AMATH 482 HW 2

The Use of Gabor Transforms in Analyzing Time-varying Signals

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Abstract: This paper aims to demonstrate the usefulness of the Gabor transform in extracting frequency information from time-varying signals. As many signals vary in time, this technique has a wide range of applications including image analysis, computational neuroscience, and even the decomposition of whale songs. It will discuss the trade-offs associated to the transform and some of the options there are for defining a sampling window. As a case study, we will use this technique to study a few pieces of music.

1 Introduction and Overview

Many of the signals we come into contact with are time-varying. Often the relative order of these frequencies carries information in and of itself and thus the local frequency content at each point in time is important: take a song or speech, scramble it up, and suddenly it loses all meaning. However, the Fourier Transform only gives us a global measurement of the frequency content, and thus this meaning is lost. The Gabor Transform seeks to improve upon this short-coming of the Fourier Transform. By only looking a specific window of time, it can break the signal down into several pieces, separating the frequency contents at each point in time. As we will see, this comes with a trade-off: we lose frequency resolution.

In this paper, we use this technique to decompose several pieces of music. The first is Handel's Messiah which we will use to study the properties of different Gabor windows. The other two pieces are *Mary Had a Little Lamb* played on a piano and recorder separately. Using what we learned about Gabor windows in the first part, we will try to make inferences about the properties of the given instruments and the sounds they generate. Using this information, we will try to recreate the music score used to produce these signals.

2 Theoretical Background

The Fourier Transform is the primary tool for converting between the spatial and frequency domains of a signal. Below is the transform and its inverse:

$$F(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} e^{-ikx} f(x) dx \tag{1}$$

$$f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} e^{ikx} F(k) dk \tag{2}$$

Where k is the wavenumber. It may not be clear from **Equations 1** or **2** why this encodes frequency information, but expanding e^{-ikx} provides some insight:

$$e^{-ikx} = \cos(kx) - i\sin(kx) \tag{3}$$

Equation 3 makes it clear why then k represents the wavenumber. Thus the transform is essentially taking the inner product of our function with a bunch of periodic sines and cosines where, roughly speaking, frequencies/wavenumbers more similar to our signal get assigned higher values. In this way, spatial information is converted to the frequency domain. However, as mentioned earlier this transformation loses all spatial information in order to get complete frequency resolution, but that information can sometimes be really important. The Gabor Transform is a compromise:

$$\mathcal{G}[f](t,\omega) = \int_{-\infty}^{\infty} f(\tau)g(\tau - t)e^{-i\omega\tau}d\tau \tag{4}$$

Where $g(\tau - t)$ defines a window centered at time t. This window acts as a filter, allowing us to focus our frequency analysis on a specific window in time. As we slide t across the temporal domain, we get a picture of the frequencies at each point in time. While this gives us information on where in time each of these frequencies are occurring (increased time resolution), the width of the window limits the wavelengths of frequencies we can detect. Thus, we lose frequency resolution. However, expand the window and we lose time resolution, and thus there is always this trade-off between spatial and frequency resolution, as dictated by the Heisenberg Uncertainty Principle. Because of this trade-off, it is important we choose an appropriate window. The width of the window should be chosen based on the range of frequencies of interest: we want them all to fit in the window. Some of the potential/common window functions are shown in **Figure 1**. Their equations, with (inverse) width parameter a and centered at t are:

Gaussian:
$$g(\tau-t)=e^{-a(\tau-t)^2}$$

Super Gaussian: $g(\tau-t)=e^{-a(\tau-t)^{10}}$
Mexican Hat Wavelet: $g(\tau-t)=(1-(\tau-t)^2)e^{-a(\tau-t)^2/2}$
Shannon (Step-Function):

$$g(\tau - t) = \begin{cases} 1 & x \in [t - \frac{1}{2a}, t + \frac{1}{2a}] \\ 0 & otherwise \end{cases}$$

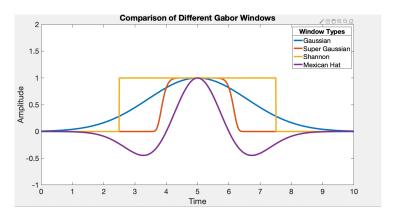


Figure 1: A few different Gabor Windows, each with the same width parameter.

These windows decay at different rates and have their own specific properties, so the window should be carefully chosen to fit the signal and task at hand. When trying to visualize the results of a Gabor Transform, spectrograms are useful tools. These are plots of frequency versus time, where color intensity highlights the prevalence of a given frequency at a specific time.

3 Algorithm Implementation and Development

3.1 Analysis of Gabor Windows and Handel's Messiah

We begin by loading a sample of Handel's music as a simple vector of amplitudes. Using the sampling rate Fs, we can reconstruct the time domain for this signal. We strip a single data point from the signal v to give it an even length. This simplification allows v to be compatible with the wavenumbers we calculate (which have an even length) and has a negligible effect on our overall data. Given the time domain we found, we reconstructed our wavenumbers, shifting them to be centered at zero (Appendix B, code 0-28). Our goal is the use a Gabor Transform to analyze the frequency content of this music as a function of time. First, we define some of the properties of our window: its sampling rate tstep (how much it slides each iteration), its width which is inversely related to a, and its function (i.e. gaussian versus Shannon). Unless otherwise specified, we used a gaussian window with width a = 30 and sampling rate tstep = 0.25 (Appendix B, code 29-52). For each sampling frame in our time domain, we plot the original signal, plot our Gabor window, apply the filter to the signal in the time domain, use the Fourier Transform to convert to the frequency domain, and plot the result in the frequency domain using our pre-calculated wavenumbers (Appendix B, code 53-98). In each sampling frame, we store the resulting frequency information. Lastly, we plot this frequency data at each point of time in a spectrogram (Appendix B, code 100 - 112). We repeat this process several times, varying the properties of the window.

3.2 Exploring Instrumental Differences: Piano vs. Recorder

Again, the signal produced by these instruments is time varying so to understand its time-dependent behavior, we aim to perform a Gabor Transform. After loading in our signal, we began this stage by repeating what we did in the previous: applying a Gabor Transform and playing with the properties of our Gabor window to find what works best (Appendix B, code 120 - 196). In the end, we decided to use a Shannon window with width parameter a=3 and tstep=0.25,. Additionally, at each window sampling frame we store the most prevalent frequencies. We do this by finding the index of the maximum magnitude in our frequency domain and mapping this index onto our wavenumbers. As our signal is in the time domain, these wavenumbers represent angular frequency. Dividing them by 2π converts them to Hz (Appendix B, code 199; 208). Once we find the most prevalent frequency at each time point, we simply find which note this frequency is most closely associated with (Appendix B, code: 212 - 233). To help visualize our results and spot some properties of the instrument (i.e. the overtones) we plot the spectrogram data we collected earlier (Appendix B, code: 237 - 248). We then repeat this entire process all over again for the recorder's signal (Appendix B, code: 256 - 366).

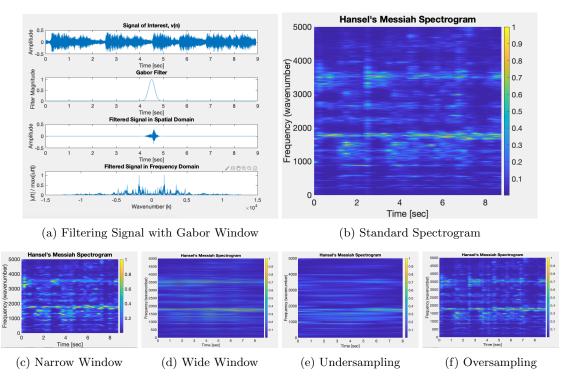


Figure 2: Analyzing effects window properties have on the Gabor Transform with a gaussian filter

4 Computational Results

4.1 Analysis of Gabor Windows and Handel's Messiah

We initially began by exploring how various properties of the window would effect the outcome, displaying these results in **Figure 2**. **2a** shows what a typical analysis would look like: we start with some signal, run the signal through our filter, get a filtered signal, and then take the Fourier Transform of that to get the prevalent frequencies. In **2b**, we show what the spectrogram output of this process would look like with middle of the road parameters. We will use this as a benchmark for comparison. We can see that there are two distinct bands of prominent frequency around wavenumber 1800 and 3500, but there is a fair bit of temporal variation in the frequency content outside of that.

When we narrow the window (a=1000), as we did in **Figure 2c**, we get a really good look at the temporal variation in the frequency content. Normally, we would lose track of our long wavelength signals (low frequency/wavenumber), but here those frequencies do not seem too prominent so that is not much of an issue. We can see there are times, like around 2 and 4 seconds, where the frequency content is much richer than other times. However, this reduction in frequency resolution makes it more difficult to differentiate long-term patterns and to determine individual frequencies (blurred areas of activity versus fine lines). When we widen the window (a=1) in **2d**, we gain frequency resolution at the cost of temporal resolution. This allows us to look for a wide range of frequencies and makes the persistent 1800/3500 wavenumbers more noticeable, but we lose a lot of

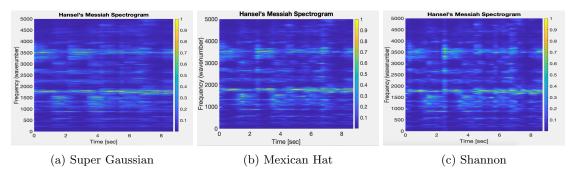


Figure 3: Analyzing the effect of different window functions on the Gabor Transform

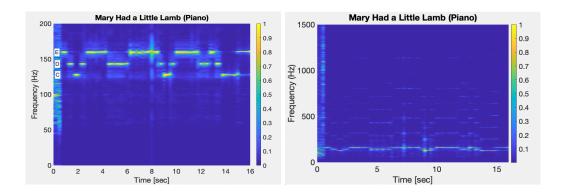


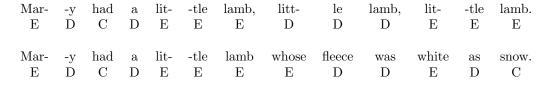
Figure 4: Piano at two different frequency ranges: left shows fundamental frequencies while the right helps show all the related overtones.

the temporal variation we observed earlier which would be important in music. We see this same issue in 2e where we lowered the sampling rate (tstep=2). When we increase the sampling rate (tstep=0.01) in 2f, we regain these areas of rich frequency content and have fairly fine areas of prominent frequencies, but there are also some times where the frequency patterns are very erratic. With such a small timestep, we are more likely to accidentally measure a rest period within the music and that is hardly what we are interested in. Furthermore, this took a considerable amount of time to run.

When we experimented with the window function, there did not seem to be too much of a difference. In **Figure 3**, we can see that the super gaussian window gave a bit clearer of a spectrogram while the Mexican Hat and Shannon windows seem to capture the areas of rich frequency content a bit better. The flanking negative areas of the Mexican Hat window may make it better at picking up changes in frequency content, but in this context all window functions seem to perform adequately.

4.2 Exploring Instrumental Differences: Piano vs. Recorder

We chose to reconstruct the music score for Mary Had a Little Lamb from the piano as we felt it likely had a smaller margin of error: the piano is a bit more precise of an instrument than the recorder. **Figure 4** shows the resulting spectrograms. In the left image, we can clearly see the dominant frequencies jump between 3 distinct levels: E (164.81 Hz), D (146.83 Hz), and C (130.81 Hz) from the top down. Sampling the note every ≈ 0.5 seconds yields the following music score:



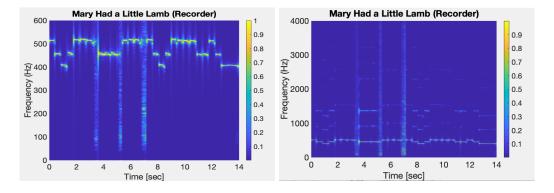


Figure 5: Recorder at two different frequency ranges: left shows fundamental frequencies while the right helps show all the related overtones.

In the right plot of **Figure 4**, we can observe several layers of overtones above the fundamental frequencies mentioned earlier - the faint bands echoing the pattern of the fundamental notes but at higher frequencies. These are a fundamental property of the instrument itself. It appears the overtones are prominent up until around the forth or fifth harmonic before disappearing. In **Figure 5**, we see the same two plots but for the recorder. In the left plot, we see the same relative trend in frequencies, moving between three frequency levels. These frequencies are playing the same three notes – E,D,C – but at a higher octave. However, these frequencies are much choppier. In comparison to the piano, there is much more variation around each of these three frequency levels, blurring the spectrogram. This suggests the recorder is a less precise instrument than the piano: it is harder to maintain a consistent note. Furthermore, observing the right plot of **Figure 5**, we can see the recorder has much fewer overtones. This could be why the noise produced by the recorder is not as rich as that produced by the piano.

5 Summary and Conclusions

The Fourier Transform is a powerful tool for analyzing the frequency content of a signal, but because it maps from the spatial domain to the frequency domain, all spatial information is lost. This is problematic when the signal is time-varying as often the relative order of frequencies is important in and of itself. As a compromise to this, the Gabor Transform is a tool that can be used to gain some spatial resolution, but it comes at the cost of frequency resolution. There is no way to avoid this trade-off between frequency and spatial resolution so it is crucial to carefully consider how to apply the Gabor Transform. Most prominently, it is important to choose a proper window, deciding what window function, width, and sampling rate fits the task at hand best. This paper highlighted the different roles each of these window parameters have. Using these ideas, we were able to successfully reconstruct the music score of Mary Had a Little Lamb from a time-varying signal (a piano). Furthermore, taking the signal from playing this same music score on a recorder, we were able to make inferences about differences between these two instruments and the overtones they produce.

6 Appendix A

Below is a brief summary of the MATLAB functions I used during this project and their functions.

pcolor: Takes a two vectors which define a grid, setting the x and y tick marks, and a 2D matrix. At each point in the grid, it colors the figure based on the corresponding (normalized) value in the 2D matrix. This gives a nice visualization on where the maximum and minimum values of the dataset are occurring. We used it to make the spectrograms.

fftn/ifftn: The former computes the n-dimensional Fourier Transform of the given n-dimensional array. Generates frequency contents in the given n-dimensions. ifftn is its inverse.

fftshift/ifftshift: The former shifts zero frequency component to the center of the frequency spectrum. Works for data in any number of dimensions. Mostly helpful for plotting in the frequency domain, and may not be necessary otherwise. ifftshift is its inverse.

7 Appendix B

```
% AMATH 482 HW 2: Gabor Transforms
   close all; clear all; clc
  load handel
  v = y'/2;
   plot((1:length(v))/Fs,v);
   xlabel('Time [sec]');
   ylabel ('Amplitude');
   title ('Signal of Interest, v(n)');
10
  %p8 = audioplayer(v,Fs);
  %playblocking (p8);
13
14
15
   close all; clc;
16
17
  v = v'/2;
18
  v = v(1:end-1); % remove last point so we have even number of points
19
  L = 9; % tspan(end); length of our signal in time (or space) domain
21
22
  tspan = (1: length(v))/Fs;
23
  n = length(v); % ALWAYS has to be even
25
  % rescale [-pi pi] domain fft assumes to fit our domain [-L, L]
26
  k = 2*pi / (2*L) .* [0:(n/2 - 1) -n/2:-1];
27
  ks = fftshift(k);
28
29
   tstep = 0.25; % how much to slide the window after each measurement to
30
      new measurement
  a = 3; %width parameter of our gabor filter; actually width is prop to
      1/a
32
  % list of values to center filter at; collect frequencies for each
  % one of the windows centered at this point for spectrogram
   tau = 0: tstep: tspan(end);
35
36
   spectrogram_data = zeros(n, length(tau));
38
   figure (1)
39
40
   for j = 1: length(tau)
       % Gaussian Window
42
       \%gabor_filter = exp(-a*(tspan - tau(j)).^2);
43
44
       % Super Gaussian Window
       \%gabor_filter = exp(-a*(tspan - tau(j)).^10);
46
47
       % Mexican Hat Window
       \%gabor_filter = (1-(tspan - tau(j)).^2).*exp(-a*(tspan - tau(j)).^2
50
       % Shannon (Step-function) Window
51
       gabor_filter = abs(tspan - tau(j)) \ll 1/(2*a);
52
53
       % plot our original signal
54
       subplot (411)
```

```
plot((1:length(v))/Fs,v);
56
        ylim ([-0.5, 0.5])
57
        ylabel('Amplitude');
        xlabel('Time [sec]');
         title ('Signal of Interest, v(n)');
60
        set (gca, 'fontsize', 18);
61
62
63
        % Plot our gabor filter
64
        subplot (412)
65
        plot(tspan, gabor_filter);
        \%ylim([-1, 2]) % mexican hat wavelet
        ylim ([0,1.1]) %gaussian limits
68
        x \lim ([0, 9])
69
        title ('Gabor Filter')
70
        ylabel('Filter Magnitude');
        xlabel('Time [sec]');
72
        set (gca, 'fontsize', 18);
73
        % plot our filtered signal
76
        v_filtered = gabor_filter .* v;
77
78
        subplot (413)
79
        plot(tspan, v_filtered);
80
        xlabel('Time [sec]');
81
        ylabel('Amplitude');
         title ('Filtered Signal in Spatial Domain')
83
        ylim ([-0.5, 0.5])
84
        set (gca, 'fontsize', 18);
85
        % plot our filtered signals frequency domain
87
        vft = fft (v_filtered);
89
        subplot (414)
        plot(ks, abs(fftshift(vft)) ./ max(abs(vft)));
91
        xlabel('Wavenumber (k)');
92
        ylabel('|uft| / max(|uft|)');
93
         title ('Filtered Signal in Frequency Domain')
        ylim ([0, 1.2])
95
        set (gca, 'fontsize', 18);
96
        %drawnow
99
        \operatorname{spectrogram\_data}(:, j) = (\operatorname{abs}(\operatorname{fftshift}(\operatorname{vft})) . / \max(\operatorname{abs}(\operatorname{vft}))).
100
101
   end
102
   %%
103
       % make spectogram plot
104
       figure (2)
       pcolor(tau, ks, spectrogram_data), shading interp;
106
       colorbar
107
       y \lim ([0, 5000])
108
       ylabel('Frequency (wavenumber)')
109
       xlabel ('Time [sec]')
110
       title ("Hansel's Messiah Spectrogram")
111
       set (gca, 'fontsize', 25)
112
114
115
```

116

```
118
   % Part 2: Piano
119
    clear all; close all; clc;
121
122
   % LOAD MUSIC
123
     figure (1)
124
     tr_piano=16; % record time in seconds
125
    m1 = audioread('music1.wav').';
126
     Fs1=length(m1)/tr_piano;
127
     plot((1:length(m1))/Fs1,m1);
     xlabel('Time [sec]'); ylabel('Amplitude');
title('Mary had a little lamb (piano)'); drawnow
129
130
    %p8 = audioplayer (m1, Fs1); playblocking (p8);
131
133
    %%
134
135
     tspan1 = (1: length(m1))/Fs1;
     L1 = tr_piano;
137
     n1 = length(m1);
138
     k1 \, = \, 2*\,pi \ / \ (2*L1) \ .* \ [\,0:(\,n1/2\,\, - \,\, 1) \ -n1/2:\, -1\,];
139
     ks1 = fftshift(k1);
140
     tstep1 = 0.25; % how much to slide the window after each measurement
141
         to new measurement
     a1 = 2; %width parameter of our gabor filter; higher a = lower width
142
143
   % list of values to center filter/gabor window at
144
    tau1 = 0: tstep1: tspan1(end);
145
   % stores the max wavenumber for each gabor window (centered at tau(j))
147
   \max_{k} 1 = \operatorname{zeros}(1, \operatorname{length}(\operatorname{tau1}));
148
    spectrogram_data1 = zeros(n1, length(tau1));
149
    for j = 1: length (tau1)
151
        \%gabor_filter1 = exp(-a1*(tspan1 - tau1(j)).^2); \% Gaussian
152
        \%gabor_filter1 = exp(-a1*(tspan1 - tau1(j)).^10); \% Gaussian
153
        % Shannon (Step-function) Window
155
        gabor_filter1 = abs(tspan1 - tau1(j)) \ll 1/(2*a1);
156
157
        % plot our original signal
        subplot (411)
159
        plot ((1:length (m1))/Fs1,m1);
160
        xlabel('Time [sec]');
161
        ylabel ('Amplitude');
162
         title ('Mary Had a little lamb (piano)');
163
164
        % Plot our gabor filter
        subplot (412)
166
        plot(tspan1, gabor_filter1);
167
        \%ylim([-2, 3]) % mexican hat wavelet
168
        ylim ([0,1.1]) %gaussian limits
169
170
        xlim ([0,L1])
171
        xlabel('Time [sec]');
172
         title ('Gabor Filter')
        ylabel ('Filter Magnitude');
174
175
        % plot our filtered signal
176
```

```
m1_filtered = gabor_filter1 .* m1;
178
         subplot (413)
179
         plot(tspan1, m1_filtered);
         xlabel('Time [sec]');
ylabel('Amplitude');
181
182
         title ('Filtered Signal in Spatial Domain')
183
         ylim ([-0.7, 0.7]);
184
         xlim ([0, L1])
185
186
187
        % plot our filtered signals frequency domain
         m1ft = fft (m1\_filtered);
189
190
         subplot (414)
191
         plot(ks1, abs(fftshift(m1ft)) ./ max(abs(m1ft)));
         xlabel('Wavenumber (k)');
193
         ylabel('|uft| / max(|uft|)');
194
         title ('Filtered Signal in Frequency Domain')
195
         x \lim ([-20000, 20000])
197
        % outside "max" call is to filter out the negative wavenumber
198
         \max_{k} k1(j) = \max(k1(abs(m1ft)) = \max(abs(m1ft)));
199
200
         \operatorname{spectrogram\_data1}(:, j) = (\operatorname{abs}(\operatorname{fftshift}(\operatorname{m1ft})) ./ \operatorname{max}(\operatorname{abs}(\operatorname{m1ft})))
201
             . ';
        %drawnow
202
203
    end
204
205
   % convert wavenumber to hertz!
207
    \max_{\text{freqs1}} = \max_{\text{k1}} . / (2*pi);
208
209
   %%
210
211
    notes = ["A", "A#", "B", "C", "C#", "D", "D#", "E", "F", "F#", "G", "G
212
    fund_freqs = [27.5, 29.135, 30.863, 32.703, 34.648, 36.708, 38.891,
        41.203, 43.654, 46.249, 48.99, 51.913;
214
    all_notes = [];
215
    all_freqs = [];
216
217
    for j = 1:8
218
         for note = notes
219
              all\_notes = [all\_notes \ strcat(note, num2str(j))];
220
         end
221
         for freq = fund_freqs
222
             % freqs double each time you move up scale
              all\_freqs = [all\_freqs freq*(2^(j-1))];
224
         end
225
    end
226
227
    music\_score1 = [];
228
229
    for freq = max_freqs1
230
         [min_val, index] = min(abs(freq - all_freqs));
         music_score1 = [music_score1 all_notes(index)];
232
    end
233
234
```

```
% make spectogram plot for Piano!
236
   close all; clc;
237
   figure (3)
239
   % divide by 2pi to convert wavenumber to Hz
240
   pcolor(tau1, ks1./(2*pi), spectrogram_data1), shading interp;
   ylim([0, 1000]) \% to see overtones
   \%ylim([0, 200]) % to see base notes
244
   ylabel ('Frequency (Hz)')
   xlabel('Time [sec]')
   title ('Mary Had a Little Lamb (Piano)')
247
   set (gca, 'fontsize', 25);
248
249
251
   % Part 2: Recorder
252
   clearvars -except all_notes all_freqs; close all; clc;
253
255
    figure (2)
256
    tr_rec=14; % record time in seconds
257
    m2=audioread('music2.wav').';
258
    Fs2=length(m2)/tr_rec;
259
    plot((1:length(m2))/Fs2,m2);
260
    xlabel('Time [sec]'); ylabel('Amplitude');
     title ('Mary had a little lamb (recorder)');
262
    p8 = audioplayer (m2, Fs2); playblocking (p8);
263
264
    %%
265
266
    tspan2 = (1: length(m2))/Fs2;
267
    L2 = tr_rec:
268
    n2 = length(m2);
    k2 = 2*pi / (2*L2) .* [0:(n2/2 - 1) -n2/2:-1];
270
    ks2 = fftshift(k2);
271
    tstep2 = 0.5; % how much to slide the window after each measurement to
272
         new measurement
    a2 = 2; %width parameter of our gabor filter
273
274
   % list of values to center filter at; collect frequencies for each
275
   % one of the windows centered at this point for spectrogram
   tau2 = 0: tstep2: tspan2(end);
277
278
   % stores the max wavenumber for each gabor window (centered at tau(j))
   \max_{k} 2 = \operatorname{zeros}(1, \operatorname{length}(\operatorname{tau}2));
280
   spectrogram_data2 = zeros(n2, length(tau2));
281
282
   for j = 1: length(tau2)
       % Shannon (Step-function) Window
284
        gabor_filter2 = abs(tspan2 - tau2(j)) \ll 1/(2*a2);
285
286
       % plot our original signal
287
        subplot (411)
288
        plot((1:length(m2))/Fs2,m2);
289
        xlabel('Time [sec]');
290
        ylabel('Amplitude');
        title ('Mary Had a little lamb (recorder)');
292
293
       % Plot our gabor filter
294
```

```
subplot (412)
295
         plot(tspan2, gabor_filter2);
296
         ylim ([0,1.1]) %gaussian limits
297
         xlim ([0, L2])
299
         xlabel('Time [sec]');
300
         title ('Gabor Filter')
301
         ylabel('Filter Magnitude');
302
303
        % plot our filtered signal
304
         m2_filtered = gabor_filter2 .* m2;
305
         subplot (413)
307
         plot(tspan2, m2_filtered);
308
         xlabel('Time [sec]');
309
         ylabel ('Amplitude');
310
         title ('Filtered Signal in Spatial Domain')
311
         ylim ([-0.7, 0.7]);
312
         xlim ([0, L2])
313
314
315
        % plot our filtered signals frequency domain
316
         m2ft = fft (m2\_filtered);
317
318
        % filter overtones?
319
320
         subplot (414)
         plot (ks2, abs (fftshift (m2ft)) ./ max(abs (m2ft)));
322
         xlabel('Wavenumber (k)');
323
         ylabel('|uft| / max(|uft|)');
324
         title ('Filtered Signal in Frequency Domain')
325
         x \lim ([-20000, 20000])
326
327
        % outside "max" call is to filter out the negative wavenumber
328
         \max_{k} 2(j) = \max(k2(abs(m2ft)) = \max(abs(m2ft)));
330
         \operatorname{spectrogram\_data2}(:, j) = (\operatorname{abs}(\operatorname{fftshift}(\operatorname{m2ft}))) . / \operatorname{max}(\operatorname{abs}(\operatorname{m2ft})))
331
             . ';
         drawnow
333
334
    end
335
337
338
   % convert wavenumber to hertz!
339
   % dividing by an extra pi makes it exactly what I would expect?
340
341
    \max_{\text{freqs2}} = \max_{\text{k2}} . / (2*pi);
342
   % Find Music Score!
344
    music\_score2 = [];
345
346
    for freq = max_freqs2
347
         [\min_{val}, index] = \min(abs(freq - all_freqs));
348
         music_score2 = [music_score2 all_notes(index)];
349
    end
350
352
   % make spectogram plot for Recorder!
353
    close all; clc;
354
```

```
figure(4)
356
357
   \% divide by 2pi to convert wavenumber to \rm Hz
   pcolor(tau2, ks2./(2*pi), spectrogram_data2), shading interp;
359
   {\tt colorbar}
360
   hold on;
361
   ylim([0, 1000]) \% to see overtones
363
   \%ylim([0, 200]) % to see base notes
364
   ylabel('Frequency (Hz)')
365
  xlabel ('Time [sec]')
```