

COSMOLOGY LIII PART I

CMB discovery: 1965

killed steady state proposals
proved hot big bang cosmology beyond any reasonable doubt

Penzias & Wilson

much bigger than expected noise
isotropic \Rightarrow cosmic origin

Another group predicted ~ 3 K background radiation.

They couldn't measure the spectrum properly
in 1965.

COBE satellite 1990: measured spectrum.

Found very nice black body radiation curve.
can estimate temperature from this

$$T \approx 2.7 \text{ K}$$

Black body radiation:

$$S_{\text{BB}} c^2 = \underbrace{\int_0^\infty \frac{x^3 dx}{e^x - 1}}_{\substack{\text{Riemann} \\ \text{Zeta} \\ \text{function}}} \quad \Gamma(4) \zeta(4) = \frac{\pi^4}{15}$$

early times were in thermal equilibrium then

$$\boxed{S_{\text{BB}} c^2 = a T^4} \quad a = \frac{8\pi^5 \hbar^9}{150^3 c^3}$$

black body radiation exerts isotropic pressure:

$$P_{\text{r}} = \frac{1}{3} S_{\text{BB}} c^2$$

Energy conservation: $\frac{d(SR^3)}{dt} = -3PR^2$

$$\Rightarrow S_g \propto 1/R^4 \quad S_g c^2 = a T_g^4$$

note that for dust,
 $S \propto 1/R^3$, unlike
 radiation which is
 considered here

$$T_g \propto 1/R$$

$$S_g(\nu, t) d\nu = \frac{1}{c^3} \frac{8\pi h \nu^3 d\nu}{\exp\left(\frac{h\nu}{kT_g}\right) - 1}$$

$$T_g \propto 1/R, \quad \nu \propto 1/R \text{ (redshift)} \Rightarrow$$

radiation
 maintains
 b. b. form
 as univ.
 expands.

number density of BB photons

$$n_g(\nu) d\nu = \frac{8\pi}{c^3} \frac{\nu^2 d\nu}{\exp\left(\frac{h\nu}{kT_g}\right) - 1}$$

integrate up to

$$n_g = 16\pi^5 \left(\frac{kT_g}{hc} \right)^3 = 410.6 \left(\frac{T_g}{2.726 \text{ K}} \right)^3 \text{ cm}^{-3}$$

critical density is $\sim 10^8 \text{ cm}^{-3}$

\uparrow
 ~ 400 photons per cm^3

$$\eta = \frac{n_b}{n_g} = 2.75 \times 10^8 \Omega_b h^2$$

very many more photons
 than baryons. The smallness of this
 is not yet well understood.

must have been an asymmetry in
matter-antimatter, so after $p^+ \bar{p}$ annihilation
we have left with \bar{p} a bit.

$$\Omega_r(0) = \frac{S_r}{S_c} = 2.4 \cdot 10^{-5} h^{-2}$$

↳ hubble parameter

CMB plays negligible control to E density
of univ. present day.

$$S_{\text{matter}} \propto t^{-3} \quad \& \quad S_r \propto t^{-4}$$

$\Rightarrow E$ density of radiation dominates early times.

Furthermore: $T_r \propto t^{-1} \Rightarrow$ early univ must
have been hot & very
dense.

~~it proves~~

i.e CMB proves
hot big bang
cosmology

• The Issues after CMB discovery

• is it BB \rightarrow CMB

• is it anisotropic?

i.e different directions do have different
temperatures or not

Discoveries of Dipole Anisotropy

Dipole Anisotropy

if $v \ll c$:

$$T'(\theta') \approx T_0 \left[1 + \frac{v}{c} \cos \theta' \right]$$

i.e. motion of observer through the CMB induces dipole anisotropy.

Put radiometer to UZ spy plane, fly over Strile.

This was measured by Planck satellite as well.

Solar dipole, correct for motion of sun,
motion compared to local group centre.

Anisotropy still present.

Questions arising:

CMB defines preferred reference frame. SR violation?
no, preferred ref frame in FRW metric. (3)

What causes this motion?

inhomogeneity of matter around us

COSMOLOGY LIII PII

de Sitter space

recall FRW: max. ly sym space
characterized by 1 number k .
3D

de Sitter space:
max. symmetric 4D spacetime.

curvature tensor

$$R_{\kappa\lambda\sigma\rho} = K(g_{\lambda\sigma}g_{\kappa\rho} - g_{\lambda\rho}g_{\kappa\sigma})$$

Ricci tensor

$$R_{\lambda\sigma} = g^{\kappa\rho} R_{\kappa\lambda\sigma\rho} = K(N-1)g_{\lambda\sigma}$$

Ricci scalar

$$R = KN(N-1)$$

$N=4$ for 4D.

Rewrite Einstein field equations:

$$R_{\mu\nu} - \Lambda g_{\mu\nu} = -8\pi G(T_{\mu\nu} - \frac{1}{2}g_{\mu\nu}T^\sigma_\sigma)$$

for empty universe:

$$R_{\mu\nu} = \Lambda g_{\mu\nu}$$

de Sitter geometry satisfies:

$$R_{\mu\nu} = 3K g_{\mu\nu}$$

de Sitter geometry is

sol of EFE if

there is no matter

$$T_{\mu\nu} = 0 \text{ \& } \Lambda \neq 0$$

$$\Rightarrow K = \frac{1}{3}\Lambda$$

ie curvature of spacetime is related
to cosm. const.

Geometry of de Sitter space?

consider constant curvature hyperboloid embedded in 2+1 dimensional Minkowski space.

$$-z^2 + x^2 + y^2 = H^{-2} \quad (\text{eq of hyperboloid})$$

Similar to FRW: there we embedded sphere

metric:

$$d\tau^2 = dz^2 - dx^2 - dy^2$$

eliminate dz :

$$d\tau^2 = \frac{(x dx + y dy)^2}{x^2 + y^2 - H^{-2}} - dx^2 - dy^2$$

make coordinate transform:

$$\begin{aligned} x &= H^{-1} \cos \chi (H \epsilon) \cos \chi \\ y &= H^{-1} \cos \chi (H \epsilon) \sin \chi \end{aligned}$$

metric reduces to:

$$d\tau^2 = d\epsilon^2 - \frac{1}{H^2} \cos^2 \chi (H \epsilon) d\chi^2$$

recognize form as FRW with 1 spatial coordinate

corresponds to closed universe



entire surface covered by:

$$-\infty < \epsilon < \infty,$$

$$0 \leq \chi \leq 2\pi$$

This universe has minimum size reached at $\epsilon = 0$.

lets make another coord. transform.

[ALGEBRA]

$$\begin{aligned} z &= z(t, \xi) \\ \eta &= \eta(t, \xi) \\ \chi &= \chi(t, \xi) \end{aligned}$$

arrive to:

$$ds^2 = dt^2 - H^{-2} e^{(2Ht)} d\xi^2$$

spatially flat universe expanding exponentially spatially flat, just a line



"time-sliced" hyperboloid in a different way

Now we are equipped to do embedding in 5D Minkowski space.

loads of algebra.

FRW metric with closed spatial sections:

$$ds^2 = dt^2 - R^2 \cos^2(\eta) [d\chi^2 + \sin^2 \chi (d\theta^2 + \sin^2 \theta d\phi^2)]$$

analogue to contracting - expanding universe.
closed spatial sections

can choose
time coord.,
change time

slicing: $ds^2 = dt^2 - R(t)^2 (dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2)$

analogous to top of this page.

$$R \propto \exp(Ht) \quad H = \sqrt{\frac{1}{3}} \quad \kappa = H^2$$

KEY POINTS

if we can time slice resulting in
closed open flat sections
what does it mean to say that our Universe is approaching
the de Sitter solution with flat spatial geometry?

Keep in mind:

de Sitter sol is a vacuum solution

\Rightarrow no unique way of time slicing
(why?)

existence of matter

fixes time slicing.

no arbitrary choice,

matter distrib defines natural time slicing

de Sitter sol is a max. ly symmetric spacetime (i.e. 4D)
with curvature K .

Think of it as a thing in 4D.

~~But~~ de Sitter is important for early & late time
universe.