# COMPUTATIONAL METHODS IN FREEFORM OPTICAL TECHNOLOGIES

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EE 575 PROJECT - COMPUTATIONAL DIFFERENTIAL GEOMETRY

### OUTLINE

- I. INTRODUCTION & DESCRIPTION
- II. THEORETICAL SETUP & FRAMEWORK
- III. IMPLEMENTATION
- IV. CONCLUSIONS & FUTURE DIRECTIONS



### INTRODUCTION

- OPTICAL APPLICATIONS
  - CORRECTIVE LENSES
  - IMAGE FORMATION
  - AR/VR SYSTEMS
  - TELESCOPES
  - PRECISION CAMERAS
  - LED DISPLAY





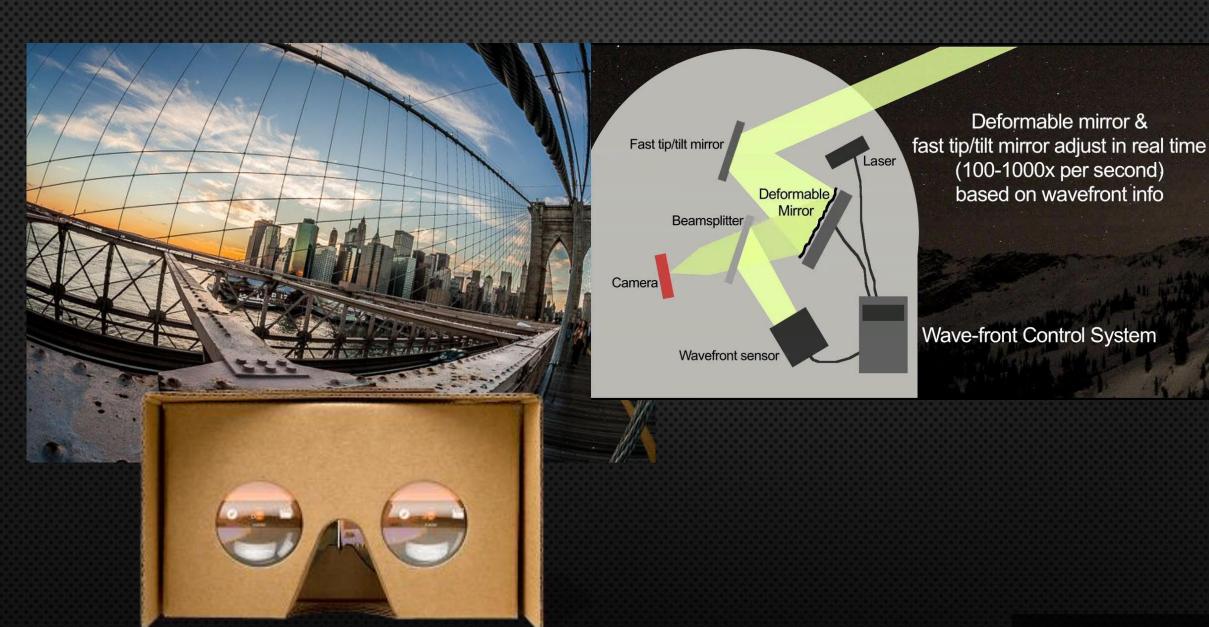




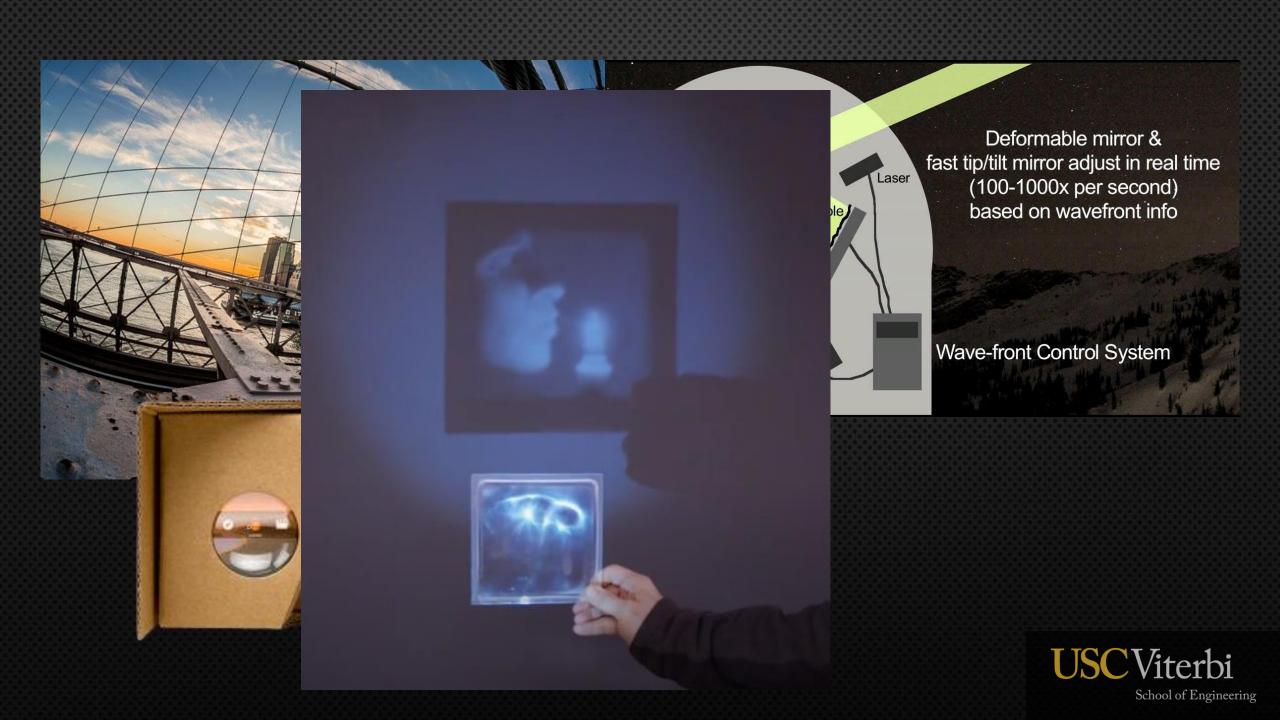








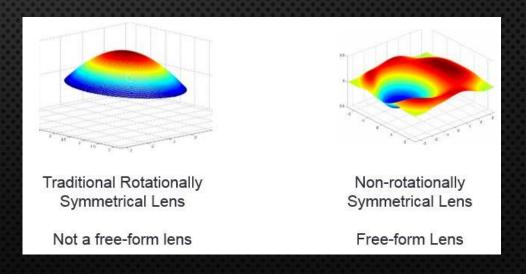




### INTRODUCTION

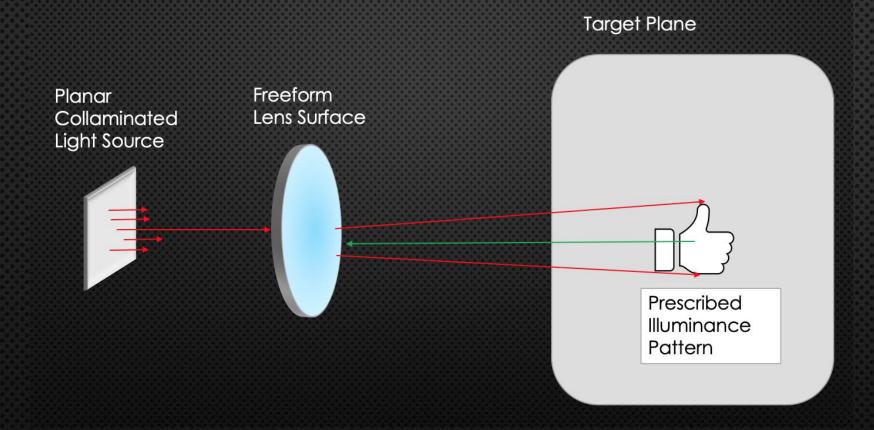
- Freeform Lenses
  - IN OPTICS, FREEFORM LENSES ARE OPTICAL SURFACES THAT ARE DESIGNED WITHOUT RIGID RADIAL DIMENSIONS SUCH AS TRANSLATIONAL OR ROTATIONAL SYMMETRY ABOUT AXES NORMAL TO ANY MEAN PLANE.







### GENERAL SETUP



GIVEN AN ILLUMINATION SOURCE AND AN IRRADIANCE DISTRIBUTION ON A PLANE. HOW CAN WE COMPUTE THE SURFACE OF THE LENS NEEDED TO GENERATE THE TARGET DISTRIBUTION?



EIKONAL EQUATION:

$$\nabla^2 E = c^2(x, y, z)$$

SOLUTIONS TO THE 3D WAVE EQUATION THAT REPRESENT LIGHT WAVEFRONT AT DIFFERENT TIMES.

Integrating the Eikonal at different spatial points yields an optical ray path length

$$\int_{A}^{B} \nabla^{2} E \ de \approx \sum_{i \in [A,B]} n_{i} d_{i}$$

FIND A REFRACTIVE OPTICAL SURFACE WHICH GIVES US RAY PATHS THAT SATISFY THE CONDITIONS FROM THE SOURCE (A) AND TARGET ENDPOINT LOCATIONS (B).



The coordinates of a Light ray on the target plane are related to the source points (x,y) by a vector valued function z(x,y) so that:

$$x_t = z_1(x, y) ; y_t = z_2(x, y)$$

Snell's Law and the Jacobian defines a conformal mapping from the freeform lens surface to the target plane:

$$dx_t dy_t = |J(z)| DXDY$$

WHERE:

$$J(z) = \begin{bmatrix} \frac{\partial z_1}{\partial x} & \frac{\partial z_1}{\partial y} \\ \frac{\partial z_2}{\partial x} & \frac{\partial z_2}{\partial x} \end{bmatrix}$$

MORE PRECISELY, THE PHYSICS INTERPRETATION DESCRIBES THE TRANSFORM BETWEEN IRRADIANCE DISTRIBUTIONS OF THE TARGET AND FREEFORM SURFACES:

$$dI_t(x_t, y_t) = |J(z)|dI(x, Y)$$



GIVEN:

$$dI_t(x_t, y_t) = |J(z)| \, \mathrm{DI}(X, Y)$$
 (Conformal Map of Irradiance Distributions) 
$$\iint I_t(x_t, y_t) \, dx_t dy_t = \iint I(x, y) \, dx dy$$
 (Conservation of Energy)

THESE CONDITIONS IMPLY A NONLINEAR MONGE-AMPERE ELLIPTICAL PDE:

DET 
$$D^2z = K(\bar{x})(1 + |Dz|^2)^{\frac{n+2}{2}}$$
 (PRESCRIBED CURVATURE PROBLEM)

ALSO EXPRESSED IN A LOCALLY LINEARIZED FORM AS:

$$A(z_{xx}z_{yy} - z_{xy}^{2}) + Bz_{xx} + Cz_{yy} + Dz_{xy} + E = 0$$

$$BC: \begin{cases} x_{t} = f_{1}(x, y, z, z_{x}, z_{y}) \\ y_{t} = f_{2}(x, y, z, z_{x}, z_{y}) \end{cases} : \partial S \rightarrow \partial T$$

Where the target coordinates are re-expressed as functions of the ray source and their incidence at the optical surface.

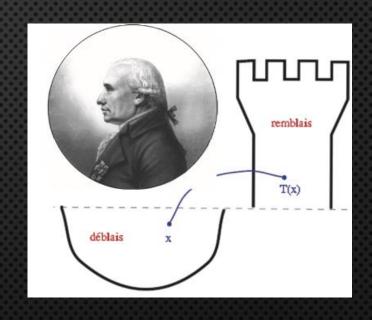


Taking the view of light as a photon mass, Energy conservation also implies An Optimal Mass Transport problem (Monge-Kantorovich Formulation):

$$\int c(\bar{x}, T(\bar{x})) I(\bar{x}) d\bar{x}$$
  
S/T:  
$$T: I(\bar{x}) \to I_t(\bar{x}_t)$$

Where the mass-preserving transform  $T(\bar{x})$  expresses a set of constraints for the Irradiance correspondence between each source to target spatial pairing.

 $T(\bar{x})$  HAS ALSO BEEN SHOWN TO BE THE GRADIENT OF THE MA PDE SOLUTIONS.





### EXISTING METHODS IN LITERATURE

- SIMULTANEOUS MULTIPLE SURFACE METHOD
- MA PDE AND NEWTON METHODS
- MK TRANSPORT AS LINEAR PROGRAM
- NURBS APPROACHES



The Goal is to compute a chart consisting of local Monge Patches defined as  $m: U \to \mathbb{R}^3$ ,  $U \subseteq \mathbb{R}^2$  where m(x,y) = (x,y,z(x,y)) uniquely satisfies the boundary conditions. From the Elliptical MA PDE and the Fundamental forms, the Monge patches should have:

DET 
$$D^2 z = K(\bar{x})(1 + |Dz|^2)^{\frac{n+2}{2}} \iff K_{Gaussian} = \frac{z_{xx}z_{yy} - z_{xy}^2}{(1 + z_x^2 + z_y^2)^2}$$

In our case, we found an explicit expression of Monge-Ampere used in Lens Design:

$$|z_{Lxx}z_{Lyy} - z_{Lxy}|^2 = |J(z)| \frac{\sqrt{1 + (z_{Lx}^2 + z_{Ly}^2)(1 - n^2)}}{K(z_{Lx}, z_{Ly})^2}$$



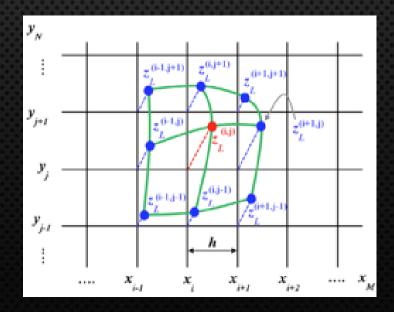
Using the Lens Expression of Monge-Ampère we compute the surface as a function of Finite Difference Approximations:

$$z_L^{(i,j)} = \frac{av_1^{(i,j)} + av_2^{(i,j)} - \sqrt{(av_1^{(i,j)} - av_2^{(i,j)})^2 + (\frac{av_3^{(i,j)} - av_4^{(i,j)}}{2})^2 + h^4g^{(i,j)}}}{2}$$

SOLVING EACH POINT USING AN ITERATIVE METHOD BASED ON GAUSS-SEIDEL.

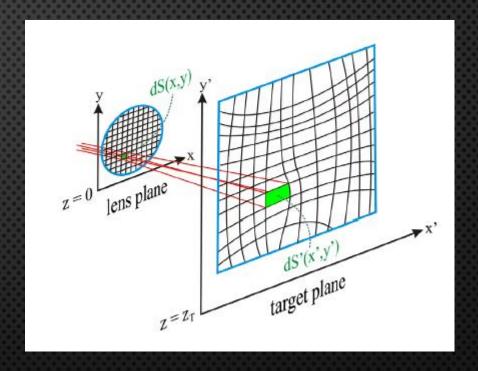
For 
$$t = (1:ITERS)$$
  $z_L^{t+1} = f(z_L^t, z_L^{t-1});$ 

END

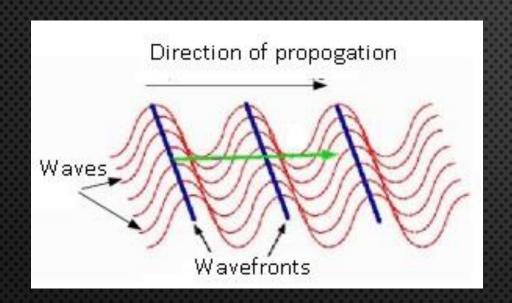


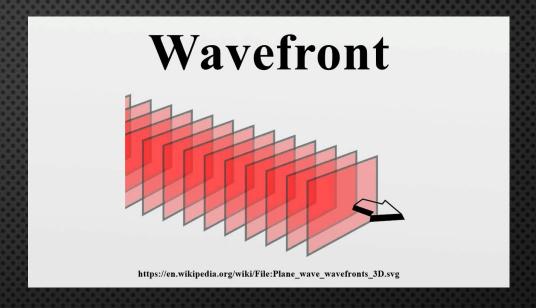


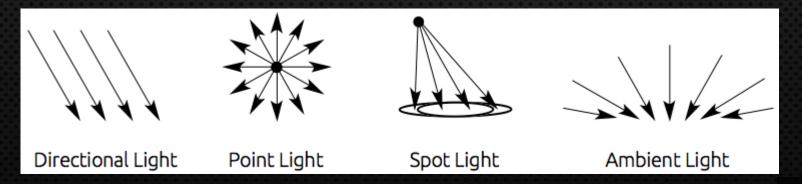




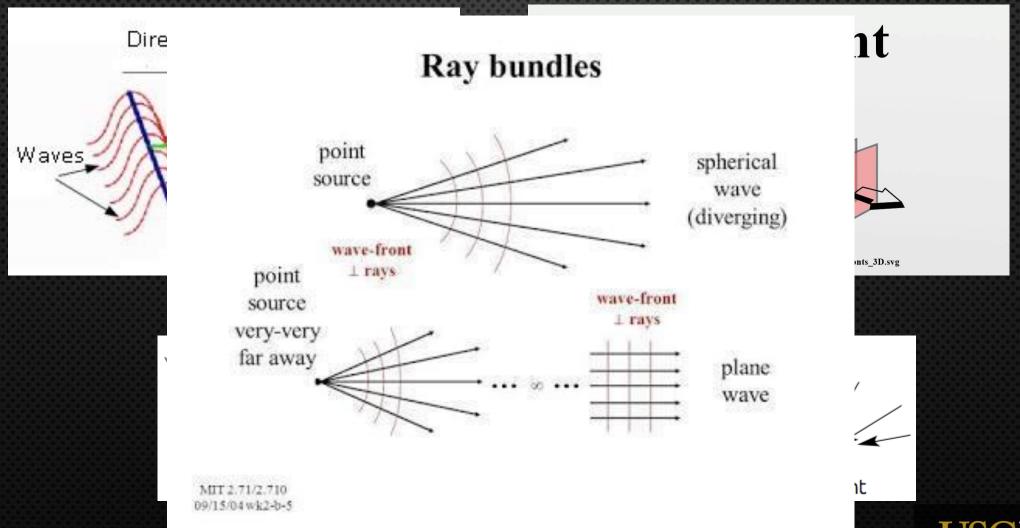




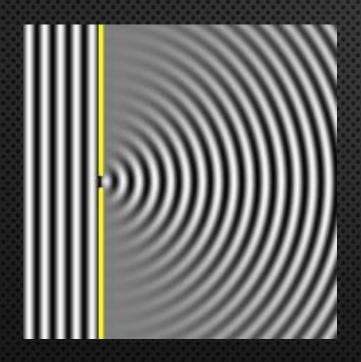


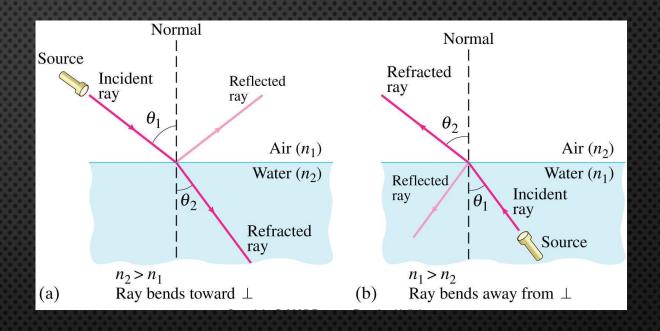






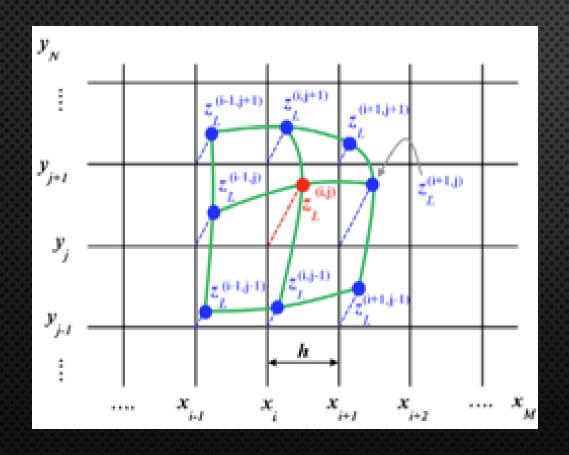


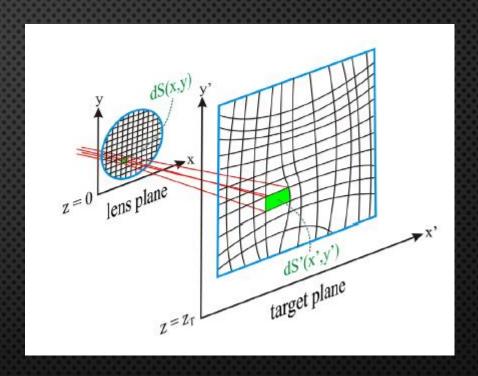




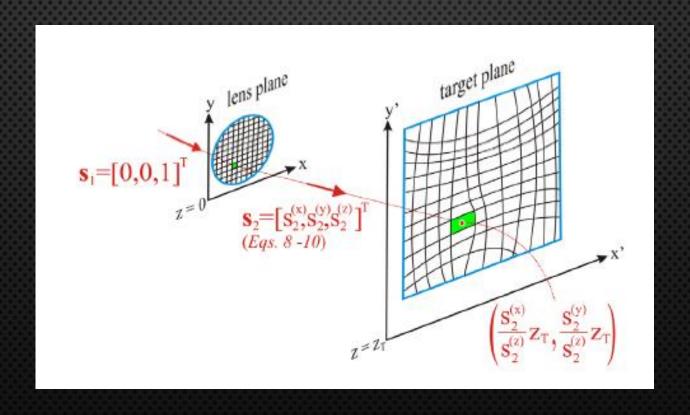
Snell's Law:  $n_1 \sin \theta_1 = n_2 \sin \theta_2$ 















```
clf;
clear all;
img = imread("USC_logo2.jpg");
imshow(img);
```



**N**XN X 3



# USC



# USC

```
img_bw = im2bw(img);
global target_img;
target_img = imcomplement(img_bw);

%Turn the following section on to lower
%resolution of target image
downSampleScalar = 100;
target_img = imresize(target_img, [floor(size(target_img, 1)/downSampleScalar), floor(size(target_img, 2)/downSampleScalar)]);
target_img = imcomplement(target_img);
imshow(target_img);
```





MATRIX VALUES 0.0

MATRIX VALUES 1.0



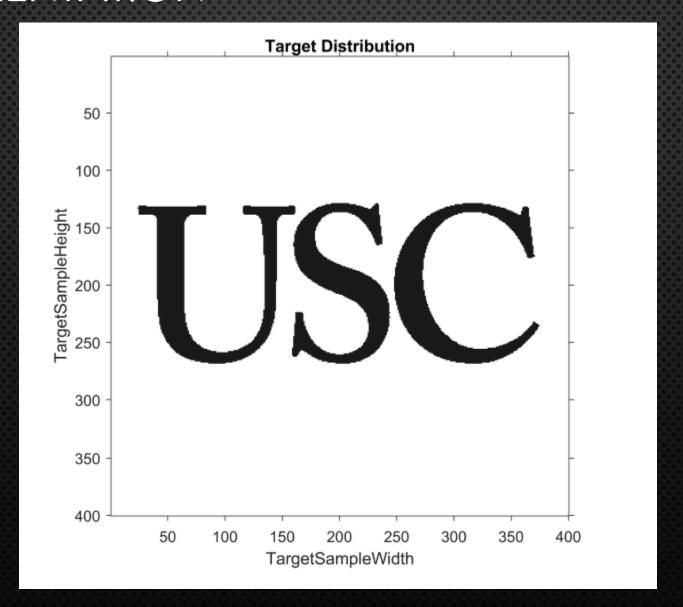
# USC

MATRIX VALUES 0.10

MATRIX VALUES 1.0

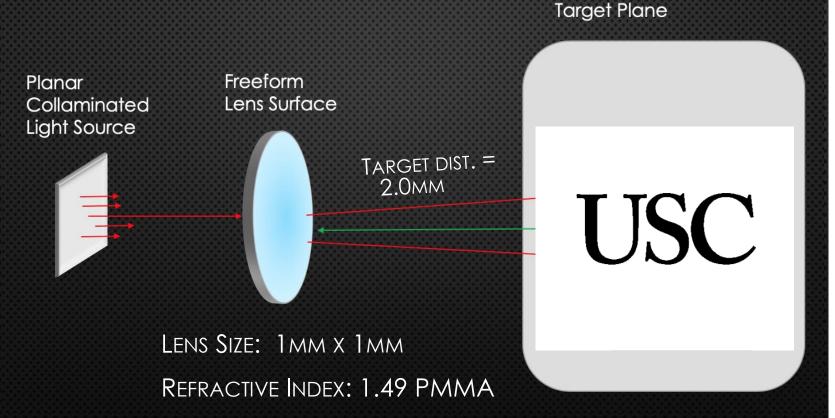
```
%Turn this on if you wish to add noise to the target image:
target_img = target_img + 0.1;
target_img(target_img>1) = 1.0;
imshow(target_img)
```







RESULTS:



LENS SIZE: 20MM X 20MM

Next we define our hyperparameters:

stepsize:

```
global h;
h = 1;
```

lens surface sample dimensions:

```
global sampleSize;
sampleSize = 99;
global stepSizeH;
%stepSizeH = 1;
stepSizeH = double(0.01);
```

#### targetDistance

```
global zT;
%zT = sampleSize*2;
zT = 2.0;
%zT = 20;
```

#### targetSize

```
global Tmax;
Tmax = 10;
%Tmax = 500;
```

#### Refractive Indicies:

```
refractiveIndexIn = 1.49; %PMMA
refractiveIndexOut = 1.0; %Vacuum
global n;
n = refractiveIndexIn/refractiveIndexOut;
```



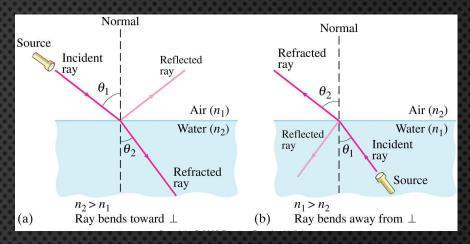
```
global i1;
global j1;
for itr = 1: total_itr
    zLensPrev = zLens;
   for i1 = 2:sampleSize-1
       for j1 = 2:sampleSize-1
            av1 = avfunc(zLens(i1+1,j1), zLens(i1-1,j1));
            av2 = avfunc(zLens(i1,j1+1), zLens(i1,j1-1));
            av3 = avfunc(zLens(i1+1,j1+1), zLens(i1-1,j1-1));
            av4 = avfunc(zLens(i1-1,j1+1), zLens(i1+1,j1-1));
            g = gfunc(i1, j1);
            zLens(i1,j1) = 0.5*(av1 + av2 - sqrt((av1-av2)^2) + (((av3-av4)*0.5)^2) + (stepSizeH^4) * (g));
        end
    end
    stabilityContainer(itr) = stabilityfunc(zLensPrev, zLens);
end
```

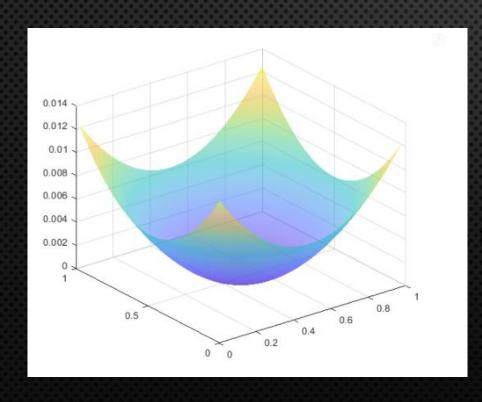


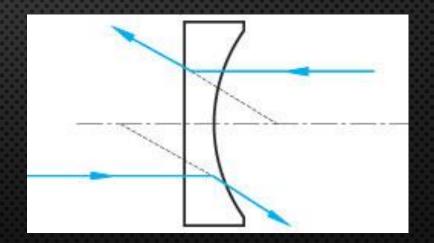
LENS SIZE = 1MM x 1MM | | | DO 1MM

TARGET DISTANCE ZT= 2.0MM

Target max dimensions = 20mm x 20mm | 1max = 10mm







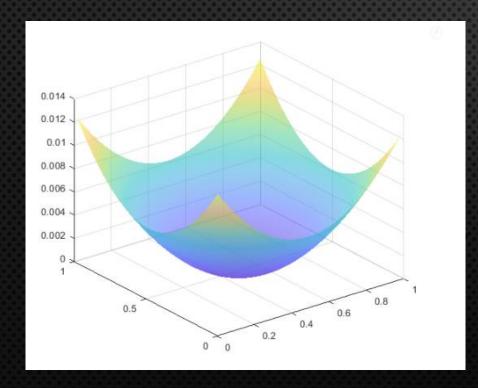
Snell's Law:  $n_1 \sin \theta_1 = n_2 \sin \theta_2$ 

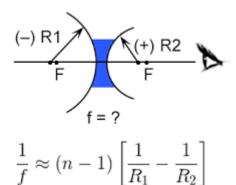


LENS SIZE = 1MM x 1MM | | | DO 1MM

TARGET DISTANCE ZT= 2.0MM

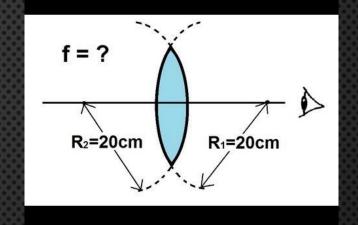
Target max dimensions = 20mm x 20mm | 1max = 10mm





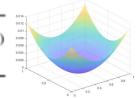
$$\frac{1}{f} \approx (n-1) \left[ \frac{1}{R_1} - \frac{1}{R_2} \right]$$

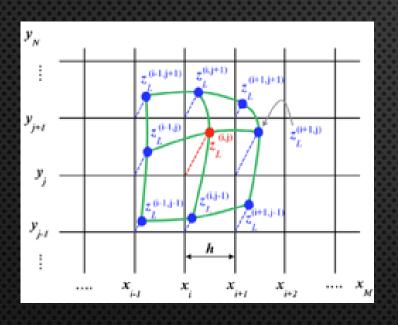
$$z_{L0}^{(i,j)} = \frac{x_{i,j}^2 + y_{i,j}^2}{\frac{D_0(n-1)z_T}{2} + \sqrt{\left(\frac{D_0(n-1)z_T}{2}\right)^2 - (x_{i,j}^2 + y_{i,j}^2)}}$$



$$R_c = \frac{D_0(n-1)z_T}{2T_{\text{max}}}$$

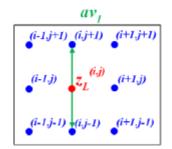
$$z_L^{(i,j)} = \frac{av_1^{(i,j)} + av_2^{(i,j)} - \sqrt{\left(av_1^{(i,j)} - av_2^{(i,j)}\right)^2 + \left(\frac{av_3^{(i,j)} - av_4^{(i,j)}}{2}\right)^2 + h^4g^{(i,j)}}}{2}$$

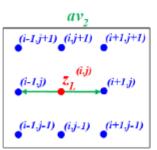


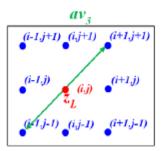


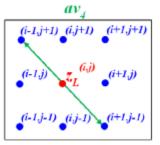
### Average:

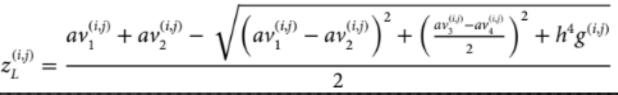
function avCalc= avfunc(input1, input2)
avCalc = double(input1 + input2)/2;
end

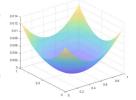


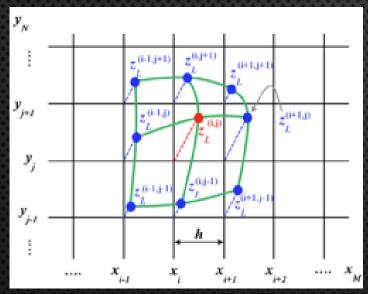


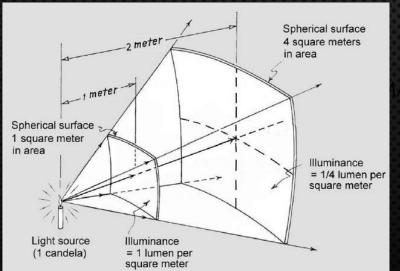




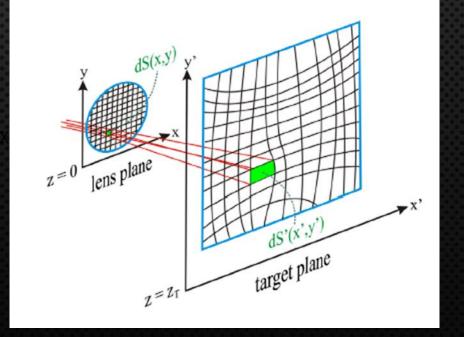




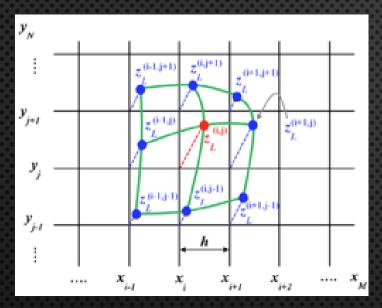




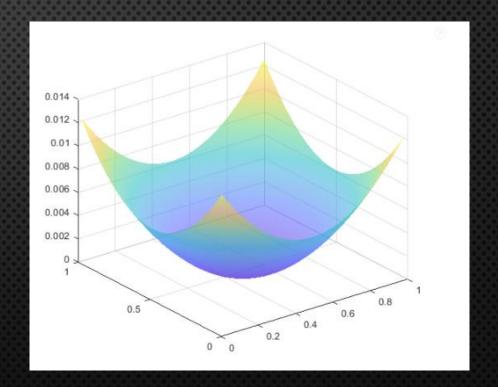
$$g^{(i,j)} = \frac{I_0(x_i, y_i)}{I_T(f_1(x_i, y_i), f_2(x_i, y_i))} \frac{\sqrt{1 + \left(z_{Lx}^{(i,j)2} + z_{Ly}^{(i,j)2}\right)\left(1 - n^2\right)}}{K(z_{Lx}^{(i,j)}, z_{Ly}^{(i,j))2}}$$



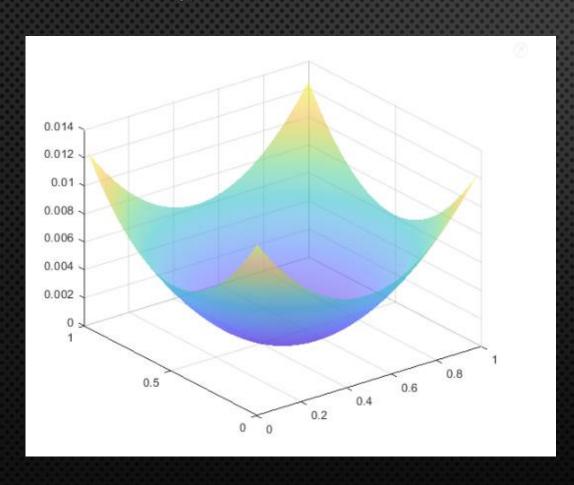




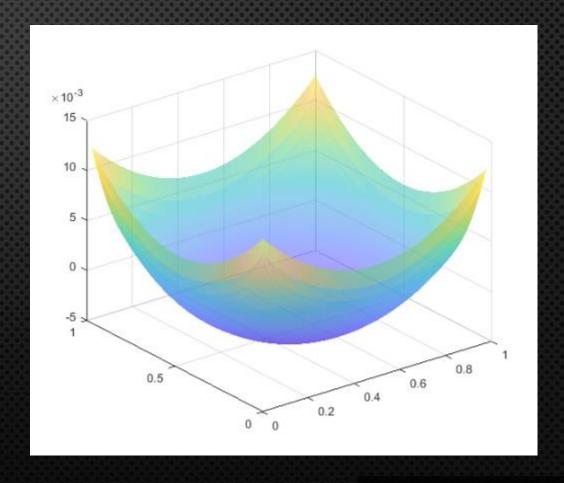
$$z_L^{(i,j)} = \frac{av_1^{(i,j)} + av_2^{(i,j)} - \sqrt{\left(av_1^{(i,j)} - av_2^{(i,j)}\right)^2 + \left(\frac{av_3^{(i,j)} - av_4^{(i,j)}}{2}\right)^2 + h^4g^{(i,j)}}}{2}$$



### INITIAL LENS:

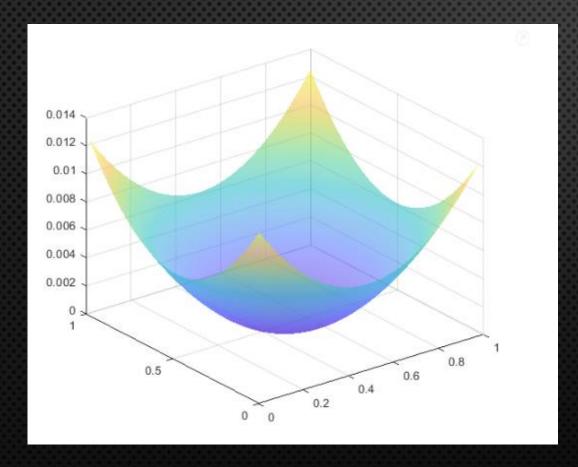


### LENS AFTER 25 ITERATIONS:

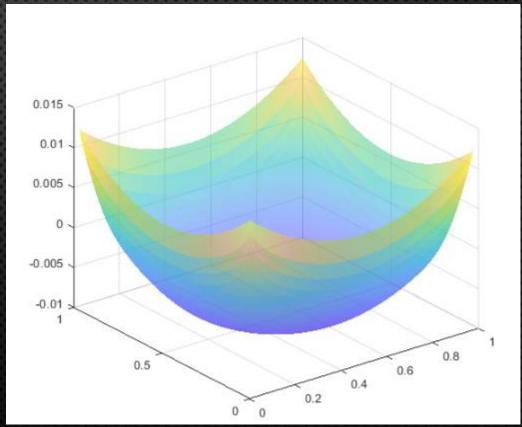




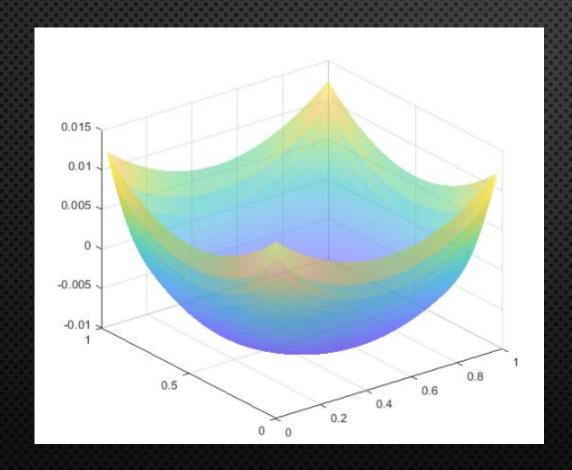
INITIAL LENS:

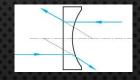


#### LENS AFTER 50 ITERATIONS:



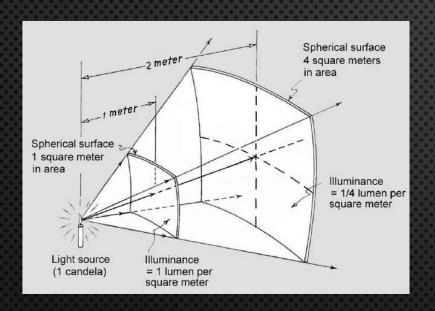








# IMPLEMENTATION - Challenges



$$z_L^{(i,j)} = \frac{av_1^{(i,j)} + av_2^{(i,j)} - \sqrt{\left(av_1^{(i,j)} - av_2^{(i,j)}\right)^2 + \left(\frac{av_3^{(i,j)} - av_4^{(i,j)}}{2}\right)^2 + h^4g^{(i,j)}}}{2}$$

$$g^{(i,j)} = \frac{I_0(x_i,y_i)}{I_T(f_1(x_i,y_i),f_2(x_i,y_i))} \frac{\sqrt{1+\left(z_{Lx}^{(i,j)2}+z_{Ly}^{(i,j)2}\right)\left(1-n^2\right)}}{K(z_{Lx}^{(i,j)},z_{Ly}^{(i,j))2}}$$

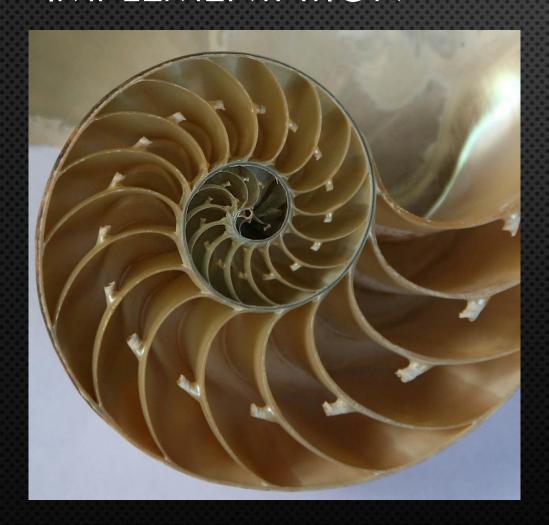
$$I_T(x',y')\cdot \big|J(x,y)\big|=I(x,y)$$

Matrix values 1.0



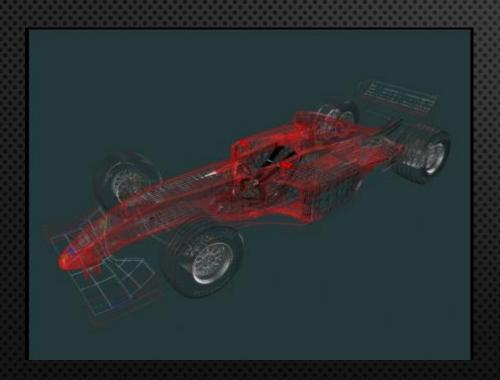


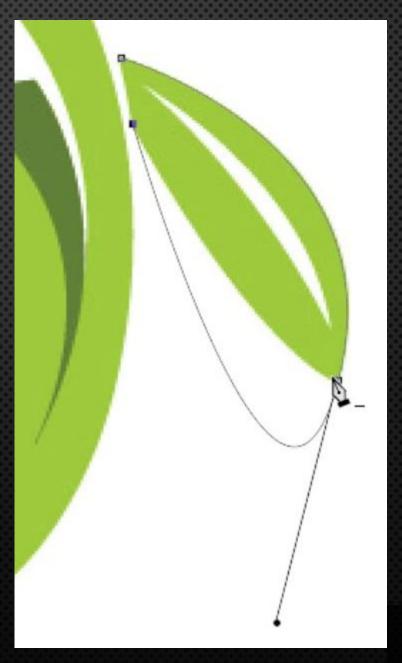






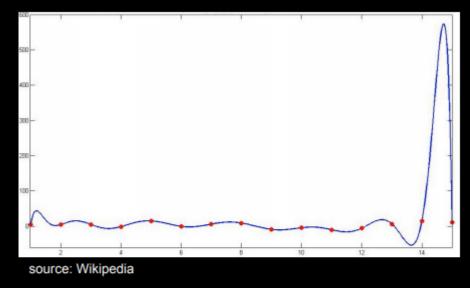






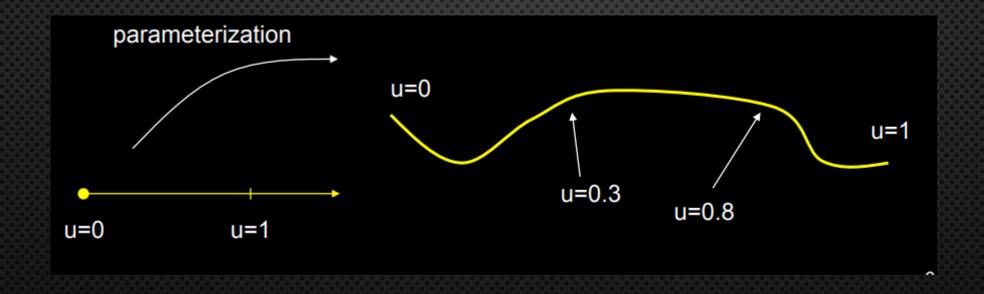


- An n-th degree polynomial fits a curve to n+1 points
  - called Lagrange Interpolation
  - result is a curve that is too wiggly, change to any control point affects entire curve (non-local)
  - this method is poor
- We usually want the curve to be as smooth as possible
  - minimize the wiggles
  - high-degree polynomials are bad

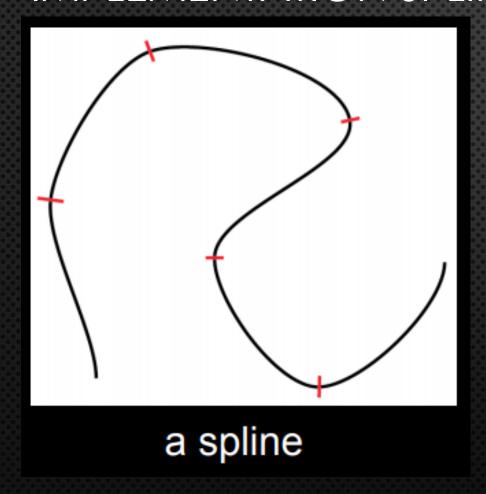


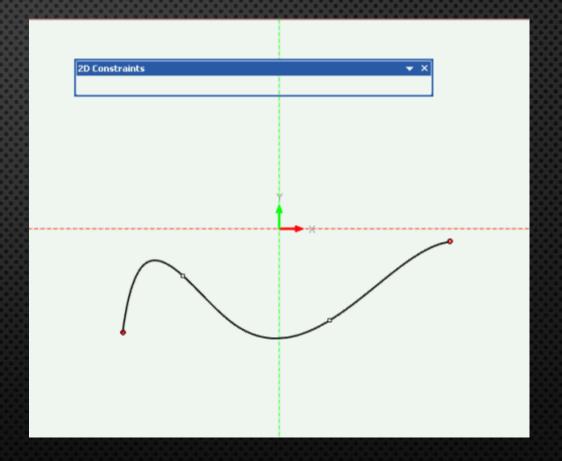
Lagrange interpolation, degree=15













- Cubic polynomial:
  - $p(u) = au^3 + bu^2 + cu + d = [u^3 \ u^2 \ u \ 1] [a \ b \ c \ d]^T$
  - a,b,c,d are 3-vectors, u is a scalar
- Three cubic polynomials, one for each coordinate:

$$- x(u) = a_x u^3 + b_x u^2 + c_x u + d_x$$

$$- y(u) = a_v u^3 + b_v u^2 + c_v u + d_v$$

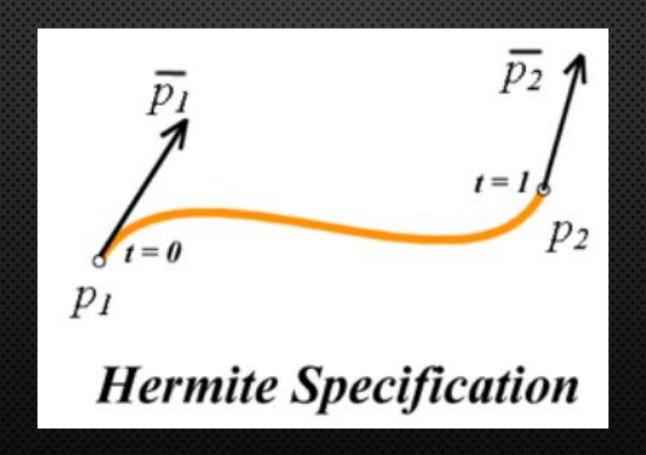
$$-z(u) = a_z u^3 + b_z u^2 + c_z u + d_z$$

In matrix notation:

$$[x(u) \quad y(u) \quad z(u)] = [u^3 \quad u^2 \quad u \quad 1] \begin{bmatrix} a_x & a_y & a_z \\ b_x & b_y & b_z \\ c_x & c_y & c_z \\ d_x & d_y & d_z \end{bmatrix}$$

• Or simply:  $p = [u^3 u^2 u 1] A$ 







 Four constraints: value and slope (in 3-D, position and tangent vector) at beginning and end of interval [0,1]:

$$p(0) = p_1 = (x_1, y_1, z_1)$$

$$p(1) = p_2 = (x_2, y_2, z_2)$$

$$p'(0) = \overline{p}_1 = (\overline{x}_1, \overline{y}_1, \overline{z}_1)$$

$$p'(1) = \overline{p}_2 = (\overline{x}_2, \overline{y}_2, \overline{z}_2)$$
the user constraints

- Assume cubic form:  $p(u) = au^3 + bu^2 + cu + d$
- Four unknowns: a, b, c, d



• Assume cubic form:  $p(u) = au^3 + bu^2 + cu + d$ 

$$p_1 = p(0) = d$$
 $p_2 = p(1) = a + b + c + d$ 
 $\overline{p_1} = p'(0) = c$ 
 $\overline{p_2} = p'(1) = 3a + 2b + c$ 

- Linear system: 12 equations for 12 unknowns (however, can be simplified to 4 equations for 4 unknowns)
- Unknowns: a, b, c, d (each of a, b, c, d is a 3-vector)



$$d = p_1$$

$$a + b + c + d = p_2$$

$$c = \overline{p_1}$$

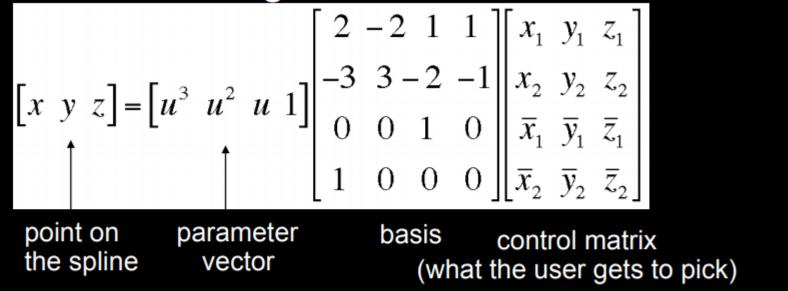
$$3a + 2b + c = \overline{p_2}$$

Rewrite this 12x12 system as a 4x4 system:

$$\begin{bmatrix} 0 & 0 & 0 & 1 \\ 1 & 1 & 1 & 1 \\ 0 & 0 & 1 & 0 \\ 3 & 2 & 1 & 0 \end{bmatrix} \begin{bmatrix} a_x & a_y & a_z \\ b_x & b_y & b_z \\ c_x & c_y & c_z \\ d_x & d_y & d_z \end{bmatrix} = \begin{bmatrix} x_1 & y_1 & z_1 \\ x_2 & y_2 & z_2 \\ \overline{x}_1 & \overline{y}_1 & \overline{z}_1 \\ \overline{x}_2 & \overline{y}_2 & \overline{z}_2 \end{bmatrix}$$

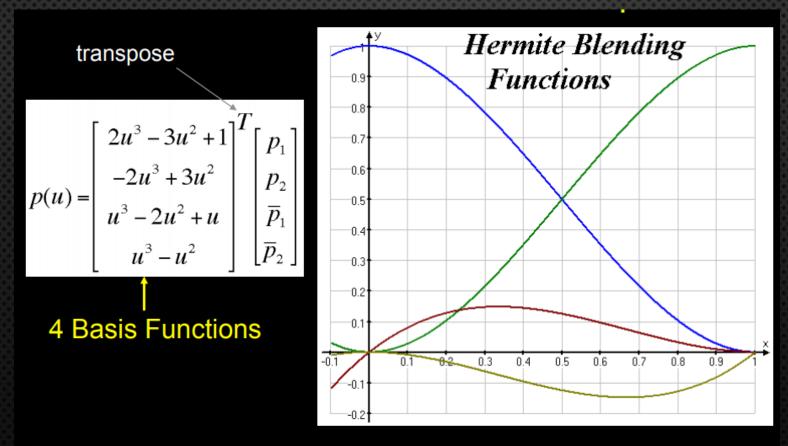


After inverting the 4x4 matrix, we obtain:



- This form is typical for splines
  - basis matrix and meaning of control matrix change with the spline type

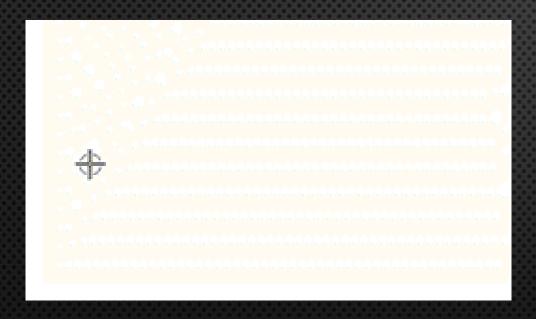


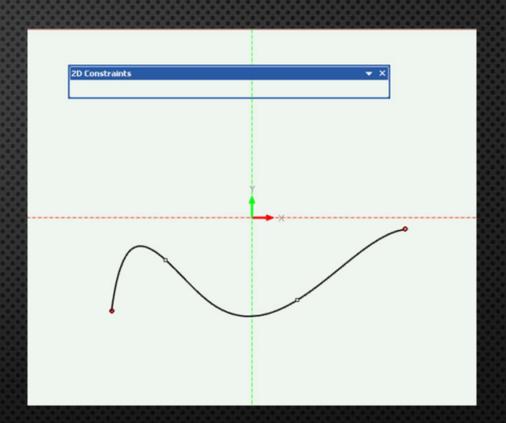


Every cubic Hermite spline is a linear combination (blend) of these 4 functions.



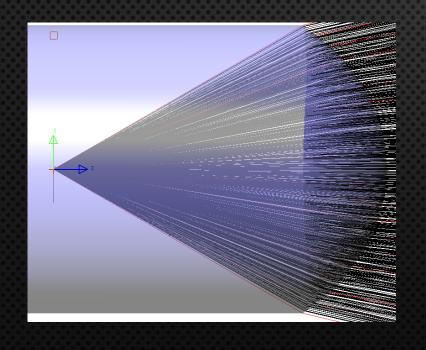
# IMPLEMENTATION – B-Splines

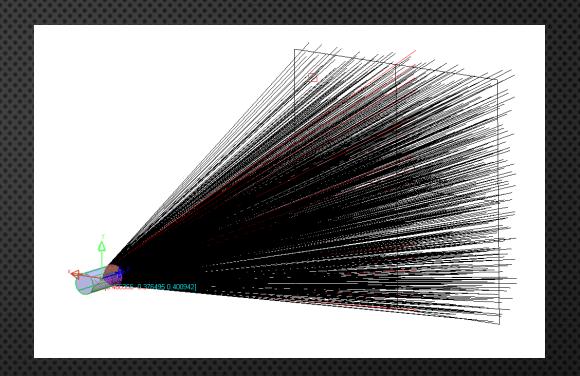






# IMPLEMENTATION LIGHT TOOLS





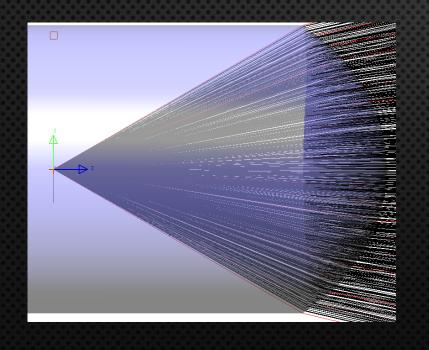
Lens Size = 1mm x 1mm | | | D0 1MM

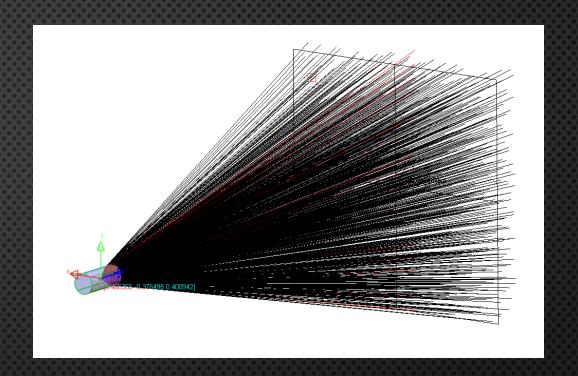
Target distance Zt= 2.0mm

TARGET MAX DIMENSIONS =  $20MM \times 20MM$  | TMAX = 10MM



# IMPLEMENTATION LIGHT TOOLS



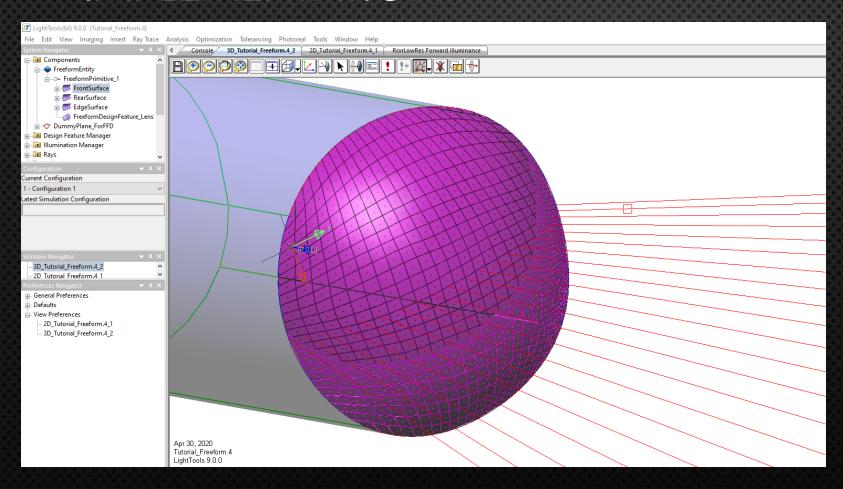


Lens Size = 1mm x 1mm | | DO 1MM

TARGET DISTANCE ZT= 2.0mm

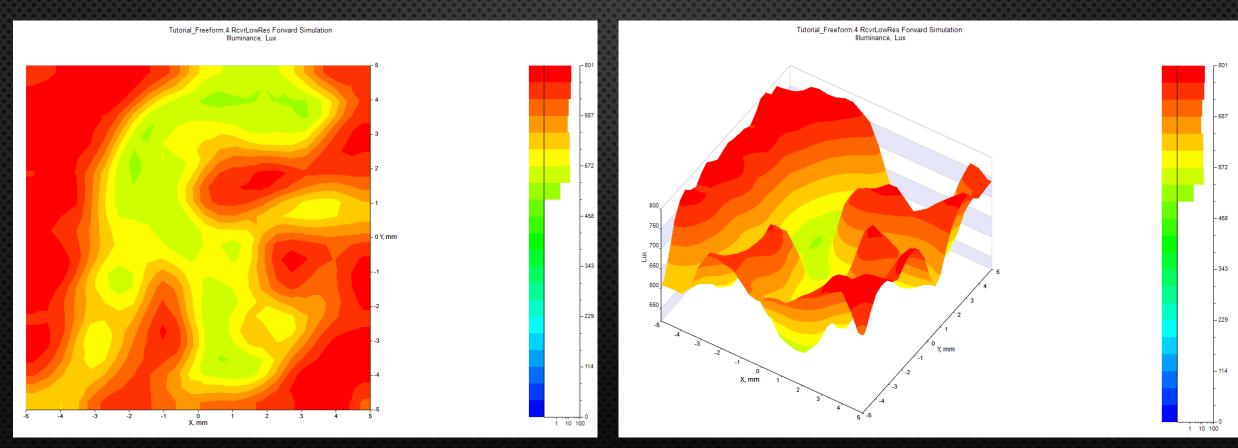
TARGET MAX DIMENSIONS =  $20MM \times 20MM$  | TMAX = 10MM

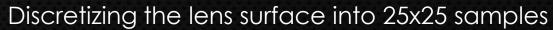




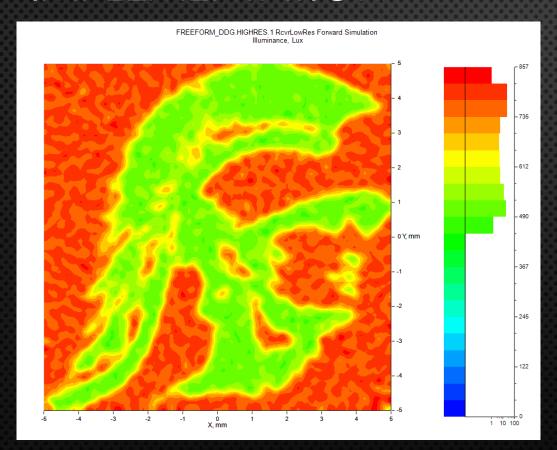


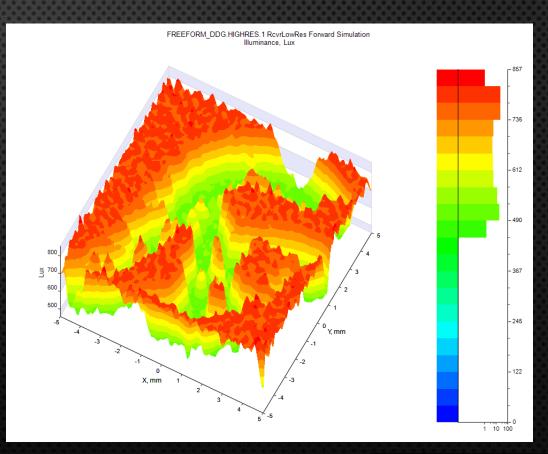












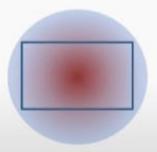
Discretizing the lens surface into 100x100 samples



Fact: Rotationally symmetrical lenses optimize image quality in a circle

The Opportunity: Optimize image quality almost any way you want

(within the laws of physics of course...)



Today: Image Quality from Conventional Rotationally Symmetric Lens

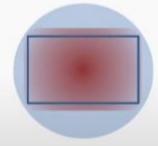
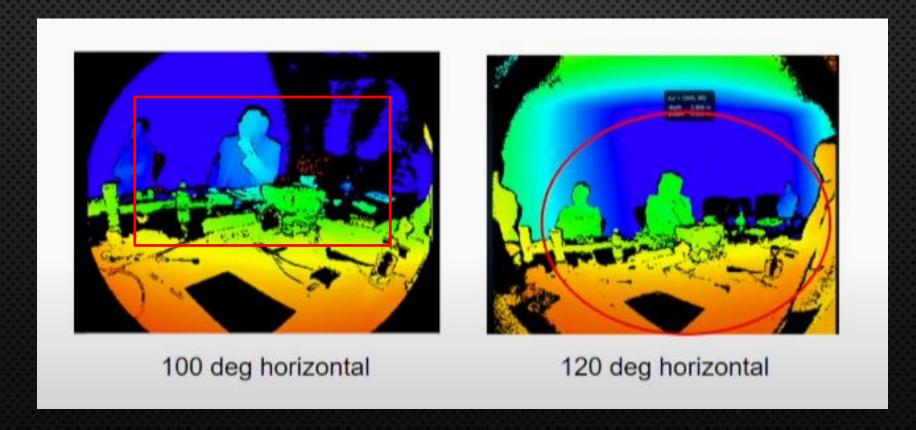


Image Quality
Optimized with using
free-form lenses



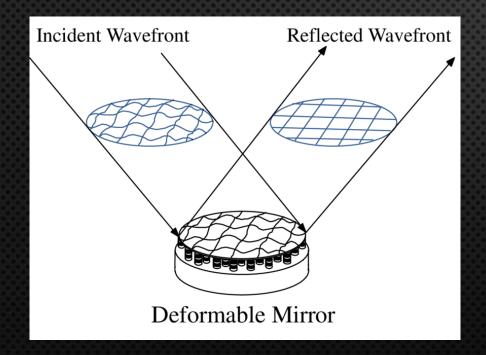




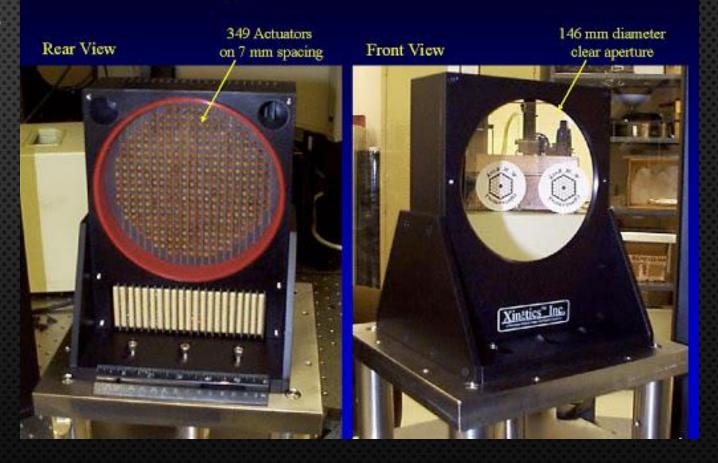




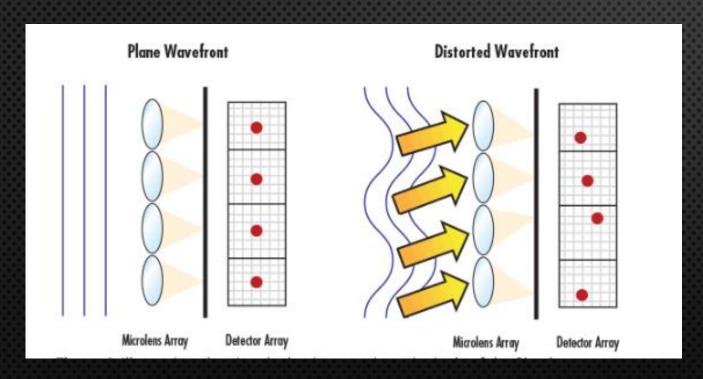


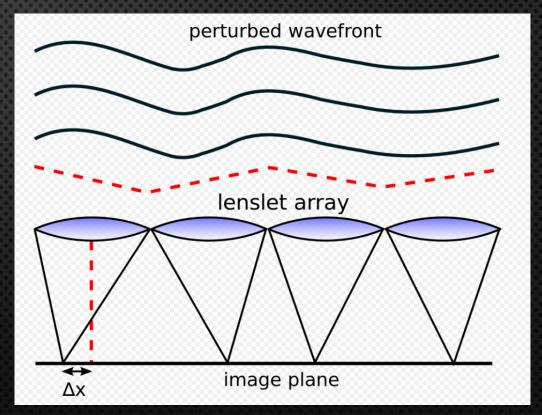


# Deformable Mirror

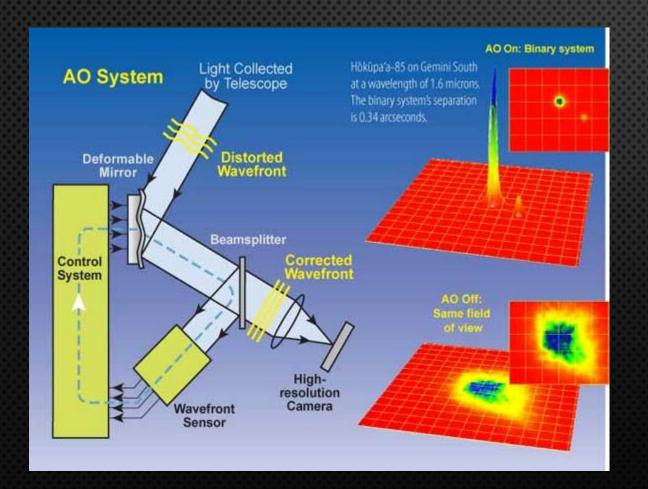


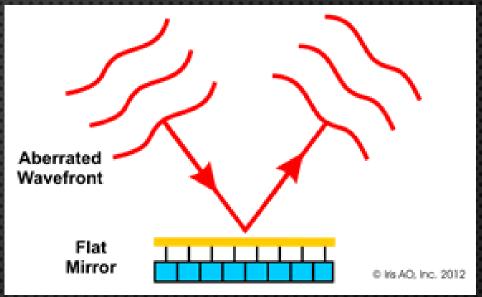


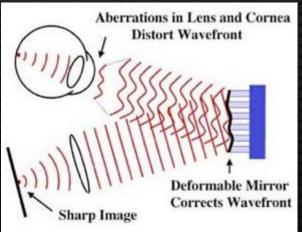




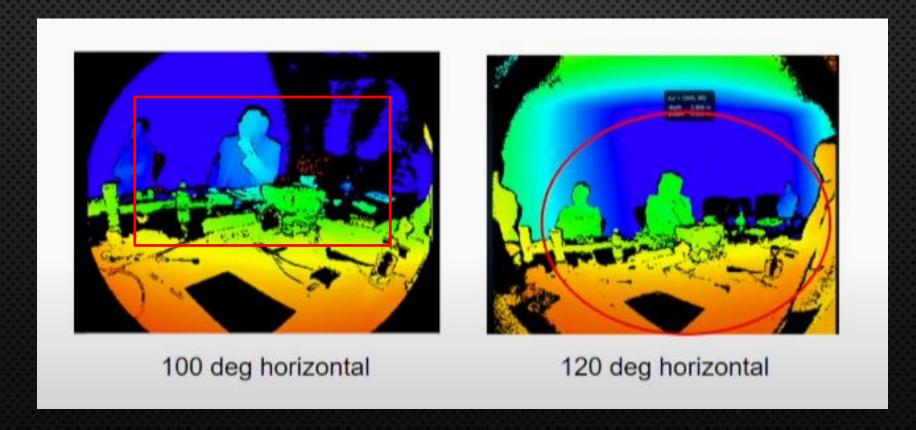






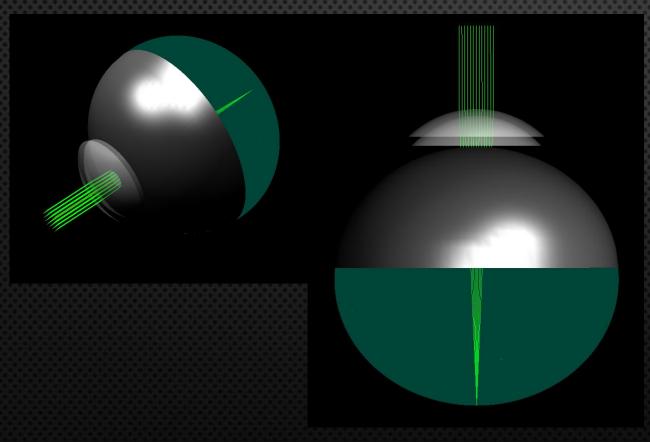


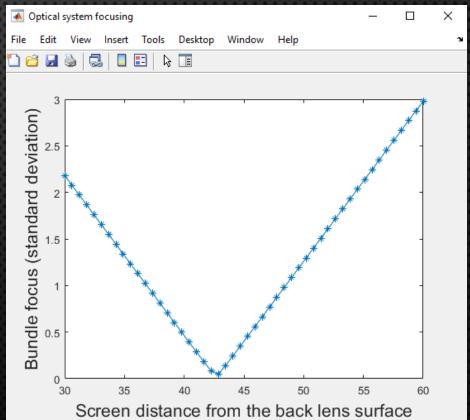






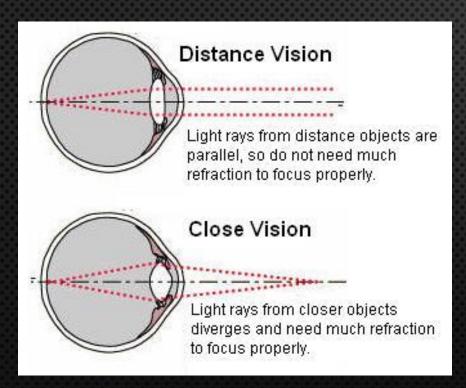
# CONCLUSIONS





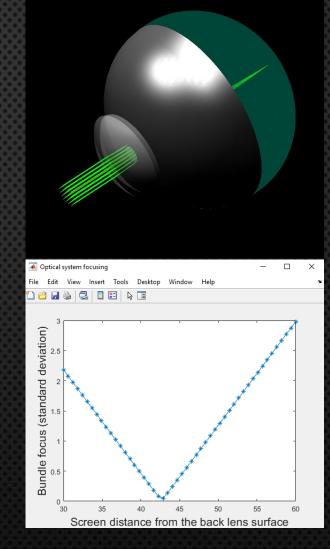


## CONCLUSIONS













END

