

GR report

①

我們要求 spherical symmetric solution.

所以我們知道我們的 metric 的形式如下

$$dt^2 = -C(r,t) dt^2 + D(r,t) dr^2 + E(r,t) dr dt + F(r,t) r^2 d\theta^2$$

$\left(\because \text{若有 } dr d\theta \text{ or } dt d\theta \text{ 的 term we can make } \theta' = -\theta \text{ 的變數} \right)$
 代換得到 $dr d\theta$ 、 $dt d\theta$ 二 term 只能 vanishing

首先我們作 $r' = \sqrt{F} r$ 的變換

$$dt^2 = -C'(r',t) dt^2 + D'(r',t) dr'^2 + E'(r',t) dr' dt + r'^2 d\theta^2$$

為了消掉 E' , $dr' dt$ by term, we need to define t'

$$dt' = \eta(r',t)[C(r',t) dt - E'(r',t) dr'].$$

where $\eta(r',t)$ 是積分因子 讓我們真的能找到 t' or 這樣
 $\eta(r',t)[C(r',t) dt - E'(r',t) dr']$ 是 exact form.

$$\left(\frac{\partial}{\partial r'} [\eta(r',t) C(r',t)] = -\frac{\partial}{\partial t} [\eta(r',t) E'(r',t)] \right)$$

$$\text{Simplifying } dt^2 = -1^2 C^{-1} dt'^2 + (D + C^{-1} E^2) dr'^2 - r'^2 d\theta^2$$

整理後我們知道 在經過一定的變數代換後

$$dt^2 = -B dt_{\text{new}}^2 + A dr_{\text{new}}^2 - r_{\text{new}}^2 d\theta^2$$

(注意 A, B depend on t and r)

而上述是最廣的結果，我們可以考慮 stationary by case
則是我們的 metric 的係數只 depend on r 的 case

$$ds^2 = -F(r)dt^2 + 2rE(r)drdt + r^2B(r)dr^2 + C(r)(dr^2 + r^2d\theta^2)$$

define $t' = t + \Phi(r)$ 其中 $\Phi(r)$ satisfy $\frac{d\Phi}{dr} = -\frac{rE(r)}{F(r)}$

則可以消到 $dr dt$ term.

$$\begin{aligned} \Rightarrow ds^2 &= -F(r)dt'^2 + 2F(r)\frac{d\Phi}{dr}drdt + \left(\frac{d\Phi}{dr}\right)^2 F(r)dr^2 + 2rE(r)drdt + r^2B(r)dr^2 \\ &\quad + C(r)(dr^2 + r^2d\theta^2) \\ &= -F(r)dt'^2 + r^2\left(D + \frac{E^2}{F^2}\right)dr^2 + C(r)(dr^2 + r^2d\theta^2) \end{aligned}$$

We denote $r^2(D + \frac{E^2}{F^2}) = G(r)$

We define $r' = \exp\left\{\left[1 + \frac{G(r)}{C(r)}\right]^{\frac{1}{2}}\right\} \frac{dr}{r}$

$$\Rightarrow dr' = r'\left(\frac{1}{r}\left[1 + \frac{G(r)}{C(r)}\right]^{\frac{1}{2}}\right)dr \Rightarrow dr'^2 = r'^2\left(\frac{1}{r^2}\left(1 + \frac{G}{C}\right)\right)dr^2$$

$$\Rightarrow ds^2 = -F'(r)dt'^2 + \frac{C(r)r^2}{r'^2}(dr'^2 + r'^2d\theta^2)$$

$$= -F'(r)dt'^2 + H(r')(dr'^2 + r'^2d\theta^2)$$

所以我們的 metric 可以寫成以上形式。

接下來我們來 Reproduce 文章的結果!!! (for static case)

$$② \quad g_{ij} = \begin{pmatrix} -N^2(r) & 0 & 0 \\ 0 & \gamma_{ij}(\vec{r}) \\ 0 & 0 \end{pmatrix}$$

$$\sqrt{-g} = N\sqrt{\gamma}$$

γ = determinants γ_{ij}

g = determinants g_{NV} .

$$\text{Let } S := S_{E-H} = \int dx \sqrt{\gamma} NR$$

R is only consider spatial 2 dimensions
so scalar curvature.

$$\delta S = \int dx^2 [\delta N] \sqrt{\gamma} R + N (\delta \sqrt{\gamma}) R + N \sqrt{\gamma} \delta R$$

$$= \int dx^2 \left[\delta \sqrt{-g_{00}} \sqrt{\gamma} + N \frac{\sqrt{\gamma}}{2} \gamma^{ij} (\delta \gamma_{ij}) \right] R + N \sqrt{\gamma} (\delta g^{ij} R_{ij} + N \sqrt{\gamma} g^{ij} (\delta R_{ij}))$$

$$= \int dx^2 \left(\sqrt{\gamma} \left[\frac{N}{2} g^{ij} (\delta g_{ij}) + \left[-\frac{\delta g_{00}}{2N} \right] R + N \sqrt{\gamma} \delta R \right] \right)$$

$$N = \sqrt{-g_{00}} \Rightarrow \delta N = -\frac{\delta g_{00}}{2N} \quad \gamma_{ij} = g_{ij} \quad \delta \sqrt{\gamma} = \frac{1}{2} \sqrt{\gamma} \gamma^{ij} (\delta \gamma_{ij})$$

$$\delta g_{ij} = -g_{ik} g_{jl} (\delta g^{kl})$$

$$\delta g_{00} = \delta (-N) = -2N \delta N$$

$$\delta g^{00} = \delta \left(-\frac{1}{N^2}\right) = \frac{2}{N^3} (\delta N)$$

$$\delta g_{00} = -N^4 (\delta g^{00})$$

$$\delta R_{ij} = \delta (R^\rho_{\ i\rho j}) = \nabla_\rho (\delta I^\rho_{j|i}) - \nabla_j (\delta I^\rho_{\rho|i})$$

$$\delta I^\lambda_{\mu\nu} = \delta g^{\lambda\rho} g_{\alpha} I^\alpha_{\mu\nu} + \frac{1}{2} g^{\lambda\rho} (\partial_\mu \delta g_{\nu\rho} + \partial_\nu \delta g_{\mu\rho} - \partial_\rho \delta g_{\mu\nu})$$

$$= \frac{1}{2} g^{\lambda\rho} (\nabla_\mu \delta g_{\nu\rho} + \nabla_\nu \delta g_{\mu\rho} - \nabla_\rho \delta g_{\mu\nu})$$

$$= -\frac{1}{2} (g_{\nu\alpha} \nabla_\mu \delta g^{\alpha\lambda} + g_{\mu\alpha} \nabla_\nu \delta g^{\alpha\lambda} - g_{\alpha\beta} g_{\nu\beta} \nabla^\lambda \delta g^{\alpha\beta})$$

$$g^{\mu\nu} \delta I^\sigma_{\mu\nu} = -\nabla_\alpha \delta g^{\alpha\sigma} + \frac{1}{2} g_{\alpha\beta} \nabla^\sigma \delta g^{\alpha\beta}$$

$$g^{\mu\nu} \delta I^\lambda_{\lambda\mu} = -\frac{1}{2} g_{\alpha\beta} \nabla^\lambda \delta g^{\alpha\beta}$$

$$\delta R = R_{ij} \delta g^{ij} - D_i D_j \delta g^{ij} + g_{ij} D^2 \delta g^{ij}$$

$$D_i = \nabla_i$$

$$D^2 = \Delta$$

$$\begin{aligned}
\Rightarrow g_{ij} &= \int d^2x \left[N \sqrt{g} (Sg^{00}) \frac{RN^2}{2} - \frac{N\sqrt{g}}{2} g_{ij} R Sg^{ij} + N\sqrt{g} SR \right] \\
&= \int d^2x \sqrt{g} \left[\frac{RN^2}{2} (Sg^{00}) - \frac{g_{ij}R}{2} Sg^{ij} + R_{ij} Sg^{ij} \right] \quad (i,j=1,2) \\
&\quad + \sqrt{N} \left(\frac{1}{N} (g_{ij} D^2 N - D_i D_j N) Sg^{ij} \right)^D \\
&\quad (\because (g_{ij} D^2 - D_i D_j) \gamma = 0) \\
&= \int d^2x \sqrt{g} \left[\frac{RN^2}{2} (Sg^{00}) - \frac{1}{N} (D_i D_j - \gamma_{ij} D^2) N Sg^{ij} \right] \quad \text{且 } \frac{g_{ij}R}{2} + R_{ij} = 0 \quad \text{直接計算}
\end{aligned}$$

$$\Rightarrow G_{00} = \frac{RN^2}{2} \quad G_{i0} = 0 \quad G_{ij} = -\frac{1}{N} (D_i D_j - \gamma_{ij} D^2) N$$

→ This term 在以前會消失因為平時考慮正常的 action 時這一項會是全微分所以是 0.

其中 $R = \frac{\gamma_{11} \left(\frac{\partial \gamma_{22}}{\partial r} \right)^2 + \left(\gamma_{22} \frac{\partial \gamma_{11}}{\partial r} - 2\gamma_{12} \frac{\partial \gamma_{12}}{\partial r} \right) \frac{\partial \gamma_{22}}{\partial r} + 2 \left(\gamma_{12}^2 - \gamma_{11} \gamma_{22} \right) \frac{\partial^2 \gamma_{22}}{\partial r^2}}{\gamma_{11} \gamma_{22}}$

是 $\begin{pmatrix} \gamma_{11} & \gamma_{12} \\ \gamma_{12} & \gamma_{22} \end{pmatrix}$ 的 scalar curvature

值得注意的是在 2 級時我們的 $R_{ij} - \frac{g_{ij}}{2} R = 0$ 換之在 2 級時我們的 Einstein tensor 會 vanishing, 所以前面 G_{ij} 計算中的 $R_{ij} - \frac{g_{ij}}{2} R = R_{ij} - \frac{g_{ij}}{2} R = 0 \quad g_{ij} = \gamma_{ij}$ for $i,j=1,2$
所以 G_{ij} 只乘 $-\frac{1}{N} (D_i D_j - \gamma_{ij} D^2) N$.

G_{ij} 就是正常的 2+1 級中, $g_{ij} = \begin{pmatrix} -N^2(r) & 0 & 0 \\ 0 & \gamma_{11}(r) & 0 \\ 0 & 0 & \gamma_{22}(r) \end{pmatrix}$

的 Einstein tensor 因此好會等於 $\frac{RN^2}{2} = g_{00}$

③ 由我們解出 point mass 在 2+1 維度的 spherical symmetric solution.

Einstein equation

$$G_{\alpha\beta} = 8\pi G T_{\alpha\beta} \quad \begin{aligned} G &\text{ is Einstein tensor} \\ T &\text{ is stress-energy tensor} \end{aligned}$$

而考慮 static 的情況 so $g_{0i}=0=g_{i0}$

So we let $g_{ij} = \begin{pmatrix} N^2(r) & 0 & 0 \\ 0 & \gamma_{11} & \gamma_{12} \\ 0 & \gamma_{21} & \gamma_{22} \end{pmatrix}$

但我們由 spherical symmetry, we consider g as following

$$g_{ij} = \begin{pmatrix} -N^2 & 0 & 0 \\ 0 & \phi & 0 \\ 0 & 0 & \phi r^2 \end{pmatrix} \quad \begin{aligned} \text{而考慮} \\ \text{a point mass at origin.} \\ \text{with mass } M. \end{aligned}$$

$$\Rightarrow T^{00} \propto M S^2(r) \quad T^{0i} = T^{i0} = T^{ij} = 0 \quad \forall i, j \in \{1, 2\}$$

之後計算可以先算出在此 metric 下的

$$I^t_{tr} = \frac{\partial N}{\partial r} \quad I^r_{tt} = \frac{N \frac{\partial N}{\partial r}}{\phi(r)} \quad I^r_{rr} = \frac{\partial \phi}{\partial r} \quad I^r_{\theta\theta} = -\frac{r^2 \frac{\partial \phi}{\partial r} + 2r\phi}{2\phi}$$

$$I^{\theta}_{r\theta} = \frac{r \frac{\partial \phi}{\partial r} + 2\phi}{2r\phi}$$

$$I^i_{kl} = \frac{1}{2} g^{ij} \left[\frac{\partial g_{jk}}{\partial x_l} + \frac{\partial g_{jl}}{\partial x_k} - \frac{\partial g_{kl}}{\partial x_j} \right]$$

接下來我們用 2 種方法計算出我們的 G_{ij}

1. 我們用第一章得到的 $G_{00} = \frac{RN^2}{2}$ $G_{0i}=0$ $G_{ij} = \frac{1}{N}(\delta_{ij}D^2 - D_i D_j)N$
來計算。

直接用 2 級的 Scalar curvature 的計算可以得到

$$R = \frac{r(\phi_r)^2 - r\phi\phi_{rr} - \phi\phi_r}{r\phi^3}$$

(這邊主要是對 (ϕ, ϕ_r) 作
計算)

$$\text{而 } D_i D_j N = \partial_i \partial_j N - I_{ij}^k \partial_k N \Rightarrow D_i D_j N = \partial_r \partial_r N - I_{rr}^r \partial_r N$$

$$= \frac{\partial N}{\partial r^2} - \frac{\phi_r}{2\phi} N_r$$

$$D^2 = g^{ab} D_a D_b$$

$$D_2 D_2 N = \partial_\theta \partial_\theta N - I_{\theta\theta}^\theta \partial_\theta N$$

$$= \left(\frac{r^2 \phi_r + 2r\phi}{2\phi} \right) N_r$$

$$\Rightarrow \frac{1}{N} (g_{ij} D^2 - D_i D_j) N = G_{ij}$$

$$D^2 N = \frac{1}{\phi} \left(N_{rr} - \frac{\phi_r N_r}{2\phi} \right) + \frac{N_r}{\phi r^2} \left(\frac{r^2 \phi_r + 2r\phi}{2\phi} \right)$$

$$\Rightarrow G_{11} = \frac{1}{N} \left(\phi \cdot \frac{1}{\phi} (N_{rr} + \frac{1}{r} N_r) - N_{rr} + \frac{\phi_r N_r}{2\phi} \right) = \frac{1}{\phi} (N_{rr} + \frac{1}{r} N_r)$$

$$= \frac{N_r}{rN} + \frac{\phi_r N_r}{2\phi N}$$

$$G_{22} = \frac{1}{N} \left(r^2 (N_{rr} + \frac{N_r}{r}) - \frac{r^2 \phi_r + 2r\phi}{2\phi} N_r \right)$$

$$= \frac{1}{N} \left(r^2 N_{rr} - \frac{r^2 \phi_r N_r}{2\phi} \right)$$

$$G_{00} = \frac{rN^2 \phi_r^2 - rN^2 \phi \phi_r - N^2 \phi \phi_r}{2r\phi^3} = \frac{N^2 R}{2}$$

2. 可以發現以下我們是用最正規 $(2+1)D$ 的 metric 然後直接
計算其 G_{ij} , 會發現結果一樣

$$G_{ij} = R_{ij}^{(3)} - \frac{1}{2} g_{ij} R^{(3)} \quad (\text{這裡 } R_{ij} \text{ 和 } R \text{ 是考慮 } 2+1 \text{ 級的曲率})$$

這是使用程式計算 $2+1$ 級的

$$G_{ij} = \begin{cases} \frac{-rN^2(\frac{\partial\phi}{\partial r})^2 - rN^2\phi\frac{\partial^2\phi}{\partial r^2} - N^2\phi\frac{\partial\phi}{\partial r}}{2r\phi^3} & 0 \\ 0 & \frac{r\frac{\partial N}{\partial r}\frac{\partial\phi}{\partial r} + 2\phi\frac{\partial N}{\partial r}}{2rN\phi} \\ 0 & \frac{2r^2\phi\frac{\partial^2 N}{\partial r^2} - r^2\frac{\partial N}{\partial r}\frac{\partial\phi}{\partial r}}{2N\phi} \end{cases}$$

$$= 8\pi G T_{ij}$$

而除了 T_{00} 以外 $T_{ij} = 0$

$$\Rightarrow \left(r\frac{\partial\phi}{\partial r} + 2\phi\right)\frac{\partial N}{\partial r} = 0, \quad r^2\left(2\phi\frac{\partial^2 N}{\partial r^2} - \frac{\partial N}{\partial r}\frac{\partial\phi}{\partial r}\right) = 0$$

$$\text{若 } \frac{\partial N}{\partial r} \neq 0 \Rightarrow r\frac{\partial\phi}{\partial r} + 2\phi = 0 \Rightarrow \frac{\partial\phi}{\partial r} = -\frac{2\phi}{r} \Rightarrow \ln\phi = -2\ln r + C \Rightarrow \phi = \frac{C}{r^2}$$

$$2\phi\frac{\partial^2 N}{\partial r^2} - \frac{\partial N}{\partial r}\frac{\partial\phi}{\partial r} = \frac{2C}{r^2}\frac{\partial^2 N}{\partial r^2} + \frac{2C}{r^3}\frac{\partial N}{\partial r} = \frac{2C}{r^2}\left(\frac{\partial^2 N}{\partial r^2} + \frac{1}{r}\frac{\partial N}{\partial r}\right)$$

$$\Rightarrow \Delta N = 0 \Rightarrow \text{Let } \frac{dN}{dr} = \alpha \quad \frac{d\alpha}{dr} + \frac{\alpha}{r} = 0 \Rightarrow \frac{d\alpha}{dr} = -\frac{\alpha}{r}$$

$$\frac{N^2\left(r \cdot \frac{4C^2}{r^6} - r \frac{C}{r^2} \cdot \frac{6C}{r^4} + \frac{C^2}{r^2} \frac{2C}{r^3}\right)}{2r\left(\frac{C}{r^2}\right)^3} = 0 \quad \Rightarrow \ln\alpha = -\ln r + C_1 \quad \Rightarrow \alpha = \frac{C_1}{r} \quad \frac{dN}{dr} = \frac{C_1}{r} \quad N = C_1 \ln r + B$$

$$G_{22} = \frac{2r^2\frac{C}{r^2}\frac{\partial^2 N}{\partial r^2} - r^2 \cdot (\frac{C}{r^2})^2 \frac{\partial N}{\partial r}}{2(C_1 \ln r + B) \frac{C}{r^2}} \quad \left(\because \Delta N = C_1 2\pi S^2 F \right)$$

$$= \frac{4\pi r^2 C_1 S^2 F}{2(C_1 \ln r + B)} \quad \text{Then } G_{22} \text{ is not identical } 0, \rightarrow \Leftarrow$$

So we have $N \equiv C$ for some constant $C \in \mathbb{R}$.

而另一個簡單的看法可以由

$$G_{ij} \equiv 0 = -\frac{1}{2N} (D_i D_j - \gamma_{ij} D^2) N \Rightarrow D_i D_j N \equiv 0 \quad D^2 N \equiv 0$$

$\Rightarrow N$ is a constant.

So WLOG $N \equiv 1$

$$g_{ij} = \begin{pmatrix} -1 & 0 & 0 \\ 0 & \phi & 0 \\ 0 & 0 & \phi r^2 \end{pmatrix}$$

$$\begin{aligned} T_{rr}^r &= \frac{\frac{\partial \phi}{\partial r}}{2\phi} \\ T_{\theta\theta}^r &= -\frac{\frac{\partial \phi}{\partial r} r^2 - 2r\phi}{2\phi} \\ T_{r\theta}^\theta &= \frac{\frac{\partial \phi}{\partial r} r + 2\phi}{2\phi r} \end{aligned}$$

$$G_{ij} = \begin{pmatrix} \frac{r(\frac{\partial \phi}{\partial r})^2 - r\phi \frac{\partial^2 \phi}{\partial r^2} - \phi \frac{\partial \phi}{\partial r}}{2r\phi^3} & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

而 $\delta_{\text{polar}}^2(\vec{r})$ 代表在 polar coordinate (r, θ) 下的 delta function.

$$\int_{\mathbb{R}^2} \delta_{\text{polar}}^2(\vec{r}) dx dy = 1$$

$$= \int_0^\infty \int_0^{2\pi} \sqrt{g'_{\text{polar}}} \delta_{\text{polar}}^2(\vec{r}) dr d\theta$$

$$\text{而 } T_{\theta\theta}^{\text{IR}} = M \delta_{\text{IR}}^2(\vec{r}) = \frac{M s'(r)}{2\pi r} \text{ on一般地 } \mathbb{R}^2 = \int_0^\infty \int_0^{2\pi} \sqrt{g'_{\text{now}}} \delta_{\text{now}}^2(\vec{r}) dr d\theta$$

$$\text{而 } T_{\theta\theta}^{\text{(now)}} = M \delta_{\text{(now)}}^2(\vec{r}) = \frac{M}{\phi} \delta_{\text{IR}}^2(\vec{r})$$

其中 $g'_{\text{polar}} = \begin{pmatrix} 1 & 0 \\ 0 & r^2 \end{pmatrix}$ $g'_{\text{now}} = \begin{pmatrix} \phi & 0 \\ 0 & \phi r^2 \end{pmatrix}$ 而 $\frac{s'(r)}{2\pi r} = \delta_{\text{IR}}^2(\vec{r})$

$$\Rightarrow \delta_{\text{now}}^2(\vec{r}) = \frac{1}{\phi} \delta_{\text{IR}}^2(\vec{r})$$

$$\Rightarrow \sqrt{g'_{\text{polar}}} = r \quad \sqrt{g'_{\text{now}}} = \phi r \quad \text{而 } \sqrt{g'_{\text{now}}} \delta_{\text{now}}^2(\vec{r}) = \sqrt{g'_{\text{polar}}} \delta_{\text{polar}}^2(\vec{r})$$

$$T_{00} = M \frac{r}{\Phi r} S^2_{\mathbb{R}^2}(\vec{r}) = \frac{M}{\Phi} S^2_{\mathbb{R}^2}(\vec{r})$$

$$G_{00} = -\frac{1}{2} \frac{\Delta(\ln \Phi)}{\Phi} = 8\pi G M \frac{S^2_{\mathbb{R}^2}(\vec{r})}{\Phi} = 4GM \Delta_{(\mathbb{R}^2)} / \ln r$$

$$\Rightarrow \ln \Phi = -8GM / \ln r + \ln C$$

$$\Rightarrow \Phi = C r^{-8GM}. \text{ Then we can rescale } r \\ \Phi = r'^{-8GM}$$

$$\Rightarrow g_{ij} = \begin{pmatrix} -1 & 0 & 0 \\ 0 & r^{-8GM} & 0 \\ 0 & 0 & r^{-8GM+2} \end{pmatrix}$$

$$\text{Let } \alpha = 1 - 4GM \quad \theta' = \alpha \theta \quad \rho = \alpha^{-1} r^\alpha$$

$$d\theta' = \alpha d\theta \quad d\rho = r^{-4GM} dr$$

$$ds^2 = -dt^2 + d\rho^2 + \rho^2 d\theta'^2$$

Now $0 \leq \theta' \leq 2\pi\alpha$ and the space is a cone,
the space which is metrically flat except at
one point, 原點.

再分幾個 case

1. $\alpha > 0 \Leftrightarrow m < \frac{1}{4G} \Leftrightarrow$ metric near the origin is

regular $\int_0^R r^{-4GM} dr = \frac{r^{-4GM+1}}{1-4GM} \Big|_0^R = \frac{R^\alpha}{1-4GM}$

$\because \alpha > 0$ so it's finite and well-defined.

2. $\alpha < 0, m > \frac{1}{4G}$, then metric near origin is singular.
 $\left(\because \int_0^R r^{-4GM} dr = \frac{r^\alpha}{\alpha} \Big|_0^R \text{ But } \alpha < 0 \quad \frac{r^\alpha}{\alpha} \Big|_0^R = \frac{R^\alpha}{\alpha} - \frac{\varepsilon^\alpha}{\alpha} \right)$
 as $\varepsilon \rightarrow 0 \Rightarrow \text{distance} \uparrow \infty$.

注意到 when $\alpha \rightarrow -\alpha$ 時 他們 equivalent.
 $r \rightarrow \frac{1}{r}$

$$\text{When } m > \frac{1}{4G}, \alpha < 0, -4GM = \alpha \Leftrightarrow \alpha' = 4GM - 1 = 1 - 4GM' > 0$$

$$\Rightarrow M' = \frac{1}{2G} - M$$

$$\Rightarrow r' = \frac{1}{r}$$

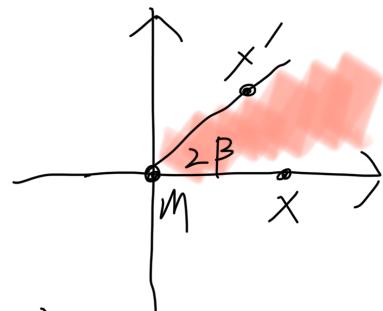
Case 3.
 So 當 $\frac{1}{2G} > m > \frac{1}{4G}$ 在原點和 $M' = \frac{1}{2G} - M$ 在無限遠點
 是一樣的效果.

$$\text{Case 4. } 4GM = 1 \quad (\alpha = 0) \quad g_{ij} = \begin{pmatrix} 0 & 0 \\ 0 & r^{-2} \\ 0 & 0 \end{pmatrix}$$

$$\Rightarrow dt^2 = -dt^2 + d(\ln r)^2 + d\theta^2 \quad -\infty < \ln r < \infty \quad 0 \leq \theta \leq 2\pi$$

④ This a strip on the Cartesian coordinate $(\ln r, \theta)$
 而當有兩個以上的粒子時，也可以透過一些等價坐标的
 技巧來求解

在一個粒子的 case 時



x 和 x' 被
 identify 成同一點
 紅色區域會被
 移除

$(\because \beta = \alpha\theta \text{ 所以 } \theta \text{ 只從 } (0, 2\pi\alpha))$ 所以有 $2\pi(1-\alpha)$ 的角度被
 堆掉 其中 $\beta = \pi(1-\alpha) = 4\pi GM$

$$f_{hk} \quad x' = \Omega(\beta) x, \quad t' = t \quad \Omega(\beta) = \begin{pmatrix} \cos \beta & \sin \beta \\ -\sin \beta & \cos \beta \end{pmatrix}$$

static 的
✓

$$\beta = 4\pi GM = \pi(-\alpha)$$

而兩個粒子的情況則如下，考慮

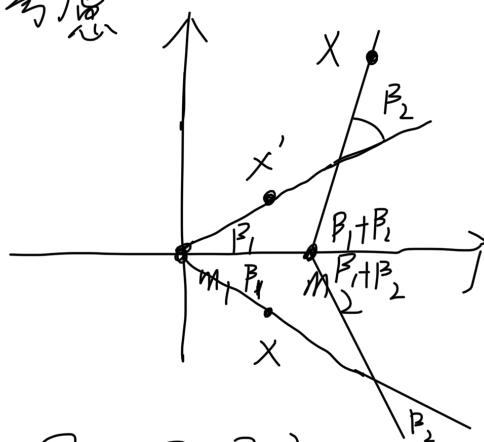
m_1 在 origin m_2 在 a 點

$$\beta_1 = 4\pi GM_1, \quad \beta_2 = 4\pi GM_2$$

$$x' = \Omega(\beta) x$$

$$x'' = \Omega(\beta_1)(a + \Omega(\beta_2)(x-a)) \quad \text{Let } \Omega_i = \Omega(\beta_i)$$

$$= \Omega_1 a + \Omega_1 \Omega_2 x - \Omega_1 \Omega_2 a$$



$$\text{Let } b = \frac{\sin \beta_2}{\sin(\beta_1 + \beta_2)} \Omega_1^{\frac{1}{2}} a$$

$$\begin{aligned} b - \Omega_1 \Omega_2 b &= \frac{\sin \beta_2}{\sin(\beta_1 + \beta_2)} \begin{pmatrix} \cos \beta_1 - \cos(3\beta_1 + 2\beta_2) & \sin \beta_1 - \sin(3\beta_1 + 2\beta_2) \\ -\sin \beta_1 + \sin(3\beta_1 + 2\beta_2) & \cos \beta_1 - \cos(3\beta_1 + 2\beta_2) \end{pmatrix} a \\ &= \frac{\sin \beta_2}{\sin(\beta_1 + \beta_2)} \begin{pmatrix} 2\sin(2\beta_1 + \beta_2)\sin(\beta_1 + \beta_2) & -2\cos(2\beta_1 + \beta_2)\sin(\beta_1 + \beta_2) \\ 2\cos(2\beta_1 + \beta_2)\sin(\beta_1 + \beta_2) & 2\sin(2\beta_1 + \beta_2)\sin(\beta_1 + \beta_2) \end{pmatrix} a \\ &= \begin{pmatrix} 2\sin(2\beta_1 + \beta_2)\sin \beta_2 & -2\cos(2\beta_1 + \beta_2)\sin \beta_2 \\ 2\cos(2\beta_1 + \beta_2)\sin \beta_2 & 2\sin(2\beta_1 + \beta_2)\sin \beta_2 \end{pmatrix} a \\ &= \begin{pmatrix} \cos 2\beta_1 - \cos(2\beta_1 + 2\beta_2) & \sin 2\beta_1 - \sin(2\beta_1 + 2\beta_2) \\ -\sin 2\beta_1 + \sin(2\beta_1 + 2\beta_2) & \cos 2\beta_1 - \cos(2\beta_1 + 2\beta_2) \end{pmatrix} a \\ &= (\Omega_1 - \Omega_1 \Omega_2) a \end{aligned}$$

$\Rightarrow x'' = b - \Omega_1 \Omega_2 b$ so we can 計算上面 particle 的總系統
等價於有一個 $M_1 + M_2$ 的粒子在 b 點的系統。

⑤ 而 horizon event 不存在的原因主要是
我們的 metric 除了 O 點是一個 singularity 以外
其餘地方都是良好定義的，所以不會有 horizon
event 發生。

而 test particle 的 加速度

$$\frac{d^2 x^\mu}{dt^2} + T_{\alpha\beta}^\mu \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = 0 \quad \text{and } g_{ij} = \begin{pmatrix} -1 & 0 & 0 \\ 0 & r^{-8GM} & 0 \\ 0 & 0 & r^{-8GM+2} \end{pmatrix}$$

$$\phi = r^{-8GM}$$

$$T_{rr}^r = \frac{\partial \phi}{\partial r} = -\frac{8GM}{2r}$$

$$T_{\theta\theta}^r = -\frac{\partial \phi}{\partial r} r^2 - 2r\phi = \frac{r(8GM+2)}{2} \quad \text{其餘都是 } 0,$$

$$T_{r\theta}^\theta = \frac{\partial \phi}{\partial r} r + 2\phi = \frac{8GM+2}{2r}$$

$$\Rightarrow \frac{d^2 x^r}{dt^2} + T_{rr}^r \frac{dx^r}{dt} \frac{dx^r}{dt} + T_{\theta\theta}^r \frac{dx^\theta}{dt} \frac{dx^\theta}{dt} = 0$$

$$\Rightarrow \frac{d^2 x^r}{dt^2} - \frac{4GM}{r} \left(\frac{dx^r}{dt} \right)^2 + 4(GM+1)r \left(\frac{dx^\theta}{dt} \right)^2 = 0$$

$$\Rightarrow \frac{d^2 x^\theta}{dt^2} + 2 \cdot \frac{(GM+1)}{r} \frac{dx^r}{dt} \frac{dx^\theta}{dt} = 0 \quad \text{So If test particle is at rest at first then test particle 不會有任何加速度}$$

而如果是一個一開始靜止的
他的 geodesic equation 只有 $T_{\theta\theta}^r$ 的項

有奇點 ($x^i = 0 \quad i=1,2$)

$$x^i + T_{\theta\theta}^i = 0 \quad \text{但 } g_{\theta\theta} = -1 \quad \text{且 } g_{i\theta} = 0 \Rightarrow T_{\theta\theta}^i = 0 \quad \forall i=1,2,$$

總結

1. 我們先許言我們一般情況下我們 metric
可以經過適當變數代換調整成好看形式

2. Reproduce some result of paper, 從 Einstein action
出發照文章計算出與原文有一點差異的 G_{ij} , 而我
們用 $S_{dS/NR^{(3)}}$ 作變分得到的 G_{ij} 和直接對

$$g_{ij} = \begin{pmatrix} -N^2 & 0 & 0 \\ 0 & \gamma_{11} & \gamma_{12} \\ 0 & \gamma_{21} & \gamma_{22} \end{pmatrix} \text{ 用 } G_{ij} = R_{ij}^{(3)} - \frac{1}{2} g_{ij} R^{(3)} \text{ 會得到一樣結果}$$

3. - 開始先驗證正在 $g_{ij} = \begin{pmatrix} -N^2 & 0 & 0 \\ 0 & \gamma_{11} & 0 \\ 0 & 0 & \gamma_{22} \end{pmatrix}$ 下第 2 部分的結果是正確的。
後來用和原文相似作法 (把細節補齊後) 得到我們 metric
的解 for one point mass 的情況。

4. 解釋若有多一個 point mass 的情況我們該如何處理

5. 解釋 horizon event 不存在以及 conical singular 手口 —
一個 test particle 的加速度。