

## TOPICS II

1.

$$J_{\text{EARTH SPIN}} = I_{\text{SPIN}} \omega = \frac{2}{5} M_{\text{EARTH}} R_{\text{EARTH}}^2 \frac{2\pi}{P_{\text{ROTATION}}}$$

$$= \frac{2}{5} \cdot 6 \cdot 10^{24} \cdot (6.4 \cdot 10^6)^2 \frac{2\pi}{24 \cdot 60 \cdot 60}$$

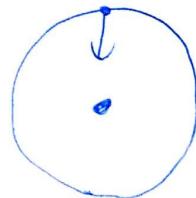
$$= \underline{\underline{7 \cdot 1 \cdot 10^{33} \text{ kg m}^2 \text{ sec}^{-1}}}$$

$$J_{\text{MOON}} = I_{\text{SPIN}} \omega = \frac{2}{5} M_{\text{MOON}} R_{\text{MOON}}^2 \frac{2\pi}{P_{\text{ROTATION MOON}}}$$

$$= \frac{2}{5} \cdot (7.35 \cdot 10^{22}) \cdot (1740)^2 \cdot \frac{2\pi}{27.3 \cdot 24 \cdot 60 \cdot 60}$$

LENGTH OF  
SIDEREAL MONTH  
IN DAYS

$$= \underline{\underline{2 \cdot 37 \cdot 10^{29} \text{ kg m}^2 \text{ sec}^{-1}}}$$



$$\text{L}_{\text{ORBITAL}} = M_{\text{MOON}} \alpha_{\text{EARTH-MOON}} v_{\text{MOON}}$$

$$\frac{v^2}{r} = \frac{GM_{\text{EARTH}}}{r^2}$$

$$= M_{\text{MOON}} \alpha \sqrt{\frac{GM_{\text{EARTH}}}{a^3}} = M_{\text{MOON}} \sqrt{GM_{\text{EARTH}} a}$$

$$= 7.35 \cdot 10^{22} \cdot \left[ G \cdot 6 \cdot 10^{24} \cdot 3.84 \cdot 10^8 \right]$$

$$= 2 \cdot 9 \cdot 10^{34} \text{ kg m}^2 \text{s}^{-2}$$

i)

$$\frac{J_{\text{EARTH}}}{L_{\text{ORBITAL}}} = \frac{7.1}{29} \approx 0.24$$

$$\frac{J_{\text{MOON}}}{L_{\text{ORBITAL}}} = \frac{2.37}{2.9 \cdot 10^5} \approx \cancel{8} \cdot 10^{-6} \quad \text{IE NEGIGIBLE.}$$

WHEN SYSTEM BECOMES FULLY SYNCED:

~~$\Omega_E \approx \omega$~~   $\omega_{\text{EARTH}}^{\text{SPIN}} = \omega_{\text{ORBITAL}}^{\text{MOON}} \rightarrow \Omega_L$

TO CONSERVE ANGULAR MOM:

~~$I\omega + L_{\text{ORBITAL}} = \text{CONSTANT}$~~

$$I\omega + M_u \sqrt{GM_E} \sqrt{a} = C$$

NOTING THAT:

~~$\Omega_L^2 a = \frac{GM_{\text{EARTH}}}{a}$~~

$$\sqrt{a} = \left( \frac{GM_{\text{EARTH}}}{\Omega_L^2} \right)^{\frac{1}{4}}$$

SUBSTITUTE THIS IN:

$$I\omega + M_u \sqrt{GM_E} \left( \frac{GM_{\text{EARTH}}}{\Omega_L^2} \right)^{\frac{1}{4}} = C$$

AT FULL SYNC,  $\omega = \Omega = \omega_c$

$$I \propto + M_m \sqrt{GM_E} \left( GM_{\text{EARTH}} \right)^{\frac{1}{4}} x^{-\frac{1}{2}} = C$$

THE NUMBERS:

$$I = \frac{2}{5} M_{\text{EARTH}} T_{\text{EARTH}}^2 = \frac{2}{5} \cdot 6 \cdot 10^{24} \cdot (6.4 \cdot 10^6)^2 = 9.8 \cdot 10^{37} \text{ kg m}^2$$

$$M_m \sqrt{GM_E} \left( GM_{\text{EARTH}} \right)^{\frac{1}{4}} = \cancel{5.9 \cdot 10^{33} \text{ kg m}^2}$$

$$C = 7.1 \cdot 10^{33} + 2.9 \cdot 10^{34} = 3.6 \cdot 10^{34}$$

REWRITE EQ:-

~~$$9.8 \cdot 10^3 x + 5.9 \cdot 10^{-2} x^{-\frac{1}{2}} = 3.6$$~~

WHICH DOESN'T HAVE REAL SOLUTIONS.  $\Rightarrow$  STH IS WRONG.

~~NOTE: WE EXPED X TO BE ON THE ORDER OF  $10^{30}$~~

~~IGNORE FIRST TERM IN EQUATION~~

~~$$x^{-\frac{1}{2}} = \frac{3.6}{5.9}$$~~

$$|x| = 1.54 \cdot 10^{-3}$$

$$\frac{2\pi}{|x|} = 4 \cdot 10^3 \text{ sec} \approx 68 \text{ HOURS}$$

WHICH IS VERY OFF.

$$E_{\text{EARTH SPIN}} = \frac{1}{2} I \omega^2$$

$$= \frac{1}{2} \cdot \frac{2}{5} \cdot 6 \cdot 10^{24} \left( 6.9 \cdot 10^6 \right)^2 \left( \frac{2\pi}{27.84 \cdot 60^2} \right)^2 = 6.5 \cdot 10^{23} J$$

$$E_{\text{ORBITAL POTENTIAL}} = - \frac{GM_{\text{MOON}} M_{\text{EARTH}}}{a}$$

$$= - \frac{G \cdot 7.35 \cdot 10^{22} \cdot 6 \cdot 10^{24}}{3.84 \cdot 10^8}$$

~~$$= -7.66 \cdot 10^{28}$$~~

$$E_{\text{ORBITAL KINETIC}} = \frac{1}{2} M_{\text{MOON}} v^2 = \frac{1}{2} M \left( \frac{2\pi r}{P} \right)^2$$

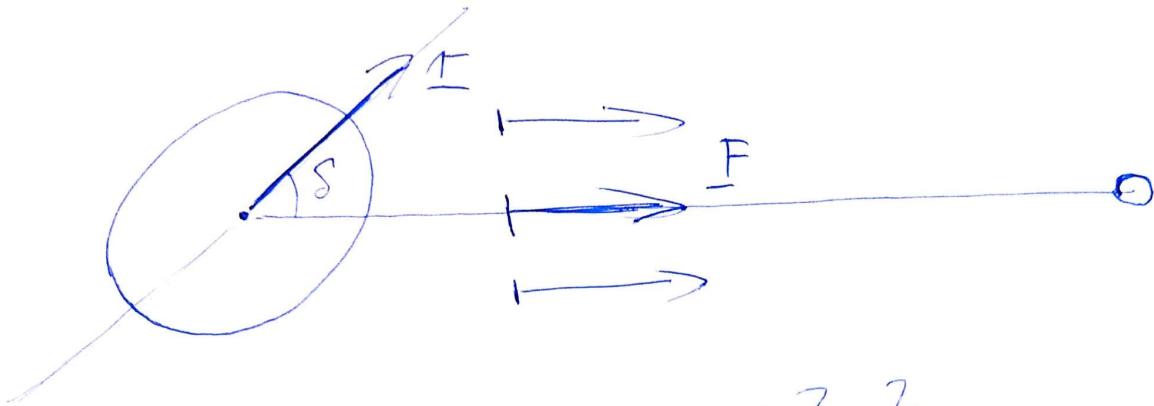
$$= \frac{1}{2} (7.35 \cdot 10^{22}) \cdot \left( \frac{2\pi \cdot 3.84 \cdot 10^8}{27.84 \cdot 60^2} \right)^2 = 3.93 \cdot 10^{28}$$

$$E_{\text{TOTAL}} = -3.9 \cdot 10^{28}$$

$$\frac{E_{\text{EARTH SPIN}}}{E_{\text{ORBITAL TOTAL}}} = \frac{6.5 \cdot 10^{23}}{3.93 \cdot 10^{28}} \approx 16.5$$

COULD'VE  
JUST  
USED  
VIRIAL  
THM.

$E_{\text{ORBITAL TOTAL}}$  WILL NOT CHANGE TENFOLD, SO  
MOST E IS DISSIPATED.



NOTES 2 PAGE 5:  $E_{\text{TIDAL}} \sim \frac{GM_m^2 R_e^2}{R_{0m}^6}$

TORQUE =  $\tau \times F$

ENERGY IN  
"TORSION",  
IE ~~FROM~~  
~~FROM~~ TORQUE

= TORQUE  $\cdot \Delta\theta$

[ HOW MUCH  
EARTH IS  
TURNED  
BY TORQUE ]

$\propto$  TORQUE  $\propto \vec{F} \times \vec{r} \propto \sin\delta$



~~$\propto \frac{GM_m^2 R_e^2}{R_{0m}^6}$~~

$\Rightarrow$  TORQUE  $\propto \sin\delta$

$\Rightarrow$  TORQUE  $\propto$  ENERGY IN TORSION

[ SAME CONSTANT ]

$\Rightarrow$  TORQUE =  $\frac{GM_m^2 R_e^2}{R_{0m}^6} \sin\delta$

( I THINK THIS IS A VERY DODGY DERIVATION )

"ASSUMING THAT---"

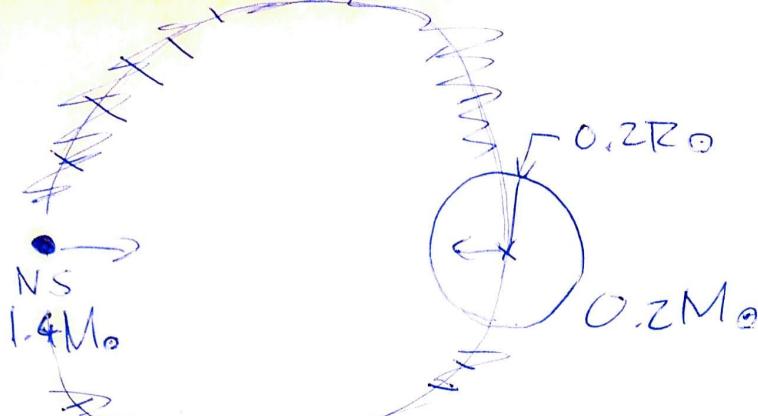
$$\chi_e \propto \left( \frac{\text{TORQUE}}{\text{MOMENT OF INERTIA}} \right)^{-1}$$

$$\frac{\chi_e}{\chi_m} \propto \left( \frac{\cancel{\text{TORQUE}}^5 R_{\text{e}}^6 M_m^2}{M_e^2 R_m^5 R_{\text{em}}^6} / \frac{\cancel{\text{I}}^2 M_e R_{\text{e}}}{M_m R_m^2} \right)^{-1}$$

$$\propto \left( \frac{M_m^3 R_e^3}{M_e^3 R_m^3} \right)^{-1} \propto \left( \frac{M_m^3 M_e}{M_e^3 M_m} \right)^{-1} = \left( \frac{M_m}{M_e} \right)^{-2}$$

$$\propto \left( \frac{M_e}{M_m} \right)^2 \Rightarrow \text{AGREES WITH OBSERVATION.}$$

WE ALWAYS SEE  
SAME SIDE OF MOON,  
MOON SEES ALL SIDES  
OF EARTH.



WITH A BIT OF HOPE:  $0.2 \ll 1.4$

$$\text{NOTES 2 PAGE 4: } R_{\text{ROCHE}} = \left( \frac{M_2}{3M_1} \right)^{\frac{1}{3}} a \\ = \left( \frac{0.2}{3 \cdot 1.4} \right)^{\frac{1}{3}} a \\ = 0.362 a$$

"FILLS ITS ROCHE LOBE"  $\Rightarrow R_{\text{ROCHE}} = 0.2 R_\odot$

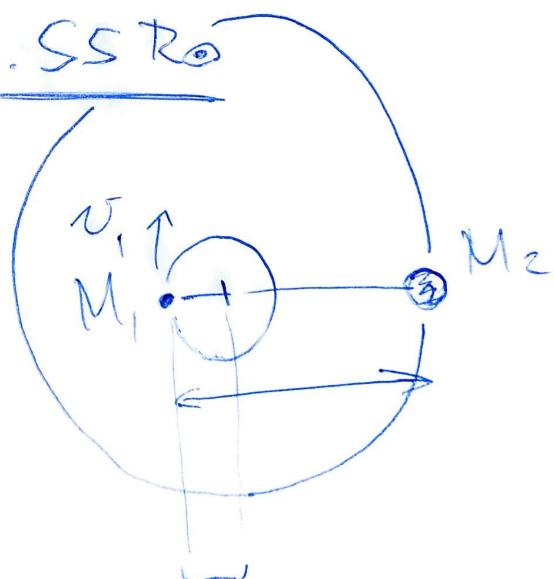
$$\Rightarrow a = \frac{0.2 R_\odot}{R_{\text{ROCHE}} / 0.362} = \frac{0.55 R_\odot}{}$$

~~WHAT'S THE PERIOD?~~

$$\frac{GM_1 M_2}{d^2} = \frac{M_1 \omega^2}{d}$$

or

$$\frac{GM_2}{d^2} = \frac{\omega^2}{M_1}$$



$$\frac{M_2}{M_1 + M_2} a = \frac{0.2}{1.4 + 0.2} a \\ = 0.125 a$$

WITHOUT LOSS OF GENERALITY, WE CAN SAY:

$$\frac{GM_i M_j}{a^2} = \frac{v_i^2}{\frac{M_j}{M_i + M_j}} M_i$$

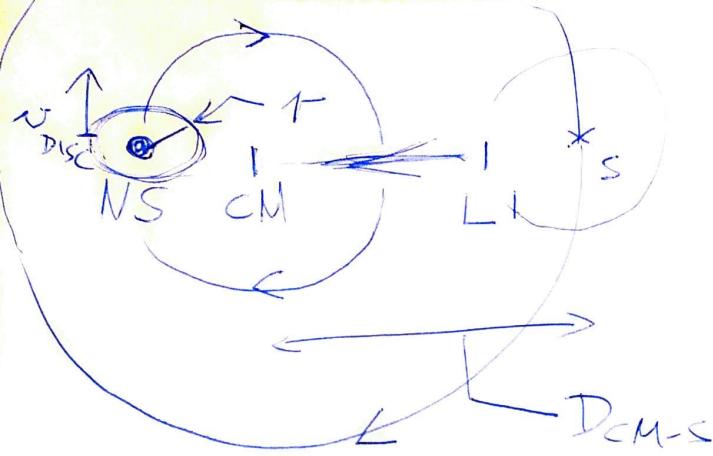
$$v_i^2 = GM_i M_j \frac{1}{M_i + M_j} \frac{1}{a}$$

$$P = \frac{2\pi a \frac{M_j}{M_i + M_j}}{v_i} = \frac{2\pi a}{\sqrt{G}} \frac{\frac{M_j}{M_i + M_j} \sqrt{M_i + M_j}}{\sqrt{M_i M_j}} a$$

$$= \frac{2\pi a^{3/2}}{\sqrt{G}} \frac{1}{\sqrt{M_i + M_j}} \cancel{\frac{M_j}{\sqrt{M_i M_j}}} = \frac{2\pi a^{3/2}}{\sqrt{G(M_i + M_j)}}$$

$$= \frac{2\pi (0.55 R_\odot)^{3/2}}{\sqrt{G(0.2 + 1.4) M_\odot}} = \cancel{2.45 \cdot 10^{16} \text{ SEC}}$$

$$\underline{\underline{= 3.2 \cdot 10^3 \text{ SEC}}}$$



ANGULAR MOMENTUM LOSS

FROM SECONDARY:

$$\approx M_{\text{STREAM}} \cdot v_{\text{SECONDARY}} D_{\text{CM-S}}$$

$$\approx M_{\text{STREAM}} \underbrace{\sqrt{G M_{\text{NS}} \frac{1}{M_s + M_{\text{NS}}}}}_{v_{\text{SECONDARY}}} \underbrace{\frac{1}{\sqrt{a}}}_{D_{\text{CM-S}}} \underbrace{\frac{M_{\text{NS}}}{M_{\text{NS}} + M_s} \alpha}_{\alpha}$$

$$\approx M_{\text{STREAM}} \sqrt{G} M_{\text{NS}}^2 \sqrt{a} \frac{1}{(M_s + M_{\text{NS}})^{3/2}}$$

THIS GOES TO ANGULAR MOM OF DISC:

$$m_{\text{STREAM}} v_{\text{DISC}} \tau$$

NOTE:  $\frac{v_{\text{DISC}}^2}{\tau} = \frac{GM_{\text{NS}}}{\tau^2} \Rightarrow v_{\text{DISC}} = \sqrt{\frac{GM_{\text{NS}}}{\tau}}$

EQUATE LOSS & GAIN:

$$\cancel{M_{STREAM}^2 N_{DISC}} = \cancel{M_{STREAM}^2} \sqrt{G} M_{NS}^2 \sqrt{a} \frac{1}{\sqrt{(M_S + M_{NS})^{\frac{3}{2}}}}$$

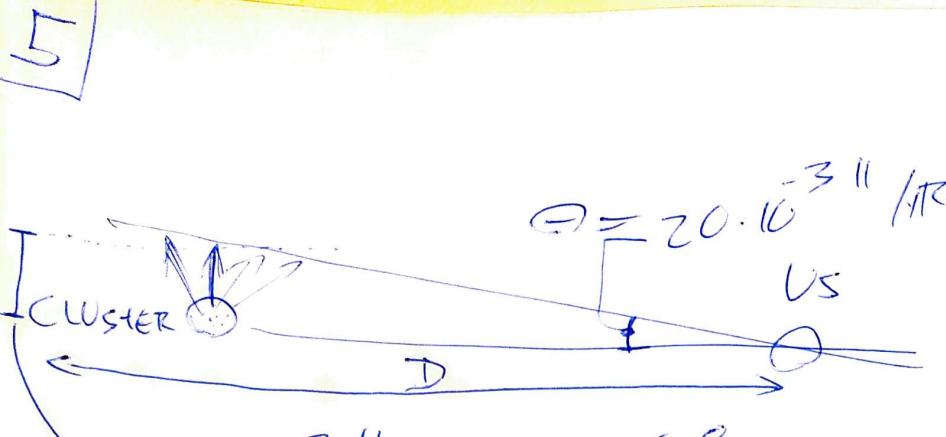
$$\frac{\cancel{\sqrt{G} \sqrt{M_{NS}}}}{\sqrt{a}} r = \cancel{\sqrt{G} M_{NS}^2} \sqrt{a} \frac{1}{\sqrt{(M_S + M_{NS})^{\frac{3}{2}}}}$$

$$r = \left( \frac{M_{NS}^{\frac{3}{2}}}{M_S + M_{NS}} \right)^{\frac{1}{2}} a$$

$$r = \left( \frac{1.4}{1.4 + 0.2} \right)^{\frac{1}{2}} a$$

$$r \approx 0.67 a$$

GIVEN THAT LI IS  $a - 0.362a = 0.638a$  AWAY FROM NS, THIS RESULT SEEMS TOO BIG.



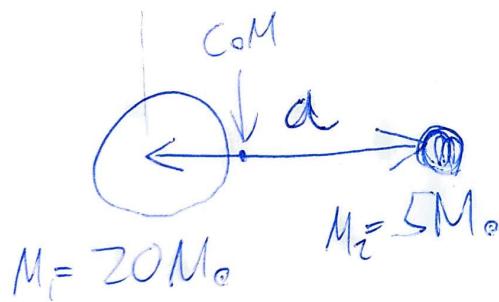
$$20 \cdot 10^{-3} \text{ rad} = 5.55 \cdot 10^{-6} \text{ rad}$$

$$d = D \theta = 500 \cdot 5.55 \cdot 10^{-6} \cancel{\frac{\pi}{180}} = \underline{\underline{4.84 \cdot 10^{-5} \frac{\text{pc}}{\text{yr}}}}$$

$$= 4.75 \cdot 10^9 \frac{\text{m}}{\text{s}}$$

THIS IS LOWER LIMIT.

REDSHIFT MEASUREMENTS OF CLUSTERS & STARS  
WOULD MAKE THIS MORE ACCURATE, BY ALLOWING  
US TO CALCULATE PERPENDICULAR-TO-SKY  
VELOCITY COMPONENT.



$$\frac{GM_1 M_2}{d^2} = \frac{v_1^2 M_1}{d(M_1, \text{CoM})} = \frac{v_2^2 M_2}{d(M_2, \text{CoM})}$$

$$d(M_i, \text{CoM}) = \frac{M_j}{M_i + M_j} d$$

$$\frac{GM_1M_2}{a^2} = \frac{v_1^2 M_1}{\frac{M_1 M_2}{M_1 + M_2} a} = \frac{v_2^2 M_2}{\frac{M_1}{M_1 + M_2} a}$$

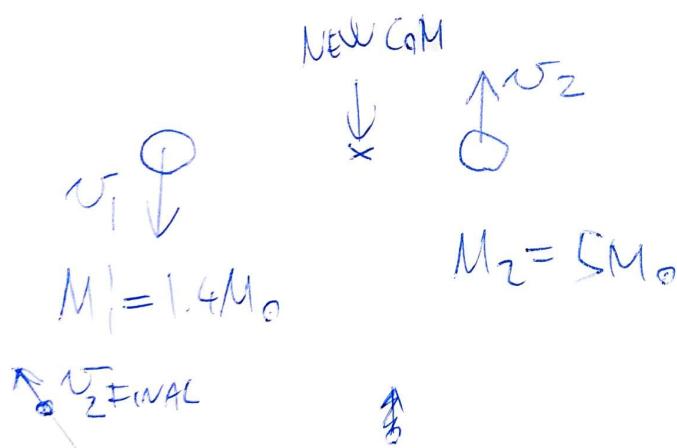
$$= v_1^2 \frac{M_1}{M_2} (M_1 + M_2) \frac{1}{a} = v_2^2 \frac{M_2}{M_1} (M_1 + M_2) \frac{1}{a}$$

↓

$$v_1^2 = \frac{GM_2^2}{M_1 + M_2} \frac{1}{a} \left| \begin{array}{l} M_1 = 20M_\odot \\ M_2 = 5M_\odot \end{array} \right| = \frac{G}{a} M_\odot$$

$$v_2^2 = \frac{GM_1^2}{M_1 + M_2} \frac{1}{a} \left| \begin{array}{l} M_1 = 20M_\odot \\ M_2 = 5M_\odot \end{array} \right| = \frac{G}{a} 16M_\odot$$

AFTER EXPLOSION:



ENERGY CONSERVATION:

$\frac{1}{2} M_1 v_1^2$

$$\frac{1}{2} M_1 v_1^2 + \frac{1}{2} M_2 v_2^2 = \frac{1}{2} M_1 v_{\text{final}}^2$$

$\downarrow$

$v_{\text{final}}$

## ENERGY CONSERVATION:

$$KE + PE \Big|_{\text{EXPLOSION}} = KE + PE \Big|_{\text{LATER}}$$

$$PE \Big|_{\text{EXPLOSION}} = \frac{GM_1 M_2}{a} \quad PE \Big|_{\text{LATER}} \approx 0$$

$$\frac{1}{2} M_1 v_1^2 + \frac{1}{2} M_2 v_2^2 + \frac{GM_1 M_2}{a} = \frac{1}{2} M_1 v_1^{\text{FINAL}} + \frac{1}{2} M_2 v_2^{\text{FINAL}}$$

## CONSERVATION OF LINEAR MOMENTUM

$$M_1 v_1^{\text{FINAL}} = M_2 v_2^{\text{FINAL}} \Rightarrow v_2^{\text{FINAL}} = \frac{M_1}{M_2} v_1^{\text{FINAL}}$$

SUB IN FOR  $v_1$  &  $v_2$  in  $M_1, M_2$

$$v_1^{\text{FINAL}} = \frac{M_2}{m_1} v_2^{\text{FINAL}}$$

$$\frac{1}{2} (1.4M_\odot) \frac{GM_\odot}{a} + \frac{1}{2} (5M_\odot) \frac{GM_\odot}{a} + \frac{G 1.4 \cdot 5 M_\odot^2}{a} =$$
 ~~$\frac{1}{2} 1.4M_\odot \frac{GM_\odot}{(5M_\odot)} \frac{1.4M_\odot}{5M_\odot}$~~

$$= \frac{1}{2} (1.4M_\odot) \left( \frac{5M_\odot}{1.4M_\odot} v_2^{\text{FINAL}} \right)^2 + \frac{1}{2} (5M_\odot) v_2^{\text{FINAL}}^2$$

~~$47.7 \frac{GM_\odot}{a} = 11.43 M_\odot v_2^{\text{FINAL}}^2$~~

$$4.173 \frac{GM_\odot}{v_2^{\text{FINAL}}^2} = a$$

CHECK UNITS:  $\frac{N \cdot m}{(\frac{kg}{s})^2} = \frac{\frac{kg \cdot m}{s^2} \frac{1}{m^2}}{(\frac{m}{s})^2} = m$  GOOD.

$$d = 4.173 \cdot \frac{G \cdot 2 \cdot 10^{30}}{(4.75 \cdot 10^9)^2} = 2.47 \cdot 10^{11}$$

$\approx \underline{\underline{2.5 \cdot 10^{11} \text{ m}}}$

$\approx 1.6 \text{ AU}$

$$N_1^{\text{FINAL}} = \frac{M_2}{M_1'} N_2^{\text{FINAL}}$$

$$\boxed{L} = \frac{5}{1.4} \approx 3.6$$

$$N_{\text{NS}} = 3.6 \times N_{\text{SECONDARY}}$$

I WOULD COOK 3.6 ARCMIN IN THE  
OPPOSITE DIRECTION.