

# Histogram

Due Date: November 9 at 11:59 pm

## Objective

The purpose of this lab is to implement an efficient histogramming equalization algorithm for an input image. Like the image convolution MP, the image is represented as `RGB float` values. You will convert that to `GrayScale unsigned char` values and compute the histogram. Based on the histogram, you will compute a histogram equalization function which you will then apply to the original image to get the color corrected image.

## Instruction

Edit the code in the code tab to perform the following:

- Cast the image to `unsigned char`
- Convert the image from RGB to Gray Scale. You will find one of the lectures and textbook chapters helpful.
- Compute the histogram of the image
- Compute the scan (prefix sum) of the histogram to arrive at the histogram equalization function
- Apply the equalization function to the input image to get the color corrected image

## Background

In this section we discuss some of the background details of the histogram equalization algorithm. For images that represent the full color space, we expect an image's histogram to be evenly distributed. This means that we expect the bin values in the histogram to be `pixel_count/.256`, assuming that we scale the pixel luminous values so that they fit between 0 and 256. This algorithm adjusts an image's histogram so that all bins have equal probability.

To get the host ptrs to the image, you want the following code:

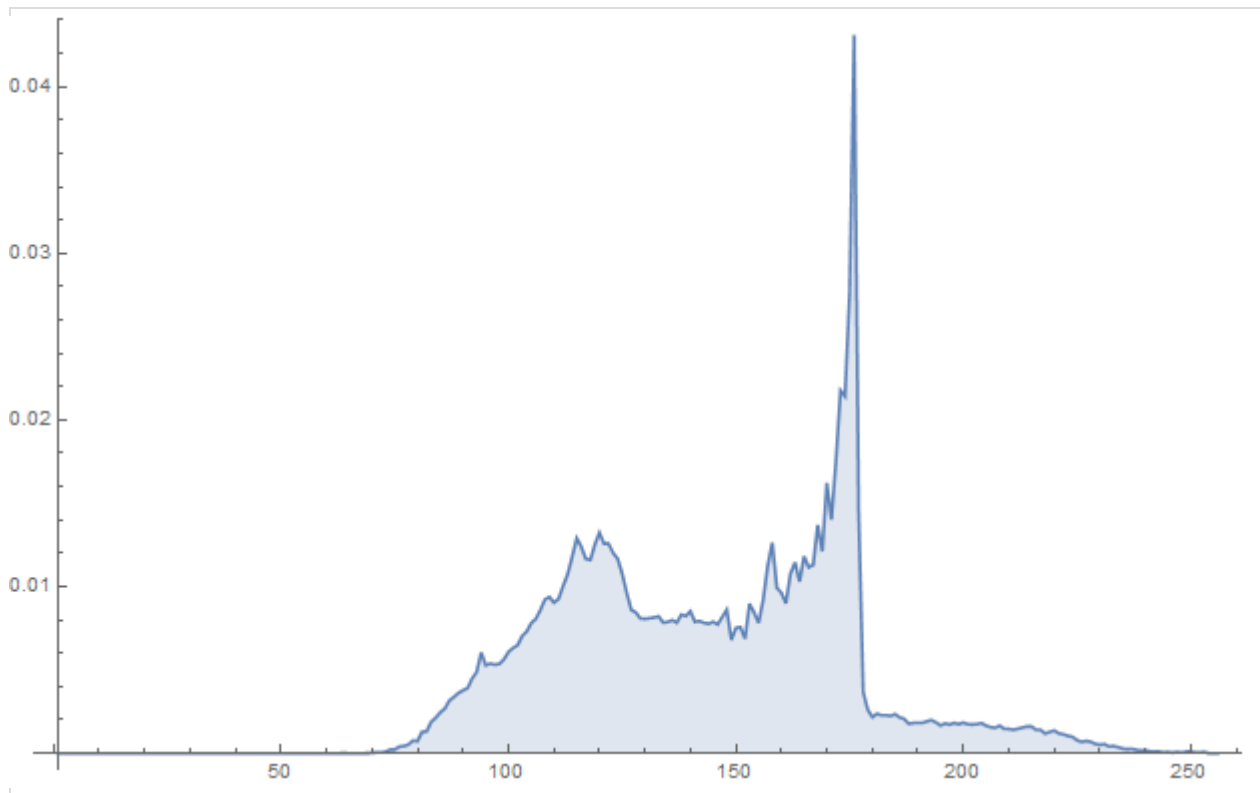
```
hostInputImageData = wblmage_getData(inputImage);  
hostOutputImageData = wblmage_getData(outputImage);
```



We first need to convert the image to gray scale by computing it's luminosity values that represent the brightness of the image and would allow us to simplify the histogram computation.

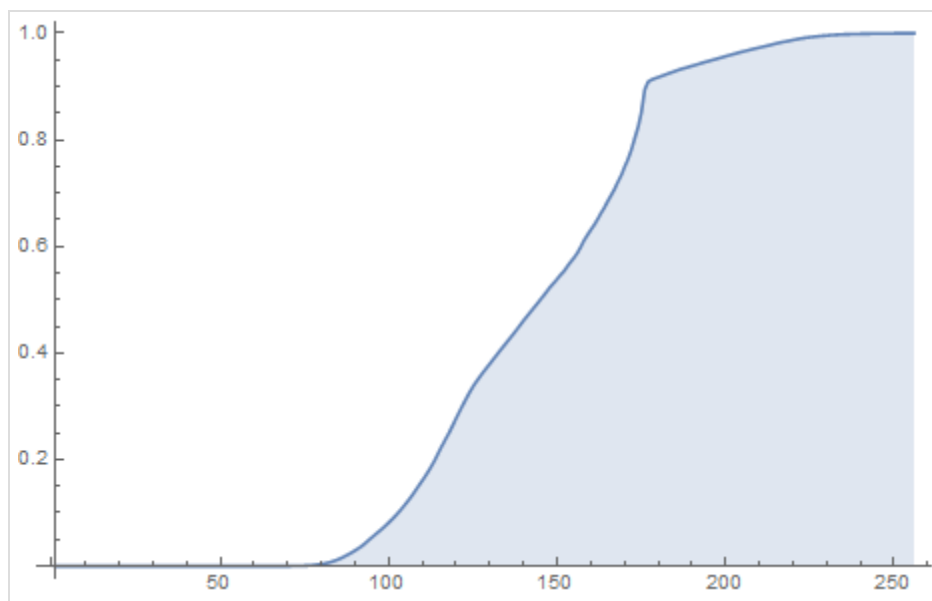


The histogram computes the number of pixels having a specific brightness value. Dividing by the number of pixels (width \* height) gives us the probability of a luminosity value to occur in an image.

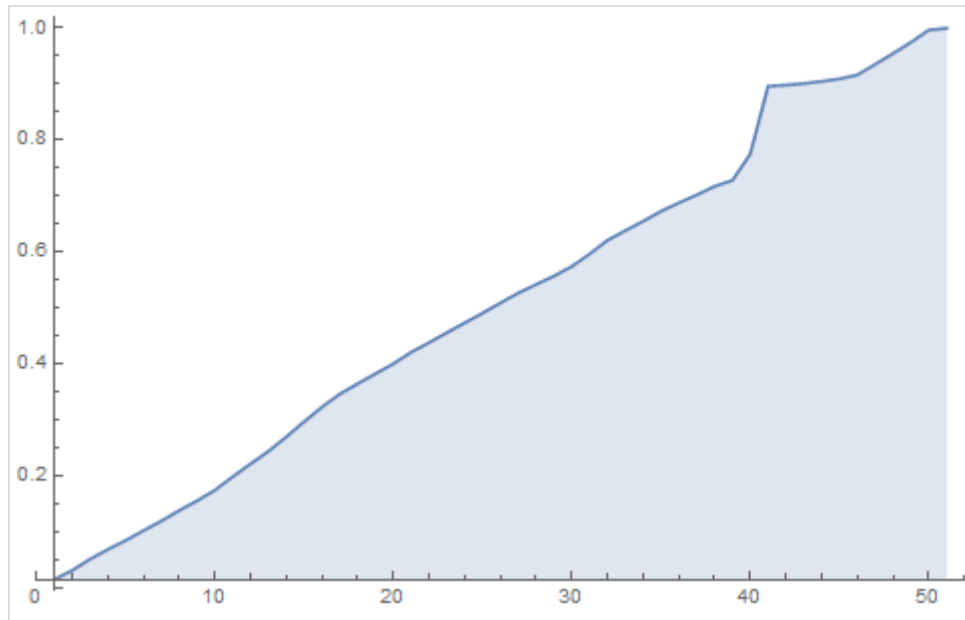


A color balanced image is expected to have a uniform distribution of the luminosity values.

This means that if we compute the Cumulative Distribution Function (CDF) we expect a linear curve for a color equalized image. For images that are not color equalized, we expect the curve to be non-linear.



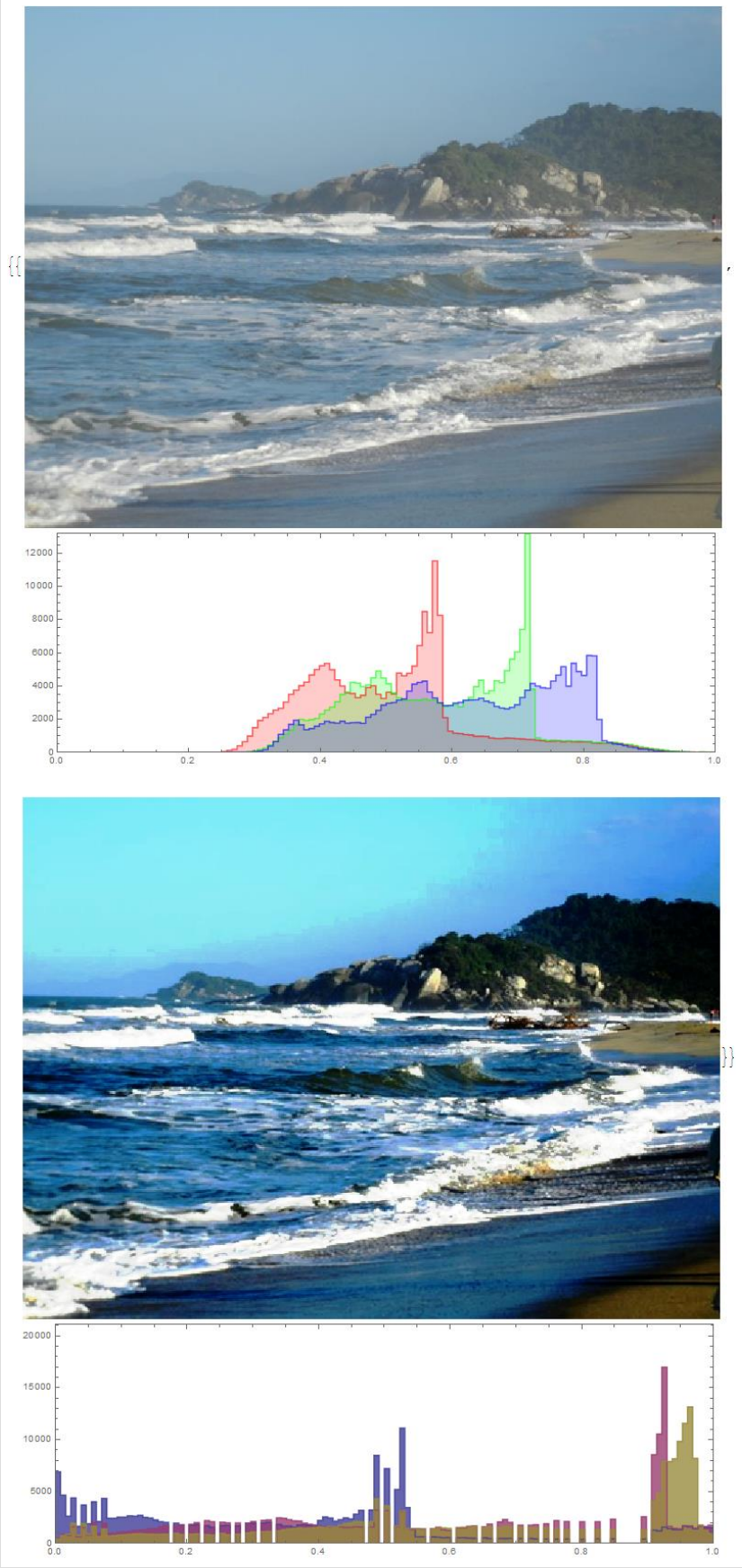
The algorithm equalizes the curve by computing a transformation function to map the original CDF to the desired CDF (the desired CDF being an almost linear function).



The computed transformation is applied to the original image to produce the equalized image.



Note that the CDF of the histogram of the new image has been transformed into an almost linear curve.



# Implementation Steps

Here we show the steps to be performed. The computation to be performed by each kernel is illustrated with serial pseudo code.

## Cast the image from float to unsigned char

Implement a kernel that casts the image from float \* to unsigned char \*.

```
for ii from 0 to (width * height * channels) do
    ucharImage[ii] = (unsigned char) (255 * inputImage[ii])
end
```

## Convert the image from RGB to GrayScale

Implement a kernel that converts the RGB image to GrayScale. A sample sequential pseudo code is shown below. You will find one the lectures and one of the textbook chapters helpful.

```
for ii from 0 to height do
    for jj from 0 to width do
        idx = ii * width + jj
        # here channels is 3
        r = ucharImage[3*idx]
        g = ucharImage[3*idx + 1]
        b = ucharImage[3*idx + 2]
        grayImage[idx] = (unsigned char) (0.21*r + 0.71*g + 0.07*b)
    end
end
```

## Compute the histogram of grayImage

Implement a kernel that computes the histogram (like in the lectures) of the image. A sample pseudo code is shown below. You will find one of the lectures and one of the textbook chapters helpful.

```
histogram = [0, ....., 0] # here len(histogram) = 256
for ii from 0 to width * height do
    histogram[grayImage[ii]]++
end
```

## Compute the Cumulative Distribution Function of histogram

This is a scan operation like you have done in the previous lab. A sample sequential pseudo code is shown below.

```
cdf[0] = p(histogram[0])
for ii from 1 to 256 do
    cdf[ii] = cdf[ii - 1] + p(histogram[ii])
end
```

Where `p()` calculates the probability of a pixel to be in a histogram bin

```
def p(x):
    return x / (width * height)
end
```

Compute the minimum value of the CDF. The maximal value of the CDF should be 1.0.

## Define the histogram equalization function

The histogram equalization function (`correct`) remaps the cdf of the histogram of the image to a linear function and is defined as

```
def correct_color(val)
    return clamp(255*(cdf[val] - cdfmin)/(1.0 - cdfmin), 0, 255.0)
end

def clamp(x, start, end)
    return min(max(x, start), end)
end
```

## Apply the histogram equalization function

Once you have implemented all of the above, then you are ready to correct the input image. This can be done by writing a kernel to apply the `correct_color()` function to the RGB pixel values in parallel.

```
for ii from 0 to (width * height * channels) do
    ucharImage[ii] = correct_color(ucharImage[ii])
end
```



## Cast back to float

```
for ii from 0 to (width * height * channels) do
    outputImage[ii] = (float) (ucharImage[ii]/255.0)
end
```

And you're done

## Image Format

For people who are developing on their own system. The images are stored in PPM (P6) format, this means that you can (if you want) create your own input images. The easiest way to create image is via external tools. You can use tools such as `bmptoppm`.

## Suggestions (for all labs)

- Backup your code regularly.
- Do not modify the template code provided -- only insert code where the `/**` demarcation is placed
- Develop your solution incrementally and test each version thoroughly before moving on to the next version
- Do not wait until the last minute to attempt the lab.
- If you get stuck with boundary conditions, grab a pen and paper. It is much easier to figure out the boundary conditions there.
- Implement the serial CPU version first, this will give you an understanding of the loops
- Get the first dataset working first. The datasets are ordered so the first one is the easiest to handle
- Make sure that your algorithm handles non-regular dimensional inputs (not square or multiples of 2). The slides may present the algorithm with nice inputs, since it minimizes the conditions. The datasets reflect different sizes of input that you are expected to handle
- Make sure that you test your program using all the datasets provided (the datasets can be selected using the dropdown next to the submission button)
- Check for errors: for example, when developing CUDA code, one can check for if the function call succeeded and print an error if not via the following macro:

```
#define wbCheck(stmt) do { \
    cudaError_t err = stmt; \
    if (err != cudaSuccess) { \
        wbLog(ERROR, "Failed to run stmt ", #stmt); \
        wbLog(ERROR, "Got CUDA error ... ", \
        cudaGetErrorString(err)); \
        return -1; \
    } \
} while(0)
```

An example usage is `wbCheck(cudaMalloc(...))`.

## Plagiarism

Plagiarism will not be tolerated. The first offense will result in the two parties getting a 0 for the machine problem. Second offense results in a 0 for the course.