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Conversion factors and general equations applied in agricultural and forest meteorology

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ABSTRACT

Scientific communication in agricultural and forest meteorology requires a correct conversion of units and the adequate use of equations. In general, most of conversion and equations are found separately in a specific book, paper or document. Thus, here we propose a single document with conversion units of solar radiation, evapotranspiration, resource use efficiency, carbon balance, ecophysiology, trace gases and water; and general equations: Penman-Monteith, canopy conductance, decoupling factor and Bowen ratio. This document has a lot of value for teaching, for scientists and as a quick reference.

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Introduction

The knowledge of the interactions between the ecosystem and atmosphere play an important role in the agricultural and forest meteorology studies (Allen et al., 1998). Hence, a lot of data from field is required to understanding all these system interactions. The correct use of all factors and equations is the key of the agricultural and forest meteorology applications and results (Nobel, 2009).

In general, most of conversion factors and equations are found separately in a specific book, paper or document.

Thus, here we propose a single document with conversion units of solar radiation, evapotranspiration, resource use efficiency, carbon balance, ecophysiology, trace gases and water; and general equations: Penman-Monteith, canopy conductance, decoupling factor and Bowen ratio.

The idealization and preparation of this document is coming from the demand on the part of students and researchers to convert units and equations according to constants, values and time. We expect that this document can be support all members of the agrometeorological communit, such as professors, scientists and professionals.

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List of abbreviations and symbols

Constant and values

Solar radiation

 $R_s (W m^{-2}) \leftrightarrow R_s (kJ m^{-2} time^{-1})$ $R_c (W m^{-2}) \leftrightarrow R_c (MJ m^{-2} time^{-1})$ R_c (W m⁻²) \Leftrightarrow PAR (µmol m⁻² s⁻¹) PAR (μ mol m⁻² s⁻¹) \Leftrightarrow PAR (mol m⁻² time⁻¹)

Evapotranspiration

LE (W m⁻²) \rightarrow ET (mm time⁻¹) $R_s (MJ m^{-2} time^{-1}) \rightarrow ET (mm time^{-1})$

General equations

Evapotranspiration - Penman-Monteith (1965) Evapotranspiration - Penman-Monteith (FAO 56) Canopy Conductance (inversion of Penman-Monteith equation) Canopy Conductance (inversion of Penman-Monteith equation) - Simplified Decoupling Factor (Ω) Bowen Ratio (β)

Resource use efficiency

WUE → Water-use efficiency (GPP/ET) WUE, \rightarrow Intrinsic water-use efficiency (GPP/ g_c) $CUE \rightarrow Carbon$ -use efficiency (NPP/GPP) LUE → Light-use efficiency (GPP/PAR) RUE → Rain-use efficiency (GPP/PPT)

Carbon balance

 CO_2 (µmol m⁻² s⁻¹) \rightarrow C (g m⁻² time⁻¹) $C (g m^{-2} yr^{-1}) \rightarrow C (t ha^{-1} yr^{-1})$ CO_2 storage (ppm \rightarrow µmol m⁻² s⁻¹)

Ecophysiology conversion units

Assimilation rate (A) $A \text{ (}\mu\text{mol CO}_{2} \text{ m}^{-2} \text{ s}^{-1}\text{)} \rightarrow A \text{ (}mg \text{ CO}_{2} \text{ m}^{-2} \text{ s}^{-1}\text{)}$

Transpiration rate (E) $E \text{ (mmol H}_{2}O \text{ m}^{-2} \text{ s}^{-1}) \rightarrow E \text{ (mg H}_{2}O \text{ m}^{-2} \text{ s}^{-1})$ $E \text{ (mmol H}_2\text{O m}^{-2}\text{ s}^{-1}) \rightarrow E \text{ (W m}^{-2})$

Stomatal conductance (g) $g_s \pmod{m^{-2} s^{-1}} \Leftrightarrow g_s \pmod{s^{-1}}$ g_{s} (mol m⁻² s⁻¹) \Leftrightarrow g_{s} (cm s⁻¹)

 $g_s \text{ (mol m}^{-2} \text{ s}^{-1}\text{)} \leftrightarrow g_s \text{ (mm s}^{-1}\text{)}$

General conversion units for trace gases and water

 CO_{2} (mmol m⁻³ \leftrightarrow mg m⁻³) $H_2O \text{ (mmol m}^{-3} \leftrightarrow \text{g m}^{-3})$ CH_4 (mmol m⁻³ \leftrightarrow mg m⁻³) NO_2 (mmol m⁻³ \leftrightarrow mg m⁻³) O_3 (mmol m⁻³ \leftrightarrow mg m⁻³)

Acknowledgments

References

LIST OF ABREVIATIONS AND SYMBOLS

A Assimilation Rate

°C Degree Celsius

C Carbon

c CO₂ Concentration

cal Calorie

CH₄ Methane

CO₂ Carbon Dioxide

 c_p Air Specific Heat

CUE Carbon-use efficiency

dz Height of the Profile

E Transpiration Rate

e_a Actual Vapor Pressure

e Saturation Vapor Pressure

ET Evapotranspiration

FAO Food and Agriculture Organization of the United

Nations

g Gram

G Soil Heat Flux

 g_a Aerodynamic Conductance

g Canopy Conductance

GPP Gross Primary Production

g_s Stomatal Conductance

h Crop Height

H Sensible Heat Flux

H₂O Water

ha Hectare

hm Maximum Measurement Height

J Joule

K Kelvin

k Kilo

k Turbulent Transfer Coefficient for Water;

 $\mathbf{k}_{_{\mathrm{H}}}$ Turbulent Transfer Coefficient for Heat;

L Liter

LAI Leaf Area Index

LAI_{acitve} Active Leaf Area Index

LE Latent Heat Flux

LUE Light-use efficiency

M Mega

m Meter

min Minute

mm Millimeter

n Number of Moles

NO₂ Nitrogen Dioxide

NPP Net Primary Production

O₃ Ozone

P Atmospheric Pressure

Pa Pascal

PAR Photosynthetically Active Radiation

ppm Parts Per Million

PPT Precipitation

R Universal Gas Law Constant

r_a Aerodynamic Resistance

r Canopy Resistance

R_d Universal Gas Constant Specific for Dry Air

R Solar Radiation

RH Relative Humidity

r₁ Stomatal resistance of the Well-Illuminated Leaf

R_n Net Radiation

r_s Surface Resistance

RUE Rain-use efficiency

R_w Universal Gas Constant Specific for Water Vapor

s Second

Sc Storage Flux

T Air Temperature

t Ton

t Time

T_{kv} Virtual Temperature

u₂ Wind Speed at 2 m height

u, Wind Speed at z m above ground surface

V Volume of the Gas

VPD Vapor Pressure Deficit

W Watt

WUE Water-use efficiency

WUE, Intrinsic water-use efficiency

yr Year

z Elevation; Height Above Sea Level

 z_{m} Height of Wind Measurements

z_L Height of Humidity Measurements

d Zero Plane Displacement Height

 $z_{_{om}}$ Roughness Length Governing Momentum Transfer

z_{oh} Roughness Lenth Governing Transfer of Heat and

Vapor k von Karman's constant

u, Wind Speed at Height z

β Bowen Ratio

γ Psychrometric Constant

Δ Slope Vapor Pressure

ε Ratio Molecular Weight of Water Vapor/Dry Air

λ Latent Heat of Vaporization for Water

 ρ_a Air density

Ω Decoupling Factor

CONSTANTS AND VALUES

1 minute = 60 seconds $1 \text{ mm H}_2\text{O} = 1 \text{ kg m}^{-2} \text{ H}_2\text{O} = 1 \text{ L m}^{-2} \text{ H}_2\text{O}$ 30 minutes = 1800 seconds $1 \text{ W} = \text{J s}^{-1}$ 60 minutes (1 hour) = 3600 seconds 1 cal = 4.184 J $\lambda (J \text{ kg}^{-1}) = 10^3 (2500 - 2.37\text{T}), T = {}^{\circ}\text{C}$ 24 hours (1 day) = 86400 seconds $\lambda (J \text{ kg}^{-1}) = 10^3 (3147 - 2.37T), T = K$ $1 \text{ mol H}_{2}O = 18 \text{ g H}_{2}O$ λ (cal g⁻¹) = 10³(597 – 0.564T), T = °C $1 \text{ mol CO}_2 = 44 \text{ g CO}_2$ 1 mol CO₂ = 12 g Carbon $\lambda = 2.45$ (MJ kg⁻¹), T = 20 °C; $1 \text{ mol CH}_{4} = 16 \text{ g CH}_{4}$ 1 W m^{-2} of total solar radiation ~ 2.02 μ mol m^{-2} s⁻¹ $1 \text{ mol NO}_2 = 46 \text{ g NO}_2$ photons of PAR $Pa = J m^{-3}$ 1 mol $O_3 = 48 g O_3$ $1 g = 10^{-6} t$ $J = L/10^3(2500 - 2.37T)$ $1 \text{ MJ m}^2 \text{ day}^{-1} = 0.408 \text{ mm day}^{-1}$ $1 \text{ kg} = 10^3 \text{ g}$ $R = 8.314510 \text{ J K}^{-1} \text{ mol}^{-1} \text{ or m}^{3} \text{ Pa K}^{-1} \text{ mol}^{-1}$ $1 \text{ kJ} = 10^3 \text{ J}$ $1 M = 10^6$ $R_d = 287.058 \text{ J kg}^{-1} \text{ K}^{-1}$ $1 \text{ ha} = 10000 \text{ m}^2$ $R_w = 461.495 \text{ J kg}^{-1} \text{ K}^{-1}$ 1 K = °C + 273.16 $\varepsilon = 0.622;$ $1 \text{ kg H}_{2}\text{O} = 1 \text{ L H}_{2}\text{O} = 10^{-3} \text{ m}^{3}\text{ H}_{2}\text{O}$ $c_p = 1.013 \times 10^{-3} \text{ (MJ kg}^{-1} \, ^{\circ}\text{C}^{-1}\text{);}$ $1 \text{ mm H}_{2}\text{O} = 10^{-3} \text{ m}^{3} \text{ H}_{2}\text{O m}^{-2} \text{ area}$ k = 0.41

Solar radiation

Conversion factor for solar radiation: (W m⁻²) to (kJ m⁻² time⁻¹)

$$\frac{W}{m^2} \to \frac{kJ}{m^2 \cdot time} \tag{1}$$

$$\frac{W}{m^2} \to \frac{J}{m^2 \cdot s} \to 10^{-3} \frac{kJ}{m^2 \cdot s}$$
 (2)

Conversion of W m⁻² to kJ m⁻² considering the time of 30 min (1800 s):

$$\frac{W}{m^2} \to 10^{-3} \frac{kJ}{m^2 \cdot \frac{1}{1800}} \to 1800 * 10^{-3} \frac{kJ}{m^2 \cdot 30min} \to 1.8 \frac{kJ}{m^2 \cdot 30min}$$
(3)

Conversion of W m^{-2} to kJ m^{-2} considering the time of 1 hour (3600 s):

$$\frac{W}{m^2} \to 10^{-3} \frac{kJ}{m^2 \cdot \frac{1}{3600}} \to 3600 * 10^{-3} \frac{kJ}{m^2 \cdot hour} \to 3.6 \frac{kJ}{m^2 \cdot hour}$$
(4)

Conversion of W m⁻² to kJ m⁻² considering the time of 1 day (86400 s):

$$\frac{W}{m^2} \to 10^{-3} \frac{kJ}{m^2 \cdot \frac{1}{86400}} \to 86400 * 10^{-3} \frac{kJ}{m^2 \cdot day} \to 86.4 \frac{kJ}{m^2 \cdot day}$$
 (5)

Examples:

$$700 \frac{W}{m^2} \to 700 * 1.8 \frac{kJ}{m^2 \cdot 30min} \to 1260 \frac{kJ}{m^2 \cdot 30min}$$
 (6)

$$700 \frac{W}{m^2} \to 700 * 3.6 \frac{kJ}{m^2 \cdot hour} \to 2520 \frac{kJ}{m^2 \cdot hour}$$
 (7)

$$300 \frac{W}{m^2} \to 300 * 86.4 \frac{kJ}{m^2 \cdot day} \to 25920 \frac{kJ}{m^2 \cdot day}$$
(8)

Conversion factor for solar radiation: (W m⁻²) to (MJ m⁻² time⁻¹)

$$\frac{W}{m^2} \to \frac{MJ}{m^2 \cdot time} \tag{9}$$

$$\frac{W}{m^2} \to \frac{J}{m^2 \cdot s} \to 10^{-6} \frac{MJ}{m^2 \cdot s} \tag{10}$$

Conversion of W m^{-2} to MJ m^{-2} considering the time of 30 min (1800 s):

$$\frac{W}{m^2} \to 10^{-6} \frac{MJ}{m^2 \cdot \frac{1}{1800}} \to 1800 * 10^{-6} \frac{MJ}{m^2 \cdot 30min} \to 0.0018 \frac{MJ}{m^2 \cdot 30min}$$
 (11)

Conversion of W m⁻² to MJ m⁻² considering the time of 1 hour (3600 s):

$$\frac{W}{m^2} \to 10^{-6} \frac{MJ}{m^2 \cdot \frac{1}{3600}} \to 3600 * 10^{-6} \frac{MJ}{m^2 \cdot hour} \to 0.0036 \frac{MJ}{m^2 \cdot hour}$$
 (12)

Conversion of W m⁻² to MJ m⁻² considering the time of 1 day (86400 s):

$$\frac{W}{m^2} \to 10^{-6} \frac{MJ}{m^2 \cdot \frac{1}{86400}} \to 86400 * 10^{-6} \frac{MJ}{m^2 \cdot day} \to 0.0864 \frac{MJ}{m^2 \cdot day}$$
 (13)

Examples:

$$700 \frac{W}{m^2} \to 700 * 0.0018 \frac{MJ}{m^2 \cdot 30min} \to 1.26 \frac{MJ}{m^2 \cdot 30min}$$
(14)

$$700 \frac{W}{m^2} \to 700 * 0.0036 \frac{MJ}{m^2 \cdot hour} \to 2.52 \frac{MJ}{m^2 \cdot hour}$$
(15)

$$300 \frac{W}{m^2} \to 300 * 0.0864 \frac{MJ}{m^2 \cdot day} \to 25.92 \frac{MJ}{m^2 \cdot day}$$
(16)

Conversion factor for solar radiation: (kJ m-2 time-1) to (W m-2)

$$\frac{kJ}{m^2.time} \to \frac{W}{m^2} \tag{17}$$

$$\frac{kJ}{m^2.s} \to 10^3 \frac{J}{m^2 \cdot s} \to 10^3 \frac{W}{m^2} \tag{18}$$

Conversion of kJ m⁻² to W m⁻² considering the time of 30 min (1800 s):

$$\frac{kJ}{m^2. time} \to \frac{kJ}{m^2. 30min} \to \frac{1}{1800} * 10^3 \frac{W}{m^2} \to 0.5556 \frac{W}{m^2}$$
(19)

Conversion of kJ m⁻² to W m⁻² considering the time of 1 hour (3600 s):

$$\frac{kJ}{m^2.time} \to \frac{kJ}{m^2.hour} \to \frac{1}{3600} * 10^3 \frac{W}{m^2} \to 0.2778 \frac{W}{m^2}$$
(20)

Conversion of kJ m⁻² to W m⁻² considering the time of 1 day (86400 s):

$$\frac{kJ}{m^2.time} \to \frac{kJ}{m^2.day} \to \frac{1}{86400} * 10^3 \frac{W}{m^2} \to 0.01157 \frac{W}{m^2}$$
(21)

Examples:

$$1260 \frac{kJ}{m^2.30min} \to 1260 * 0.5556 \frac{W}{m^2} \to 700 \frac{W}{m^2}$$
(22)

$$2520 \frac{kJ}{m^2. hour} \to 2520 * 0.2778 \frac{W}{m^2} \to 700 \frac{W}{m^2}$$
(23)

$$25920 \frac{kJ}{m^2. day} \to 25920 * 0.01157 \frac{W}{m^2} \to 300 \frac{W}{m^2}$$
(24)

Conversion factor for solar radiation: (MJ m⁻² time⁻¹) to (W m⁻²)

$$\frac{MJ}{m^2.time} \to \frac{W}{m^2} \tag{25}$$

$$\frac{MJ}{m^2.s} \to 10^6 \frac{J}{m^2 \cdot s} \to 10^6 \frac{W}{m^2} \tag{26}$$

Conversion of MJ m⁻² to W m⁻² considering the time of 30 min (1800 s):

$$\frac{MJ}{m^2.time} \to \frac{MJ}{m^2.30min} \to \frac{1}{1800} * 10^6 \frac{W}{m^2} \to 555.56 \frac{W}{m^2}$$
(27)

Conversion of MJ m⁻² to W m⁻² considering the time of 1 hour (3600 s):

$$\frac{MJ}{m^2.time} \to \frac{MJ}{m^2.hour} \to \frac{1}{3600} * 10^6 \frac{W}{m^2} \to 277.78 \frac{W}{m^2}$$
(28)

Conversion of MJ m⁻² to W m⁻² considering the time of 1 day (86400 s):

$$\frac{MJ}{m^2. time} \to \frac{MJ}{m^2. day} \to \frac{1}{86400} * 10^6 \frac{W}{m^2} \to 11.57 \frac{W}{m^2}$$
(29)

Examples:

$$1.26 \frac{MJ}{m^2.30min} \to 555.56 * 1.26 \frac{W}{m^2} \to 700 \frac{W}{m^2}$$
(30)

$$2.52 \frac{MJ}{m^2.hour} \to 277.78 * 2.52 \frac{W}{m^2} \to 700 \frac{W}{m^2}$$
(31)

$$25.92 \frac{MJ}{m^2.day} \to 11.57 * 25.92 \frac{W}{m^2} \to 300 \frac{W}{m^2}$$
(32)

Conversion factor for solar radiation: R_s (W m⁻²) to PAR (μ mol m⁻² s⁻¹)

$$R_{s} - \frac{W}{m^{2}} \rightarrow PAR - \frac{\mu mol}{m^{2}.s} \tag{33}$$

$$1 \frac{W}{m^2} \to 2.02 \frac{\mu mol}{m^2.s}$$
 (34)

Examples:

$$500 \frac{W}{m^2} \to 500 * 2.02 \frac{\mu mol}{m^2.s} \to 1010 \frac{\mu mol}{m^2.s}$$
(35)

$$800 \frac{W}{m^2} \to 800 * 2.02 \frac{\mu mol}{m^2.s} \to 1616 \frac{\mu mol}{m^2.s}$$
 (36)

The approximation 1 W m⁻² \approx 4.57 μ mol m⁻² s⁻¹ (Thimijan & Heins, 1983) is assuming that the W m⁻² is for photosynthetically active radiation (PAR) from 4.0 to 7.0 μ m. But, in general pyranometers measure total solar radiation and only \sim 45% of the energy of solar radiation is actually in the 4.0 to 7.0 μ m range. Thus, the conversion of solar radiation (R_s) to photosynthetically active radiation (PAR) is about \sim 2.02, rather than 4.57.

The photosynthetically active radiation (PAR) depends of the light color (spectrum, clouds, etc). Here, the \sim 2.02 conversion factor is described in Mavi & Tupper (2004), but there are many others conversion factors for different region around the world described in Foken (2008).

Conversion factor for solar radiation: PAR (μ mol m $^{-2}$ s $^{-1}$) to R $_s$ (W m $^{-2}$)

$$PAR - \frac{\mu mol}{m^2.s} \to R_s - \frac{W}{m^2} \tag{37}$$

$$1\frac{\mu mol}{m^2.s} \rightarrow \frac{1}{2.02} \frac{W}{m^2} \tag{38}$$

Examples:

$$1010 \frac{\mu mol}{m^2.s} \to 1010 * \frac{1}{2.02} \frac{W}{m^2} \to 500 \frac{W}{m^2}$$
(39)

$$1616 \frac{\mu mol}{m^2.s} \to 1616 * \frac{1}{2.02} \frac{W}{m^2} \to 800 \frac{W}{m^2}$$
(40)

Conversion factor for solar radiation: PAR (µmol m⁻² s⁻¹) to PAR (mol m⁻² time⁻¹)

$$\frac{\mu mol}{m^2.s} \to \frac{mol}{m^2.time} \tag{41}$$

$$\frac{\mu mol}{m^2.s} \to 10^{-6} \frac{mol}{m^2 \cdot time} \tag{42}$$

Conversion of µmol m⁻² s⁻¹ to mol m⁻² considering the time of 30 min (1800 s):

$$\frac{\mu mol}{m^2.s} \to 10^{-6} \frac{mol}{m^2.\frac{1}{1800}} \to 1800 * 10^{-6} \frac{mol}{m^2.30min} \to 0.0018 \frac{mol}{m^2.30min}$$
(43)

Conversion of µmol m⁻² s⁻¹ to mol m⁻² considering the time of 1 hour (3600 s):

$$\frac{\mu mol}{m^2.s} \to 10^{-6} \frac{mol}{m^2.\frac{1}{3600}} \to 3600 * 10^{-6} \frac{mol}{m^2.hour} \to 0.0036 \frac{mol}{m^2.hour}$$
(44)

Conversion of umol m⁻² s⁻¹ to mol m⁻² considering the time of 1 day (86400 s):

$$\frac{\mu mol}{m^2.s} \to 10^{-6} \frac{mol}{m^2. \frac{1}{86400}} \to 86400 * 10^{-6} \frac{mol}{m^2. day} \to 0.0864 \frac{mol}{m^2. day}$$
(45)

Examples:

$$1500 \frac{\mu mol}{m^2.s} \to 1500 * 0.0018 \frac{mol}{m^2.30min} \to 2.7 \frac{mol}{m^2.30min}$$
(46)

$$1500 \frac{\mu mol}{m^2.s} \to 1500 * 0.0036 \frac{mol}{m^2.hour} \to 5.4 \frac{mol}{m^2.hour}$$
(47)

$$500 \frac{\mu mol}{m^2, s} \to 500 * 0.0864 \frac{mol}{m^2, day} \to 43.2 \frac{mol}{m^2, day}$$
 (48)

Conversion factor for solar radiation: PAR (mol m⁻² time⁻¹) to PAR (µmol m⁻² s⁻¹)

$$\frac{mol}{m^2.time} \to \frac{\mu mol}{m^2.s} \tag{49}$$

$$\frac{mol}{m^2.time} \to 10^6 \frac{\mu mol}{m^2 \cdot time} \tag{50}$$

Conversion of mol m^{-2} s⁻¹ to μ mol m^{-2} considering the time of 30 min (1800 s):

$$\frac{mol}{m^2.time} \rightarrow \frac{mol}{m^2.30min} \rightarrow \frac{1}{1800} * 10^6 \frac{\mu mol}{m^2.s} \rightarrow 555.56 \frac{\mu mol}{m^2.s}$$

$$(51)$$

Conversion of mol m^{-2} s⁻¹ to μ mol m^{-2} considering the time of 1 hour (3600 s):

$$\frac{mol}{m^2.time} \rightarrow \frac{mol}{m^2.hour} \rightarrow \frac{1}{3600} * 10^6 \frac{\mu mol}{m^2.s} \rightarrow 277.78 \frac{\mu mol}{m^2.s} \tag{52}$$

Conversion of mol m^{-2} s⁻¹ to μ mol m^{-2} considering the time of 1 day (86400 s):

$$\frac{mol}{m^{2}.time} \to \frac{mol}{m^{2}.day} \to \frac{1}{86400} * 10^{6} \frac{\mu mol}{m^{2}.s} \to 11.57 \frac{\mu mol}{m^{2}.s}$$
(53)

Examples:

$$2.7 \frac{mol}{m^2.30min} \rightarrow 2.7 * 555.56 \frac{\mu mol}{m^2.s} \rightarrow 1500 \frac{\mu mol}{m^2.s}$$
(54)

$$5.4 \frac{mol}{m^2.hour} \rightarrow 5.4 * 277.78 \frac{mol}{m^2.s} \rightarrow 1500 \frac{\mu mol}{m^2.s}$$
 (55)

$$43.2 \frac{mol}{m^2. day} \to 43.2 * 11.57 \frac{\mu mol}{m^2. s} \to 500 \frac{\mu mol}{m^2. s}$$
(56)

Evapotranspiration

Conversion factor for evapotranspiration: LE (W m⁻²) to ET (mm time⁻¹)

$$LE = ET.\lambda \rightarrow ET = \frac{LE}{\lambda}$$
 (57)

Conversion value of LE to ET considering the time of 30 minutes (1800 seconds), and:

Equation (58)

 λ (J kg⁻¹) = 10³(3147.5 – 2.37T), T = Air Temperature in K (Kelvin) (Aubinet et al., 2012):

$$ET\left(\frac{mm}{30min}\right) = \frac{1800}{10^3(3147.5 - 2.37T)} LE \to \frac{1.8}{(3147.5 - 2.37T)} LE \tag{58}$$

Equation (59)

 λ (J kg⁻¹) = 10³(2500 – 2.37T), T = Air Temperature in °C (Celsius) (Pereira et al., 2013):

$$ET\left(\frac{mm}{30min}\right) = \frac{1800}{10^3(2500 - 2.37T)}LE \rightarrow \frac{1.8}{(2500 - 2.37T)}LE \tag{59}$$

Equation (59)

 λ (cal g⁻¹) = 10³(597 – 0.564T), T = Air Temperature in °C (Celsius), 1 cal = 4.184 J:

$$ET\left(\frac{mm}{30min}\right) = \frac{\frac{1.8}{4.184}}{\left(\frac{2500}{4.184} - \frac{2.37}{4.184}T\right)} LE \to \frac{0.43}{(597 - 0.564T)} LE \tag{60}$$

Conversion factor for evapotranspiration: LE (W m⁻²) to ET (mm time⁻¹)

$$LE = ET.\lambda \rightarrow ET = \frac{LE}{\lambda}$$
 (61)

$$ET\left(\frac{mm}{time}\right) = \frac{LE}{\lambda} = \frac{\left(\frac{W}{m^2}\right)}{\left(\frac{J}{kg}\right)} \to \frac{\left(\frac{J}{m^2.s}\right)}{\left(\frac{J}{kg}\right)} \to \left(\frac{J}{m^2.s}\right) * \frac{kg}{J} \to \frac{kg}{m^2.s} \to \frac{mm}{s} \tag{62}$$

$$ET\left(\frac{mm}{30min}\right) = \frac{mm}{\frac{1}{1800}} \to 1800 \frac{mm}{30min} \to 1800 \frac{LE}{\lambda} \to 1800 \left(\frac{LE}{10^3(3147.5 - 2.37T)}\right) \tag{63}$$

Examples:

Considering:

- . LE (latent heat flux) = 500 W m⁻²
- . T (air temperature) = 25 °C

$$ET\left(\frac{mm}{30min}\right) = \frac{1.8}{\left(3147.5 - 2.37.(25 + 273.16)\right)} * 500 \rightarrow 0.3687 \frac{mm}{30min}$$
(64)

$$ET\left(\frac{mm}{30min}\right) = \frac{1.8}{(2500 - 2.37.(25))} * 500 \rightarrow 0.3687 \frac{mm}{30min}$$
(65)

$$ET\left(\frac{mm}{30min}\right) = \frac{0.43}{(597 - 0.564.(25))} * 500 \to 0.3688 \frac{mm}{30min}$$
(66)

Conversion factor for evapotranspiration: R_s (MJ m⁻² time⁻¹) to ET (mm time⁻¹)

$$R_{s} - \frac{MJ}{m^{2}.time} \rightarrow ET - \frac{mm}{time}$$
(67)

Considering:

1 mm = 1 L m⁻²

$$1 L H_{2}O = 1 kg H_{2}O$$

To evaporate 1 L H_2 O is necessary 10^3 (2500 – 2.37T) J

 $J = L/10^3(2500 - 2.37T)$

$$\frac{MJ}{m^2.time} \to 10^6 \frac{J}{m^2.time} \tag{68}$$

$$10^{6} \frac{L}{10^{3}(2500-2.37T).m^{2}.time} \rightarrow \frac{1mm}{(2.5-0.00237T).time} \rightarrow \frac{1}{(2.5-0.00237T)} \frac{mm}{time}$$
(69)

Example:

Considering:

- . T (air temperature) = 25 °C
- . Energy to evaporate = 12 MJ m⁻² day⁻¹

$$ET\left(\frac{mm}{day}\right) = \frac{12}{(2.5 - 0.00237 * (25))} \to ET = 4.91 \frac{mm}{day}$$
 (70)

General equations

Evapotranspiration - Penman-Monteith (1965)

$$\lambda ET = \frac{\Delta (R_n - G) + \rho_a c_p \frac{(e_s - e_a)}{r_a}}{\Delta + \gamma (1 + \frac{r_s}{r_a})} = \frac{\Delta (R_n - G) + \rho_a c_p \frac{VPD}{r_a}}{\Delta + \gamma (1 + \frac{r_s}{r_a})}$$
(71)

Where:

 λ is latent heat of vaporization of water (MJ kg⁻¹);

ET is evapotranspiration (mm day⁻¹);

 R_n is net radiation at the crop surface (MJ m⁻² day⁻¹);

G is soil heat flux (MJ m⁻² day⁻¹);

 ρ_a is mean air density at constant pressure (kg m⁻³);

 $c_{_{D}}$ is air specific heat at constant pressure (MJ kg $^{-1}$ °C $^{-1}$);

e is saturation vapor pressure (kPa);

e is actual vapor pressure (kPa);

 $(e_s - e_a)$ is saturation vapor pressure deficit (kPa);

VPD is vapor pressure deficit (kPa);

Δ is slope vapor pressure curve (kPa °C⁻¹);

γ is psychrometric constant (kPa °C⁻¹);

r_s is surface resistance (s m⁻¹);

r_a is aerodynamic resistance (s m⁻¹).

The Penman-Monteith is a combined method that included the energy balance and mass transfer. The meteorological variables, solar radiation, air temperature, relative humidity and wind speed, are used to compute the evapotranspiration (Allen et al., 1998).

Evapotranspiration - Penman-Monteith (FAO 56)

$$ET_{o} = \frac{0.408\Delta(R_{n} - G) + \gamma \frac{900}{T + 273}u_{2}(e_{s} - e_{a})}{\Delta + \gamma(1 + 0.34u_{2})} = \frac{0.408\Delta(R_{n} - G) + \gamma \frac{900}{T + 273}u_{2}VPD}{\Delta + \gamma(1 + 0.34u_{2})}$$

$$(72)$$

Where:

 ET_0 is reference evapotranspiration (mm day⁻¹);

R_n is net radiation at the crop surface (MJ m⁻² day⁻¹);

G is soil heat flux density (MJ m⁻² day⁻¹);

T is mean daily air temperature at 2 m height (m s⁻¹); u_2 is wind speed at 2 m height (m s⁻¹); e_s is saturation vapor pressure (kPa); e_a is actual vapor pressure (kPa); $e_s - e_a$ is saturation vapor pressure deficit (kPa); VPD is vapor pressure deficit (kPa); Δ is slope vapor pressure curve (kPa °C⁻¹); γ is psychrometric constant (kPa °C⁻¹).

The FAO Penman-Monteith method provides the evapotranspiration of the reference surface (ET_o). This method is defined according to the following definition for the reference surface: "A hypothetical reference crop with an assumed crop height of 0.12 m, a fixed surface resistance of 70 s m^{-1} and an albedo of 0.23" (Allen et al., 1998).

Evapotranspiration-Parameters

The evapotranspiration parameters are described in Allen et al. (1998).

Saturation Vapor Pressure (es)

$$e_s = 0.6108 exp^{\left(\frac{17.27T}{T + 237.3}\right)} \tag{73}$$

Where:

 e_s is saturation vapor pressure (kPa); T is mean air temperature (°C).

Actual Vapor Pressure (e₂)

$$e_a = \frac{e_S * RH}{100} \tag{74}$$

Where:

 e_a is actual vapor pressure (kPa); e_s is saturation vapor pressure (kPa); RH is relative humidity (%).

Relative Humidity (RH)

$$RH = \frac{e_a}{e_s} * 100 \tag{75}$$

Where:

RH is relative humidity (%); e_a is actual vapor pressure (kPa); e_s is saturation vapor pressure (kPa).

Vapor Pressure Deficit (VPD)

$$VPD = e_s - e_a \tag{76}$$

Where:

VPD is vapor pressure deficit (kPa); e_s is saturation vapor pressure (kPa); e_s is actual vapor pressure (kPa).

Latent Heat of Vaporization (λ)

$$\lambda = \frac{10^3 (2500 - 2.37T)}{10^6} \tag{77}$$

Where:

 λ is latent heat of vaporization (MJ kg⁻¹);

T is mean air temperature (°C).

"As λ varies only slightly over normal temperature ranges a single value of 2.45 MJ kg $^{-1}$ is taken in the simplification of the FAO Penman-Monteith equation. This is the latent heat flux for an air temperature of about 20 °C" (Allen et al., 1998).

Slope of saturation vapor pressure curve (Δ)

$$\Delta = \frac{4098[0.6108exp(\frac{17.27T}{T+237.3})]}{(T+237.3)^2}$$
(78)

Where:

 Δ is slope of saturation vapor pressure curve at air temperature T (kPa $^{\circ}C^{\text{-1}});$

T is mean air temperature (°C).

Psychrometric Constant (γ):

$$\gamma = \frac{c_p P}{\varepsilon \lambda} \tag{79}$$

Where:

 γ is psychrometric constant (kPa °C⁻¹);

P is atmospheric pressure (kPa);

 λ is latent heat of vaporization (MJ kg⁻¹);

c_n is air specific heat at constant pressure, 1.013 x 10^{-3} (MJ kg⁻¹ °C⁻¹);

 ε ratio molecular weight of water vapor/dry air, (ε = 0.622).

Atmospheric Pressure (P):

$$P = 101.3 \left(\frac{293 - 0,0065z}{293} \right)^{5.26} \tag{80}$$

Where:

P is atmospheric pressure (kPa);

z is elevation above sea level (m).

Wind Profile Relationship (u₂)

$$u_2 = u_z \frac{4.87}{\ln(67.8z - 5.42)} \tag{81}$$

Where

u, is wind speed at 2 m above ground surface (m s⁻¹);

 u_z is measured wind speed at z m above ground surface (m s⁻¹);

z is height of measurement above ground surface (m).

Air Density (ρ_a)

$$\rho_a = \frac{P}{T_{kv}R_d} \tag{82}$$

Where:

 ρ_a is air density (kg m⁻³);

P is atmospheric pressure (kPa);

 T_{vk} is virtual temperature (K);

 R_d is universal gas constant specific for dry air = 0.287 kJ kg⁻¹ K⁻¹.

Virtual Temperature (T_{vk})

$$T_{kv} = 1.01(T + 273) \tag{83}$$

Where:

 T_{ky} is virtual temperature (K);

T is mean air temperature (°C).

Surface resistance (r_.):

$$r_{\rm S} = \frac{r_l}{LAI_{active}} \tag{84}$$

Where:

r_s is surface resistance (s m⁻¹);

 r_1 is stomatal resistance of the well-illuminated leaf (s m^{-1});

LAI_{active} is active (sunlit) leaf area index (m² (leaf area) m⁻² (soil surface).

Leaf Area Index Active (LAI_{active})

$$LAI_{active} = 0.5LAI (85)$$

Where:

LAI_{active} * is active (sunlit) leaf area index (m² (leaf area) m⁻² (soil surface);

LAI is leaf area index (m² m⁻²).

*The LAI active is also depends on the leaf angle distribution (extinction coefficient).

Aerodynamic resistance (ra)

$$r_a = \frac{\ln\left[\frac{z_m - d}{z_{om}}\right] \cdot \ln\left[\frac{z_h - d}{z_{oh}}\right]}{k^2 \cdot u_z} \tag{86}$$

Where:

r_a is aerodynamic resistance (s m⁻¹);

 z_m is height of wind measurements (m);

 z_h is height of humidity measurements (m);

d is zero plane displacement height (m);

 z_{om} is roughness length governing momentum transfer (m);

 z_{oh} is roughness length governing transfer of heat and vapor (m);

k is von Karman's constant, 0.41 (dimensionless);

 u_z is wind speed at height z (m s⁻¹).

Zero Plane Displacement Height (d)

$$d = \frac{2}{3}h\tag{87}$$

Where:

d* is zero plane displacement height (m);

h is crop height (m).

*The d is also depends on canopy structure and openness.

Roughness Length (z)

$$z_{om} = 0.123h \tag{88}$$

$$z_{oh} = 0.1 z_{om} \tag{89}$$

Where:

h is crop height (m);

 z_{om} is roughness length governing momentum transfer (m);

 z_{oh} is roughness length governing transfer of heat and vapor (m).

Canopy conductance (inversion of Penman-Monteith)

Canopy conductance (g_c , m s⁻¹) was calculated by inverting the Penman-Monteith equation (Monteith, 1965):

$$\lambda ET = \frac{\Delta (Rn - G) + \rho_a c_p \frac{VPD}{r_a}}{\Delta + \gamma (1 + \frac{r_s}{r_a})}$$
(90)

Considering:

$$r_a = \frac{1}{g_a} \to g_a = \frac{1}{r_a} \tag{91}$$

$$r_{\rm S} = r_{\rm C} = \frac{1}{g_{\rm C}} \rightarrow g_{\rm C} = \frac{1}{r_{\rm C}} \tag{92}$$

$$\frac{r_s}{r_a} = \frac{r_c}{r_a} = \frac{\frac{1}{g_c}}{\frac{1}{g_a}} \to \frac{1}{g_c} \cdot \frac{g_a}{1} \to \frac{g_a}{g_c}$$
(93)

"The surface resistance (r_s) describes the resistance of vapor flow through stomata openings, total leaf area and soil surface" (Allen et al., 1998). Thus, here we are considering the approach of surface resistance (r_s) ~ canopy resistance (r_s) .

Then:

$$\lambda ET = \frac{\Delta(Rn - G) + \rho_a c_p VPD g_a}{\Delta + \gamma (1 + \frac{g_a}{g_c})}$$
(94)

$$\lambda ET\Delta + \lambda ET\gamma + \lambda ET\gamma \frac{g_a}{g_c} = \Delta (Rn - G) + \rho_a c_p VPD g_a$$
 (95)

$$\lambda ET(\Delta + \gamma) + \lambda ET\gamma \frac{g_a}{g_c} = \Delta (Rn - G) + \rho_a c_p VPD g_a$$
(96)

$$\lambda ET\gamma \frac{g_a}{g_c} = \Delta (Rn - G) + \rho_a c_p VPD g_a - \lambda ET(\Delta + \gamma)$$
(97)

$$g_c = \frac{\lambda ET\gamma g_a}{\Delta (Rn - G) + \rho_a c_p VPD g_a - \lambda ET(\Delta + \gamma)}$$
(98)

Canopy conductance (inversion of Penman-Monteith) - Simplified

Canopy conductance (g_c , m s⁻¹) was calculated by inverting the Penman-Monteith equation (Monteith, 1965). It can be expressed in its simplest form as follows:

First Equation (Noormets et al., 2010):

$$g_c = \frac{PETR_w}{\rho_a VPDR_d} \tag{99}$$

$$g_{c} = \frac{kPa * \frac{kg}{m^{2}.s} * \frac{J}{kg.K}}{\frac{J}{kg.K}} \to \frac{kPa * \frac{1}{m^{2}.s} * \frac{J}{1.K}}{\frac{1}{m^{3}} * kPa * \frac{J}{1.K}} \to \frac{kPa * J}{m^{2}.s * K} * \frac{m^{3} * K}{kPa * J} \to \frac{m}{s}$$
(100)

Where:

 g_c is canopy conductance (m s⁻¹);

P is atmospheric pressure (kPa);

ET is evapotranspiration (kg m⁻² s⁻¹);

 R_{w} is universal gas constant specific for water vapor (461.495 J kg⁻¹ K⁻¹);

 ρ_a is air density (kg m⁻³);

VPD is vapor pressure deficit (kPa);

 R_A is universal gas constant specific for dry air (287.058 J kg⁻¹ K⁻¹).

Second Equation (Mallick et al., 2013):

$$g_c = \frac{\gamma LE}{\rho_a c_p VPD} \tag{101}$$

$$g_{c} = \frac{\frac{kPa}{\circ C} * \frac{W}{m^{2}}}{\frac{kg}{m^{3}} * \frac{J}{kg \cdot C} * kPa} \rightarrow \frac{\frac{kPa}{\circ C} * \frac{J}{m^{2} \cdot S}}{\frac{1}{m^{3}} * \frac{J}{1 \cdot \circ C} * kPa} \rightarrow \frac{kPa * J}{\circ C * m^{2} \cdot S} * \frac{m^{3} * \circ C}{J * kPa} \rightarrow \frac{m}{s}$$

$$(102)$$

Where:

 g_c is canopy conductance (m s⁻¹);

γ is psychrometric constant (kPa °C⁻¹);

LE is latent heat flux (W m⁻²);

 ρ_a is air density (kg m⁻³);

 c_p is specific heat at constant pressure, 1.013 x 10^{-3} (MJ kg $^{-1}$ °C $^{-1}$) = 1013 (J kg $^{-1}$ °C $^{-1}$); VPD is vapor pressure deficit (kPa).

Considerations:

$$\frac{R_w}{R_d} = \frac{1}{\varepsilon} \tag{103}$$

$$g_c = \frac{PETR_w}{\rho_a VPDR_d} = \frac{\gamma LE}{\rho_a c_p VPD} \tag{104}$$

So, these two canopy conductance simplified equations are similarity.

Decoupling Factor (Ω)

The dimensionless decoupling factor (Ω) was calculated according to Jarvis & McNaughton (1986):

$$\Omega = \frac{\frac{\Delta}{\gamma} + 1}{\frac{\Delta}{\gamma} + 1 + \frac{g_a}{g_c}} \tag{105}$$

$$\Omega = \frac{\frac{kPa}{\frac{\circ C}{kPa}} + 1}{\frac{kPa}{\frac{\circ C}{c}} + 1 + \frac{m}{\frac{s}{s}}} \rightarrow dimensionless$$

$$\frac{\frac{kPa}{\frac{kPa}{\circ C}} + 1 + \frac{m}{\frac{s}{s}}}{\frac{kPa}{\frac{\circ C}{c}} + 1 + \frac{m}{\frac{m}{s}}}$$
(106)

Where:

 Ω is decoupling factor (dimensionless); Δ is slope vapor pressure curve (kPa °C⁻¹); γ is psychrometric constant (kPa °C⁻¹); g_a is aerodynamic conductance (m s⁻¹); g_c is canopy conductance (m s⁻¹).

The decoupling factor ranges from zero to one.

Bowen Ratio (β)

The Bowen ratio (β) is calculated by the ratio of sensible heat flux (H) and latent heat flux (LE) (Perez et al., 1999):

$$\beta = \frac{H}{LE} \tag{107}$$

Where:

 β is Bowen ration (dimensionless); H is sensible heat flux (W $m^{\text{-}2}$);

LE is latent heat flux (W m⁻²).

Considering the simplified energy balance equation:

$$R_n = LE + H + G \tag{108}$$

Where:

 R_n is net radiation (W m⁻²);

LE is latent heat flux (W m⁻²);

H is sensible heat flux (W m⁻²);

G is soil heat flux (W m^{-2}).

The H and LE can be expression according to energy balance equation:

$$\lambda LE = \frac{R_n - G}{1 + \beta} \tag{109}$$

$$H = \frac{\beta}{1+\beta} (R_n - G) \tag{110}$$

Considering T and VPD in two levels:

Measurements of the air temperature and vapor pressure gradients between two levels (1 and 2) within the surface layer, β can be obtained by following equation (Perez et al. 1999):

$$\beta = \gamma \frac{\delta T}{\delta e} \tag{111}$$

Considering:

$$\lambda LE = \frac{-\rho_a c_p}{\gamma} k_e \frac{\delta e}{\delta z} \tag{112}$$

$$H = -\rho_a c_p k_H \frac{\delta T}{\delta z} \tag{113}$$

$$\beta = \frac{-\rho_{a}c_{p}k_{H}\frac{\delta T}{\delta z}}{\frac{-\rho_{a}c_{p}}{\gamma}k_{e}\frac{\delta e}{\delta z}}$$
(114)

Assuming $k_{H} = k_{e}$

$$\beta = \frac{-\rho_a c_p k_H \frac{\delta T}{\delta z}}{\frac{-\rho_a c_p}{\nu} k_e \frac{\delta e}{\delta z}} \to \gamma \frac{\delta T}{\delta z} \frac{\delta z}{\delta e} \to \gamma \frac{\delta T}{\delta e} \to \gamma \frac{(T_2 - T_1)}{(e_{a1} - e_{a2})}$$
(115)

Where:

 $k_{_{\!\scriptscriptstyle H}}$ is turbulent transfer coefficient for heat;

 $k_{_{\it e}}$ is turbulent transfer coefficient for water;

 T_1 is air temperature level 1 (°C);

T₂ is air temperature level 2 (°C);

e_{a1} is air temperature level 1 (kPa);

e_{a2} is air temperature level 2 (kPa).

RESOURCE USE EFFICIENCY

WUE - Water-use efficiency (GPP/ET)

$$WUE = \frac{GPP}{ET} = \frac{gC}{kg.H_2O} \tag{116}$$

$$WUE = \frac{GPP}{ET} = \frac{gC}{L.H_2O} \tag{117}$$

Considering:

GPP = $g C m^{-2} time^{-1}$

ET = mm time⁻¹

 $1 \text{ mm} = 1 \text{ L m}^{-2}$

 $1 L H_2 O = 1 kg H_2 O$

$$WUE = \frac{GPP}{ET} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{mm}{time}} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{L.H_2O}{m^2.time}} \rightarrow \frac{gC}{m^2.time} * \frac{m^2.time}{L.H_2O} \rightarrow \frac{gC}{kg.H_2O}$$

$$(118)$$

$$WUE = \frac{GPP}{ET} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{mm}{time}} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{L.H_2O}{m^2.time}} \rightarrow \frac{gC}{m^2.time} * \frac{m^2.time}{L.H_2O} \rightarrow \frac{gC}{L.H_2O}$$

$$(119)$$

Examples:

Considering the time of 1 hour:

GPP = 1.5 g C m⁻² hour⁻¹

 $ET = 0.5 \text{ mm hour}^{-1}$

$$WUE = \frac{GPP}{ET} \to \frac{1.5}{0.5} \to 3 \frac{gC}{kg.H_2O}$$
(120)

$$WUE = \frac{GPP}{ET} \rightarrow \frac{1.5}{0.5} \rightarrow 3\frac{gC}{L \cdot H_2O}$$
(121)

 WUE_i - Intrinsic water-use efficiency (GPP/ g_i)

$$WUE_i = \frac{GPP}{g_c} = \frac{gC}{kg.H_2O} \tag{122}$$

$$WUE_i = \frac{GPP}{g_c} = \frac{gC}{L \cdot H_2 O} \tag{123}$$

Considering:

GPP = $g C m^{-2} time^{-1}$

 $g_c = m \text{ time}^{-1}$

 $1 \text{ kg H}_2\text{O} = 1 \text{ L H}_2\text{O} \ 10^{-3} \text{ m}^3 \text{ H}_2\text{O}$

$$WUE_{i} = \frac{GPP}{g_{c}} \rightarrow \frac{\frac{gC}{m^{2}, time}}{\frac{m}{time}} \rightarrow \frac{gC}{m^{2}, time} * \frac{time}{m, H_{2}O} \rightarrow \frac{gC}{m^{3}, H_{2}O} \rightarrow 10^{3} \frac{gC}{kg, H_{2}O}$$

$$(124)$$

$$WUE_{i} = \frac{GPP}{g_{c}} \rightarrow \frac{\frac{gC}{m^{2}.time}}{\frac{m}{time}} \rightarrow \frac{gC}{m^{2}.time} * \frac{time}{m.H_{2}O} \rightarrow \frac{gC}{m^{3}.H_{2}O} \rightarrow 10^{3} \frac{gC}{L.H_{2}O}$$

$$(125)$$

Examples:

Considering the time of 1 hour:

GPP = 1.5 g C m⁻² hour⁻¹

 $g_c = 0.01 \text{ m s}^{-1} \rightarrow 36 \text{ m hour}^{-1}$

$$WUE_i = \frac{GPP}{g_c} \to \frac{1.5}{36} \to 0.0416 * 10^3 \frac{gC}{kg.H_2O} \to 41.6 \frac{gC}{kg.H_2O}$$
 (126)

$$WUE_i = \frac{GPP}{g_c} \to \frac{1.5}{36} \to 0.0416 * 10^3 \frac{gC}{L.H_2O} \to 41.6 \frac{gC}{L.H_2O}$$
(127)

CUE - Carbon-use efficiency (NPP/GPP)

$$CUE = \frac{NPP}{GPP} \tag{128}$$

Considering:

GPP = $g C m^{-2} time^{-1}$

NPP = $g C m^{-2} time^{-1}$

$$CUE = \frac{NPP}{GPP} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{gC}{m^2.time}} \rightarrow \frac{gC}{m^2.time} * \frac{m^2.time}{gC} \rightarrow dimensionless$$
(129)

Example:

Considering the time of 1 hour:

GPP = 1.5 g C m⁻² hour⁻¹

NPP = 1.2 g C m^{-2} hour⁻¹

$$CUE = \frac{NPP}{GPP} \to \frac{1.2}{1.5} \to 0.8 \tag{130}$$

LUE - Light-use efficiency (GPP/PAR)

$$LUE = \frac{GPP}{PAR} \to \frac{gC}{molphotons}$$
 (131)

$$LUE = \frac{GPP}{PAR} \to \frac{molC}{molphotons}$$
(132)

Considering:

GPP = $mol m^{-2} time^{-1} CO_2$

PAR = mol m⁻² time⁻¹ photons

1 mol $CO_2 = 12 g C$

$$LUE = \frac{GPP}{PAR} \rightarrow \frac{\frac{mol}{m^2.time}CO_2}{\frac{mol}{m^2.time}photons} \rightarrow \frac{mol}{m^2.time}CO_2 * \frac{mol}{m^2.time}photons$$
(133)

$$\rightarrow \frac{molCO_2}{molphotons} \rightarrow 12 \cdot \frac{gC}{molphotons}$$
(134)

$$LUE = \frac{GPP}{PAR} \rightarrow \frac{\frac{mol}{m^2.time}CO_2}{\frac{mol}{m^2.time}photons} \rightarrow \frac{mol}{m^2.time}CO_2 * \frac{mol}{m^2.time}photons \rightarrow \frac{molCO_2}{molphotons}$$

$$(135)$$

Examples:

Considering the time of 1 hour:

GPP = $0.15 \text{ mol m}^{-2} \text{ time}^{-1} \text{ CO}_{2}$

PAR = 100 mol m⁻² time⁻¹ photons

$$LUE = \frac{GPP}{PAR} \to \frac{0.15}{100} * 12 \to 0.018 \frac{gC}{molphotons}$$
 (136)

$$LUE = \frac{GPP}{PAR} \rightarrow \frac{1.5}{100} \rightarrow 0.0015 \frac{molCO_2}{molphotons}$$
 (137)

RUE - Rain-use efficiency (GPP/PPT)

$$RUE = \frac{GPP}{PPT} \to \frac{gC}{kg.H_{20}}$$
(138)

$$RUE = \frac{GPP}{PPT} \to \frac{gC}{L.H_{20}} \tag{139}$$

Considering:

GPP = $g C m^{-2} time^{-1}$

 $ET = mm time^{-1}$

 $1 \text{ mm} = 1 \text{ L m}^{-2}$

 $1 L H_{2}O = 1 kg H_{2}O$

$$RUE = \frac{GPP}{PPT} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{mm}{time}} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{L.H_2O}{m^2.time}} \rightarrow \frac{gC}{m^2.time} * \frac{m^2.time}{L.H_2O} \rightarrow \frac{gC}{kg.H_2O}$$

$$(140)$$

$$RUE = \frac{GPP}{PTT} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{mm}{time}} \rightarrow \frac{\frac{gC}{m^2.time}}{\frac{L.H_2O}{m^2.time}} \rightarrow \frac{gC}{m^2.time} * \frac{m^2.time}{L.H_2O} \rightarrow \frac{gC}{L.H_2O}$$

$$(141)$$

Examples:

Considering the time of 1 hour:

GPP = 1.5 g C m⁻² hour⁻¹

PPT = 5 mm hour⁻¹

$$RUE = \frac{GPP}{PPT} \to \frac{1.5}{5} \to 0.3 \frac{gC}{kg.H_2O}$$
 (142)

$$RUE = \frac{GPP}{PTT} \to \frac{1.5}{5} \to 0.3 \frac{gC}{L \cdot H_2 O}$$
(143)

CARBON BALANCE

Conversion factor for carbon balance: (µmol CO₂ m⁻² s⁻¹) to (g C m⁻² time⁻¹)

$$\frac{\mu molCO_2}{m^2.s} \to \frac{gC}{m^2} \tag{144}$$

Conversion value of CO₂ considering the interval of 30 minutes (1800 seconds):

$$\frac{\mu molCO_2}{m^2.s} \rightarrow \frac{10^{-6} molCO_2}{m^2.\frac{1}{1800}} \rightarrow \frac{10^{-6} * 12gC}{m^2.\frac{1}{1800}} \rightarrow 10^{-6} * 12 * 1800 \frac{gC}{m^2} \rightarrow 0.0216 \frac{gC}{m^2}$$
(145)

Example:

$$10 \frac{\mu molCO_2}{m^2.s} \to 10 * 0.0216 \frac{gC}{m^2} \to 0.216 \frac{gC}{m^2}$$
 (146)

Conversion annual value of CO_2 (g m⁻² yr⁻¹ C) to (t ha⁻¹ yr⁻¹ C), 1 ha = 10000 m²:

$$\frac{gC}{m^2, year} \to \frac{tC}{ha, year} \tag{147}$$

$$\frac{gC}{m^2. year} \to 10^{-6} \frac{tC}{\frac{1}{10000}. year} \to \frac{1}{10^{-4}} * 10^{-6} \frac{tC}{ha. year} \to 10^{-2} \frac{tC}{ha. year}$$
(148)

Examples:

If the ecosystem was a carbon source:

$$800 \frac{gC}{m^2.year} \rightarrow 800 \cdot 10^{-2} \frac{tC}{ha.year} \rightarrow 8 \frac{tC}{ha.year}$$
(149)

If the ecosystem was a carbon sink:

$$-800\frac{gC}{m^2.year} \rightarrow -800 \cdot 10^{-2} \frac{tC}{ha.year} \rightarrow -8\frac{tC}{ha.year}$$
(150)

CO, storage

The CO₂ storage is deduced from CO₂ concentration (c) profile as describe in Aubinet et al. (2001):

$$Sc = \frac{Pa}{RT} \int_0^{hm} \frac{dc(z)}{dt} dz \tag{151}$$

$$Sc = \frac{Pa}{\frac{J}{K.mol}.K} * \frac{\mu mol}{mol} * \frac{m}{s} \rightarrow \frac{J}{m^3} * \frac{\mu mol}{J} * \frac{m}{s} \rightarrow \frac{\mu mol}{m^2.s}$$
(152)

Where:

Sc is storage flux (µmol m⁻² s⁻¹);

Pa is atmospheric pressure (Pa);

R is universal gas law constant (J mol⁻¹ K⁻¹);

T is air temperature (K);

hm is maximum measurement height (m);

c is CO₂ concentration (ppm or µmol mol⁻¹);

t is time (s);

z is height above ground surface (m);

dc/dt refers to the temporal change in ${\rm CO_2}$ concentration over a period of time;

dz is the height of the profile.

ECOPHYSIOLOGY CONVERSION UNITS

Assimilation rate (A) – (μ mol CO $_2$ m $^{-2}$ s $^{-1}$ \rightarrow mg CO $_2$ m $^{-2}$ s $^{-1}$)

$$\frac{\mu molCO_2}{m^2.s} \to \frac{mgCO_2}{m^2.s} \tag{153}$$

$$\frac{\mu molCO_2}{m^2.s} \to 10^{-6} \frac{molCO_2}{m^2.s} \to 44 * 10^{-6} \frac{gCO_2}{m^2} \to 44 * 10^{-6} * 10^3 \frac{mgCO_2}{m^2} \to 0.044 \frac{mgCO_2}{m^2.s}$$
(154)

Example:

$$20 \frac{\mu molCO_2}{m^2.s} \to 20 * 0.044 \frac{mgCO_2}{m^2.s} \to 0.88 \frac{mgCO_2}{m^2.s}$$
(155)

Transpiration rate (E) - (mmol $H_2O m^{-2} s^{-1} \rightarrow mg H_2O m^{-2} s^{-1}$)

$$\frac{mmolH_2O}{m^2.s} \to \frac{mgH_2O}{m^2.s} \tag{156}$$

$$\frac{mmolH_2O}{m^2.s} \to 18 \frac{mgH_2O}{m^2.s}$$
 (157)

Example:

$$5\frac{mmolH_2O}{m^2.s} \to 5*18\frac{mgH_2O}{m^2.s} \to 90\frac{mgH_2O}{m^2.s}$$
(158)

Transpiration rate (E) - (mmol $H_2O m^{-2} s^{-1} \rightarrow W m^{-2}$)

$$\frac{mmolH_2O}{m^2 \cdot s} \to \frac{W}{m^2} \tag{159}$$

$$\frac{mmolH_2O}{m^2.s} \to 10^{-3} \frac{molH_2O}{m^2.s} \to 10^{-3} * 18 \frac{gH_2O}{m^2.s} \to 10^{-3} * 18 * 2450 \frac{J}{m^2.s} \to 44.1 \frac{W}{m^2} \tag{160}$$

Example:

Considering:

- $.1 \text{ mol H}_{2}\text{O} = 18 \text{ g H}_{2}\text{O}$
- . To evaporate 1 g H₂O is necessary ~ 2450 J

$$5\frac{mmolH_2O}{m^2.s} \to 5*44.1 \frac{W}{m^2} \to 220.5 \frac{W}{m^2}$$
 (161)

Stomatal Conductance (g_s) - (mol m⁻² s⁻¹ \leftrightarrow m s⁻¹)

$$g_s\left(\frac{mol}{m^2s}\right) \leftrightarrow g_s\left(\frac{m}{s}\right)$$
 (162)

Simplified or approximation conversion

Through the stomata there are input and output of many gases, including: CO_2 , O_2 , water vapor and other gases. Thus, considering air temperature of 0 °C and atmospheric pressure of 101 325 Pa the molar density of gases is 44.6 mol m⁻³. Considering the ideal gas equation (Nobel 2009):

$$PV = nRT (163)$$

$$\frac{n}{V} = \frac{P}{RT} \tag{164}$$

$$\frac{mol}{m^3} = \frac{P}{RT} \tag{165}$$

$$\frac{n}{V} = \frac{101325}{8.314510 * (273.16 + 0)} \rightarrow \frac{101325}{2271.1915516} \rightarrow 44.6 \frac{mol}{m^3}$$
(166)

Then:

$$\frac{mol}{m^3} = \frac{P}{RT} = \frac{g_s\left(\frac{mol}{m^2s}\right)}{g_s\left(\frac{m}{s}\right)} \tag{167}$$

$$\frac{g_s\left(\frac{mol}{m^2s}\right)}{g_s\left(\frac{m}{s}\right)} = \frac{P}{RT} \tag{168}$$

$$\frac{g_s\left(\frac{mol}{m^2s}\right)}{g_s\left(\frac{m}{s}\right)} = 44.6 \frac{mol}{m^3} \tag{169}$$

$$g_s\left(\frac{mol}{m^2s}\right) = g_s\left(\frac{m}{s}\right) * 44.6\frac{mol}{m^3} \tag{170}$$

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2s}\right) * \frac{1}{44.6} \frac{m^3}{mol} \tag{171}$$

Examples:

(1) Stomatal conductance = $0.8 \text{ mol m}^{-2} \text{ s}^{-1}$

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2 \cdot s}\right) * \frac{1}{44.6} \frac{m^3}{mol} \tag{172}$$

$$g_s\left(\frac{m}{s}\right) = 0.8 * \frac{1}{44.6} = 0.018 \tag{173}$$

$$g_s\left(\frac{cm}{s}\right) = 0.8 * \frac{1}{44.6} = 0.018 * 100 = 1.8$$
 (174)

$$g_s\left(\frac{mm}{s}\right) = 0.8 * \frac{1}{44.6} = 0.018 * 1000 = 18$$
 (175)

(2) Stomatal conductance = 0.015 m s⁻¹

$$gs\left(\frac{mol}{m^2.s}\right) = gs\left(\frac{m}{s}\right) * 44.6 \frac{mol}{m^3} \tag{176}$$

$$gs\left(\frac{mol}{m^2.s}\right) = 0.015 * 44.6 = 0.669 \tag{177}$$

(3) Stomatal conductance = 1.5 cm s⁻¹

$$gs\left(\frac{mol}{m^2.s}\right) = gs\left(\frac{m}{s}\right) * \frac{1}{100} * 44.6 \frac{mol}{m^3}$$
(178)

$$gs\left(\frac{mol}{m^2.s}\right) = 1.5 * \frac{1}{100} * 44.6 = 0.669$$
(179)

(4) Stomatal conductance = 15 mm s⁻¹

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * \frac{1}{1000} * 44.6 \frac{mol}{m^3}$$
(180)

$$g_s\left(\frac{mol}{m^2.\,s}\right) = 15 * \frac{1}{1000} * 44.6 = 0.669 \tag{181}$$

Stomatal Conductance (g_s) - (mol m⁻² s⁻¹ \leftrightarrow m s⁻¹)

To calculate the stomatal conductance in metric dimensions, with other air temperature and atmospheric pressure conditions the ideal gas law comes in handy (Nobel, 2009):

$$PV = nRT (182)$$

$$\frac{n}{V} = \frac{P}{RT} \tag{183}$$

$$\frac{mol}{m^3} = \frac{P}{RT} \tag{184}$$

$$\frac{mol}{m^3} = \frac{P}{RT} \to \frac{Pa}{\frac{J}{mol,K} * K} \to \frac{Pa}{\frac{J}{mol}} \to Pa * \frac{mol}{J} \to \frac{J}{m^3} * \frac{mol}{J} \to \frac{mol}{m^3}$$
(185)

Where:

P is atmospheric pressure (Pa);

R is universal gas law constant (J mol⁻¹ K⁻¹);

T is air temperature (K).

Considering:

$$\frac{mol}{m^3} = \frac{g_s\left(\frac{mol}{m^2.s}\right)}{g_s\left(\frac{m}{s}\right)} \to \frac{mol}{m^2.s} * \frac{s}{m} \to \frac{mol}{m^3}$$
(186)

Then:

$$\frac{g_s\left(\frac{mol}{m^2 \cdot s}\right)}{g_s\left(\frac{m}{s}\right)} = \frac{P}{RT} \tag{187}$$

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * \frac{P}{RT} \tag{188}$$

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2 \cdot s}\right) * \frac{RT}{P} \tag{189}$$

Stomatal Conductance (g_e) - (mol m⁻² s⁻¹ \leftrightarrow cm s⁻¹)

$$g_s\left(\frac{mol}{m^2 \cdot s}\right) = g_s\left(\frac{cm}{s}\right) * \frac{1}{100} * \frac{P}{RT} \tag{190}$$

$$g_s\left(\frac{cm}{s}\right) = g_s\left(\frac{mol}{m^2 \cdot s}\right) * 100 * \frac{RT}{P}$$
(191)

Stomatal Conductance (g_s) - $(mol m^{-2} s^{-1} \leftrightarrow mm s^{-1})$

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{mm}{s}\right) * \frac{1}{1000} * \frac{P}{RT}$$
(192)

$$g_s\left(\frac{mm}{s}\right) = g_s\left(\frac{mol}{m^2 \cdot s}\right) * 1000 * \frac{RT}{P} \tag{193}$$

Example:

Considering:

- . Air temperature (T) = 25 °C;
- . Atmospheric pressure (P) = 101 325 Pa;
- . Universal gas law constant (R) = $8.314510 \, \text{J mol}^{-1} \, \text{K}^{-1}$.

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * \frac{P}{RT} \tag{194}$$

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * \frac{101325}{8.314510 * (273.16 + 25)}$$
(195)

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * \frac{101325}{2479.0543016} \tag{196}$$

$$g_s\left(\frac{mol}{m^2.s}\right) = g_s\left(\frac{m}{s}\right) * 40.87 \tag{197}$$

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2, s}\right) * \frac{RT}{P} \tag{198}$$

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2.s}\right) * \frac{8.314510 * (273.16 + 25)}{101325}$$
(199)

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2.s}\right) * \frac{2479.0543016}{101325}$$
(200)

$$g_s\left(\frac{m}{s}\right) = g_s\left(\frac{mol}{m^2 \cdot s}\right) * 0.0245 \tag{201}$$

GENERAL CONVERSION UNITS FOR TRACE GASES AND WATER Conversion factor for ${\rm CO_2}$ variable: molar to mass density

$$\frac{mmol}{m^3}CO_2 \to \frac{mg}{m^3}CO_2 \tag{202}$$

$$\frac{mmol}{m^3}CO_2 \to 44\frac{mg}{m^3}CO_2 \tag{203}$$

Example:

$$15\frac{mmol}{m^3}CO_2 \to 15*44\frac{mg}{m^3}CO_2 \to 660\frac{mg}{m^3}CO_2 \tag{204}$$

Conversion factor for CO₂ variable: mass to molar density

$$\frac{mg}{m^3}CO_2 \to \frac{mmol}{m^3}CO_2 \tag{205}$$

$$\frac{mg}{m^3}CO_2 \to \frac{1}{44} \frac{mmol}{m^3}CO_2 \to 0.0227 \frac{mmol}{m^3}$$
(206)

Example:

$$660 \frac{mg}{m^3} CO_2 \to 660 * 0.0227 \frac{mmol}{m^3} CO_2 \to 15 \frac{mmol}{m^3} CO_2$$
 (207)

Conversion factor for H₂O variable: molar to mass density

$$\frac{mmol}{m^3}H_2O \to \frac{g}{m^3}H_2O \tag{208}$$

$$\frac{mmol}{m^3}H_2O \to \frac{18}{1000}\frac{g}{m^3}H_2O \to 0.018\frac{g}{m^3}H_2O \tag{209}$$

Example:

$$1000 \frac{mmol}{m^3} H_2 O \to 1000 * 0.018 \frac{g}{m^3} H_2 O \to 18 \frac{g}{m^3} H_2 O$$
 (210)

Conversion factor for H₂O variable: mass to molar density

$$\frac{g}{m^3}H_2O \to \frac{mmol}{m^3}H_2O \tag{211}$$

$$\frac{g}{m^3} H_2 O \to \frac{1000}{18} \frac{mmol}{m^3} H_2 O \to 55.55 \frac{mmol}{m^3} H_2 O \tag{212}$$

Example:

$$18\frac{g}{m^3}H_2O \to 18*55.55\frac{mmol}{m^3}H_2O \to 1000\frac{mmol}{m^3}H_2O \tag{213}$$

Conversion factor for CH_4 variable: molar to mass density

$$\frac{mmol}{m^3}CH_4 \to \frac{mg}{m^3}CH_4 \tag{214}$$

$$\frac{mmol}{m^3}CH_4 \to 16\frac{mg}{m^3}CH_4 \tag{215}$$

Example:

$$0.15 \frac{mmol}{m^3} CH_4 \to 0.15 * 16 \frac{mg}{m^3} CH_4 \to 2.4 \frac{mg}{m^3} CH_4$$
 (216)

Conversion factor for CH₄ variable: mass to molar density

$$\frac{mg}{m^3}CH_4 \to \frac{mmol}{m^3}CH_4 \tag{217}$$

$$\frac{mg}{m^3}CH_4 \to \frac{1}{16}\frac{mmol}{m^3}CH_4 \to 0.0625\frac{mmol}{m^3}CH_4$$
 (218)

Example:

$$2.4 \frac{mg}{m^3} CH_4 \rightarrow 2.4 * 0.0625 \frac{mmol}{m^3} CH_4 \rightarrow 0.15 \frac{mmol}{m^3} CH_4$$
 (219)

Conversion factor for NO2 variable: molar to mass density

$$\frac{mmol}{m^3}NO_2 \to \frac{mg}{m^3}NO_2 \tag{220}$$

$$\frac{mmol}{m^3}NO_2 \to 46\frac{mg}{m^3}NO_2 \tag{221}$$

Example:

$$0.0250 \frac{mmol}{m^3} NO_2 \to 0.0250 * 46 \frac{mg}{m^3} NO_2 \to 1.15 \frac{mg}{m^3} NO_2$$
 (222)

Conversion factor for NO2 variable: mass to molar density

$$\frac{mg}{m^3}NO_2 \to \frac{mmol}{m^3}NO_2 \tag{223}$$

$$\frac{mg}{m^3}NO_2 \to \frac{1}{46} \frac{mmol}{m^3}NO_2 \to 0.02174 \frac{mmol}{m^3}NO_2 \tag{224}$$

Example:

$$1.15 \frac{mg}{m^3} NO_2 \rightarrow 1.15 * 0.02174 \frac{mmol}{m^3} NO_2 \rightarrow 0.0250 \frac{mmol}{m^3} NO_2$$
 (225)

Conversion factor for O₃ variable: molar to mass density

$$\frac{mmol}{m^3}O_3 \to \frac{mg}{m^3}O_3 \tag{226}$$

$$\frac{mmol}{m^3}O_3 \to 48\frac{mg}{m^3}O_3 \tag{227}$$

Example:

$$0.015 \frac{mmol}{m^3} O_3 \to 0.015 * 48 \frac{mg}{m^3} O_3 \to 0.72 \frac{mg}{m^3} O_3$$
(228)

Conversion factor for O₃ variable: mass to molar density

$$\frac{mg}{m^3}O_3 \to \frac{mmol}{m^3}O_3 \tag{229}$$

$$\frac{mg}{m^3}O_3 \to \frac{1}{48} \frac{mmol}{m^3}O_3 \to 0.02083 \frac{mmol}{m^3}O_3 \tag{230}$$

Example:

$$0.72 \frac{mg}{m^3} O_3 \to 0.72 * 0.02083 \frac{mmol}{m^3} O_3 \to 0.015 \frac{mmol}{m^3} O_3 \tag{231}$$

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Fatores de conversão e equações gerais aplicadas em meteorologia agrícola e florestal

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INFORMAÇÕES

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RESUMO

A comunicação científica em meteorologia agrícola e florestal exige uma conversão correta de unidades e o uso adequado de equações. Em geral, a maioria das conversões e equações são encontradas separadamente em um livro, artigo ou documento específico. Assim, propomos aqui um único documento com unidades de conversão de radiação solar, evapotranspiração, eficiência no uso de recursos, balanço de carbono, ecofisiologia, gases traços e água; e equações gerais: Penman-Monteith, condutância do dossel, fator de desacoplamento e razão de Bowen. Este documento tem muito valor para o ensino, para os cientistas e como uma referência rápida.

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