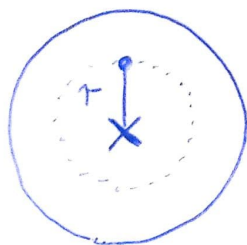


SDSG2

2.1.

POTENTIAL FROM PART OF SPHERE \rightarrow ↑
 NEWTON'S FIRST THM: NO FORCE DUE TO
 EXTERNAL SPHERICALLY
 SYMMETRIC MASS
 DISTRIBUTION



$\Rightarrow \Phi = \text{CONSTANT CONTRIB.}$
 FROM THIS BIT.

POTENTIAL FROM INSIDE SPHERE:

NEWTON SECOND THM: TREAT INSIDE AS
 POINT MASS.

$$f = -\nabla \Phi \Rightarrow \Phi = - \int_{\infty}^r f dr$$

$$= - \int_{\infty}^r \frac{m_{<r'} G}{r'^2} dr'$$

$$= - \int_{\infty}^r \frac{4}{3} \pi r'^3 \rho \frac{1}{r'^2} G dr' = \int_{\infty}^r \frac{4}{3} \pi \rho G r' dr' = \frac{4}{3} \pi \rho G \left[\frac{r'^2}{2} \right]_{\infty}^r$$

$$= \frac{2}{3} \pi \rho G r^2 = \frac{1}{2} \Omega^2 r^2 + C$$

$$\text{WHERE } \Omega^2 = \frac{4}{3} \pi \rho G$$

THE NOTATION

Ω SUGGESTS

THAT $\frac{4}{3} \pi \rho G$
 SHOULD HAVE

DIMENSIONS SEC^{-2}

CHECK: $\sqrt{\frac{\text{kg}}{\text{m}^3}} \sqrt{\text{N} \frac{\text{m}^2}{\text{kg}^2}} = \sqrt{\frac{\text{N}}{\text{m kg}}} = \sqrt{\frac{\text{kg} \frac{\text{m}}{\text{kg}^2 \text{sec}^2}}{\text{m kg}}} = \frac{1}{\text{SEC}} \quad \text{GOOD.}$

THE FACT WHICH WORKS

UNSTABLE IF: $\frac{d}{dr}(R^3 \ell) > 0$

ARISING QUESTIONS

RECALL: FOR CIRCULAR ORBIT:

$$-R\dot{\phi}^2 = \ell$$

• WHEN TAKING SECOND DERIVATIVE OF $V_{\text{EFFECTIVE}}$ DOESN'T WORK? IE

$$V_{\text{EFF}}(r) = U(r) + \frac{L^2}{2r^2}$$

$$\frac{dV_{\text{EFF}}}{dr} = \frac{dU(r)}{dr} + \frac{L}{r^2} L' - L^2 r^{-3}$$

CONSIDER:

$$\frac{dU(r)}{dr} = -f(r) = r\dot{\phi}^2 - \frac{L^2}{r^3} \rightarrow 0 \text{ FOR CIRCULAR}$$

$$\frac{dV_{\text{EFF}}}{dr} = r\dot{\phi}^2 + \frac{L L'}{r^2} - L^2 r^{-3} \rightarrow 0 \text{ BEC } L' = 0$$

ORBIT IS WHERE $\frac{dV_{\text{EFF}}}{dr} = 0$, IE $r\dot{\phi}^2 - L^2 r^{-3} = 0$

$$\frac{dV_{\text{EFF}}}{dr} = \cancel{r\dot{\phi}^2} - L^2 r^{-3} = 0$$

$$\dot{\phi}^2 = \frac{L^2}{r^4}$$

$$L^2 = \dot{\phi}^2 r^4 \quad (\text{AS EXPECTED})$$

OH I SEE. SHOULDN'T THROW OUT THE $\frac{L L'}{r^2}$ TERM BEFORE TAKING SECOND DERIVATIVE.

• ~~IS $\frac{d}{dr}(L^2)$ EVER SATISFIED?~~ IS THE METHOD CORRECT THOUGH IN PRINCIPLE?
 $\frac{d}{dr}(L^2) = 2L L'$. L IS CONSTANT
 SO $L' = 0 \Rightarrow \frac{d}{dr}(L^2) = 0$ IF ANGULAR MOMENTUM IS CONSERVED IE ALWAYS
 SO HOW ARE THERE UNSTABLE ORBITS?

$$\frac{d}{dr}(R^3(-R\dot{\phi}^2)) > 0$$

$$\frac{d}{dr}(R^4 \dot{\phi}^2) < 0$$

RENAME $R \rightarrow r$
 USE: $L = r^2 \dot{\phi}$

$$\frac{d}{dr}(L^2) < 0$$

AS REQUIRED.

RECALL ORBIT EQ.:

$$\frac{d^2 u}{d\phi^2} + u = -\frac{f(r)}{h^2 u^2}$$

$$\text{with } u = \frac{1}{r}$$

$$h = r^2 \dot{\phi}$$

$$u = \frac{1}{r} = a^{-1} \exp(-b\phi)$$

$$\frac{1}{u^2 h^2} = \frac{1}{u^2 \dot{\phi}^2} = r^4 \dot{\phi}^{-2}$$

$$\frac{b^2}{a} \exp(-b\phi) + \frac{1}{a} \exp(-b\phi) = \nabla \Phi u^2 \frac{1}{\dot{\phi}^2}$$

$$\frac{1}{a} (b^2 + 1) \exp(-b\phi) = \nabla \Phi \frac{u^2}{\dot{\phi}^2}$$

$$= \nabla \Phi \frac{1}{\dot{\phi}^2} a^{-2} \exp(-2b\phi)$$

$$\dot{\phi}^2 a (b^2 + 1) \exp(b\phi) = \nabla \Phi$$

$$\frac{h^2}{r^4} \overbrace{a (b^2 + 1) \exp(b\phi)}^r = \nabla \Phi$$

$$h^2 (b^2 + 1) r^{-3} = \nabla \Phi$$

$$\Rightarrow \Phi \propto \frac{1}{r^2}$$

CONSTANT OF
PROPORTIONALITY
CAN BE OBTAINED
IF WE KNOW h
OF THE ORBIT IS
WELL, CORRECT?

2.4/

$$u^2 = 2 \frac{E - \Phi}{h^2}$$

$$\Rightarrow \frac{d^2 \Phi}{du^2} = \frac{d^2}{du^2} \left(E - \frac{1}{2} u^2 h^2 \right) = -h^2$$

$$S(r) > 0 \quad \forall r \Rightarrow \Phi < 0 \quad \forall u$$

(USING: $\Phi(r) = - \int \frac{G S(r') dr'}{\sqrt{|r-r'|}}$)

WANT TO PROVE: THERE ARE 0 OR 2 ROOTS.

IE DISPROVE THAT THERE CAN BE 1 ROOT.

$$1 \text{ ROOT IF: } E - \Phi = 0$$

$$\frac{d^2 \Phi}{du^2} = -h^2 \Rightarrow \frac{d\Phi}{du} \text{ IS DECREASING AS } u \text{ INCREASES}$$

$$\Rightarrow \Phi \text{ IS CONCAVE. (AND THEN WHAT?)}$$

$$E = \underbrace{KE}_{\text{KINETIC ENERGY}} + \Phi \quad \Rightarrow E > \Phi \Rightarrow E - \Phi > 0$$

$$KE > 0$$

(IE PARTICLE IN
CENTRAL FORCE
FIELD DOES NOT
STAY STATIONARY)

$$\Downarrow$$

$$E - \Phi \neq 0$$

\Downarrow
CANNOT BE 1 ROOT

$$E - \Phi > 0 \Rightarrow u^2 = \frac{2}{h^2} (E - \Phi) \text{ MUST HAVE 2 ROOTS.}$$

THIS IS NOT WHAT WE WERE SUPPOSED TO BE PROVING.

SDSGZ

2.5 (I)

$$\nabla^2 \Phi = 4\pi G \rho$$

$$\text{TOTAL MASS} = \int_{\text{ALL SPACE}} \rho dV = \int_{\text{ALL SPACE}} \frac{1}{4\pi G} \nabla^2 \Phi dV$$

$$= \int_{r=0}^{\infty} \int_{\theta=0}^{\pi} \int_{\phi=0}^{2\pi} \left[\frac{1}{4\pi G} \nabla^2 \Phi \right] r^2 \sin \theta d\theta d\phi dr$$

$$= \frac{1}{G} \int_0^{\infty} \nabla^2 \Phi r^2 dr = \frac{1}{G} \int_0^{\infty} \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \Phi}{\partial r} \right) r^2 dr$$

$$= \int_0^{\infty} \frac{\partial}{\partial r} \left[r^2 \frac{\partial}{\partial r} \left(\frac{-M}{b + \sqrt{b^2 + r^2}} \right) \right] dr$$

$$= M \int_0^{\infty} - \frac{\partial}{\partial r} \left[r^2 \frac{\partial}{\partial r} \left(\frac{1}{b + \sqrt{b^2 + r^2}} \right) \right] dr$$

— TRUST OURSELVES THAT WE'RE GOOD IN SETTING UP INTEGRALS

OR

— BRUTE FORCE ALGEBRA

OR

— PUT IT INTO COMPUTER
(THIS WAS MY CHOICE)

ARRIVE AT

$$= M$$

SDSG2
2.5(II)

$$S(0) = \frac{1}{4\pi G} \nabla^2 \Phi \Big|_{r=0}$$

$$= \frac{1}{4\pi G} \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \left(\frac{-M}{b + \sqrt{b^2 + r^2}} \right) \right) \Big|_{r=0}$$

①

$$= \frac{3M}{16\pi b^3}$$

[DONE BY COMPUTER,
RESULT CAN BE VERIFIED ON DESMOS:
[DESMOS.COM/CALCULATOR/ECNFBTM4MC](https://www.desmos.com/calculator/ECNFBTM4MC)]

$$S(r \gg b) = \frac{1}{4\pi G} \nabla^2 \Phi \Big|_{r \gg b}$$

using: $\frac{1}{b + \sqrt{b^2 + r^2}} \approx \frac{1}{2r}$ WHEN $r \gg b$

$$\approx \frac{1}{4\pi} \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \left(\frac{-M}{r} \right) \right) \Big|_{r \gg b}$$

$$= \frac{1}{4\pi} \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \left(\frac{-M}{r} \right) \right)$$

$$= \frac{bM}{2\pi r^4}$$

[VERIFIED BY DESMOS]
(LINK ABOVE ↑)

SDSG2

2.7.

(I)

HANDOUT 3 SLIDE 22: $K^2 = \frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} + \frac{3L_z^2}{R_c^4}$

EQUATION 3.11 $\Rightarrow \frac{L_z^2}{R_c^4} = \dot{\Phi}^2 = \Omega^2$

~~$\Rightarrow K^2 = 3\Omega_0^2$~~ $\Rightarrow K^2 = 3\Omega_0^2 + \frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)}$

HANDOUT 3 SLIDE 22: $V^2 = \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(R_c, 0)}$

~~$K^2 = 2\frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)}$~~

$$K^2 + V^2 = \frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} + \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(R_c, 0)} + \frac{3L_z^2}{R_c^4}$$

$$= \frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} + \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(R_c, 0)} + 3\Omega_0^2$$

THIS IS SUPPOSED TO BE EQUAL TO:

$= 2\Omega_0^2 \Rightarrow$ WE WANT TO PROVE:

$$\frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} + \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(R_c, 0)} = -\Omega_0^2$$

~~$\Phi(R, z) = \Phi(R, -z) \Rightarrow \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(R_c, 0)} = 0$~~

WE HAVE:

~~$\frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} = -\Omega_0^2$~~

~~$\frac{\partial^2 \Phi}{\partial R^2} \Big|_{(R_c, 0)} = -\Omega_0^2 R + C$~~

SDSG 2
2.7(±)

HANDOUT 2, EQUATION 2.3:

$$\nabla^2 \Phi(r) = 4\pi G \rho(r) \stackrel{=}{=} 0$$

↑
PROBLEM
TELLS US THAT
 $\rho(r) = 0$

IN
CRC:

$$\nabla^2 \Phi = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \Phi}{\partial r} \right) + \frac{\partial^2 \Phi}{\partial z^2} = 0$$

$$\Rightarrow -\frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \Phi}{\partial r} \right) = \frac{\partial^2 \Phi}{\partial z^2}$$

RECALL, FOR CIRCULAR ORBITS:

$$\ddot{r} - r \dot{\phi}^2 \Big|_{r=\text{const}} = -r_c \Omega_0^2 = -\frac{\partial \Phi}{\partial r} \Big|_{(r_c, 0)}$$

SUBSTITUTE:

$$-\frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \Phi}{\partial r} \right) \Big|_{r=r_c} = -\frac{1}{r} \frac{\partial}{\partial r} \left(r^2 \Omega_0^2 \right) \Big|_{(r_c, 0)}$$

$$= -\frac{1}{r} \left(2r \Omega_0^2 \right) \Big|_{(r_c, 0)} = -2 \Omega_0^2$$

$$\Rightarrow \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(r_c, 0)} = -2 \Omega_0^2 \Big|_{(r_c, 0)} = -2 \Omega_0^2$$

WE WANTED TO PROVE:

$$\frac{\partial^2 \Phi}{\partial r^2} \Big|_{(r_c, 0)} + \frac{\partial^2 \Phi}{\partial z^2} \Big|_{(r_c, 0)} = -\Omega_0^2$$

+ TRANSFORMED IT INTO:

$$\frac{\partial^2 \Phi}{\partial r^2} \Big|_{(r_c, 0)} - 2 \Omega_0^2 = -\Omega_0^2$$

SDSGZ
2.7 III

$$\Rightarrow \partial_R^2 \Phi \Big|_{(R_c, 0)} = \Omega_0^2$$

THIS IS WHAT WE WANT TO PROVE NOW.

$$\int dR$$

$$\partial_R \Phi \Big|_{(R_c, 0)} = \Omega_0^2 R + C = -f$$

FORCE ON
PARTICLE.

AT $R=0, f=0$ (SYMMETRY) $\Rightarrow C=0$

$$\Rightarrow \partial_R \Phi \Big|_{(R_c, 0)} = \Omega_0^2 R$$

§ CONSIDERING THAT THE EQUATION:

$$\ddot{R} - R \dot{\Phi}^2 \Big|_{R=R_c} = - \frac{\partial \Phi}{\partial R} \Big|_{R=R_c}$$

IN OUR CASE IS:

$$-R \dot{\Phi}^2 \Big|_{\substack{R=R_c \\ z=0}} = - \partial_R \Phi \Big|_{\substack{R=R_c \\ z=0}}$$

↓

$$R \Omega_0^2 = \partial_R \Phi \Big|_{(R_c, 0)}$$

AS DESIRED.

~~SO WHAT WE WANTED TO BE IS TRUE~~

$$\Rightarrow k^2 + v^2 = 2\Omega_0^2$$