## Octupole-order Lidov-Kozai Population Statistics

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## I. 10/22/20—INITIAL THOUGHTS

## A. Equations

The equations of motion we want to study come from LML15. Describe the inner binary by  $(a, e, I, \Omega, \omega)$  and the outer binary with "out" subscripts, and denote  $I_{\text{tot}} = I + I_{\text{out}}$ . Call the inner binary component masses  $m_1, m_2$ , and the tertiary mass  $m_3$ , define the inner binary total and reduced masses  $m_{12} = m_1 + m_2$  and  $\mu = m_1 m_2 / m_{12}$ , and define the tertiary orbit total and reduced masses  $m_{123} = m_{12} + m_3$  and  $\mu_{\text{out}} = m_{12} m_3 / m_{123}$ . The equations of motion are  $(j(e) = \sqrt{1 - e^2})$ 

$$\frac{\mathrm{d}a}{\mathrm{d}t} = -\frac{64}{5} \frac{a}{t_{\rm GW} j^7(e)} \left( 1 + \frac{73}{24} e^2 + \frac{37}{96} e^4 \right),\tag{1}$$

$$\frac{de}{dt} = \frac{j(e)}{64t_{LK}} \left\{ 120e \sin^2 I_{tot} \sin(2\omega) + \frac{15\epsilon_{oct}}{8} \cos \omega_{out} \left[ \left( 4 + 3e^2 \right) \left( 3 + 5\cos(2I_{tot}) \right) \right] \times \sin \omega + 210e^2 \sin^2 I_{tot} \sin 3\omega \\
- \frac{15\epsilon_{oct}}{4} \cos I_{tot} \cos \omega \left[ 15(2 + 5e^2) \cos(2I_{tot}) + 7 \left( 30e^2 \cos(2\omega) \sin^2 I_{tot} - 2 - 9e^2 \right) \right] \sin \omega_{out} \\
- \frac{304}{15} \frac{e}{t_{cov} i^5(e)} \left( 1 + \frac{121}{304} e^2 \right), \tag{2}$$

$$\frac{\mathrm{d}I}{\mathrm{d}t} = -\frac{3e}{32t_{\mathrm{LK}}j(e)} \left\{ 10\sin\left(2I_{\mathrm{tot}}\right) \left[ e\sin(2\omega) + \frac{5\epsilon_{\mathrm{oct}}}{8} \left(2 + 5e^2 + 7e^2\cos(2\omega)\right) \cos\omega_{\mathrm{out}}\sin\omega \right] + \frac{5\epsilon_{\mathrm{oct}}}{8} \cos\omega \left[ 26 + 37e^2 - 35e^2\cos(2\omega) - 15\cos\left(2I_{\mathrm{tot}}\right) \left( 7e^2\cos(2\omega) - 2 - 5e^2 \right) \right] \times \sin I_{\mathrm{tot}}\sin\omega_{\mathrm{out}} \right\}$$
(3)

$$\frac{\mathrm{d}\Omega}{\mathrm{d}t} = \frac{\mathrm{d}\Omega_{\mathrm{out}}}{\mathrm{d}t} = -\frac{3\csc I}{32t_{\mathrm{LK}}j(e)} \left\{ 2\left[ (2+3e^2 - 5e^2\cos(2\omega)) + \frac{25\epsilon_{\mathrm{oct}}e}{8}\cos\omega \left( 2 + 5e^2 - 7e^2\cos(2\omega) \right) \cos\omega_{\mathrm{out}} \right] \right. \\
\left. \times \sin(2I_{\mathrm{tot}}) - \frac{5\epsilon_{\mathrm{oct}}e}{8} \left[ 35e^2(1+3\cos(2I_{\mathrm{tot}})) \cos 2\omega - 46 - 17e^2 - 15\left( 6 + e_1^2 \right) \cos(2I_{\mathrm{tot}}) \right] \right. \\
\left. \times \sin I_{\mathrm{tot}} \sin\omega \sin\omega_{\mathrm{out}} \right\}, \tag{4}$$

$$\frac{d\omega}{dt} = \frac{3}{8t_{LK}} \left\{ \frac{1}{j(e)} \left[ 4\cos^2 I_{tot} + (5\cos(2\omega) - 1) \right] \right.$$

$$\times \left( 1 - e^2 - \cos^2 I_{tot} \right) \left] + \frac{L\cos I_{tot}}{L_{out} j(e_{out})} \left[ 2 + e^2 (3 - 5\cos(2\omega)) \right] \right\} + \frac{15\epsilon_{oct}}{64t_{LK}} \left\{ \left( \frac{L}{L_{out} j(e_{out})} + \frac{\cos I_{tot}}{j(e)} \right) \right.$$

$$\times e \left[ \sin \omega \sin \omega_{out} \left[ 10(3\cos^2 I_{tot} - 1)(1 - e^2) + A \right] \right.$$

$$- 5B\cos I_{tot} \cos \Theta \right] - \frac{j(e)}{e} \left[ 10\sin \omega \sin \omega_{out} \cos I_{tot} \right.$$

$$\times \sin^2 I_{tot} \left( 1 - 3e^2 \right) + \cos \Theta \left( 3A - 10\cos^2 I_{tot} + 2 \right) \right] \right\}$$

$$+ \Omega_{GR}, \tag{5}$$

$$\frac{\mathrm{d}e_{\mathrm{out}}}{\mathrm{d}t} = \frac{15eL\,j(e_{\mathrm{out}})\epsilon_{\mathrm{oct}}}{256t_{\mathrm{LK}}e_{\mathrm{out}}L_{\mathrm{out}}} \Big\{ \cos\omega \Big[6 - 13e^2 + 5(2 + 5e^2)\cos(2I_{\mathrm{tot}}) + 70e^2\cos(2\omega)\sin^2I_{\mathrm{tot}} \Big] \\
\times \sin\omega_{\mathrm{out}} - \cos I_{\mathrm{tot}}\cos\omega_{\mathrm{out}} \Big[5(6 + e^2)\cos(2I_{\mathrm{tot}}) + 7\left(10e^2\cos(2\omega)\sin^2I_{\mathrm{tot}} - 2 + e^2\right)\Big]\sin\omega \Big\}, \tag{6}$$

$$\frac{\mathrm{d}I_{\mathrm{out}}}{\mathrm{d}t} = -\frac{3eL}{32t_{\mathrm{LK}}j(e_{\mathrm{out}})L_{\mathrm{out}}} \left\{ 10 \left[ 2e\sin I_{\mathrm{tot}}\sin(2\omega) + \frac{5\epsilon_{\mathrm{oct}}}{8}\cos\omega\left(2 + 5e^2 - 7e^2\cos(2\omega)\right)\sin(2I_{\mathrm{tot}})\sin\omega_{\mathrm{out}} \right] + \frac{5\epsilon_{\mathrm{oct}}}{8} \left[ 26 + 107e^2 + 5(6 + e^2)\cos(2I_{\mathrm{tot}}) - 35e^2\left(\cos2(I_{\mathrm{tot}}) - 5\right)\cos(2\omega) \right]\cos\omega_{\mathrm{out}}\sin I_{\mathrm{tot}}\sin\omega \right\}, \tag{7}$$

$$\frac{d\omega_{\text{out}}}{dt} = \frac{3}{16t_{\text{LK}}} \left\{ \frac{2\cos I_{\text{tot}}}{j(e)} \left[ 2 + e^2 (3 - 5\cos(2\omega)) \right] \right. \\
+ \frac{L}{L_{\text{out}} j(e_{\text{out}})} \left[ 4 + 6e^2 + (5\cos^2 I_{\text{tot}} - 3) \right. \\
\times \left[ 2 + e^2 (3 - 5\cos(2\omega)) \right] \right] \right\} - \frac{15\epsilon_{\text{oct}} e}{64t_{\text{LK}} e_{\text{out}}} \\
\times \left\{ \sin \omega \sin \omega_{\text{out}} \left[ \frac{L(4e_{\text{out}}^2 + 1)}{e_{\text{out}} L_{\text{out}} j(e_{\text{out}})} 10\cos I_{\text{tot}} \sin^2 I_{\text{tot}} \right. \\
\times \left. (1 - e^2) - e_{\text{out}} \left( \frac{1}{j(e)} + \frac{L\cos I_{\text{tot}}}{L_{\text{out}} j(e_{\text{out}})} \right) \right. \\
\times \left[ A + 10 \left( 3\cos^2 I_{\text{tot}} - 1 \right) \left( 1 - e^2 \right) \right] \right] + \cos \Theta \\
\times \left[ 5B\cos I_{\text{tot}} e_{\text{out}} \left( \frac{1}{j(e)} + \frac{L\cos I_{\text{tot}}}{L_{\text{out}} j(e_{\text{out}})} \right) \right. \\
+ \frac{L(4e_{\text{out}}^2 + 1)}{e_{\text{out}} L_{\text{out}} j(e_{\text{out}})} A \right] \right\}. \tag{8}$$

where  $n=\sqrt{Gm_{12}/a^3}$  is the mean motion,  $L=\mu\sqrt{Gm_{12}a}$  and  $L_{\rm out}=\mu_{\rm out}\sqrt{Gm_{123}a_{\rm out}}$  are the circular

angular momenta, and

$$t_{\rm LK}^{-1} = n \left(\frac{m_3}{m_{12}}\right) \left(\frac{a}{a_{\rm out} j(e_{\rm out})}\right)^3,$$
 (9)

$$t_{\rm GW}^{-1} = \frac{G^3 \mu m_{12}^2}{c^5 a^4},\tag{10}$$

$$\Omega_{\rm GR} = \frac{3Gnm_{12}}{c^2aj^2(e)},\tag{11}$$

$$\epsilon_{\text{oct}} = \frac{m_2 - m_1}{m_{12}} \frac{a}{a_{\text{out}}} \frac{e_{\text{out}}}{1 - e_{\text{out}}^2},$$
(12)

$$A \equiv 4 + 3e^2 - \frac{5}{2}B\sin^2 I_{\text{tot}},\tag{13}$$

$$B \equiv 2 + 5e^2 - 7e^2 \cos(2\omega), \tag{14}$$

$$\cos\Theta \equiv -\cos\omega\cos\omega_{\text{out}} - \cos I_{\text{tot}}\sin\omega\sin\omega_{\text{out}}.$$
 (15)

These equations can be nondimensionalized via the following steps (I won't rewrite the equations): (i) multiply through by  $t_{\rm LK,0}$  ( $a=a_0$  and  $e_{\rm out}=0$ ), and call  $\tau\equiv t/t_{\rm LK,0}$  the new variable of differentiation, (ii) reexpress all of the timescales as

$$\frac{t_{\rm LK,0}}{t_{\rm LK}} = \left(\frac{a}{a_0}\right)^{3/2} j^{-3} \left(e_{\rm out}\right),\tag{16}$$

$$\frac{t_{\rm LK,0}}{t_{\rm GW}} = \frac{G^3 \mu m_{12}^3}{m_3 c^5 a^4} \frac{1}{n_0} \left(\frac{a_{\rm out}}{a_0}\right)^3,$$

$$= \epsilon_{\rm GW} \left(\frac{a_0}{a}\right)^4,\tag{17}$$

$$\epsilon_{\rm GW} \equiv \frac{G^3 \mu m_{12}^3 a_{\rm out}^3}{m_3 c^5 a_0^7 n_0},$$
(18)

$$\Omega_{\rm GR} t_{\rm LK,0} = \frac{3Gnm_{12}^2}{m_3 c^2 a} \frac{1}{n_0} \left(\frac{a_{\rm out}}{a_0}\right)^3,$$

$$= \epsilon_{\rm GR} \left(\frac{a_0}{a}\right)^{5/2},\tag{19}$$

$$\epsilon_{\rm GR} = \frac{3Gm_{12}^2 a_{\rm out}^3}{m_3 c^2 a_0^4}.$$
 (20)

(iii) re-express da/dt as

$$\frac{\mathrm{d}(a/a_0)}{\mathrm{d}\tau} = -\frac{64}{5} \frac{\epsilon_{\mathrm{GW}}}{j^{7/2}(e)} \left(\frac{a_0}{a}\right)^3 \left(1 + \frac{73}{24}e^2 + \frac{37}{96}e^4\right). \tag{21}$$

As such, the natural unit of length is  $a_0=1$ , the natural unit of time is  $t_{\rm LK,0}=1$ , and everything else is dimensionless. When computing these  $\epsilon$ , I use convention where  $1M_{\odot}=1~{\rm AU}=c=1$ , under which  $G=9.87\times 10^{-9}$ .

## B. Points of Inquiry

The goal is to understand how the merger window varies as a function of  $q \equiv m_1/m_2$  ( $m_1 < m_2$ ) when the octupole order LK effects are important.

• First, let's set  $\epsilon_{\rm GW}=0$ . It is well known that the octupole order LK is nonintegrable. What does the Fourier spectrum of the eccentricity look like? Will this help us get a delay time distribution between high-e phases?

When the octupole effect is unimportant, the spectrum falls off exponentially over scales  $\tau \simeq P_{\rm LK} j^{-1}(e_{\rm max})$ , where  $P_{\rm LK}$  is the quadrupole LK period. One imagines the tail of the spectrum gets heavier when  $\epsilon_{\rm oct}$  is increased, and this might help us get the delay time distribution.

A second way we can postprocess this is to take a histogram of e(t). If there is some regular structure, it's likely this will allow us to compute the average rate of binary coalescence due to GW radiation.

• The goal is to understand the size of the merger window,  $\Delta I$ , as a function of q. To do this, we numerically sample the merger time function  $T_{\rm m}\left(I_0,q\right)$ . At each  $I_0$ , the natural thing to do would be to try for  $\sim 5-10$  random  $\Omega,\omega$ , and define the merger window to be where  $T_{\rm m} \leq 10^{10}$  yr.