# Modern Physics Final Study Guide

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## 1 Special Relativity

## 1.1 Postulates of Special Relativity

#### Postulates:

- The laws of physics are the same in all inertial frames.
- The speed of light is constant in a vacuum for all observers, regardless of their motion relative to the source.

## 1.2 Relativity of Simultaneity

**Relativity of Simultaneity:** Two events that appear simultaneous to one observer may not appear simultaneous to another observer moving relative to the first. This is a direct consequence of the constancy of the speed of light and the Lorentz transformations.

#### 1.3 Lorentz Transformations

**Lorentz Transformations:** The Lorentz transformations relate spacetime coordinates of an event in two inertial reference frames moving at a constant velocity relative to each other. These transformations preserve the invariant interval  $(\Delta s^2)$ , a cornerstone of special relativity.

#### **Transformations:**

- For a frame S' moving with velocity v along the x-axis relative to S:
  - Time:  $t' = \gamma \left(t \frac{vx}{c^2}\right)$
  - Space (x-direction):  $x' = \gamma (x vt)$
  - Space (y- and z-directions): y' = y, z' = z
- The Lorentz factor is  $\gamma = \frac{1}{\sqrt{1 v^2/c^2}}$ .

**Lorentz Transformation Matrix:** The Lorentz transformation can be expressed using matrix multiplication:

$$x' = \Lambda x$$
,

where  $x = \begin{bmatrix} ct \\ x \\ y \\ z \end{bmatrix}$  and x' are the spacetime coordinates in S and S', respectively. For a boost

along the x-axis, the Lorentz transformation matrix  $\Lambda$  is:

$$\Lambda = egin{bmatrix} \gamma & -\gamma \beta & 0 & 0 \ -\gamma \beta & \gamma & 0 & 0 \ 0 & 0 & 1 & 0 \ 0 & 0 & 0 & 1 \end{bmatrix},$$

where  $\beta = v/c$ .

#### **Key Properties:**

• The inverse transformation is obtained by negating v, flipping the sign of  $\beta$ :

$$\Lambda^{-1} = \begin{bmatrix} \gamma & \gamma\beta & 0 & 0 \\ \gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}.$$

• The Lorentz transformations preserve the invariant interval:

$$\Delta s^2 = x \cdot x = x' \cdot x'.$$

## **Applications:**

• Lorentz transformations explain phenomena like time dilation, length contraction, and the relativity of simultaneity.

• They form the basis for defining and transforming four-vectors, such as four-momentum and four-velocity.

#### 1.4 The Relativistic Invariant Interval

**Invariant Interval:** The relativistic invariant interval is a spacetime quantity that remains unchanged under Lorentz transformations.

Invariant Interval:

$$\Delta s^2 = (c\Delta t)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2$$

#### Classifications of the Invariant Interval:

• Time-like:  $\Delta s^2 > 0$ , related to proper time:

$$\Delta s^2 = c^2 \Delta \tau^2$$

• Space-like:  $\Delta s^2 < 0$ , related to proper length:

$$-\Delta s^2 = L^2$$

• **Light-like:**  $\Delta s^2 = 0$ , events are connected by light:

$$c^2 \Delta t^2 = (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2$$

## 1.5 Relativistic Velocity Addition

Relativistic Velocity Addition: When combining velocities in different inertial frames, the relativistic formula must be used. Typically, these problems involve 3 objects and the following reference frames and velocities are defined:

- $\bullet$  S: The stationary reference frame.
- S': The moving reference frame.
- u: The velocity of an object relative to S.
- v: The velocity of S' relative to S.
- u': The velocity of the object relative to S'.

#### Relativistic Velocity Addition Formula:

$$u' = \frac{u+v}{1+\frac{uv}{c^2}}$$

Here, u' is the velocity of the object as measured in S', given the velocity u of the object relative to S and the velocity v of S' relative to S.

#### 1.6 Four-Vectors in Special Relativity

**Four-Vectors:** Physical quantities in relativity are represented as four-vectors that transform under Lorentz transformations.

#### General Form of a Four-Vector:

$$a^{\mu} = (a^0, \vec{a}) = (a^0, a^1, a^2, a^3), \quad a^0 = c\Delta t$$

**Dot Product of Four-Vectors:** The invariant dot product between two four-vectors  $a^{\mu}$  and  $b^{\mu}$  is:

$$a \cdot b = a^0 b^0 - \vec{a} \cdot \vec{b}$$

This quantity is Lorentz invariant (unchanged under Lorentz transformations).

## 1.7 Examples of Four-Vectors

#### **Common Four-Vectors:**

• Four-Position:

$$x^{\mu}=(ct,x,y,z)$$

• Four-Velocity:

$$u^{\mu} = \frac{dx^{\mu}}{d\tau} = \gamma(c, \vec{v}), \quad \gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$$

• Four-Momentum:

$$p^\mu = m u^\mu = \gamma m(c, \vec{v})$$

For massless particles, like photons:

$$p \cdot p = 0, \quad E = |\vec{p}|c$$

#### 1.8 Relativistic Conservation Laws

Relativistic Conservation Laws: The principles of conservation of energy and momentum extend to special relativity, incorporating the total energy (rest energy + kinetic energy) and relativistic momentum.

#### Relativistic Total Energy:

 $\bullet$  The total energy E includes both the rest energy and the relativistic kinetic energy:

$$E = \gamma m_0 c^2.$$

- The rest energy is given by  $E_0 = m_0 c^2$ , where  $m_0$  is the rest mass.
- In any physical process, the total energy is conserved.

#### Relativistic Momentum:

• The relativistic momentum is defined as:

$$\vec{p} = \gamma m_0 \vec{v},$$

where 
$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$$
.

• Both the magnitude and direction of total momentum in an isolated system are conserved.

#### Four-Momentum Conservation:

• Energy and momentum form components of the four-momentum vector:

$$p^{\mu} = \left(\frac{E}{c}, \vec{p}\right).$$

• In all inertial frames, the conservation of four-momentum holds:

$$\sum p^{\mu}_{\rm initial} = \sum p^{\mu}_{\rm final}.$$

#### Relativistic Collisions:

- Total energy and total momentum are conserved in collisions, including elastic and inelastic cases.
- Unlike classical mechanics, the concept of conserved "mass" is not generally applicable. For example:
  - In pair annihilation, rest mass converts to energy.
  - In particle creation, energy converts into rest mass.

## **Energy-Momentum Relation:**

$$E^2 = (pc)^2 + (m_0c^2)^2$$

This relation connects the total energy E, the momentum p, and the rest mass  $m_0$  of a particle.

## Invariant Four-Momentum Magnitude:

$$p^\mu \cdot p^\mu = m_0^2 c^2$$

This invariant magnitude holds in all inertial frames, ensuring the consistency of relativistic conservation laws.

## 2 Quantum Mechanics

## 2.1 Quantum Theory of Light

## 2.1.1 Blackbody Radiation and Planck Hypothesis

Blackbody Radiation: Classical physics predicted the "ultraviolet catastrophe," where energy radiated at high frequencies diverged. Planck resolved this by quantizing energy:

$$E = nhf, \quad n = 1, 2, 3, \dots$$

Planck's Law:

$$I(\nu,T) = \frac{8\pi h \nu^3}{c^3} \frac{1}{e^{h\nu/k_BT}-1}, \label{eq:interpolation}$$

where  $I(\nu, T)$  is the intensity of radiation at frequency  $\nu$ .

#### 2.1.2 Photoelectric Effect

**Photoelectric Effect:** Electrons are emitted from a material when light shines on it, with the energy of the electrons depending on the frequency of light.

$$hf = \phi + KE_{max}, \quad hf_c = \phi, \quad KE_{max} = eV_s$$

Threshold Frequency: No electrons are emitted if  $f < f_c$ .

## 2.1.3 Compton Scattering

**Compton Scattering:** The scattering of X-rays by electrons demonstrates particle-like behavior of light. The wavelength shift is given by:

$$\lambda' - \lambda = \frac{h}{m_e c} (1 - \cos \theta).$$

**Energy Relation:** 

$$\frac{1}{E'} - \frac{1}{E} = \frac{1 - \cos \theta}{m_e c^2}.$$

#### 2.1.4 Wave-Particle Duality

Wave-Particle Duality: Light and matter exhibit dual behavior, displaying characteristics of both waves and particles depending on the experimental conditions.

#### **Key Features:**

- Travels as a wave, interacting with itself, and demonstrates phenomena such as interference and diffraction.
- Interacts as a particle, transferring discrete packets of energy, known as photons or quanta.
- Confirmed by experiments:
  - Double-Slit Experiment: Demonstrates interference patterns for light and electrons, showing wave-like behavior.
  - Photoelectric Effect: Demonstrates the particle nature of light, as photons eject electrons from a material.
  - Blackbody Radiation: Explained by Planck's quantization hypothesis, resolving the ultraviolet catastrophe.
  - Compton Scattering: Demonstrates photon momentum through scattering with electrons.

De Broglie Wavelength:  $\lambda = \frac{h}{p} = \frac{h}{mv}$ 

Wave Number:  $k = \frac{2\pi}{\lambda}$ 

#### 2.1.5 Heisenberg Uncertainty Principle

Heisenberg Uncertainty Principle: A fundamental principle of quantum mechanics stating that certain pairs of physical observables cannot be simultaneously measured with arbitrary precision.

#### **Key Relations:**

• Position-Momentum Uncertainty:

$$\Delta x \Delta p \ge \frac{\hbar}{2},$$

where  $\Delta x$  and  $\Delta p$  are the standard deviations of position and momentum, respectively.

• Energy-Time Uncertainty:

$$\Delta E \Delta t \ge \frac{\hbar}{2},$$

where  $\Delta E$  and  $\Delta t$  are the uncertainties in energy and time, respectively.

Variance-Squared Average Relationship: The uncertainty in a measurable quantity A is defined as:

$$\Delta A = \sqrt{\langle A^2 \rangle - \langle A \rangle^2},$$

where  $\langle A \rangle$  is the expectation value, and  $\langle A^2 \rangle$  is the expectation value of  $A^2$ .

#### **Implications:**

- Establishes the probabilistic nature of quantum mechanics, replacing deterministic classical physics.
- Demonstrates the impossibility of assigning definite trajectories to particles as in classical mechanics.

## 2.1.6 Group and Phase Velocities

Group and Phase Velocities: Velocities associated with wave propagation.

• Group velocity:

$$v_g = \frac{d\omega}{dk},$$

representing the velocity of the wave packet and the velocity of the particle

• Phase velocity:

$$v_p = \frac{\omega}{k},$$

representing the velocity of individual wave crests.

## 2.2 Schrödinger Equation

#### 2.2.1 Overview

Schrödinger Equation: Governs the quantum mechanical behavior of particles. It describes how the wave function of a particle evolves under the influence of potential energy. General Forms:

• 1D Time-Independent Schrödinger Equation:

$$-\frac{\hbar^2}{2m}\frac{d^2\psi(x)}{dx^2}+V(x)\psi(x)=E\psi(x).$$

• 3D Schrödinger Equation (spherical coordinates):

$$-\frac{\hbar^2}{2m}\nabla^2\psi(r,\theta,\phi) + V(r)\psi(r,\theta,\phi) = E\psi(r,\theta,\phi).$$

The solution is separable:

$$\psi(r, \theta, \phi) = R(r)Y(\theta, \phi),$$

where  $Y(\theta, \phi)$  are spherical harmonics.

#### 2.2.2 Different Potentials

#### Free Particle:

• For V(x) = 0, solutions are plane waves:

$$\psi(x) = Ae^{i(kx - \omega t)},$$

where  $k = \frac{p}{\hbar}$  and  $\omega = \frac{E}{\hbar}$ .

#### Particle in a Box:

• For V(x) = 0 inside the box and  $V(x) = \infty$  outside:

$$\psi_n(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right), \quad E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2}.$$

#### Step Potential:

- E < U: Tunneling occurs. The wave function decays exponentially in the classically forbidden region.
- E > U: The wave function is continuous, with partial reflection and transmission.

## 2.3 Hydrogen Atom

## 2.3.1 Quantized Energy Levels

**Hydrogen Atom:** The Schrödinger equation in spherical coordinates yields quantized energy levels and quantum numbers.

**Energy Levels:** 

$$E_n = -\frac{13.6 \,\text{eV}}{n^2}.$$

#### 2.3.2 Quantum Numbers

## Quantum Numbers:

- Principal quantum number (n): Determines energy level (n = 1, 2, 3, ...).
- Orbital angular momentum quantum number (l): Determines the shape of the orbital (l = 0, 1, ..., n 1).
- Magnetic quantum number  $(m_l)$ : Determines the orientation  $(m_l = -l, -l + 1, ..., l)$ .
- Spin quantum number  $(m_s)$ : Intrinsic angular momentum  $(m_s = \pm \frac{1}{2})$ .

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#### 2.3.3 Wave Functions

Wave Functions:

$$\psi(r, \theta, \phi) = R_{nl}(r)Y_l^m(\theta, \phi),$$

where  $R_{nl}(r)$  is the radial wave function, and  $Y_l^m(\theta, \phi)$  are spherical harmonics. **Radial Probability Distribution:** The probability density P(r) is given by:

$$P(r) = r^2 |R_{nl}(r)|^2$$
.

## 3 Statistical Mechanics

## 3.1 Key Concepts

- Microstates and macrostates: Connection between microscopic and macroscopic descriptions.
- Boltzmann distribution:  $P_i = \frac{e^{-E_i/k_BT}}{Z}$ .
- Partition function: Summation of states to understand thermodynamic properties,  $Z = \sum_{i} e^{-E_i/k_BT}$ .
- Occupation numbers: Average number of particles in a state,  $n_i = \frac{1}{e^{(E_i \mu)/k_B T} \pm 1}$  (for fermions and bosons).
- Degeneracy: The number of microstates corresponding to a single energy level.
- Distinguishable vs. indistinguishable particles: Boltzmann distribution for distinguishable particles; Fermi-Dirac and Bose-Einstein distributions for indistinguishable particles.

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### 3.2 Essential Equations

Boltzmann distribution:  $P_i = \frac{e^{-E_i/k_BT}}{Z}$ 

Partition function:  $Z = \sum_{i} g_i e^{-E_i/k_B T}$ 

Fermi-Dirac distribution:  $n_i = \frac{1}{e^{(E_i - \mu)/k_B T} + 1}$ 

Bose-Einstein distribution:  $n_i = \frac{1}{e^{(E_i - \mu)/k_B T} - 1}$ 

Average energy:  $\langle E \rangle = \sum_{i} P_{i} E_{i} = -\frac{\partial \ln Z}{\partial \beta}$ 

Probability of a state:  $P(E) = \frac{g_i e^{-E/k_B T}}{Z}$ 

#### 3.3 Units and Unit Conversions

- Energy:  $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}.$
- Temperature: Convert between Kelvin and energy using  $k_B = 1.38 \times 10^{-23}$  J/K.
- Wavelength and frequency:  $E=hf,\,\lambda=\frac{c}{f},\,$  with hc=1240 eV  $\cdot$  nm.

## 4 Quick Reference

#### 4.1 Mathematical Tools

• Trigonometric Identities:

$$\cos^2 x = \frac{1 + \cos(2x)}{2}$$

$$\sin^2 x = \frac{1 - \cos(2x)}{2}$$

 $\sin(2x) = 2\sin x \cos x$ 

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- Series:
  - Taylor Series Expansion:  $\sum_{n=0}^{\infty} C_n (x-a)^n \text{ where } C_n = \frac{f^{(n)}(a)}{n!}$
  - Binomial Series Expansion:  $(1 \pm x)^r = 1 \pm rx \pm ...$
- Integration Tips:
  - For even functions:  $\int_{-a}^a f(x) dx = 2 \int_0^a f(x) dx$
  - For odd functions:  $\int_{-a}^{a} f(x)dx = 0$
  - Gaussian integrals often appear in wave packets
- Complex Numbers Review
  - Complex conjugate:  $z^* = a bi$  for z = a + bi
  - Modulus:  $|z| = \sqrt{a^2 + b^2}$
  - Euler's formula:  $e^{i\theta} = \cos \theta + i \sin \theta$
  - Argument:  $\arg(z) = \tan^{-1}\left(\frac{y}{x}\right)$
- Second Order Differential Equations
  - If ODE has this form:  $\frac{d^2f}{dx^2} = k^2f$  where k is a constant The solution is:

 $f(x) = Ae^{kx} + Be^{-kx}$ 

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## 4.2 Important Constants

- Stefan-Boltzmann constant:  $\sigma = 5.67 \times 10^{-8} \text{ W/m}^2 \cdot K^4$
- Planck's constant:  $h = 6.626 \times 10^{-34} \,\mathrm{J\cdot s} = 4.136 \times 10^{-15} \,\mathrm{eV\cdot s}$
- Electron mass:  $m_e = 9.11 \times 10^{-31} \,\mathrm{kg}$
- Speed of light:  $c = 3.0 \times 10^8 \,\mathrm{m/s}$
- Bohr radius:  $a_0 = 0.0529 \times 10^{-9} \,\mathrm{m}$
- Compton wavelength of the electron:  $\frac{h}{m_e c} = 0.002426 \ \mathrm{nm}$

## 4.3 Unit Conversions

- $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$
- $hc = 1240 \text{ eV} \cdot \text{nm}$
- $\hbar c = \frac{1240}{2\pi} \text{ eV} \cdot \text{nm}$