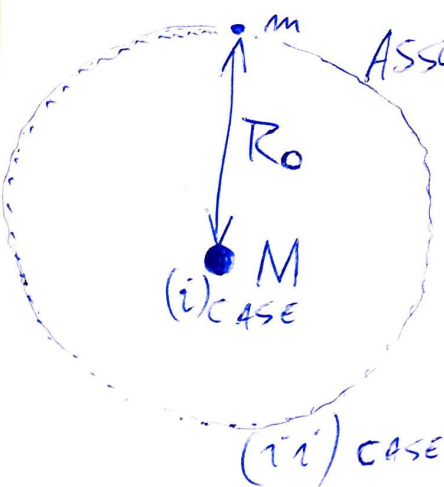


8.



(a)(i) / MINUS SIGN TO INDICATE IT IS FALLING INWARDS.

$$a_o = -\frac{GM}{R_o^2}$$

(a)(ii)

$$a_o = -\frac{GM}{R_o^2}$$

BECAUSE BY GAUSS'S LAW, THIS IS JUST CASE (i)

(b)(i) / $\ddot{R} = -\frac{GM}{R^2}$ / $\cdot \ddot{R}$

$$\ddot{R} \ddot{R} = -\frac{GM}{R^2} \ddot{R} \quad / \text{REARRANGE}$$

$$\Rightarrow \frac{d}{dt} \left[\frac{1}{2} \dot{R}^2 - \frac{GM}{R} \right] = 0 \quad / \text{REMOVE } \frac{d}{dt}$$

$$\frac{1}{2} \dot{R}^2 - \frac{GM}{R} = C \quad / \text{REARRANGE}$$

$$\frac{dR}{dt} = \pm \sqrt{\frac{2GM}{R}} + C$$

~~Wait, what? Why do I have positive \dot{R} , when I want to fall inwards?~~

We want to fall inwards, so we choose the \ominus sign.

$$\frac{dR}{dt} = -\sqrt{\frac{2GM}{R}} + C$$

Initially we're at ~~rest~~ rest, so:

$$\left. \frac{dR}{dt} \right|_{t=0} = 0 \quad \text{if} \quad \left. \frac{dR}{dt} \right|_{R=R_0} = 0$$

Rewrite constant in our equation to reflect this:

$$\frac{dR}{dt} = \cancel{\sqrt{2GM}} \cancel{R} = -\sqrt{2GM} \left(\sqrt{\frac{1}{R} - \frac{1}{R_0}} \right)$$

$$t_{\text{FALL}} = \int_{\substack{\text{rest} \\ t|_{R_0}}}^{\substack{0 \\ t|_{R=0}}} dt = \int_{R_0}^0 \frac{1}{-\sqrt{2GM}} \frac{1}{\sqrt{\frac{1}{R} - \frac{1}{R_0}}} dR$$

SUBSTITUTE: $R = R_0 \sin^2 \theta$

LIMIT GOES AS: $\theta: \frac{\pi}{2} \rightarrow 0$

$$= -\frac{1}{\sqrt{2GM}} \int_{\theta=\frac{\pi}{2}}^{\theta=0} \left(\frac{1}{R_0 \sin^2 \theta} - \frac{1}{R_0} \right)^{-\frac{1}{2}} \underbrace{R_0 2 \sin \theta \cos \theta}_{dR} d\theta$$

$$= -\frac{1}{\sqrt{2GM}} \sqrt{R_0} R_0 2 \int_{\theta=\frac{\pi}{2}}^0 \left(\underbrace{\frac{1}{\sin^2 \theta} - 1}_{\cot^2 \theta} \right)^{-\frac{1}{2}} \sin \theta \cos \theta d\theta$$

$$= -\sqrt{\frac{2R_0^3}{GM}} \int_{\theta=\frac{\pi}{2}}^0 \tan \theta \sin \theta \cos \theta d\theta$$

$$= - \left[\frac{2R_0^3}{GM} \right]_{\theta=\frac{\pi}{2}}^0 \sin^2 \theta = - \left[\frac{2R_0^3}{GM} \right] \left(-\frac{\pi}{4} \right)$$

$$= \frac{\pi}{2} \sqrt{\frac{R_0^3}{2GM}}$$

(b)(ii)

$$\ddot{R} = - \frac{G}{R^2} \underbrace{M_{\text{ENCLOSED}}(R)}_{\substack{\text{MASS OF} \\ \text{VOLUME WITHIN} \\ \text{RADIUS } R}}$$

$$M_{\text{ENCLOSED}}(R) = \cancel{\frac{4}{3}\pi R^3 \rho} \left(\frac{R}{R_0} \right)^3 M$$

REWRITE EQUATION ABOVE:

$$\ddot{R} = - \frac{G}{R^2} \frac{4}{3}\pi R^3 \rho = - \frac{4GM}{3} \frac{R}{R_0^3}$$

$$\Rightarrow \ddot{R} = - \frac{G}{R^2} \left(\frac{R}{R_0} \right)^3 M = - \frac{GM}{R_0^3} R$$

$$\Rightarrow \ddot{R} + \frac{GM}{R_0^3} R = 0$$

THIS IS JUST SHM EQ.

$$\text{FALL TIME} = \frac{\text{QUARTER OF PERIOD}}{4} = \frac{1}{4} \cdot \sqrt{\frac{R_0^3}{GM}} \cdot 2\pi = \frac{\pi}{2} \sqrt{\frac{R_0^3}{GM}}$$

(SAW THESE IN TOPICS LECTURES)

"A cluster consists initially..." PART

$$t = \frac{\pi}{2} \sqrt{\frac{R_0^3}{GM}}$$

MASS OF A GLOBULAR CLUSTER $\sim 10^6 M_\odot$

SIZE \sim few parsecs

$$t = \frac{\pi}{2} \sqrt{\frac{(3 \cdot 10^{16})^3}{G \cdot 10^6 \cdot 2 \cdot 10^{30}}}$$

~~16~~ $\sim 10^{12}$ sec

$\rightarrow \sim 2 \cdot 10^4$ yrs

This is conveniently high
~~too~~ enough to be believable.