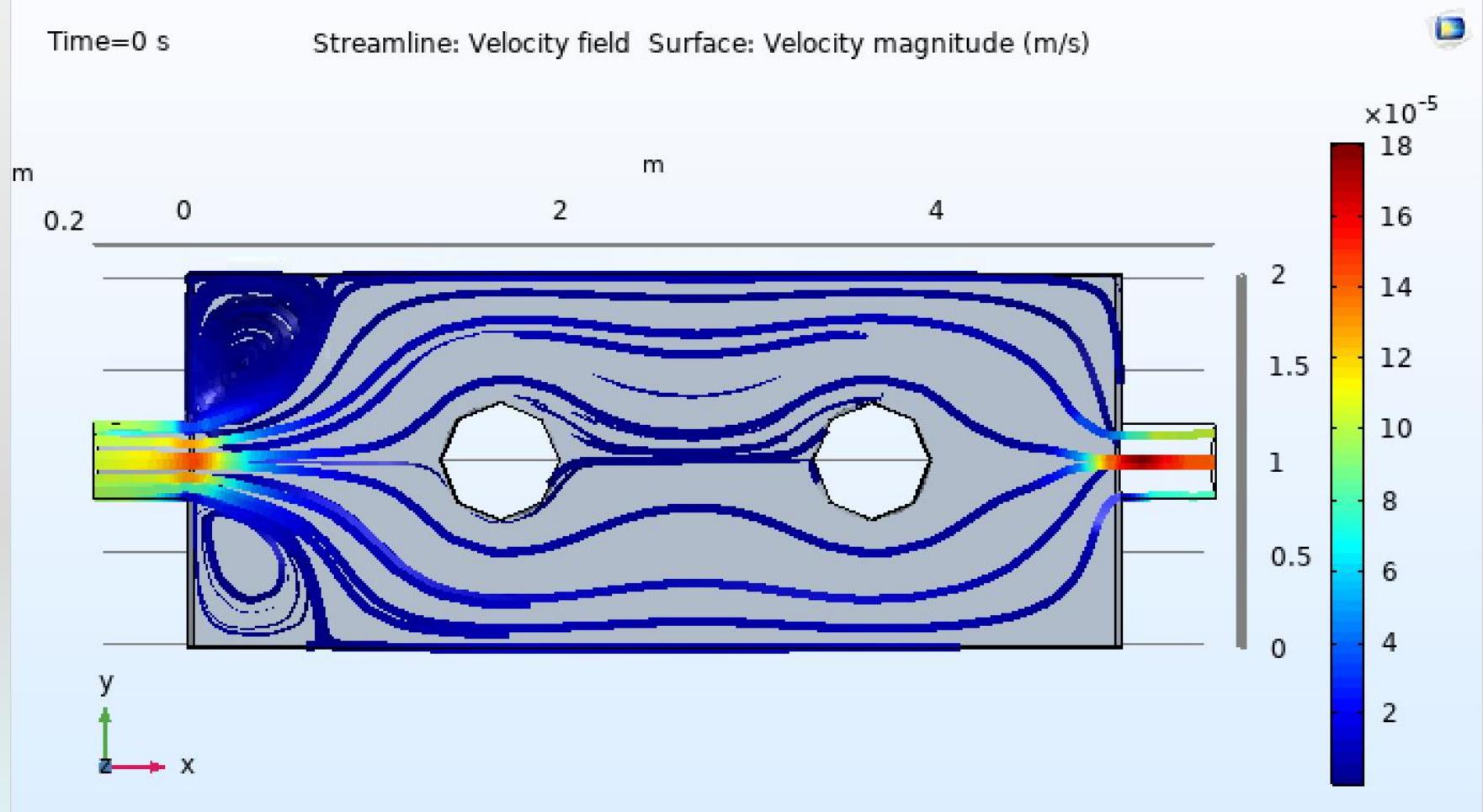
Characteristic Length: Normal

Mean Velocity: 0.0001 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



3. Parameter Experiments

EXP5: Dynamic Viscosity

Experiment 5: Experimental Group

Material: Water

Temperature: 25°C

Density: 997.05 kg/m³

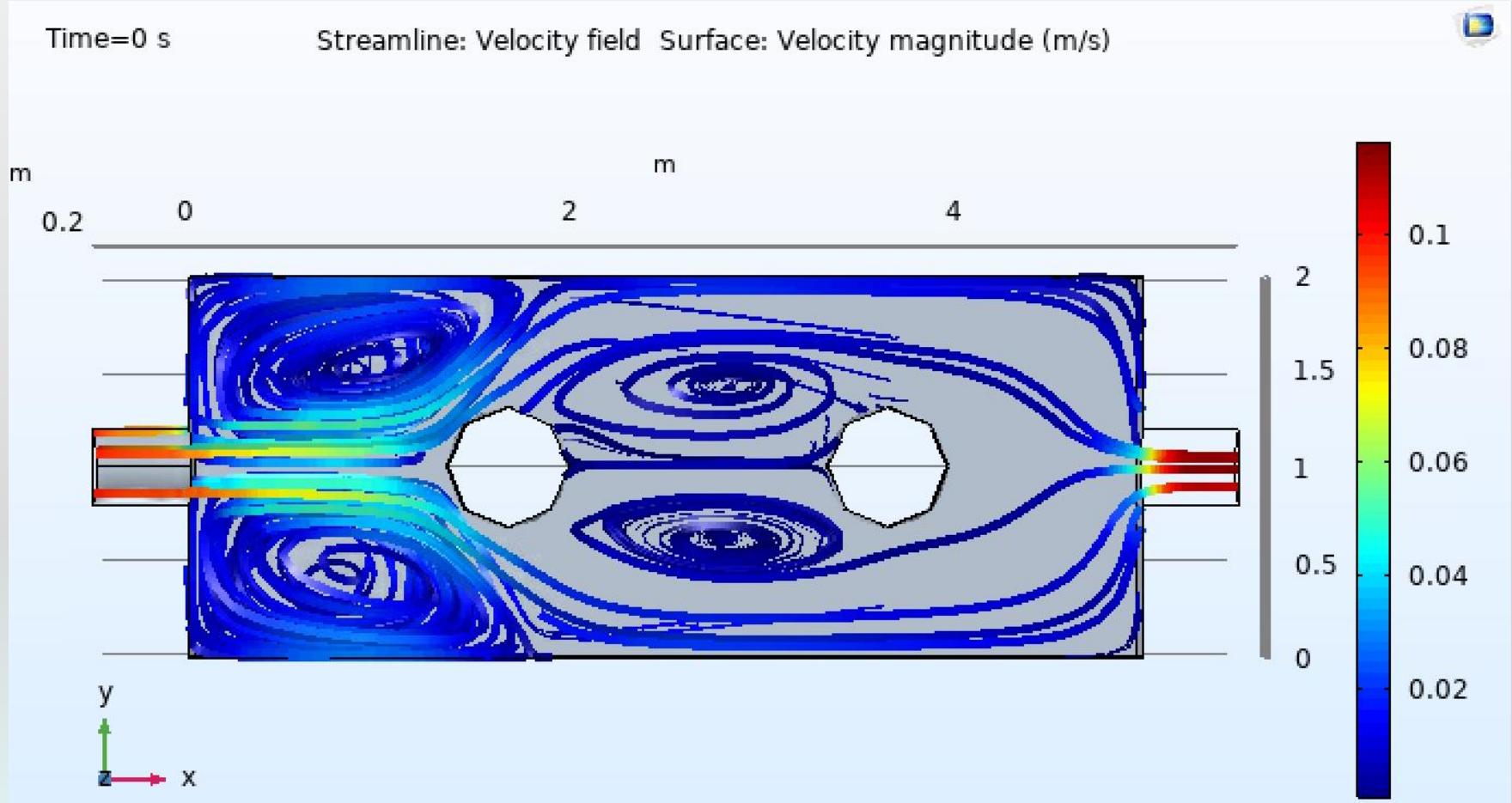
Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 8.90×10^{-4} Pa·s

Time=0 s

Streamline: Velocity field Surface: Velocity magnitude (m/s)



Experiment 5: Controlled Group 2 - Lower Speed

Material: Honey

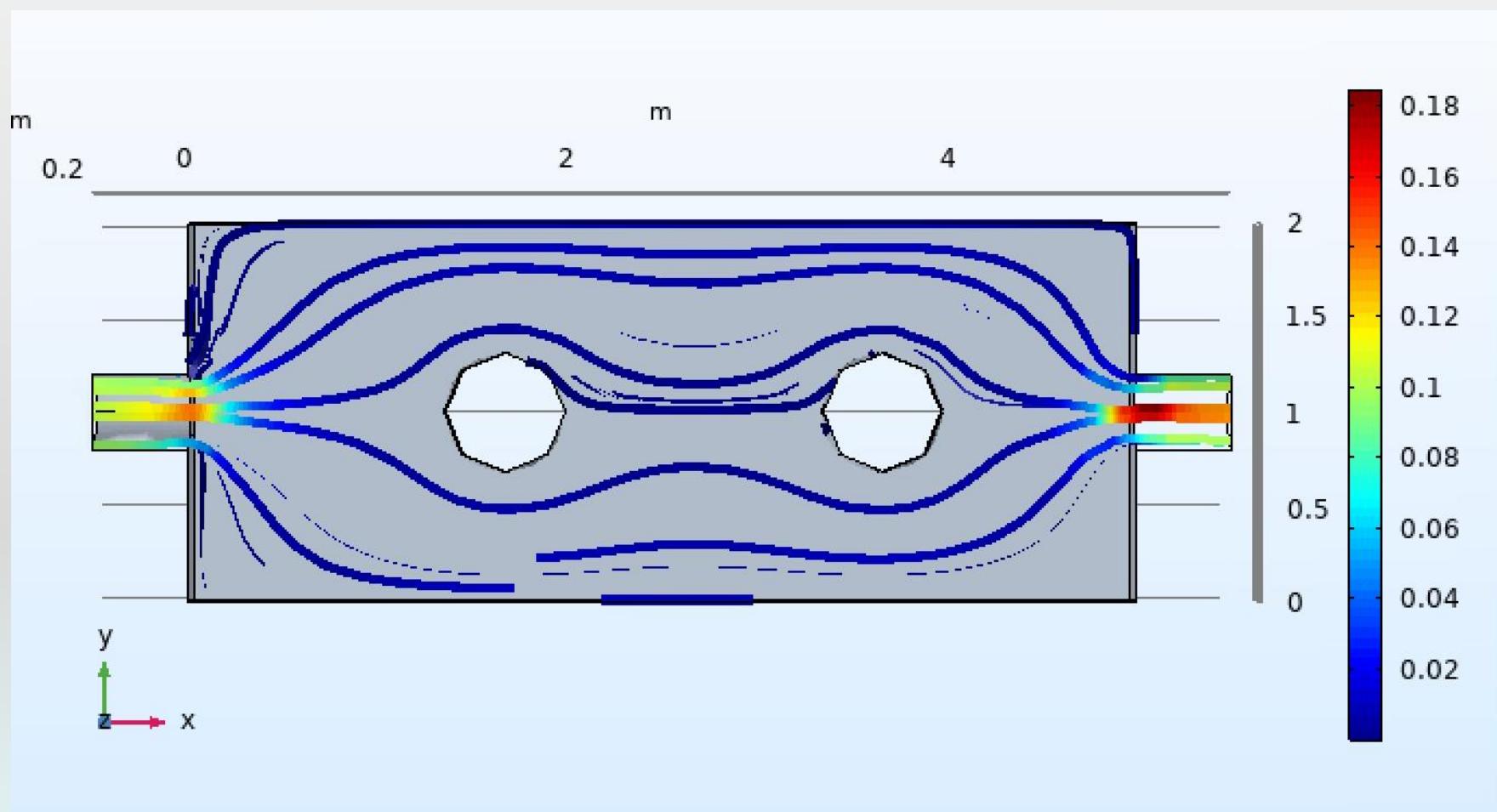
Temperature: 25°C

Density: 1450 (1400-1500) kg/m³

Characteristic Length: Normal

Mean Velocity: 0.1 m/s

Dynamic Viscosity $\mu = \nu \rho$: 7.5 (2-10) pa·s



Experiment 5: Controlled Group 2 - Lower Speed

Material: Honey

Temperature: 25°C

Density: 1450 (1400-1500) kg/m³

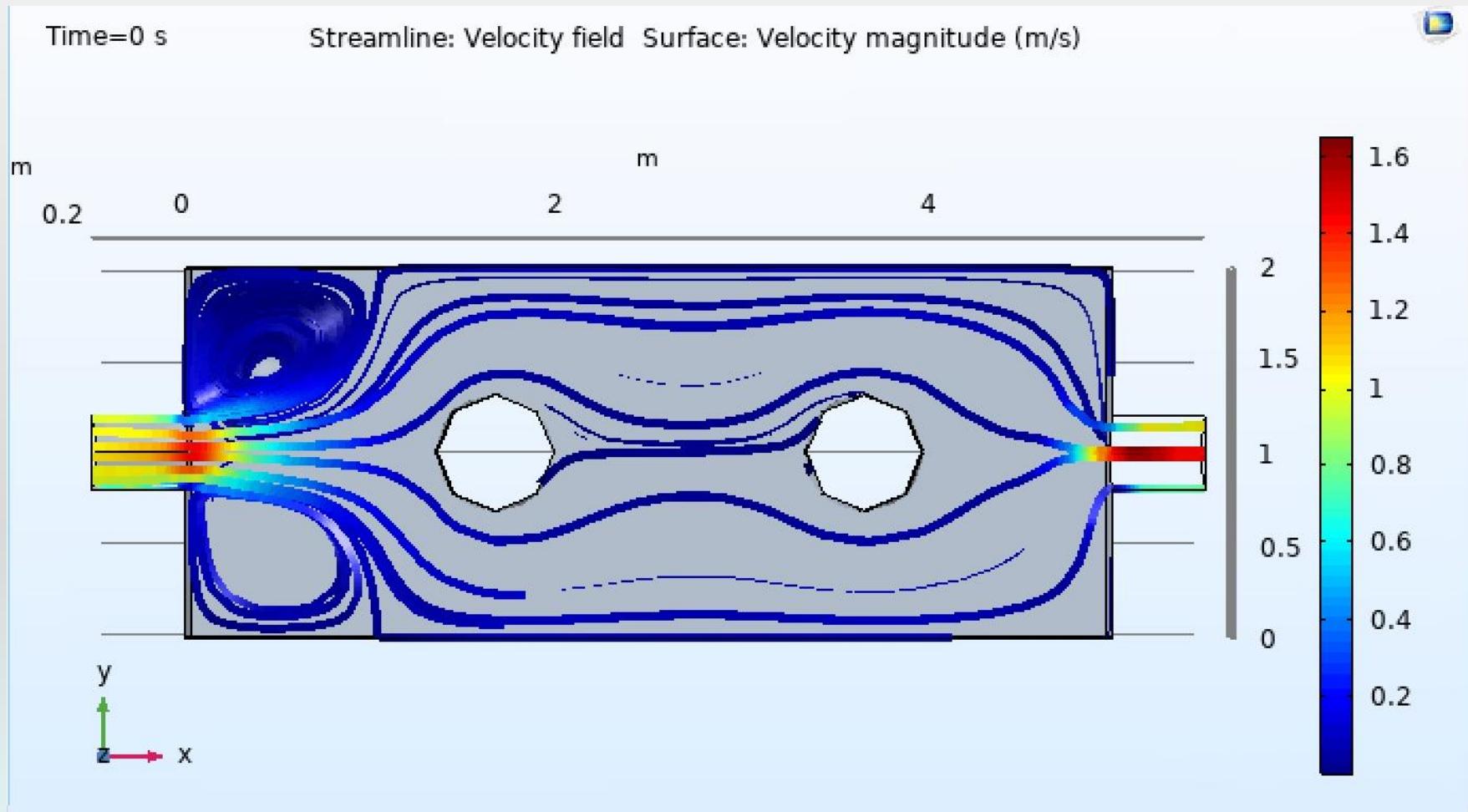
Characteristic Length: Normal

Mean Velocity: 1 m/s

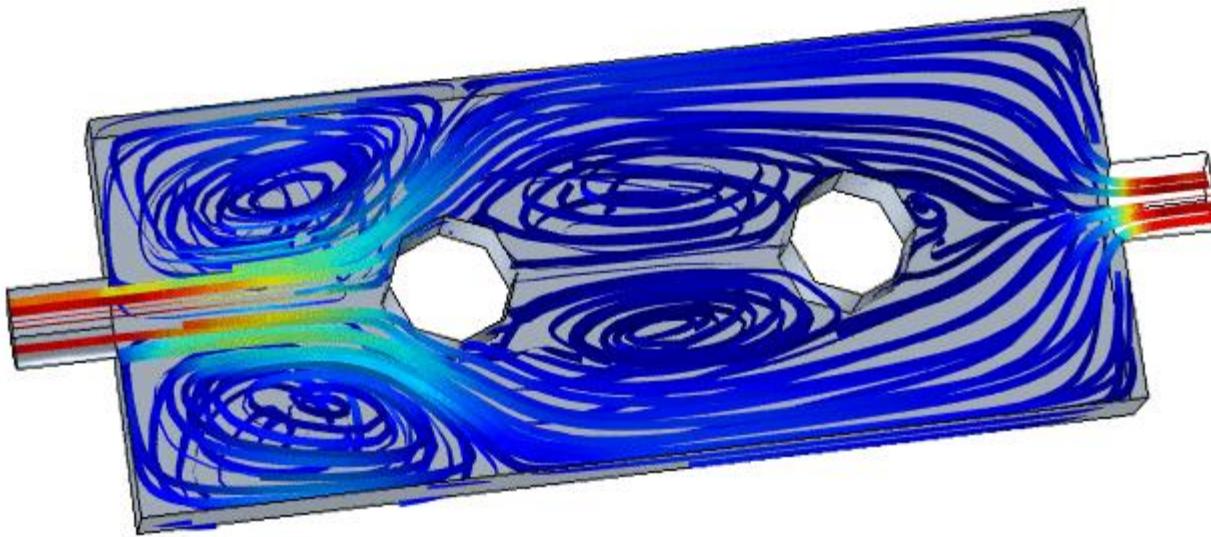
Dynamic Viscosity $\mu = \nu \rho$: 7.5 (2-10) pa·s

Time=0 s

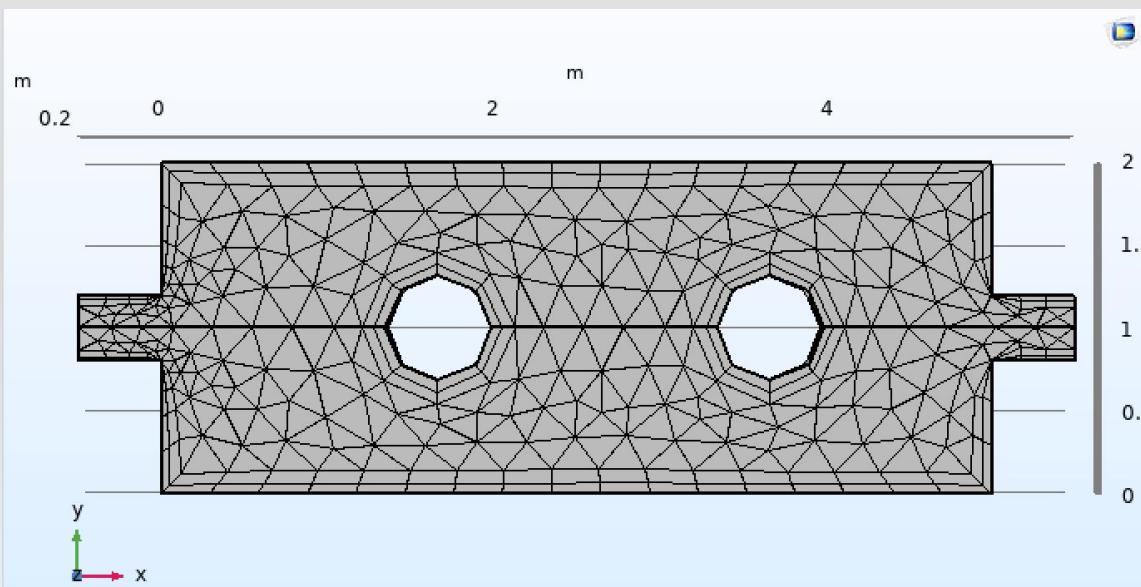
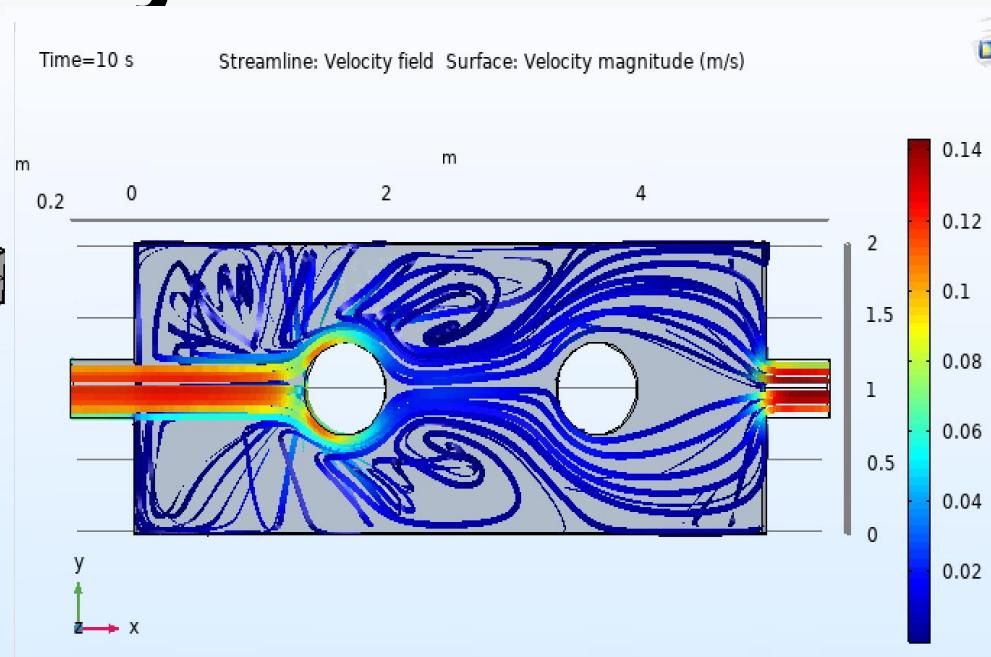
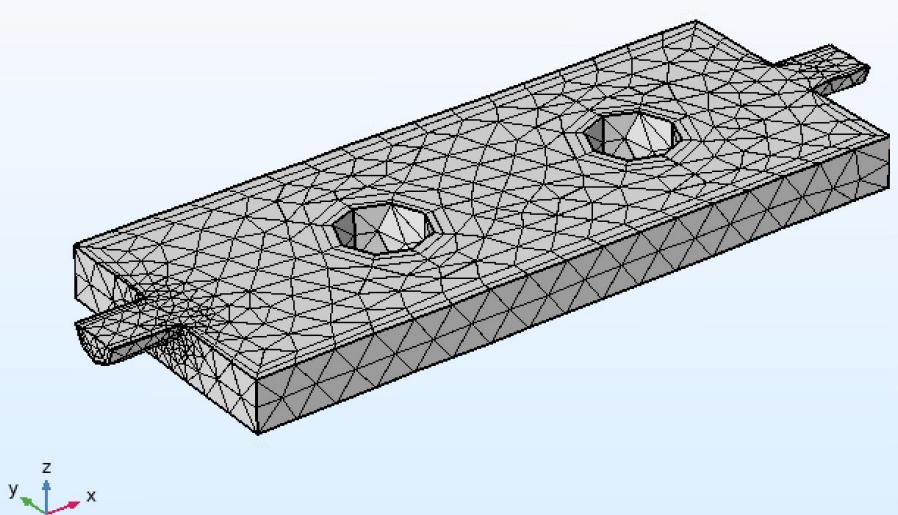
Streamline: Velocity field Surface: Velocity magnitude (m/s)



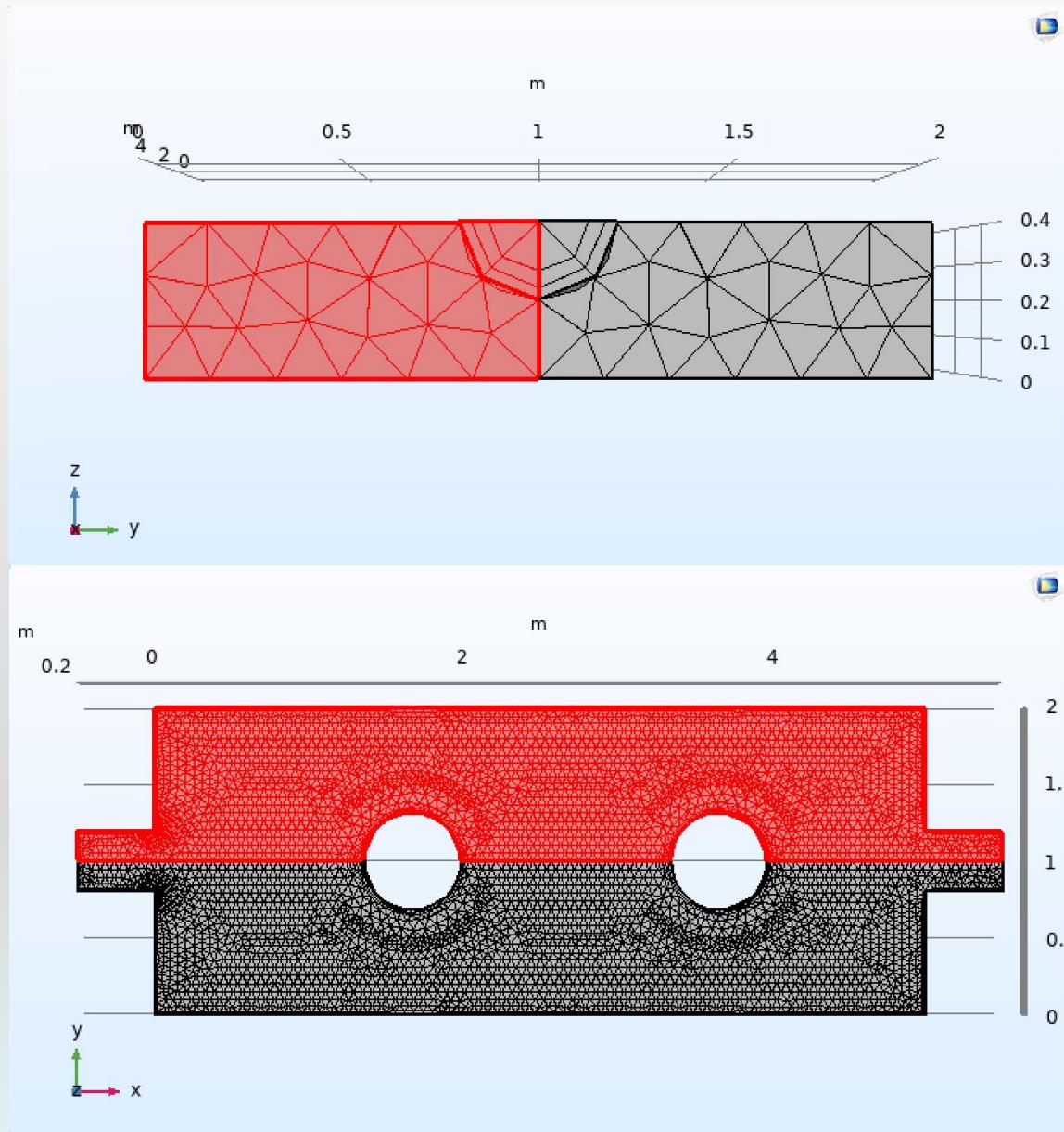
3. Results



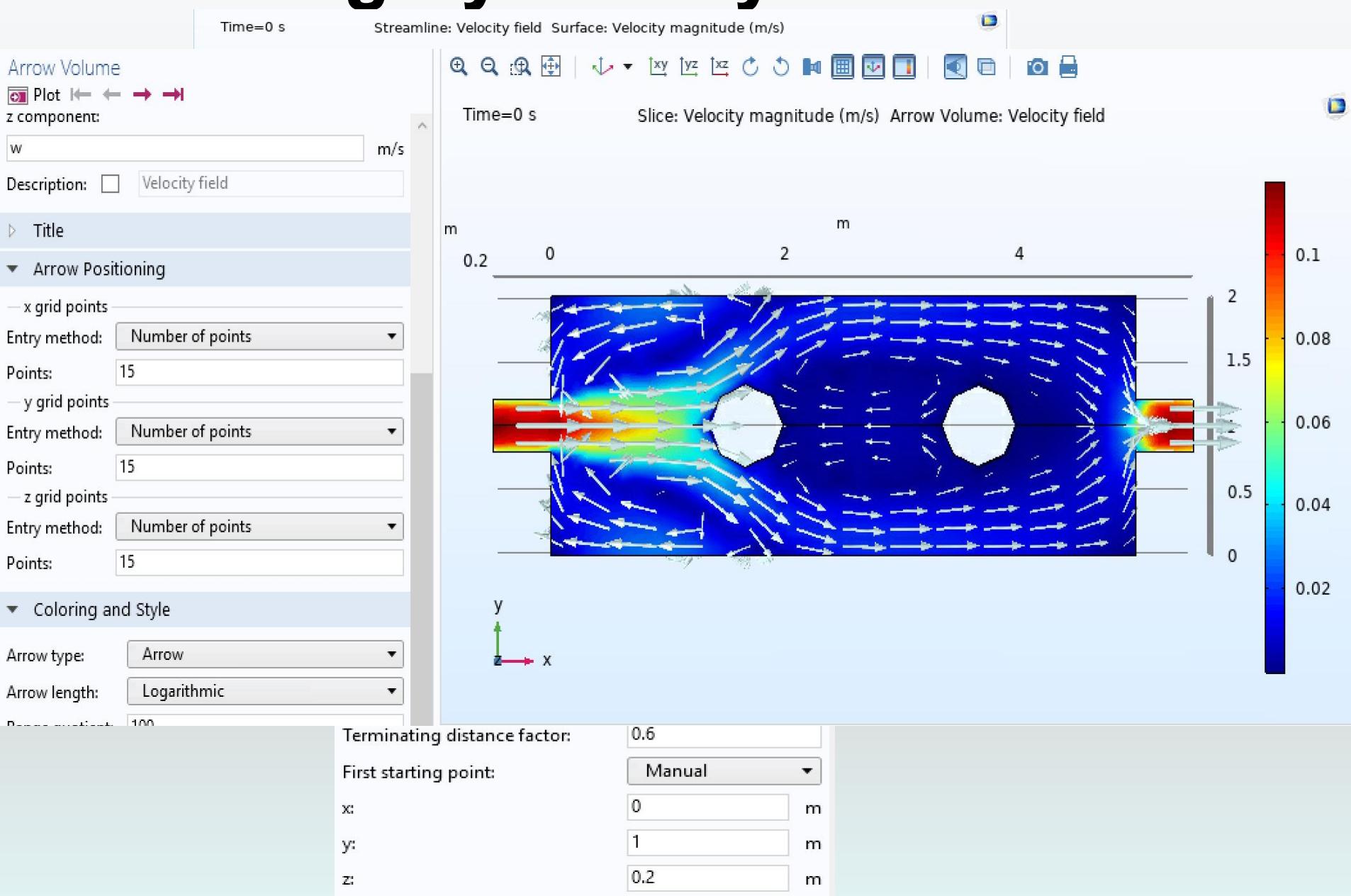
4. Plotting Symmetry



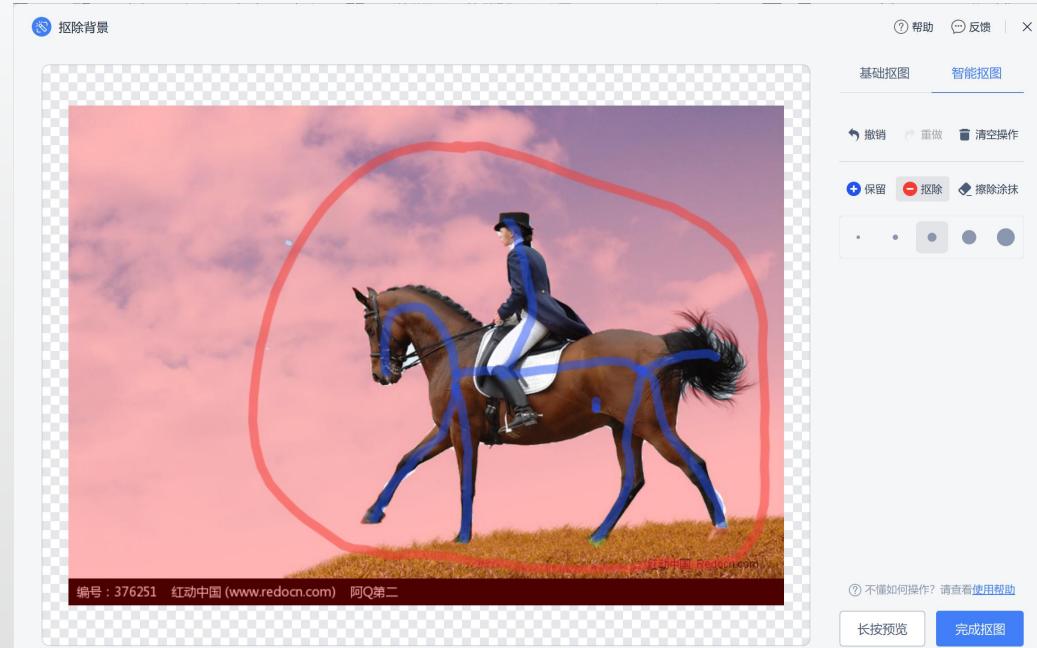
4. Plotting Symmetry



4. Plotting Symmetry

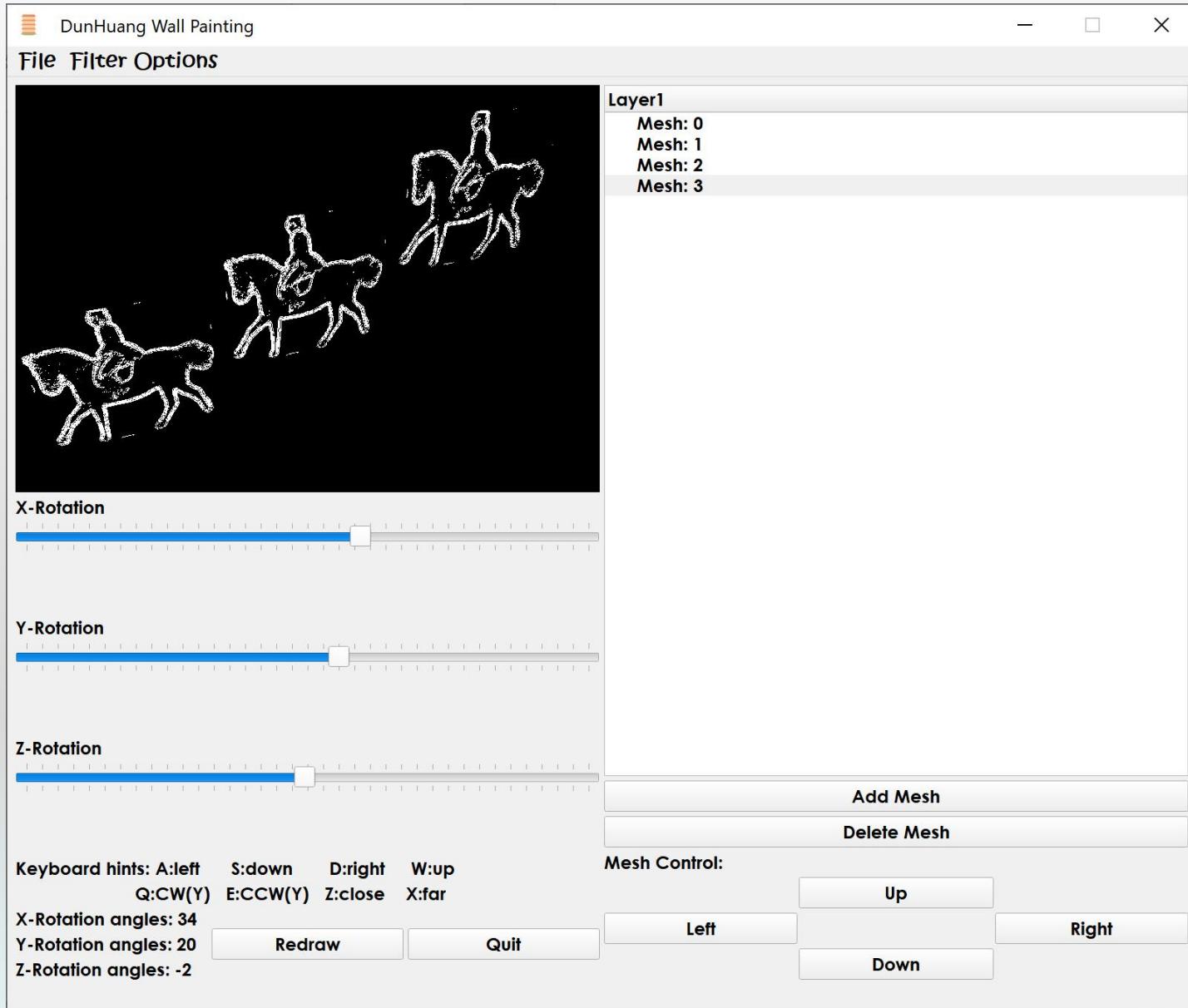


5. Future Talking



Lazy Snapping

5. Future Talking

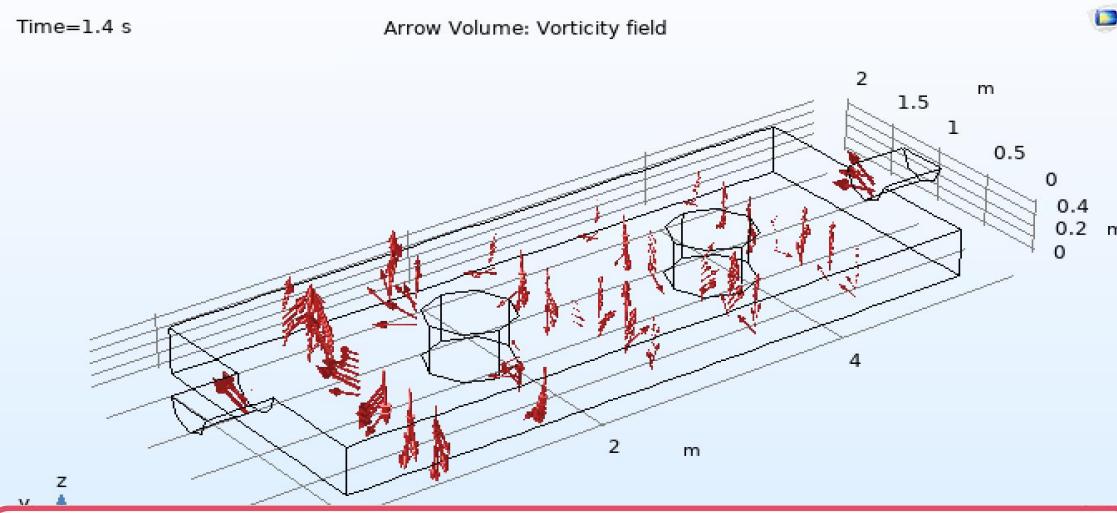


5. Future Talking



一四六 九色鹿本生之四 第二五七窟 西壁

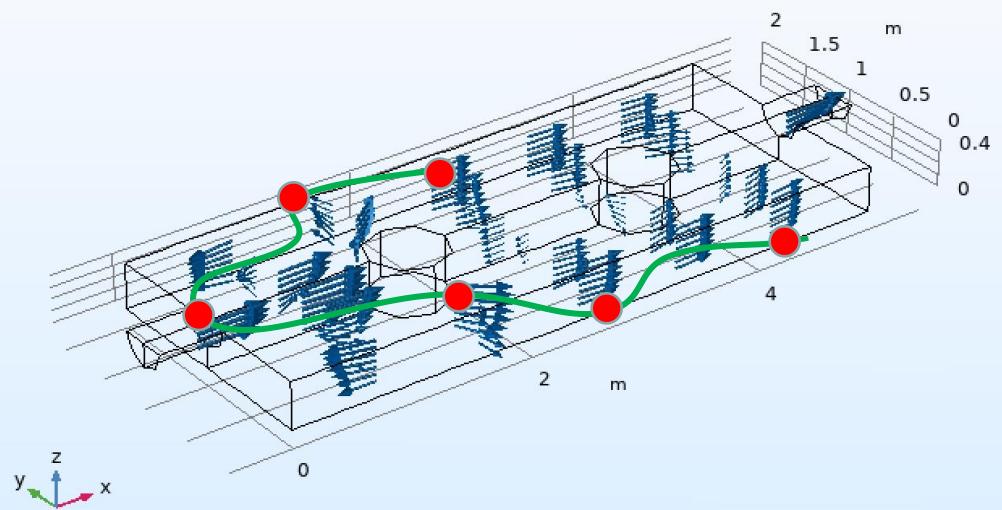
5. Future Talking



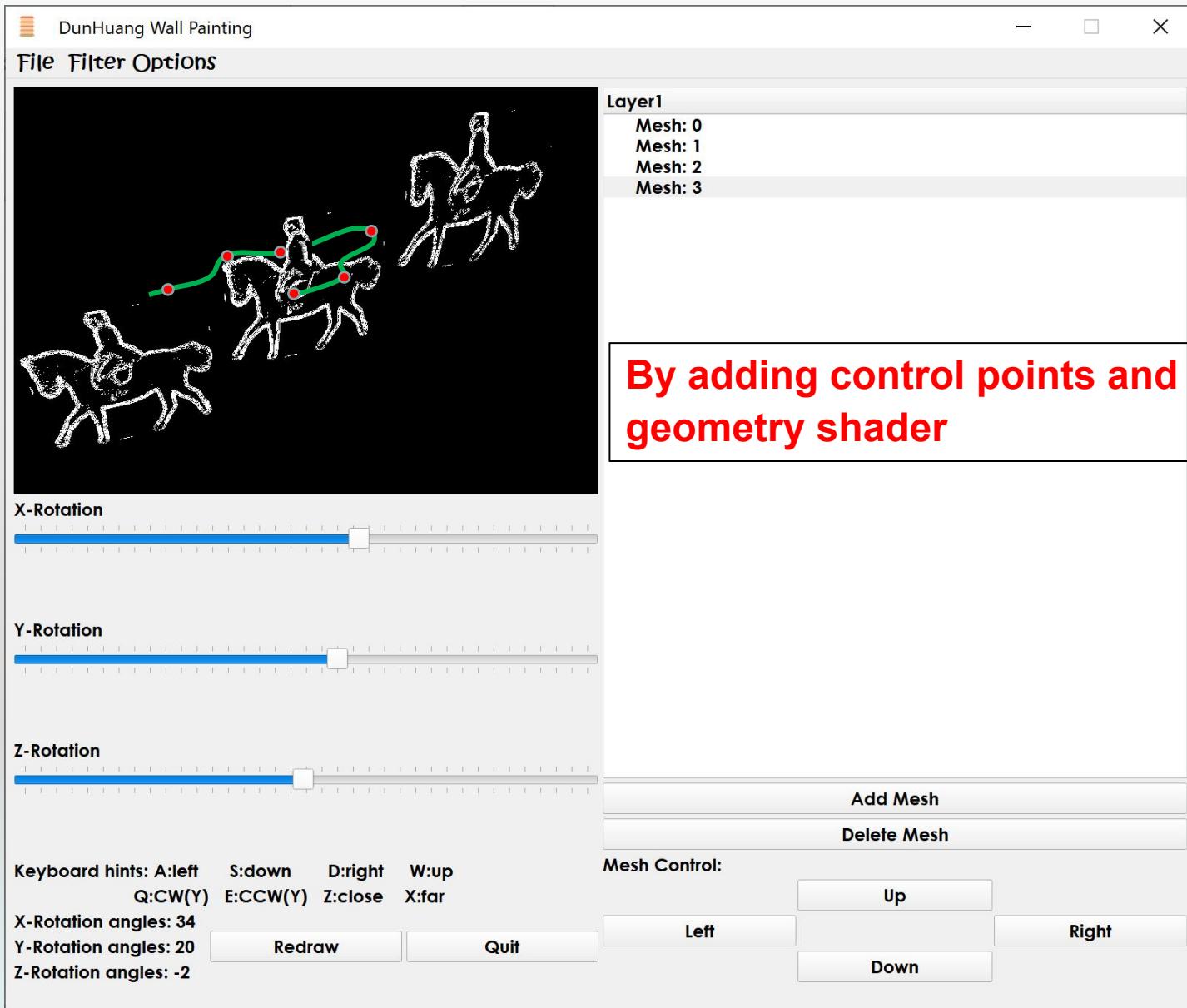
Reminder: the Catmull-Rom Sub-curve Equation is:

$$P(t) = 0.5 * [2 * P_1 + t * (-P_0 + P_2) + t^2(2 * P_0 - 5 * P_1 + 4P_2 - P_3) + t^3(-P_0 + 3P_1 - 3P_2 + P_3)]$$

$0. \leq t \leq 1.$



5. Future Talking



Thank you!