

Q1.1

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + \Lambda g_{\mu\nu} = -8\pi G T_{\mu\nu}$$

 $\cdot g^{\mu\nu}$

WE WANT: GET RID OF R
PLAN: CONTRACT W/ g , EXPRESS R,
SUBSTITUTE BACK

$$R_{\mu\nu} g^{\mu\nu} - \frac{1}{2} g_{\mu\nu} g^{\mu\nu} R + \Lambda g_{\mu\nu} g^{\mu\nu} =$$

$$= -8\pi G T_{\mu\nu} g^{\mu\nu}$$

$$R_{\mu}^{\mu} - \frac{1}{2} \underbrace{g_{\mu}^{\mu}}_4 R + \Lambda \underbrace{g_{\mu}^{\mu}}_4 = -8\pi G T_{\mu}^{\mu}$$

$$R - 2R + \Lambda \cdot 4 = -8\pi G T_{\mu}^{\mu}$$

$$R - 4\Lambda = 8\pi G T_{\mu}^{\mu}$$

$$R = 8\pi G T_{\sigma}^{\sigma} + 4\Lambda$$

SUBSTITUTE BACK:

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} (8\pi G T_{\sigma}^{\sigma} + 4\Lambda) + \Lambda g_{\mu\nu} =$$

$$= -8\pi G T_{\mu\nu}$$

REARRANGE:

$$R_{\mu\nu} - \Lambda g_{\mu\nu} = -8\pi G \left(T_{\mu\nu} - \frac{1}{2} g_{\mu\nu} T_{\sigma}^{\sigma} \right)$$

AS REQUIRED.

Q1.2

$$\text{FRZ: } \left(\frac{\dot{R}}{R} \right)^2 + \frac{K}{R^2} = \frac{8\pi G}{3} S \left(+ \frac{1}{3} \right)$$

NOTING THAT:

$$\dot{R} = \frac{dR}{dt} = \frac{dR}{R \frac{1}{R} dt} = \frac{1}{R} \frac{dR}{dt} = H$$

REWRITE FRZ:

$$\frac{H^2}{R^2} + \frac{K}{R^2} = \frac{8\pi G}{3} S$$

$$H^2 + K = \frac{8\pi G}{3} S R^2 \quad /: H^2 \quad /-1$$

$$\frac{K}{H^2} = \frac{8\pi G}{3} S \frac{R^2}{H^2} - 1$$

$= \frac{S}{S_{\text{crit}}} = \Omega$

$$\underline{\underline{\frac{K}{H^2} = \Omega - 1}}$$

WE'RE OFF BY A c^2 BUT
THAT PROBABLY HAS TO DO
WITH HOW WE DEFINE K .

Q1.2

FRI:

$$\frac{\ddot{R}}{R} = -\frac{4\pi G}{3}(S+3P) + \frac{1}{3}$$

NOTING THAT:

$$\frac{d^2 R}{dt^2} = \frac{d}{dt} \frac{d}{dt} R = \frac{d}{R dt} \left(\frac{d}{dt} R \right) = \frac{1}{R} \frac{d}{dt} H$$

REWRITE FRI:

$$\frac{1}{R^2} \frac{dH}{dt} = -\frac{4\pi G}{3}(S+3P) \quad / \cdot 2R^2$$

$$2 \frac{dH}{dt} = -\frac{8\pi G R^2}{3} (S+3\omega S c^2)$$

$$= -\frac{8\pi G R^2}{3} S (1+3\omega c^2)$$

$$= -\Omega H^2 (1+3\omega c^2)$$

USING:

$$\frac{R c^2}{H^2} = \Omega - 1$$

REWRITE:

$$= -\left(\frac{R c^2}{H^2} + 1\right) H^2 (1+3\omega c^2)$$

$$\Rightarrow 2 \frac{dH}{dt} = -\left(H^2 + R c^2\right) (1+3\omega c^2)$$

APART FROM
C² FACTORS,
WE'RE GOOD

Q1.2

USE THE RESULTS I WAS SUPPOSED TO GET:

$$\frac{\Omega c^2}{H^2} = \Omega - 1 \quad 2 \frac{dH}{dt} = -(3\omega + 1)(H^2 + \Omega c^2)$$

$$\frac{d\Omega}{dt} = \frac{d}{dt} \left(\frac{\Omega c^2}{H^2} + 1 \right) = \Omega c^2 \frac{d}{dt} \frac{1}{H^2} = \Omega c^2 \cdot -2 \frac{1}{H^3} \cdot \frac{dH}{dt}$$

$$= \Omega c^2 \cdot -2 \frac{1}{H^3} \cdot \frac{1}{2} \cdot -(3\omega + 1)(H^2 + \Omega c^2)$$

$$= \Omega c^2 \frac{1}{H} (3\omega + 1) \underbrace{\left(1 + \frac{\Omega c^2}{H^2} \right)}_{\Omega}$$

$$= (3\omega + 1) H \Omega (\Omega - 1)$$

AS REQUIRED.

$\omega = 0$ CASE:

$$\frac{d\Omega}{dt} = H \Omega (\Omega - 1) \quad \& \quad \frac{dH}{dt} = -\frac{1}{2}(H^2 + \Omega c^2)$$

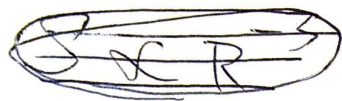
$$\Omega = \int_0^{\Omega_F} d\Omega = \left[\left(\frac{1}{3} \Omega^3 - \frac{1}{2} \Omega^2 \right) H \right]_0^{\Omega_F} - \left[\left(\frac{1}{3} \Omega^3 - \frac{1}{2} \Omega^2 \right) \cdot -\frac{1}{2}(H^2 + \Omega c^2) \right]_0^{\Omega_F}$$

~~I DON'T SEE HOW TO SKETCH THIS WITHOUT KNOWING~~

Q1.3

FRI: $\frac{\ddot{R}}{R} = -\frac{4\pi G}{3}(\rho + 3p) + \frac{1}{3}$

MATTER DOMINATED UNIVERSE $\Rightarrow P=0$



$$2 \cdot q_0 = 2 \cdot \frac{1}{H_0^2} \left(-\frac{1}{R} \frac{d^2 R}{dt^2} \right)$$

$$= 2 \cdot \frac{1}{H_0^2} \frac{4\pi \rho G}{3} (S)$$

$$= \frac{8\pi G \rho S}{3H_0^2} = \frac{S}{S_{\text{CRIT}}} \frac{1}{R^2} = \frac{\Omega_0}{R^2}$$

FRI FOR MATTER
DOMINATED
UNIVERSE

THIS IS RATHER UNFORTUNATE.

AFTER TALKING TO A FRIEND I REALIZE
THAT I AM CONFUSING CURLY H WITH HUBBLE H.

THIS IS
TRUE: $\frac{8\pi G R^2}{3H_0^2} = S_{\text{CRIT}}$

THIS IS
NOT: $\frac{8\pi G R^2}{8H_0^2} = S_{\text{CRIT}}$

$$H = \frac{dR}{dt R} = \frac{1}{R} \cdot \frac{d}{dt} R = \frac{1}{R^2} \frac{dR}{dt} R = \frac{1}{R} \mathcal{H}$$

$$\Rightarrow \frac{8\pi G}{3H^2} = S_{\text{CRIT}}$$

SO, ARMED WITH THIS NEW RELATION, TRY AGAIN:

$$\frac{8\pi G S}{3H_0^2} = \frac{8\pi G S R^2}{3H_0^2} = \frac{S}{S_{\text{CRIT}}} = \Omega_0$$

$$\Rightarrow 2q_0 = \Omega_0 \text{ AS REQUIRED.}$$

~~FRI: $\left(\frac{\dot{R}}{R}\right)^2 + \frac{K}{R^2} = \frac{8\pi G}{3} S + \frac{1}{3}$~~

~~$\rightarrow \frac{\ddot{R}}{R}$ FROM FRI~~

~~$$\left(\frac{\dot{R}}{R}\right)^2 = -2\frac{\ddot{R}}{R} - \frac{K}{R^2}$$~~

~~$$= -2q_0 H_0^2$$~~

~~$$\left(\frac{\dot{R}}{R}\right)^2 = +2q_0 H_0^2 - \frac{K}{R^2}$$~~

~~$$\left(\frac{\dot{R}}{R_0}\right)^2 = 2q_0 H_0^2 \frac{R^2}{R_0^2} - \frac{K}{R_0^2}$$~~

~~WRITE K AS: $K = \left[\frac{8\pi G}{3} S_0 - H_0^2 \right] R_0^2$~~

$$q_0 = \frac{1}{H_0^2} \left(-\frac{1}{R} \frac{d^2 R}{dt^2} \right)$$

\Downarrow

$$\frac{\ddot{R}}{R} = -q_0 H_0^2$$

$$\therefore R_0^2 \cdot R^2$$

Q 1.3

$$H_0^2 \cdot \left(1 - 2q_0 + 2q_0 \frac{R_0}{R}\right) \quad \text{DEF OF } q_0$$

$$= H_0^2 + 2 \frac{1}{R_0} \frac{d^2 R}{dt^2} \Big|_{t_0} - 2 \frac{1}{R_0} \frac{d^2 R}{dt^2} \Big|_{t_0} \frac{R_0}{R} \quad \text{/ REARRANGE}$$

$$= H_0^2 + 2 \left(\frac{1}{R_0} - \frac{1}{R} \right) \frac{d^2 R}{dt^2} \Big|_{t_0} \quad \text{USE FRI}$$

$$= H_0^2 + 2 \left(\frac{1}{R_0} - \frac{1}{R} \right) \left(-\frac{4\pi G}{3} S_0 R_0 \right) \quad \text{/ REARRANGE}$$

$$= H_0^2 - \frac{8\pi G}{3} S_0 + \frac{8\pi G}{3} S_0 \frac{R_0}{R} \quad \text{/ USE EXPRESSION FOR K (LEC1 SLIDE 23)}$$

$$= -\frac{K}{R_0^2} + \frac{8\pi G}{3} S_0 \frac{R_0}{R} \quad \text{/ REWRITE}$$

$$= \frac{1}{R_0^2} \left(-K + \frac{8\pi G}{3} S_0 \frac{R_0^3}{R} \right) \quad \text{/ LEC1 SLIDE 21}$$

$$= \frac{\dot{R}^2}{R_0^2} \quad \text{AS WANTED.}$$

Q1.4)

$$t_0 = \frac{1}{H_0} \int_0^{R_0} \frac{dR}{R E(z)}$$

$$E(z) = \sqrt{\Omega_M (1+z)^3 + \Omega_1}$$

/ SUB. IN

$$= \frac{1}{H_0} \int_0^{R_0} \frac{dR}{R \sqrt{\Omega_M (1+z)^3 + \Omega_1}}$$

/ REARRANGE,
AIMING FOR A
FORM SUGGESTED

$$= \frac{1}{H_0} \int_0^{R_0} \frac{dR}{R \Omega_1^{1/2} \sqrt{\frac{\Omega_M}{\Omega_1} (1+z)^3 + 1}}$$

$$\text{LET: } \frac{\Omega_M}{\Omega_1} (1+z)^3 = \tan^2 \theta$$

$$\text{USE: } \tan^2 \theta + 1 = \sec^2 \theta$$

$$= \frac{1}{H_0} \int_0^{R_0} \frac{\cos \theta dR}{\Omega_1^{1/2} R}$$

$$\text{USE: } R = \frac{R_0}{1+z} = \frac{R_0}{\left(\frac{\Omega_1}{\Omega_M}\right)^{1/3} \tan^{2/3} \theta}$$

$$= \frac{1}{H_0} \int_0^{R_0} \frac{\cos \theta \left(\frac{\Omega_1}{\Omega_M}\right)^{1/3} \tan^{2/3} \theta \frac{1}{R_0} dR}{\Omega_1^{1/2}}$$

GET dR :

$$\frac{dR}{d\theta} = \frac{R_0}{\left(\frac{\Omega_1}{\Omega_M}\right)^{\frac{1}{2}}} \frac{d}{d\theta} \left(\tan^{-\frac{2}{3}} \theta \right)$$

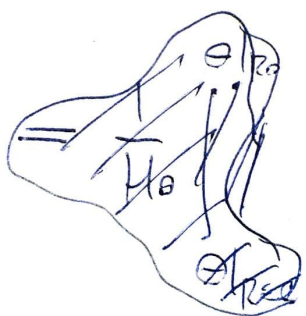
$$= \frac{R_0}{\left(\frac{\Omega_1}{\Omega_M}\right)^{\frac{1}{2}}} \cdot -\frac{2}{3} (\tan \theta)^{-\frac{5}{3}} \cdot \sec^2 \theta$$

SUBSTITUTE THIS INTO THE INTEGRAL:

$$= \frac{1}{H_0} \int_{\theta|_{R=0}}^{\theta|_{R_0}} \frac{\cos \theta \left(\frac{\Omega_1}{\Omega_M}\right)^{\frac{1}{3}} \tan^{\frac{2}{3}} \theta \cdot \frac{1}{R_0} \cdot \frac{R_0}{\left(\frac{\Omega_1}{\Omega_M}\right)^{\frac{1}{3}}} \cdot -\frac{2}{3} \tan^{-\frac{5}{3}} \theta \sec^2 \theta d\theta}{\Omega_1^{\frac{1}{2}}}$$

$$= \frac{1}{H_0} \int_{\theta|_{R=0}}^{\theta|_{R_0}} \frac{\cos \theta \cot \theta \sec^2 \theta \cdot \left(-\frac{2}{3}\right) d\theta}{\Omega_1^{\frac{1}{2}}}$$

$$= \frac{1}{H_0} \int_{\theta|_{R=0}}^{\theta|_{R_0}} -\frac{2}{3} \frac{1}{\Omega_1^{\frac{1}{2}}} \frac{1}{\sin \theta} d\theta$$



$$= \frac{1}{H_0} \cdot -\frac{2}{3} \cdot \frac{1}{\Omega_1^{\frac{1}{2}}} \cdot \left[\ln \frac{\cos \theta + 1}{\sin \theta} \right]_{\theta|_{R=0}}^{\theta|_{R_0}}$$

EVALUATE THE LIMITS:

$$R=0 \Rightarrow \lim_{R \rightarrow 0} z = \infty \Rightarrow \theta = \frac{\pi}{2}$$

$$\ln \left. \frac{\cos \theta + 1}{\sin \theta} \right|_{\theta = \frac{\pi}{2}} = 0$$

WE WANT AGE AT z :

$$\text{i.e. when } \tan \theta = \left(\frac{\Omega_m}{\Omega_\Lambda} \right)^{\frac{1}{2}} (1+z)^{\frac{3}{2}}$$

$$\cancel{t(z)} = \frac{1}{H_0} \frac{2}{3} \frac{1}{\sqrt{\Omega_\Lambda}} \ln \left. \frac{1 + \cos \theta}{\sin \theta} \right|_{\theta = \arctan \left(\sqrt{\frac{\Omega_m}{\Omega_\Lambda}} (1+z)^{\frac{3}{2}} \right)}$$

AS REQUIRED.

$$t(z) = \underline{\underline{1.4 \cdot 10^{10} \text{ yrs}}}$$

$$z=0$$

$$H_0 = 67.5 \frac{\text{km}}{\text{Mpc sec}} = 3.23 \cdot 10^{-18} \frac{1}{\text{sec}}$$

$$\Omega_m = 0.31$$

$$\Omega_\Lambda = 1 - 0.31$$

CONSISTENT
WITH LITERATURE.

1.5

MANAGED TO GET DONE ONLY THE VERY LAST BIT.

$$\frac{dS}{dR} = -3(1+\omega) \frac{S}{R}$$

$$\frac{dS}{S} = -3(1+\omega) \frac{dR}{R}$$

$$\ln S = -3(1+\omega) \ln R + C$$

$$= \ln R^{-3(1+\omega)} + \ln C$$

$$= \ln \left(C \cdot R^{-3(1+\omega)} \right)$$

$$\Rightarrow \underline{\underline{S \propto R^{-3(1+\omega)}}}$$

ATTEMPTS FOR FIRST PART (DIDN'T WORK)

$$ds^2 = c^2 dt^2 - R^2(t) \frac{dr^2}{1 - kr^2}$$

$$\frac{ds}{dt} = c$$

↳ PROPER TIME OF PARTICLE

$$\text{[scribbled out]} = c^2 dt^2 - R^2(t) dx^2$$

AND I DON'T KNOW WHAT I DO WITH THESE.

HOWEVER, IN THE SECOND PART, CAN USE:

$$g_{\mu\nu} u^\mu u^\nu = \text{CONSTANT} = c^2$$

$$u^\mu = \gamma_v (c, v, 0, 0)$$

USING METRIC:

$$ds^2 = c^2 dt^2 - R^2(t) dx^2$$

$$g_{\mu\nu} u^\mu u^\nu = \text{[scribbled out]} = c^2 - R^2 v^2$$

$$= g_{00} u^0 u^0 + g_{11} u^1 u^1$$

$$= c^2 c^2 \gamma_v^2 - R^2 \gamma_v v \gamma_v v$$

$$= c^4 \gamma_v^2 - R^2 \gamma_v^2 v^2$$

$$= \gamma_v^2 (c^4 - R^2 v^2) = \text{CONSTANT}$$

$$\sigma_{v_1}^2 (c^4 - R^2(t_1) v_1^2) = \sigma_{v_2}^2 (c^4 - R^2(t_2) v_2^2)$$

WE SHOULDN'T HAVE THE c^4 TERMS.

IF WE DIDN'T HAVE THEM, WE'D HAVE:

$$\sigma_{v_1}^2 R(t_1) v_1 = \sigma_{v_2}^2 R(t_2) v_2$$

\Downarrow

$$\frac{R(t_1)}{R(t_2)} = \frac{\sigma_{v_2}^2 v_2}{\sigma_{v_1}^2 v_1}$$

BUT WE DO HAVE THE c TERMS SO I'M PROBABLY WRONG.