

# Tenta-kit SSY085

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## 1 General

Digital signal power:  $S = E_b \cdot R_b$

Noise power:  $N = n_0 \cdot B$

Shannons formula:

$$\frac{C}{B} = \log_2(1 + SNR)$$

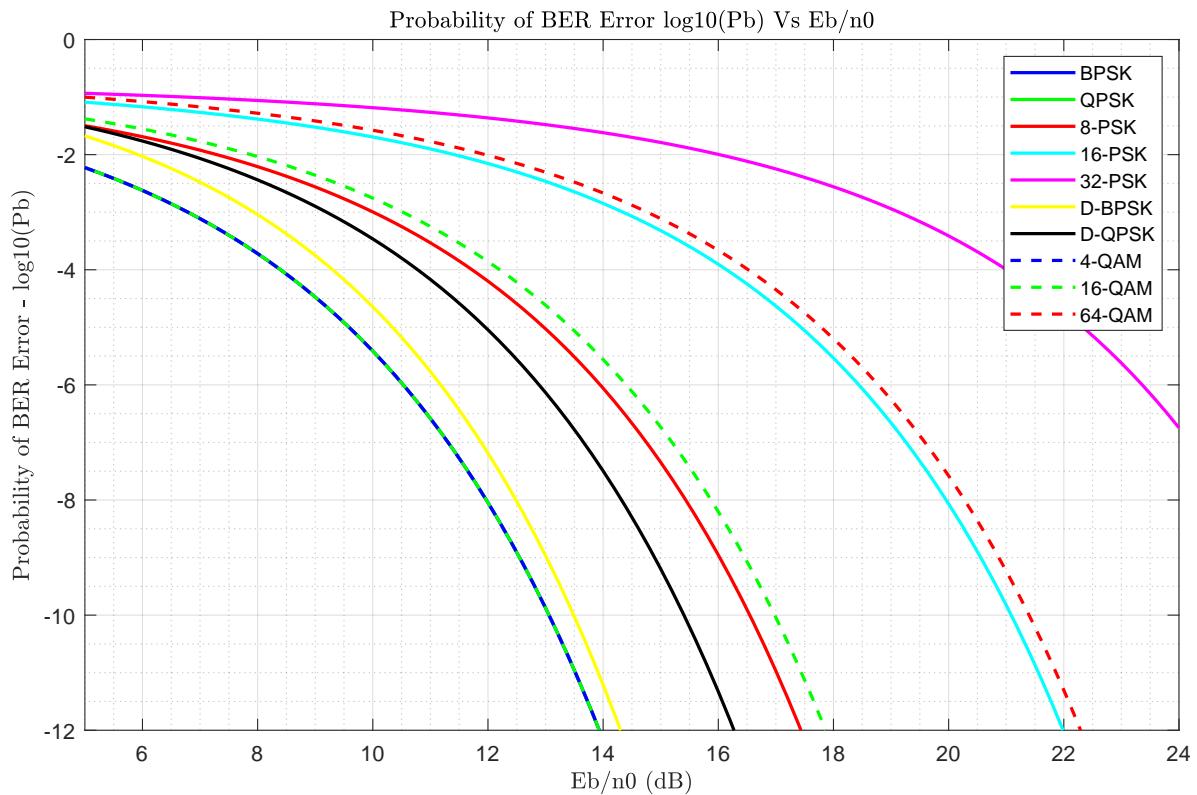
$$SNR = \frac{E_b}{n_0} \frac{R_b}{B}$$

$$\text{FSK: } 2 \cdot P_e = \operatorname{erfc} \left( \sqrt{\frac{E_b}{2n_0}} \right)$$

$$\text{PSK: } 2 \cdot P_e = \operatorname{erfc} \left( \sqrt{\frac{E_b}{n_0}} \right)$$

Bit error rates:

$$\text{ASK: } 2 \cdot P_e = \operatorname{erfc} \left( \sqrt{\frac{E_b}{4n_0}} \right)$$



## Noise Figure

- Noise figure is used to characterize the degradation in SNR for a receiver

$$F = \frac{\left(\frac{S}{N}\right)_{in}}{\left(\frac{S}{N}\right)_{out}} = \frac{S_{in} \cdot k(T_0 + T_e)B}{kT_0 B \cdot S_{in}} = 1 + \frac{T_e}{T_0} \geq 1$$

$$T_e = (F-1)T_0$$

- Noise figure depends on  $T_0$ 
  - Standard uses  $T_0 \equiv 290$  K

## S/N in digital communication systems

- Digital signal power,  $S = E_b \cdot R_b$

–  $R_b$  = Bitrate (bps)

- Noise power,  $N = n_0 \cdot B$

–  $B$  = Bandwidth (Hz)

- SNR vs.  $E_b/n_0$

$$\frac{S}{N} = \frac{E_b}{n_0} \frac{R_b}{B}$$

- Spectral efficiency

- Depends on modulation format
- Typically 1–4 (Lecture 3)

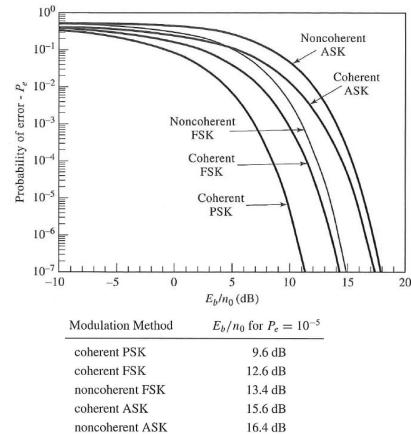
## Theoretical bit error probabilities

- Assuming Gaussian noise, the error probabilities ( $P_e = \text{BER}$ ) can be derived theoretically:

|     | Coherent demodulation  | Envelope demodulation   |
|-----|--|---|
| ASK | $P_e = \frac{1}{2} \operatorname{erfc} \left( \sqrt{\frac{E_b}{4n_0}} \right)$ | $P_e = \frac{1}{2} e^{-\frac{v^2}{2\sigma^2}} + \frac{1}{2} \int_0^{v_0} \frac{r}{\sigma^2} e^{-\frac{(v^2+r^2)}{2\sigma^2}} I_0 \left( \frac{vr}{\sigma^2} \right) dr$ |
| FSK | $P_e = \frac{1}{2} \operatorname{erfc} \left( \sqrt{\frac{E_b}{2n_0}} \right)$ | $P_e = \frac{1}{2} e^{-E_b/2n_0}$   |
| PSK | $P_e = \frac{1}{2} \operatorname{erfc} \left( \sqrt{\frac{E_b}{n_0}} \right)$  | Not possible  |

$$\operatorname{erfc} x = \frac{2}{\pi} \int_x^{\infty} e^{-t^2} dt$$

## Comparison of binary modulation techniques



- ASK
  - Very simple demodulator
  - Common in optical systems
- FSK
  - Acceptable performance even with envelope demod
- PSK
  - Highest performance
  - Requires synchronized oscillator
  - Used in GPS

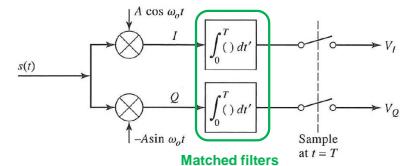
## Pulse shaping filters

- Reduces the spectral widening of a digital pulse train without distorting the message
  - More efficient spectral usage
  - Designed to have zero inter-symbol-interference
- Raised cosine filters typically used



- Roll-off factor parameter ( $\alpha$  or  $\beta$ , typ = 0.25)
- Often split into two root-raised cosine filters (at TX and RX)
- How is the transmitted signal affected by the filtering?

## QPSK demodulator



- Error probability

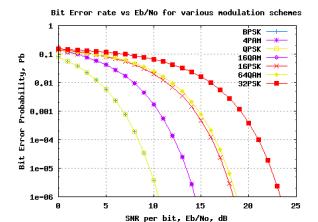
$$P_e = \frac{1}{2} \operatorname{erfc} \left( \sqrt{\frac{E_b}{n_0}} \right)$$

- Same as for BPSK

- QPSK has twice the bitrate compared to BPSK
- QPSK has twice the power (cos + sin modulation)

## Performance comparison for digital modulation formats

| Modulation Type | $E_b/n_0$ (dB) for $P_e = 10^{-5}$ | Bandwidth Efficiency |
|-----------------|------------------------------------|----------------------|
| binary ASK      | 15.6                               | 1                    |
| binary FSK      | 12.6                               | 1                    |
| binary PSK      | 9.6                                | 1                    |
| QPSK (4-QAM)    | 9.6                                | 2                    |
| 8-PSK           | 13.0                               | 3                    |
| 16-PSK          | 18.7                               | 4                    |
| 16-QAM          | 13.4                               | 4                    |
| 64-QAM          | 17.8                               | 6                    |



- There may be other tradeoffs

- PSK has constant amplitude → High transmitter efficiency and small linearity problems
- Receiver complexity

- Clearly, there is a tradeoff between bandwidth efficiency and  $E_b/n_0$  required. How can this be understood?

## Channel capacity

- Above a certain  $E_b/n_0$ , the bit error probability → 0
- The corresponding maximum bitrate for error-free transmission is given by:
- Shannon's channel capacity theorem

$$C = B \log_2 \left( 1 + \frac{S}{N} \right) = B \log_2 \left( 1 + \frac{E_b}{n_0} \frac{R_b}{B} \right)$$

C Max. bitrate for error free transmission (bps)

B Channel bandwidth (Hz)

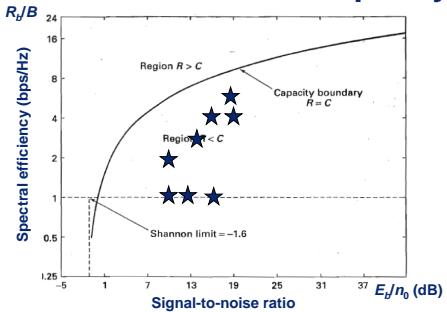
S Signal power (W)

N Noise power =  $n_0 \cdot B$  (W)

$R_b$  Bitrate (bps)

$E_b$  Bit energy (J)

## Channel capacity



| Modulation Type | $E_b/n_0$ (dB) for $P_e = 10^{-5}$ | Bandwidth Efficiency |
|-----------------|------------------------------------|----------------------|
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| 64-QAM          | 17.8                               | 6                    |

- Upper limit of spectral efficiency for a given  $E_b/n_0$
- Compare w. M-QAM for  $P_e = 10^{-5}$ 
  - Higher order modulation schemes utilize spectrum better, but require higher SNR = higher output power

## 2 Wireless

### 2.1 Antennas

Far-field:

Big antennas:

$$R_{ff} = \frac{2 \cdot D}{\lambda}$$

Small antennas:

$$R_{ff} = 2\lambda$$

Other equations:

$$P_r = S_{avg} \cdot A_e$$

( $S_{avg}$  power density,  $A_e$  effective aperture area)

$$A_e = \frac{D \cdot \lambda^2}{4\pi}$$

( $D$  directivity)

$$G = e_{rad} \cdot D$$

$$e_{rad} = \frac{P_{rad}}{P_{in}} = 1 - \frac{P_{loss}}{P_{in}}$$

Antenna temperature:

$$T_A = T_b e_{rad} + (1 - e_{rad}) T_p$$

Friis equation:

$$P_r = P_t \frac{G_t G_r \lambda^2}{(4\pi R)^2}$$

Modified Friis equation:

$$P_r = P_t \frac{G_t G_r \lambda^2}{(4\pi R_0)^2} \cdot \left(\frac{R_0}{R}\right)^N$$

( $R_0$ : first obstruction, 1-100 m in urban, 1 km in rural, for N, see table)

Reciever SNR:

$$SNR_{in} = \frac{G_r}{T_A} \cdot \frac{P_t G_t \lambda^2}{(4\pi R)^2}$$

Radar equation:

$$P_r = P_t \frac{G_1 G_2 \lambda^2}{(4\pi)^3 R^4 \sigma}$$

### 2.2 Fading

Attenuation: Frequency dependent, see graph.

- Rain
- Atmosphere

Multi-path fading:

- Reyleigh fading (small scale, no LoS):

$$\ln(1 - \mathbf{P}_{outage}) = -\frac{P_{thr}}{P_0}$$

$$-\frac{P_{thr}}{P_0} = \text{Fade depth}$$

( $P_0$ : average power)

- Rician fading (point to point link, direct path and stochastic reflections):

$$K = A_0 / 2\sigma^2$$

see graph for Fade depth.

- Log-normal fading (large scale, no LoS, multi-path):

$$\mathbf{P}_{out} = \frac{1}{2} + \frac{1}{2} erfc \left( \frac{P_{thr,dBm} - P_{0,dBm}}{\sigma_{dB} \cdot \sqrt{2}} \right)$$

### 2.3 Cascation

Noise figure:

$$F = F_1 + \frac{F_2 - 1}{G_1} + \frac{F_3 - 1}{G_1 \cdot G_2} + \dots$$

Equivalent noise temperature:

$$T = T_1 + \frac{T_2}{G_1} + \frac{T_3}{G_1 \cdot G_2} + \dots$$

Output power intercept point 3 (**OIP3 = P3!**)

$$\frac{1}{P3} = \frac{1}{P3_N} + \frac{1}{P3_{N-1} \cdot G_N} + \frac{1}{P3_{N-2} G_N G_{N-1}} + \dots$$

### 2.4 Power formulas

$$OIP3 = P_{1dB}^{out} - 10\text{dB}$$

$$OIP3 = IIP3 \cdot G$$

Linear dynamic range:

$$DR_l = \frac{P_{1dB}^{out}}{N_{out}}$$

( $N_{out} = k(T_A + T_e)BG$ , output noise power)

Spurious free dynamic range:

$$DR_{sf} = \left( \frac{OIP3}{N_{out}} \right)^{2/3}$$

### 2.5 Mixing

$$f_{RF/IM} = f_{LO} \pm f_{IF}$$

$$BW > 5\%f$$

Down or up-conversion need:

$$f_{IF} > \frac{f_{RF}}{40}$$

$$\left[ f_{LO}^{(1)}, f_{LO}^{(2)} \right] - f_{IF} \neq \left[ f_{IM}^{(1)}, f_{IM}^{(2)} \right] \text{ (no overlap)}$$

(also channels need to fit within filter bandwidths)

$$G_c = \frac{P_{out,RF}}{P_{in,IF}} \text{ or } G_c = \frac{P_{out,in}}{P_{in,RF}}$$

Overall mixers have nasty noise characteristics. Check the slides.

## 2.6 Oscillators

$$v_0 = V_0 \cos(\omega_0 t + \theta(t)), \quad \theta(t) = \theta_p \sin(\Delta\omega t)$$

$$L(\Delta\omega) = \frac{P_{noise(SSB)}}{P_{carrier}} = \frac{\theta_p^2}{4} = \frac{\theta_{rms}^2}{2}$$

$$\frac{P_{noise(SSB)}}{P_{carrier}} = \int_0^{B/2} L(\Delta\omega) d\Delta f = L \frac{B}{2} = \frac{\theta_{rms}^2}{2}$$

Multiplier step to increase oscillator frequency affects noise by:

$$L_{VCO \times N} = \left( \frac{N\theta_{rms}}{2} \right)^2 = N^2 L_{VCO}$$

## 2.7 MIMO

$$E_{tot} = E_{element} \cdot AF$$

$$AF = \sum_{n=0}^{N-1} e^{jn(kd \cos(\theta) + \beta)}$$

$$\langle P \rangle \propto \frac{E^2}{2} \propto \frac{N^2}{2}$$

## 2.8 Other

Geo-stationary orbit: 36000 km

## Far-field assumption

- E and H fields approximate a planar wave front at far distance
- Far field distance:  
*Deviation from planar wave front <22.5°*

$$R_{ff} = \frac{2D^2}{\lambda} \quad \text{Antenna "size"}$$

- For small antennas  $R_{ff} = 2\lambda$ 
  - Small dipoles and loops
- All calculations in this course will assume distances  $> R_{ff}$

## Effective aperture area, $A_e$

- A receiver antenna captures EM radiation from a certain area, the effective aperture area

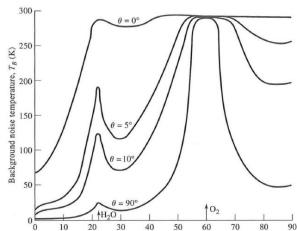
$$P_r = S_{avg} A_e$$

- Capture area and directivity are related
  - Natural if we consider reciprocity of antennas

$$A_e = \frac{D\lambda^2}{4\pi}$$

- Effective aperture area is often close to physical aperture area

## Background noise temperature, $T_B$



- Noise power
- $$N_0 = kTB$$
- K:  $1.38 \times 10^{-23}$
  - B: Bandwidth
  - T: Noise temperature

- Noise temperature vs. frequency for different antenna angles
  - $\theta = 0^\circ$ : Horizon
  - $\theta = 90^\circ$ : Towards sky
- How should the figure be interpreted?

## Properties of Friis equation

$$P_r = P_t \frac{G_t G_r \lambda^2}{(4\pi R)^2}$$

- Assumes free space propagation
  - Applicable for point-to-point links, satellite communication etc.
- Received power is proportional to  $1/R^2$ 
  - In practice,  $1/R^N$ , where  $1.5 < N < 6$ 
    - When could  $N < 2$  be possible?
  - Still better than cable losses for long distance
    - Loss proportional to  $e^{-\alpha R}$
  - More detailed and realistic models presented next lecture

## Antenna directivity, $D$

- The ability of an antenna to direct energy

$$D = \frac{\bar{S}_{max}}{\bar{S}_{avg}} = \frac{4\pi r^2 \bar{S}_{max}}{P_{rad}} = \frac{4\pi \bar{S}_{max}}{\int_{\phi=0}^{2\pi} \int_{\theta=0}^{\pi} \bar{S}(\theta, \phi) \sin \theta d\theta d\phi}$$

- Directivity of isotropic antenna

- Assume  $S(\theta, \phi) = S_{max} = 1 \Rightarrow D_{isotropic} = 1 = 0 \text{ dB}$
- Directivity usually expressed in dBi

- Narrow beam  $\leftrightarrow$  high directivity

- Approximate expression for narrow beams

$$D \cong \frac{\pi^2}{\theta_1 \theta_2} \quad \theta_1 \text{ and } \theta_2 \text{ in radians}$$

## Radiation efficiency and antenna gain

- Radiation efficiency,  $e_{rad}$

- Amount of power radiated, in relation to the power inserted into the antenna
- Resistive losses
- Dielectric losses

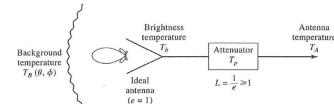
$$e_{rad} = \frac{P_{rad}}{P_{in}} = \frac{P_{in} - P_{loss}}{P_{in}} = 1 - \frac{P_{loss}}{P_{in}}$$

- Antenna gain,  $G$

- Directivity, accounting for the losses
- Suitable for system calculations

$$G = e_{rad} D$$

## Antenna noise temperature, $T_A$



- Contributions from losses and background

- Antenna sees background temperature,  $T_B$ 
  - Weighted with antenna pattern, yields brightness temp.,  $T_b$
- Antenna losses
  - Expressed by radiation efficiency,  $e_{rad}$ , where  $G = e_{rad} D$
  - Antenna held at physical temperature,  $T_p$
- Overall equivalent antenna temperature,  $T_A$

$$T_A = e_{rad} T_b + (1 - e_{rad}) T_p$$

## Modified Friis equation

- Modified Friis equation for generic path loss exponent,  $N$

$$P_r = P_t \frac{G_t G_r \lambda^2}{(4\pi R_0)^2} \frac{1}{(R/R_0)^N}$$

- Free space propagation until a distance  $R_0$

- $R_0$  = (representative) distance to first obstruction

- Urban areas and indoors:  $R_0 = 1m - 100m$
- Rural areas:  $R_0 = 1km$  is a common assumption

- $R_0$  must be in far-field region

## Realistic path loss

- Path loss exponent depends on environment and assumptions

| Environment            | Path Loss Exponent |
|------------------------|--------------------|
| Free space             | 2                  |
| Urban                  | 2.7–3.5            |
| Shadowed urban         | 3–5                |
| In-building LOS        | 1.6–1.8            |
| In-building shadowed   | 4–6                |
| Factory shadowed       | 2–3                |
| Retail store           | 2.2                |
| Office—soft partitions | 2.4                |

- Path loss exponent need often to be characterized experimentally in practical situations

$$P_r = P_t \frac{G_r G_t \lambda^2}{(4\pi R)^2}$$

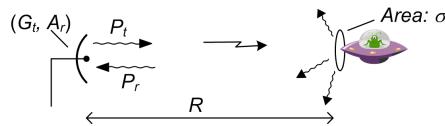
## G/T

$$\frac{S_i}{N_i} = \frac{G_r G_t P_t \lambda^2}{k T_A B (4\pi R)^2} = \frac{G_r}{T_A} \frac{G_t P_t \lambda^2}{k B (4\pi R)^2}$$

Note:  
Error in Pozar, eq (4.30)

- $G/T_A$  is the only parameter that can be controlled by the receiver design
  - Important figure-of-merit for receivers
- G/T usually expressed in units "dB/K"
  - $G/T$  (dB/K) =  $10 \cdot \log_{10}(G/T_A)$
  - For receivers:  $T_A$  may include the receiver noise temperature
- Improved by high increasing antenna gain
  - May not be allowed if an omni-directional coverage is desired

## The RADAR equation



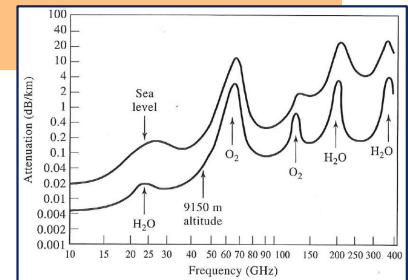
- Finally, assuming identical transmit and receive antennas (using relation between gain and area):

$$P_r = P_t \frac{G_t^2 \lambda^2}{(4\pi)^3 R^4} \sigma$$

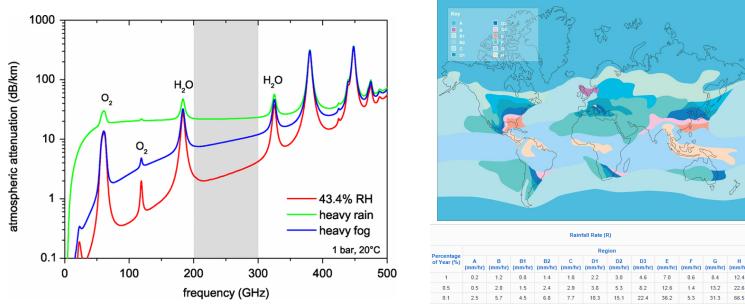
- Very strong dependence on distance,  $R$

## Attenuation effects

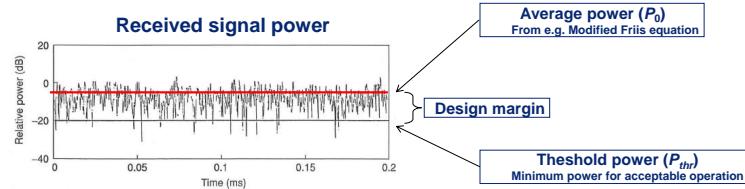
- Loss may also be caused by losses in the propagation path
  - Atmospheric losses: O<sub>2</sub> and H<sub>2</sub>O resonances
  - Precipitation
  - Walls
  - Ceilings



## Rain attenuation



## Rayleigh fading: Received power vs. time



## Rayleigh fading: Outage probability

- Practical systems have a minimum acceptable power threshold level for proper operation,  $P_{thr}$
- The outage probability,  $p_{out}$ , of the received power being  $< P_{thr}$  is then given by

$$\begin{aligned} p_{out} &= \Pr(P < P_{thr}) = \int_0^{P_{thr}} f_p(P) dP \\ &= \int_0^{P_{thr}} \frac{1}{P_0} e^{-P/P_0} dP = 1 - e^{-P_{thr}/P_0} \end{aligned}$$

## Rician fading: Received signal power

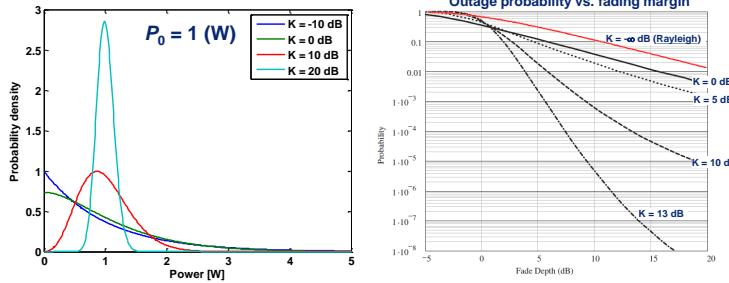
- The direct path results in the Gaussian variables  $X$  and  $Y$  in Rayleigh fading having non-zero mean
  - Received power follows in this case a "non-central chi-square" distribution:

$$f_p(P) = \frac{(1+K)e^{-K}}{P_0} e^{-(1+K)P/P_0} I_0\left(\sqrt{\frac{4K(1+K)P}{P_0}}\right)$$

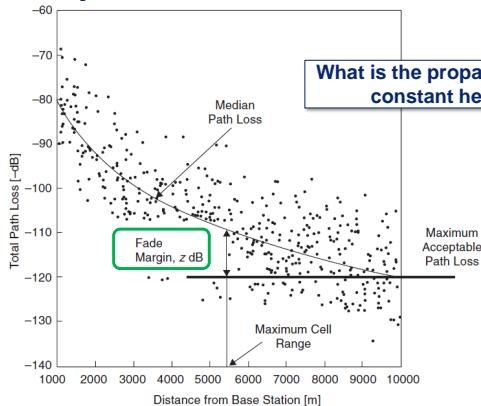
- $I_0$  is the modified Bessel function of first kind
- $P_0$  is the average, total power received
- $K$  is a parameter relating the power of the line of sight component to the stochastic components:  $K = \frac{A_0}{2\sigma^2}$

## Rician fading: Received signal power

- Probability density function for received power
- Converges to Rayleigh when  $K \rightarrow 0$  ( $= -\infty$  dB)
- Explain figure for  $K \rightarrow \infty$ !



## Log-Normal fading: Received power measurement example



Source: Saunders - "Antennas and propagation for wireless communication systems", Wiley & Sons, 1999.

## Intercept point of cascaded components

- General expression for output intercept point of cascaded stages

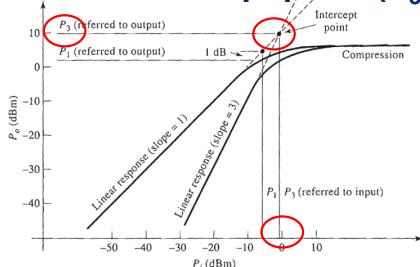
$$P_3 = \left( \frac{1}{G^{(2)} G^{(3)} \cdots G^{(N)} P_3^{(1)}} + \cdots + \frac{1}{G^{(N-1)} G^{(N)} P_3^{(N-2)}} + \frac{1}{G^{(N)} P_3^{(N-1)}} + \frac{1}{P_3^{(N)}} \right)^{-1}$$

$G^{(i)}$ : Gain of  $i^{\text{th}}$  stage

$P_3^{(i)}$ : Output intercept point of  $i^{\text{th}}$  stage

- Compare with expression for cascaded noise
- Influence scales inversely with gain in front of it
  - In transmitters, the last stage typically dominates ( $P_3 \approx P_3^{(N)}$ )

## 3<sup>rd</sup> order intercept point ( $P_3$ )



- Extrapolation of linear response and IM3 signals yield an intercept point ( $\text{IP3} = P_3$  in Pozar)
- Referenced to input power: IIP3
- Referenced to output power: OIP3
- Approximate relation:  $P_3 \approx P_1 + 10 \text{ dB}$

## Log-Normal fading: Outage probability

- The outage probability,  $p_{out}$ , of the received power being  $< P_{thr,dBm}$  is then given by

$$\begin{aligned} p_{out} &= \Pr(P_{r,dBm} < P_{thr,dBm}) = \int_{-\infty}^{P_{thr,dBm}} f(P_{r,dBm}) dP_{r,dBm} \\ &= \int_{-\infty}^{P_{thr,dBm}} \frac{1}{\sigma_{dB} \sqrt{2\pi}} \exp\left(-\frac{(P_{r,dBm} - P_{0,dBm})^2}{2\sigma_{dB}^2}\right) dP_{r,dBm} \\ &= 0.5 + 0.5 \operatorname{erf}\left(\frac{P_{thr,dBm} - P_{0,dBm}}{\sigma_{dB} \sqrt{2}}\right) \end{aligned}$$

## Noise figure of cascaded components

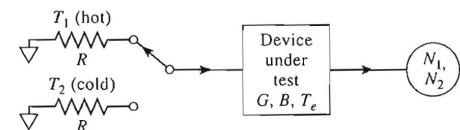
- Generalized expression for noise temperature

$$T_e = T_{e1} + \frac{T_{e2}}{G_1} + \frac{T_{e3}}{G_1 G_2} + \dots$$

- Generalized expression for noise figure

$$F = F_1 + \frac{F_2 - 1}{G_1} + \frac{F_3 - 1}{G_1 G_2} + \dots$$

## Noise temperature measurements The Y-factor method



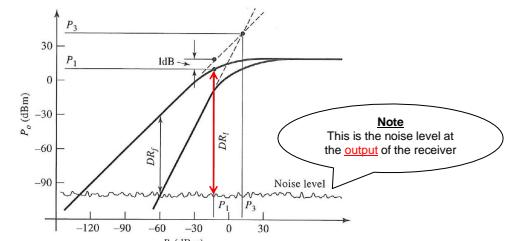
- An Y-factor is defined as

$$Y = \frac{N_1}{N_2} = \frac{T_1 + T_e}{T_2 + T_e}$$

- The unknown noise temperature is then

$$T_e = \frac{T_1 - YT_2}{Y - 1}$$

## Linear dynamic range, $DR_L$



- Linear dynamic range

- Power range between noise at low power and 1-dB compression point at high power

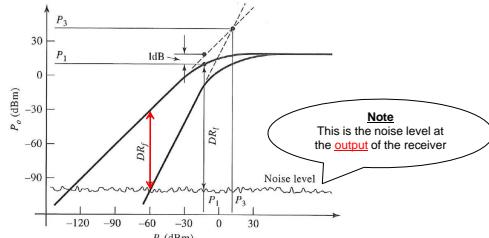
- Referenced to output:

$$DR_L = \frac{P_1}{N_o}$$

$P_1$ : 1-dB compression point (output)

$N_o$ : Noise power at output, typ.  $k(T_e + T_0)BG$

## Spurious free dynamic range, $DR_f$

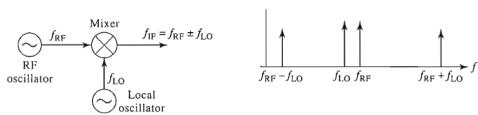


- Spurious free dynamic range

- Power range where signal is higher than noise floor and distortion levels are below noise floor

$$DR_f = \frac{P_{o1}}{P_{2\omega_1-\omega_2}} \Big|_{P_{2\omega_1-\omega_2}=N_0} = \left( \frac{P_{o1}}{P_{\omega_1}} \right)^3 / \left( \frac{P_3}{N_0} \right)^2 = N_0 \left( \frac{P_3}{N_0} \right)^{2/3}$$

## Down conversion mixer



- IF output signal

- Assuming that the mixer is a multiplying device

$$v_{IF}(t) = K \cdot v_{RF}(t) \cdot v_{LO}(t) = K \cos(2\pi f_{RF} t) \cos(2\pi f_{LO} t) \\ = \frac{K}{2} [\cos(2\pi(f_{LO} - f_{RF})t) + \cos(2\pi(f_{LO} + f_{RF})t)]$$

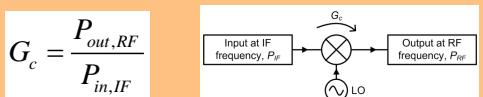
- Desired output frequency is:

$$f_{IF} = f_{RF} - f_{LO} \quad \text{The other component } (f_{RF} + f_{LO}) \text{ is easy to filter away (lowpass filter)}$$

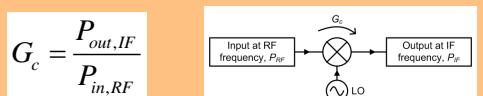
## Mixer conversion gain (or loss)

- Conversion gain (or loss) specifies the ratio of output signal power to the power available at the input

- For up conversion (transmitter):



- For down conversion (receivers):



## Phase noise representation

- Resulting output signal, assuming small  $\theta_p$

$$v_o(t) = V_0 \left\{ \cos \omega_0 t - \frac{\theta_p}{2} \underbrace{[\cos(\omega_0 + \Delta\omega)t - \cos(\omega_0 - \Delta\omega)t]}_{\text{Phase noise sidebands at } \omega_0 \pm \Delta\omega} \right\}$$

- Phase noise at  $\Delta\omega$ :  $L(\Delta\omega)$ :

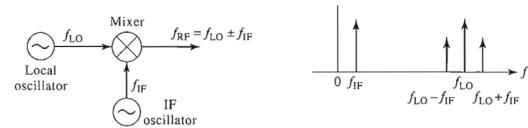
$$L(\Delta\omega) = \frac{P_{noise(SSB)}}{P_{carrier}} = \frac{1}{2} \left( \frac{V_0 \theta_p}{2} \right)^2 = \frac{\theta_p^2}{4} = \frac{\theta_{rms}^2}{2}$$

- Only one sideband is used in the definition

- Note that the phase noise power scales with the received signal power

- Compare with thermal noise and shot noise...

## Up conversion mixer



- RF output signal

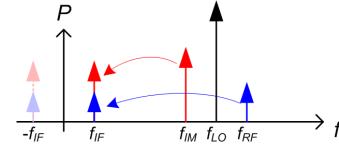
- Assuming that the mixer is a multiplying device

$$v_{RF}(t) = K \cdot v_{IF}(t) \cdot v_{LO}(t) = K \cos(2\pi f_{LO} t) \cos(2\pi f_{IF} t) \\ = \frac{K}{2} [\cos(2\pi(f_{LO} - f_{IF})t) + \cos(2\pi(f_{LO} + f_{IF})t)]$$

- Output signal has spectral components at

$$f_{RF} = f_{LO} \pm f_{IF}$$

## Image frequency in receivers

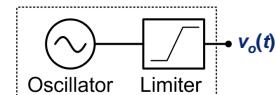


- Two RF frequencies produce the same IF output frequency in down-conversion mixers

- The desired frequency:  $f_{IF} = f_{LO} + f_{RF}$

- The image frequency:  $f_{IM} = f_{LO} - f_{RF}$

## Phase noise representation



- General oscillator output signal

$$v_o(t) = V_0 [1 + A(t)] \cos[\omega_0 t + \theta(t)] \approx V_0 \cos[\omega_0 t + \theta(t)]$$

- $A(t)$ : Amplitude modulation (AM) noise

- AM noise can be suppressed by amplitude limiter

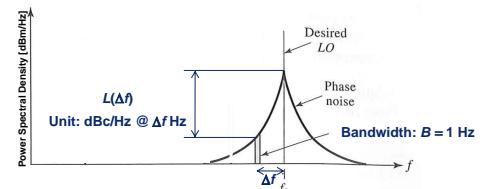
- $\theta(t)$ : Phase modulation (PM) noise = phase noise

- Phase noise at offset  $\Delta\omega$

- Represented as a frequency modulation ( $\Delta\omega$ ) of carrier frequency

$$\theta(t) = \theta_p \sin \Delta\omega t \quad \theta_p = \text{Peak phase deviation}$$

## Oscillator output spectrum



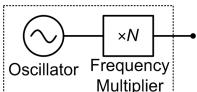
- Noise power spectral density (W/Hz) relative to total oscillator power (W)

- $L(\Delta f) = (P_{noise(SSB)} / B) / P_{carrier}$

- Phase noise within channel bandwidth (B) is added at receiver output

$$P_{noise,SSB} / P_{carrier} = \int_0^{B/2} L(\Delta f) d\Delta f = LB/2 = \frac{\theta_{rms}^2}{2}$$

## Frequency multiplication



- Frequency multiplier = Phase multiplier

$$v_o(t) = V_0 \cos[\omega_0 t + \theta(t)] \rightarrow v_o(t) = V_0 \cos[N\omega_0 t + N\theta(t)]$$

... and therefore a phase noise multiplier

$$\theta_{rms} \rightarrow N\theta_{rms}$$

- Phase noise ( $L$ ) after multiplier

$$L_{VCO \times N} = \frac{(N\theta_{rms})^2}{2} = N^2 L_{VCO}$$

– Degradation of phase noise with  $20\log_{10}(N)$  dB

## Phased array antennas

- An array of antennas transmitting phase-delayed copies

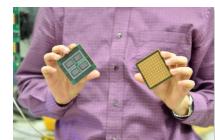
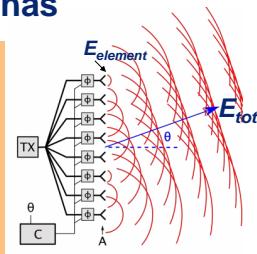
- Main beam direction controlled by phase shift
- Electrical phase-shifters → fast scanning

- Applications

- Electrically scanned radars
- Directed communication (5G)
- Satellite communication

$$\bullet E_{tot}(\theta, \phi) = E_{element}(\theta, \phi) * AF(\theta, \phi)$$

• AF = array factor



## Phased array antennas

- Example: Linear array

$$AF = \sum_{n=0}^{N-1} e^{jn(kd \cos \theta + \beta)}$$

–  $N$  = Number of antenna elements

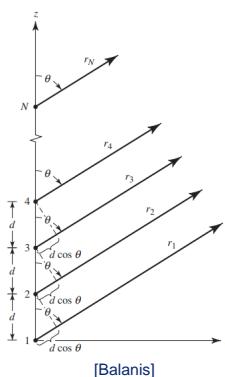
–  $k$  = wave number

–  $d$  = antenna separation

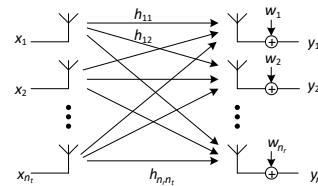
–  $\beta$  = Phase progression between antennas

•  $\max(AF) = N \rightarrow E_{tot,max} = E_{element} * N$

• Multiplication of antenna gain with  $N^2$



## MIMO system model



$$\begin{bmatrix} y_1 \\ \vdots \\ y_{n_r} \end{bmatrix} = \begin{bmatrix} h_{1,1} & \cdots & h_{1,n_t} \\ \vdots & \ddots & \vdots \\ h_{n_r,1} & \cdots & h_{n_r,n_t} \end{bmatrix} \begin{bmatrix} x_1 \\ \vdots \\ x_{n_t} \end{bmatrix} + \begin{bmatrix} w_1 \\ \vdots \\ w_{n_r} \end{bmatrix}$$

- Generic model of MIMO system

$$\mathbf{y} = \mathbf{H}\mathbf{x} + \mathbf{w}$$

$\mathbf{x}$  = Transmitted signals

$\mathbf{H}$  = Channel matrix

$\mathbf{w}$  = Receiver noise

$\mathbf{y}$  = Received signals



Fundamentals of Wireless Communication  
D. Tse and P. Viswanath  
Cambridge University Press, 2005

Full text version available for download!

[https://web.stanford.edu/~dntse/wireless\\_book.html](https://web.stanford.edu/~dntse/wireless_book.html)

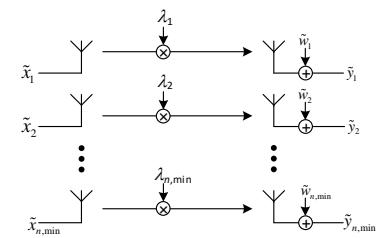
## Singular value decomposition (SVD) of H

$$\mathbf{H} = \mathbf{U}\Lambda\mathbf{V}^*$$

- $\mathbf{U}^{(nr,nr)}$  = Receiver rotation matrix ;  $\mathbf{U}\mathbf{U}^* = \mathbf{I}$
- $\mathbf{V}^{(nt,nt)}$  = Transmitter rotation matrix ;  $\mathbf{V}\mathbf{V}^* = \mathbf{I}$
- $\Lambda^{(nmin,nmin)} = \text{diag}(\lambda_1, \lambda_2, \dots, \lambda_{n_{\min}})$ 
  - diagonal matrix of eigenvalues
  - $n_{\min} = \text{rank}(\mathbf{H}) \leq \min(n_r, n_t)$

$$\begin{aligned} \mathbf{y} &= \mathbf{U}\Lambda\mathbf{V}^*\mathbf{x} + \mathbf{w} \\ \mathbf{U}^*\mathbf{y} &= \Lambda\mathbf{V}^*\mathbf{x} + \mathbf{U}^*\mathbf{w} \\ \tilde{\mathbf{x}} &= \mathbf{V}^*\mathbf{x} \\ \tilde{\mathbf{y}} &= \mathbf{U}^*\mathbf{y} \\ \tilde{\mathbf{w}} &= \mathbf{U}^*\mathbf{w} \end{aligned} \Rightarrow \tilde{\mathbf{y}} = \Lambda\tilde{\mathbf{x}} + \tilde{\mathbf{w}}$$

## Data multiplexing



- A transformed system with  $n_{\min}$  independent paths
- Power in each path scales with the eigenvalue

### 3 Photonics

( $n_1$ :core,  $n_2$ :cladding)

$$\Delta = \left( \frac{n_{core} - n_{clad}}{n_{core}} \right)$$

$$NA = \sin(\theta_{i,max}) n_0 = \sqrt{n_1^2 - n_2^2} \approx n_1 \sqrt{2\Delta}$$

Typical fiber has:  $n_1 = 1.5$ ,  $\Delta = 0.1\%$ .

#### 3.1 Dispersion

$$\sigma_{\tau,step} \simeq \frac{L\Delta n_1}{2c}$$

$$\sigma_{\tau,grin} \simeq \frac{L\Delta^2 n_1}{4c}$$

$$\sigma_{\tau,step} \simeq D\sigma_\lambda L$$

$$\sigma_\lambda = \Delta\lambda = \Delta f \frac{\lambda^2}{c}$$

$$D = 16 \text{ ps/nm/km} @ 1.55\mu\text{m}$$

$$\sigma_{sys}^2 = \sigma_{TX}^2 + \sigma_\tau^2 + \sigma_{RX}^2$$

$$\sigma_{TX/RX}^2 = \frac{1}{B^2}, \text{ BW for device!}$$

We need to have for example:

$$\sigma_{sys} < 70\% \cdot T_{bit} = 70\% \cdot \frac{1}{R_{baud}}$$

$$\sigma_\tau < \frac{T}{4} = \frac{1}{4B_{signal}}$$

#### 3.2 Noise, SNR, Q-value

$$SNR = \frac{I_p^2}{\sigma^2} = \frac{(R \cdot P_{in})^2}{\sigma_T^2 + \sigma_S^2 + \sigma_{s-sp}^2 + \sigma_{sp-sp}^2}$$

$$Q = \frac{I_1 - I_0}{\sigma_1 + \sigma_0}$$

$$Q > 6 \Leftrightarrow BER < 10^{-9}, Q > 7 \Leftrightarrow BER < 10^{-12}$$

(for OOK)

Shot noise:

$$\sigma_S^2 = 2qI_pB = 2qRP_{in}\Delta f$$

( $P_{in}$  on photo diode,  $P_{in} = P_{LO}$  for coherent)

( $R$  responsivity)

$$R = \eta \frac{q}{h\nu} = \eta \frac{\lambda [\mu\text{m}]}{1.24}$$

Thermal noise:

$$\sigma_T^2 = \frac{4kT}{R_L} B = \frac{4kT}{R_L} \Delta f$$

$$(\sigma_T^2 \approx \sigma_S^2 @ P_{in} = 0.83 \text{ mW} @ 1.55\mu\text{m})$$

Signal spontaneous noise:

$$\sigma_{s-sp}^2 = 4R^2 P_{in} S_{sp} \Delta f$$

( $S_{sp}$  spectral noise density)

$$S_{sp} = \frac{F_n}{2} G h\nu \cdot N$$

(balanced detector, for coherent  $P_{in} = 2\sqrt{P_s P_{LO}}$ )

Spontaneous spontaneous beat noise:

$$\sigma_{sp-sp}^2 = 4R^2 S_{sp}^2 \Delta\nu \Delta f$$

$$OSNR = \frac{P_{sig}}{P_{ASE}} = \frac{P_{sig}}{S_{sp} \Delta\nu \cdot N} = \frac{P_{sig}}{\frac{F_n}{2} G h\nu \Delta\nu \cdot N}$$

Coherent receiver power:

$$P(t) = P_s + P_{LO} + 2\sqrt{P_s P_{LO}} @ \text{photo diode}$$

Pre-amplified receiver (consists of a optical amplifier followed by a optical pass band filter followed by a photo diode detector for example):

$$P_{in} = h\nu F_n B \left( Q^2 + Q \sqrt{\frac{\Delta\nu}{\Delta f}} \right)$$

(or just use dominating noise)

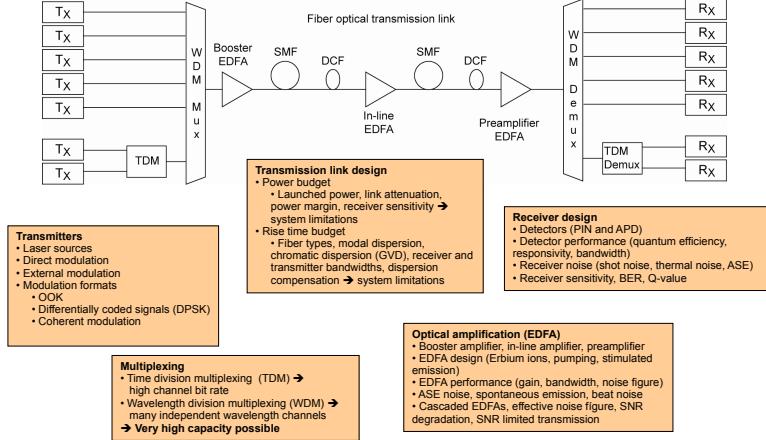
#### 3.3 Other

Bandwidth:

$$\Delta\lambda = \Delta f \frac{\lambda^2}{c}$$

C-band:  $\Delta\lambda = 35 \text{ nm}$ ,  $\Delta f = 4,37 \text{ THz}$ .

## Course summary



## Ray-optics description of step-index fiber

Apply Snell's law at the input interface:

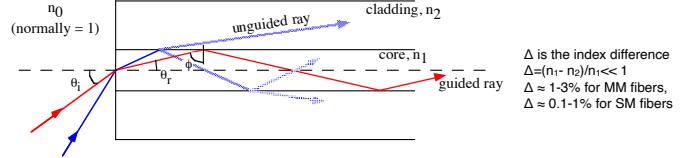
$$n_0 \sin(\theta_i) = n_1 \sin(\theta_r)$$

For total internal reflection at the core/cladding interface we have a critical, minimum, angle:

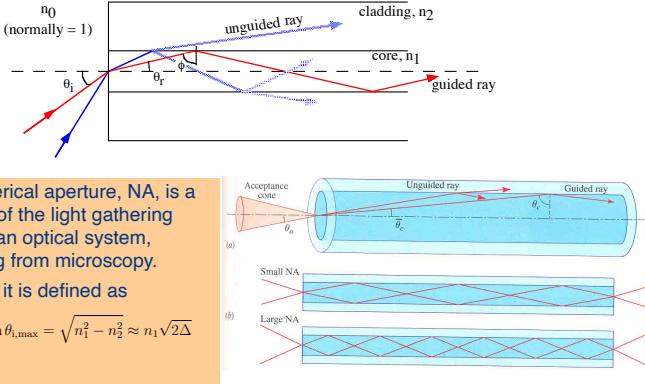
$$n_1 \sin(\phi_c) = n_2 \sin(90^\circ) \rightarrow \sin(\phi_c) = n_2/n_1$$

Relate to maximum entrance angle:

$$\begin{aligned} n_0 \sin(\theta_{i,\max}) &= n_1 \sin(\theta_{r,\max}) = n_1 \sin(90 - \theta_c) = \\ n_1 \cos(\theta_c) &= n_1 \sqrt{1 - \sin(\theta_c)^2} = \sqrt{n_1^2 - n_2^2} \end{aligned}$$

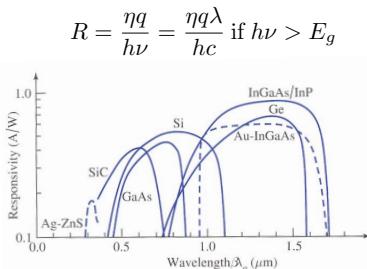


## Numerical aperture



## Photodetectors for fiber systems - overview

- Responsivity increases linearly with wavelength...
- ...until the photon energy is below the bandgap energy of the semiconductor.
- 870 nm range
  - Si p-i-n or APD
    - low cost
    - low noise
- 1300-1600 nm range
  - InGaAs, Ge, InP
    - InGaAs APD common in commercial systems (10 Gb/s per WDM channel)
    - InGaAs/InP p-i-n common in high bandwidth systems (> 40 Gb/s per channel)



| Material | $hc/E_g [\mu m]$ |
|----------|------------------|
| Si       | 1.15             |
| Ge       | 1.88             |
| GaAs     | 0.87             |
| InAs     | 3.5              |

## The p-i-n characteristics

A photo diode p-n junction has the following characteristics

$$i = i_s (\exp(\frac{eV}{k_B T}) - 1) - RP_{in}$$

$i_s$  is the "dark current"  
 $\Phi = RP_{in}$  denotes photoelectron flux

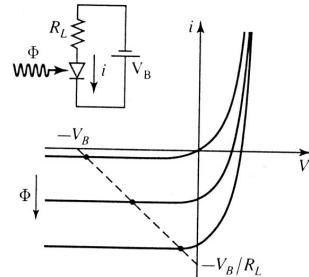


Photo diodes are usually operated strongly reverse-biased:

- reduces the carrier transit time - increases the speed (bandwidth)
- increases the linearity

The operating point is indicated by a dashed line

## Transmitters - overview

- Directly driven LED
  - Low cost
  - Incoherent (high linewidth) source
  - High coupling losses due to weak directionality
  - MMF
  - Mb/s links, OOK
- Directly modulated Semiconductor Laser
  - Higher cost
  - Coherent (low linewidth) source
  - Less coupling losses due to stronger directionality of radiation
  - MMF or SMF
  - ≤ 10-100 Gb/s link, OOK
- Externally modulated Semiconductor Laser
  - High cost
  - Coherent (low linewidth) source
  - Less coupling losses due to stronger directionality of radiation
  - SMF
  - ≥ 100 Gb/s links, QPSK

## Responsivity

A requirement is that the detector material bandgap energy ( $E_g$ ) must be smaller than the photon energy ( $h\nu$ ).

The detector responsivity  $R$  is defined as the photo-current/incident optical power:

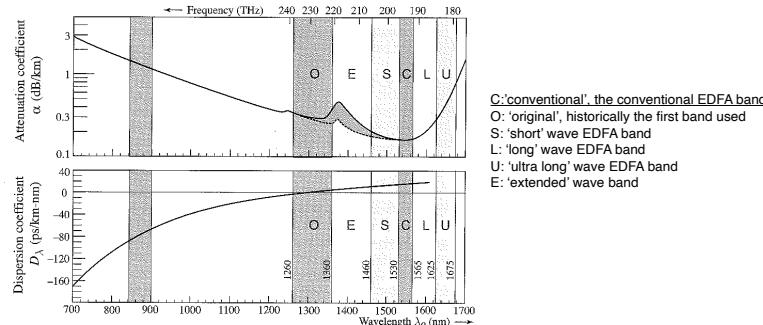
$$R = \eta \frac{q}{h\nu} = \eta \frac{\lambda}{1.24}$$

$\eta$  = the quantum efficiency = the number of e-h pairs generated per photon (ideally,  $\eta = 100\%$ ),  $\lambda$  is wavelength in  $\mu m$ .

Material used: InGaAs (sometimes Ge), with typical data:

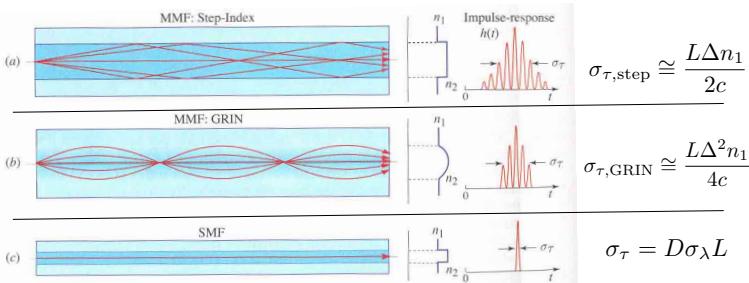
$$\eta \approx 0.8, R \approx 1 \text{ A/W}, \text{BW} \approx 10 \text{ GHz into } 50 \Omega, \text{dark current} \approx 0.1 \text{ nA}$$

## What wavelengths are used in fiber optics?



- 850 nm: common in short MMF-based links
- 1550 nm: most commonly used SMF wavelength
- 1300-1600: also used in SMF but less frequently

## Dispersive broadening summary



## Rise time budget

- A time (or rise time) budget accounts for the dispersive broadening of the fiber, as well as the response times of the transmitter and receiver,  $\sigma_{tx}$  and  $\sigma_{rx}$ .
- Assuming Gaussian responses, a quadratic summation must be used.

$$\sigma_{sys}^2 = \sigma_{tx}^2 + \sigma_{rx}^2 + \sigma_{\tau}^2 < (0.7 T)^2 = \frac{0.49}{B^2}$$

Arbitrary but common criterion

Often the fiber dispersion-induced rise time is designed to be less than:

$$\sigma_{\tau} = \frac{T}{4}$$

- How large is this in various fibers?

## Signal-to-Noise Ratio (SNR)

The total noise variance is the sum of the noise contributions (since the noise sources are independent):

$$\sigma^2 = \sigma_s^2 + \sigma_T^2$$

The SNR of an electrical signal is simply defined as:

$$SNR = \frac{\text{signal power}}{\text{noise power}} = \frac{I_p^2}{\sigma^2}$$

## Q-value

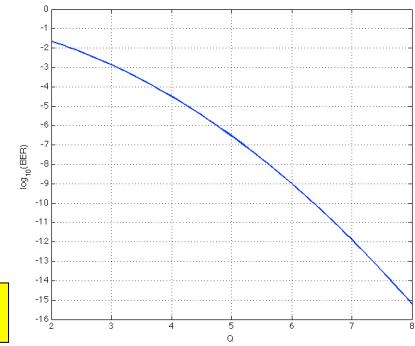
To calculate the BER in optical systems with OOK it is common to introduce the **Q-value**. Note that  $Q^2$  is (approx.) proportional to the SNR.

$$Q \equiv \frac{I_1 - I_0}{\sigma_1 + \sigma_0}$$

Note:  $\sigma_1$  and  $\sigma_0$  may not be equal, for example due to shot noise! With optimum setting of the decision point, and Gaussian noise PDFs:

$$BER \approx \frac{\exp(-Q^2/2)}{Q\sqrt{2\pi}}$$

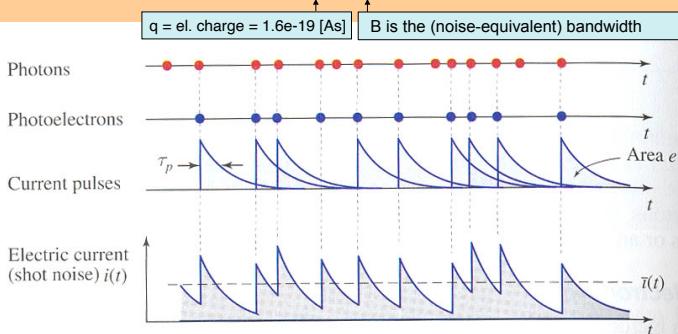
$Q > 6$  is needed for  $BER < 10^{-9}$   
 $Q > 7$  is needed for  $BER < 10^{-12}$



## Shot noise

Shot noise arises from the particle nature of light (photons) and charges (electrons). The variance of the photo current  $I_p$  is (in units of [ $A^2$ ]):

$$\sigma_s^2 = 2qI_pB$$



## The optical beat noise terms

- The signal-spontaneous (s-sp) beat noise (signal beats with the noise) is proportional to the electrical bandwidth:

$$\sigma_{s-sp}^2 = 4R^2 P_{in} S_{sp} \Delta f$$

ASE noise power spectral density [W/Hz]  
el. BW [Hz]

- The spontaneous-spontaneous (sp-sp) beat noise (noise beats with itself) is proportional to also the optical noise bandwidth

$$\sigma_{sp-sp}^2 = 4R^2 S_{sp}^2 \Delta \nu \Delta f$$

opt. BW [Hz]

- An optical bandpass filter can reduce the sp-sp part significantly.

## Thermal noise

(circuit noise, Johnson/Nyquist noise)

Thermal noise arises from random movements of conducting electrons due to the temperature. The noise current variance is (in units of [ $A^2$ ]):

$$\sigma_T^2 = \frac{4k_B T}{R_L} B$$

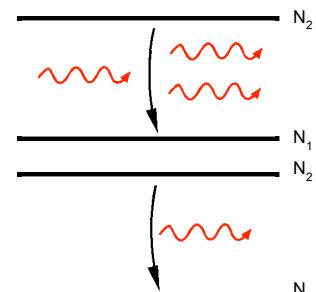
(Boltzmann's const. x Temp) [J], [Ws]  
Load resistance [Ohm]

To account for electrical amplifiers generating excess thermal noise the above expression is sometimes modified to:

$$\sigma_T^2 = \frac{4k_B T F_n}{R_L} B \quad F_n \text{ is the electric amplifier noise figure (>1)}$$

## Gain and noise in optical amplifiers

- Most amplifiers operate by the principle of **stimulated emission**
- Noise added through **spontaneous emission** is amplified along with the signal:  
**Amplified spontaneous emission (ASE)**
- The **population-inversion factor**,  $n_{sp}$  depends on the population in the ground state ( $N_1$ ) and in the excited state ( $N_2$ )

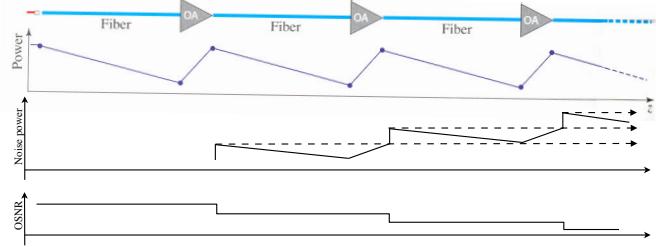


Optical power spectral density of ASE in an amplifier:  $n_{sp} = \frac{N_2}{N_2 - N_1}$

$$S_{sp}(\nu) = (G - 1)n_{sp}h\nu = \frac{F_n}{2}Gh\nu \quad [\text{W / Hz}]$$

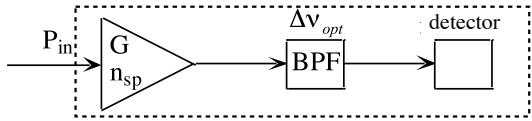
$$F_n = 2n_{sp} \frac{G - 1}{G}$$

## Optically amplified chains



- Each amplifier produces optical noise, *amplified spontaneous emission*, ASE, which propagates through the chain.
- The noise power adds incoherently and increases linearly with the number of amplifiers.  $P_{ASE} = P_{ASE,1} + P_{ASE,2} + \dots = Nhv \frac{F_n}{2} G \Delta\nu$
- The optical SNR, the OSNR, will decrease linearly.  $OSNR = \frac{P_{sig}}{P_{ASE}}$

## The optically preamplified lightwave receiver



The power entering the detector is:

$$GP_{in} + S_{sp} \Delta\nu_{opt} = GP_{in} + (G - 1)n_{sp} h\nu \Delta\nu_{opt}$$

The input average power needed to reach a certain Q is:

$$P_{in} = h\nu F_n B (Q^2 + Q \sqrt{\frac{\Delta\nu_{opt}}{B}})$$

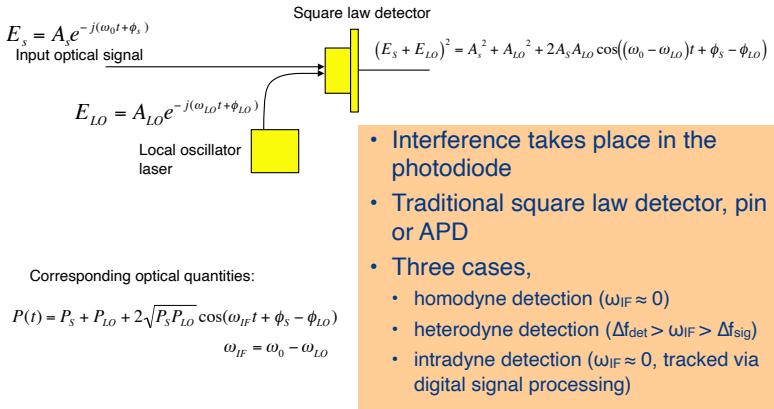
Derived from:  

$$Q = \frac{RP_1 - RP_0}{\sigma_1 + \sigma_0} = \frac{RP_1}{\sqrt{\sigma_{s-sp}^2 + \sigma_{sp-sp}^2} + \sigma_{sp-sp}}$$

Use a narrow optical filter (but not too narrow!) and a low-noise EDFA

At 10 Gb/s, BER =  $10^{-9}$  (Q=6) can be achieved with  $P_{in} = -40$  dBm (100 nW)

## Coherent detection - Basic Concept



## SNR for Homodyne and direct detection

|               | Homodyne / Intradyne  | Direct Detection   |
|---------------|---|--|
| Signal power: | $I(t)^2 = 4R^2 P_s P_{LO}$  | $I(t)^2 = R^2 P_s^2$   |
| Shot noise:   | $\sigma_s^2 = 2qR P_{LO} \Delta f$<br>(LO-limited)                                      | $\sigma_s^2 = 2qR P_s \Delta f$  |
|               | $SNR = \frac{I^2}{\sigma^2} = \frac{4R^2 P_s P_{LO}}{\sigma_T^2 + 2qR P_{LO} \Delta f}$ | $SNR = \frac{I^2}{\sigma^2} = \frac{R^2 P_s^2}{\sigma_T^2 + 2qR P_s \Delta f}$ |

For sufficiently large LO power the thermal noise can be neglected, and one finds (use  $R=q/hv$ ):

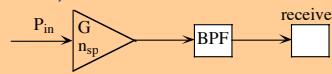
$$SNR = \frac{I^2}{\sigma_s^2} = \frac{2P_s}{h\nu \Delta f} = 4\bar{m}$$

the average photon number per bit, using a Nyquist-limited bitrate  $B=\Delta f/2$

$$SNR = \frac{I^2}{\sigma_s^2} = \frac{P_s}{2h\nu \Delta f} = \bar{m}$$

## The preamplified receiver

One of the main applications for EDFA's is as an amplifier in front of the receiver, in order to overcome thermal noise.



The amplified signal entering the detector is  $P_{amp} = GP_{in} + P_{sp}$

where  $P_{sp} = S_{sp} \Delta\nu_{opt} = (G - 1)n_{sp} h\nu \Delta\nu_{opt}$  is the spontaneous emission noise power, and  $\Delta\nu_{opt}$  is the bandwidth of the optical bandpass filter.

The corresponding photocurrent is

$$I = R|\sqrt{GE_s} + E_{sp}|^2 + i_s + i_T = R(GP_{in} + 2\sqrt{GE_s E_{sp}} + |E_{sp}|^2) + i_s + i_T$$

Legend: Signal field, ASE noise field, shot/thermal noise, signal-spontaneous beating, spontaneous-spontaneous beating

## Wavelength MUX/DEMUX

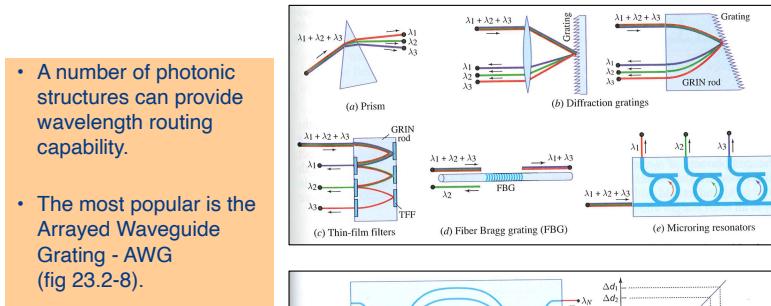


Figure 23.2-8 Wavelength-division demultiplexing by use of a wave-grating router (WGR).

## Homodyne / Intradyne receiver

Detected signal current (I-quadrature)

$$I \approx R(P_{LO} + 2Re(E_s E_{LO}^*))$$

Dominating noise (if no optical amps):

$$LO \text{ shot noise: } \sigma_s^2 = 2qR P_{LO} \Delta f$$

$$SNR = \frac{4R^2 P_s P_{LO}}{2qR P_{LO} \Delta f} = 2 \frac{P_s}{h\nu \Delta f}$$

Example: BPSK

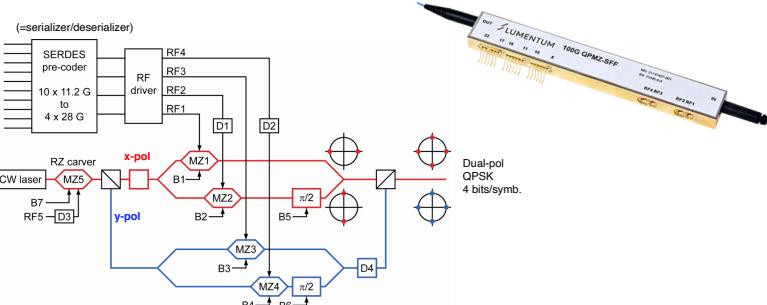
$$Q = \frac{\mu_1 - \mu_0}{\sigma_1 + \sigma_0} = 2\sqrt{\frac{P_s}{2h\nu \Delta f}}$$

$$\mu_1 = \mu_0 = 2R\sqrt{P_s P_{LO}}$$

$$\sigma_1 = \sigma_0 = 2qR P_{LO} \Delta f$$

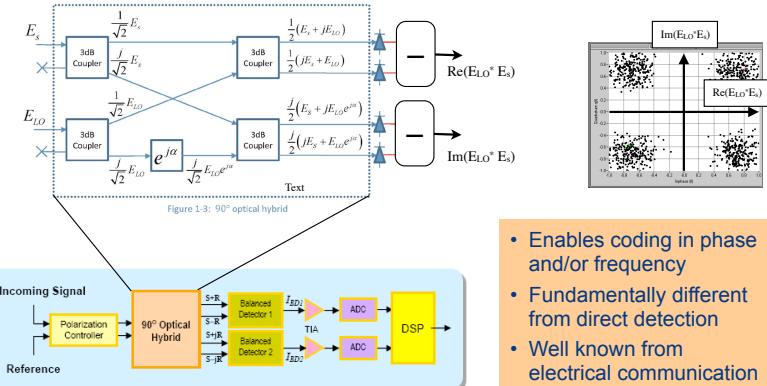
- High receiver gain is possible by using high-power LO
- Phase locking of LO required!
  - Usually done via DSP!
- Best possible receiver!

## Coherent Transmitter

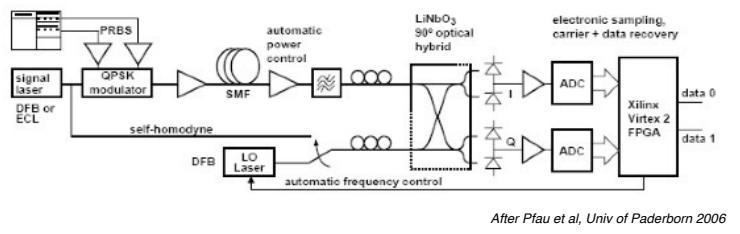


- A CW laser is split into 2 polarizations...
- ...each modulated with dual-arm Mach-Zehnder modulators.
- 4 binary data signals, 6 bias voltages

## Quadrature Detection - 90-deg hybrid



## Example of Intradyne detection



- The phase locking is done via electronic processing, in FPGA:s
- Similar performance as homodyne
- Complex electric signal is mapped to the electric domain
- Used in all modern coherent receivers (ASIC-based).

## Digital Demodulation in optical receivers

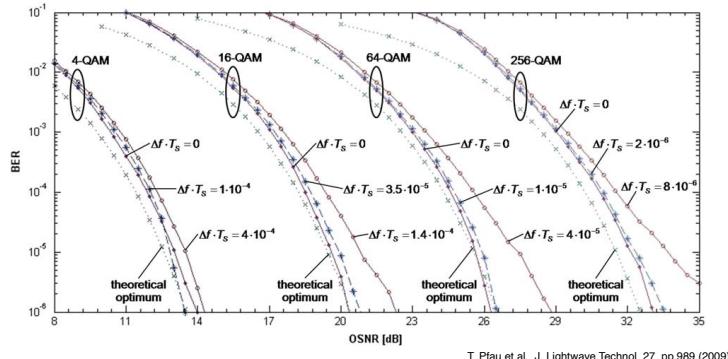
Before the decisions, “undo” the effects of the channel.

For details: Seb J. Savory, IEEE J. Select. Topics Quantum Electron., 16, 5, 2010

This is a simplified list of all signal processing done by the DSP in a modern optical coherent receiver.

| Subsystem            | Functionality   |
|----------------------|---|
| ADC                  | Sample and convert data to digital domain             |
| Static filtering     | FIR filter to compensate for the chromatic dispersion |
| Dynamic filtering    | Compensate polarization mixing and track polarization |
| Frequency estimation | Frequency offset compensation                         |
| Phase estimation     | Phase drift compensation                              |
| Symbol decisions     | Data recovery   |
| FEC decoder          | Decode words and correct errors                       |

## Linewidth Requirements



## Nonlinear limits on transmitted power

The Kerr effect (cause of FWM, and other NL distortions)

- refractive index slightly power dependent
  - induced nonlinear phase shift from power  $P$  over a fiber length  $L$  is  $\phi = \gamma PL$
  - where  $\gamma \approx 1.5 \text{ [W}^{-1} \text{ km}^{-1}\text{]}$  in SMF is the nonlinear coefficient.
  - the nonlinear phase shift arises during the first loss length,  $1/\alpha \approx 20 \text{ km}$  after each amplifier.
  - After  $N$  amplifiers accumulated nonlinear phase shift is
- $$\phi_{NL} = \int_0^L P_0 \exp(-\alpha z) \gamma dz \approx \frac{N \gamma P_0}{\alpha}$$
- total peak power      Nbr of amplifier spans
- This limits the total transmitted power to be  $< 0.1\text{-}10 \text{ mW}$  in most systems.







Part 1

$$\begin{aligned} R_b &= 200 \text{ kbit/s} \\ f_u &= 34.45 \text{ GHz} \\ f_{lo} &= 500 \text{ Hz} \\ f_c &= 32.05 \text{ GHz} \\ T_{sys} &= 28 \text{ K} \end{aligned}$$

What we need:

1. Receive from earth
2. Transmit to earth
3. Receive from terminal
4. Transmit to terminal

Terminal

5. Receive from relay

6. Transmit to earth

Part 1 Receive from Earth

downlink = 34 GHz

$$A_{eff} = \frac{\pi D^2}{4\lambda} = 9.08 \text{ m}^2$$

$$D_i = \frac{4\pi A_{eff}}{\lambda^2} = [\lambda = 8.7 \cdot 10^{-3} \text{ m}] = 150.7 \cdot 10^6$$

$$= 81.8 \text{ dB}$$

$$\text{Assume } G_r \approx 0, \text{ assume } G_t \approx 0$$

Part 2 transmit to earth

Invertible Tsys vs equilibrium temp of receiver at earth

Assume  $T_s = 50 \text{ K}$  again  
QPSK, BER =  $10^{-5}$

$$f_n = 1 + \frac{1}{2} k = 1.083 = 0.35 \text{ dB}$$

(Very good, but maybe possible at NASA?)

$$\text{SNR}_{in} = 12.8 \text{ dB} + 0.35 \approx 13 \text{ dB}$$

$$\text{SNR}_{out} = (\text{Tx Tsys}) k_B = 1.10^{17} \text{ W} = -140 \text{ dBm}$$

$$\text{Pensitivity} = -130 \text{ dBm}$$

Assume downlink same path loss as uplink.

Path loss from part 1 = 136 dB or worse

$$\therefore P_t = -130 \text{ dBm} + 136 \text{ dB} = 6 \text{ dBm}$$

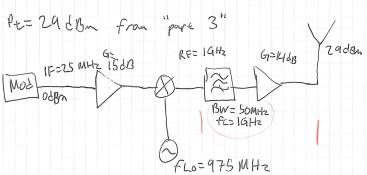
per channel, so in total

$$P_t = 0.2 \text{ W} = 23 \text{ dBm}$$

High but reasonable. — not very high

Part 6

Transmit to relay from terminal



Final words:

I shortness of time, I have left our noise figure and dynamic range calculations in part 3 and 5. They are done in the same way as in "part 1".

Very good!

$$\begin{aligned} P_{max} &= 500 \text{ W} \\ P_{channel} &= \frac{500 \text{ W}}{50} \approx 10 \text{ W/Channel} \\ \text{Received at relay} & \\ P_r &= \frac{P_c G_1 G_2 \lambda_u^2}{(4\pi R)^2} \quad (\text{Assume point-to-point link}) \end{aligned}$$

Satellite antenna size = 2 m  
So  $A_{eff} = \pi r^2$

$$\therefore D_2 = \frac{4\pi A_{eff}}{\lambda u^2} = 521.6 \cdot 10^3 = 57.2 \text{ dB}$$

$$\text{Assume } G_r \propto D_2$$

$$R_E (55700 \text{ km}, 401300 \text{ km})$$

$$\text{Received from earth: } |P_r|_E = (1.2 \cdot 10^{-7} \text{ W}, 2.34 \cdot 10^{-4} \text{ W}) = (-34.2 \text{ dBm}, -56.3 \text{ dB}$$

Assume QPSK, BER =  $10^{-5}$

$$\therefore \text{Efficiency} = 2 \Rightarrow \text{BW} = 100 \text{ kHz}$$

$$|E_b|_n = 9.5 \text{ dB}$$

$$\therefore \text{SNR}_{out} = 10^{0.45} \cdot 3 = 12.5 \text{ dB}$$

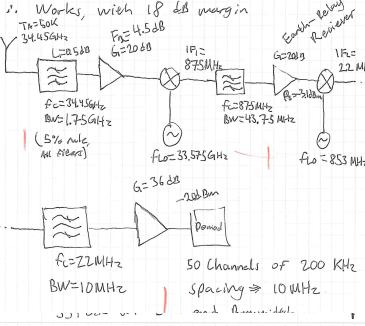
Antenna temp assumed

$$T_a = 50 \text{ K} \quad \text{with } \Delta T = 10 \text{ K}$$

$$\begin{aligned} \text{Add losses from rain:} & \\ & 2.0 \text{ dB/km} \cdot \text{height of clouds} \approx 2 \text{ km} \quad ++ \\ P_r &= (-7.92 \text{ dBm}, -96.3 \text{ dBm}) \\ \text{Receiver } f_n &= 5 \text{ dB} \\ \text{SNR in} &= 17.5 \text{ dB} \\ N_{in} &= k_{Bd} \theta = 6.9 \cdot 10^{-7} = -131.6 \text{ dBm} \end{aligned}$$

$$\text{Sensitivity} = -131.6 \text{ dBm} + 17.5 \text{ dB}$$

$$= -114 \text{ dBm}$$



Part 4 Transmitter to terminal

Assume same antennas, same path loss  
Same modulation, ~~same~~ very close frequency

$$\text{SNR}_{in} = 17.5 \text{ dB}$$

Antenna temperature on max with low gain antenna  $\approx 290 \text{ K}$

$$N_{in} = 4 \cdot 10^{-6} \text{ W} = -124 \text{ dBm}$$

$$P_{in} = N_{in} \cdot \text{SNR} = -124 \text{ dBm} + 17.5 \text{ dB}$$

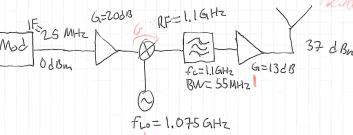
$$= -106 \text{ dBm}$$

$$\therefore P_t = -106 \text{ dBm} + 143 \text{ dB} = 37 \text{ dBm}$$

All terminals means

$$P_{t,avg} = 50 \text{ ft} = 54 \text{ dBm}$$

Very high for a satellite, must have very good solar panels or something.



We can eliminate  $OIP_{3,relay}$  by combining these expressions. This gives, after some manipulations, the following expression for  $OIP_{3,relay,now}$

$$OIP_{3,relay,now} = \frac{1}{\frac{1}{OIP_{3,relay}} - 9 \left( \frac{G_{IF}}{OIP_{3,IF}} \right)} + \frac{G_{IF} G_{IF}}{OIP_{3,IF}}$$

Inserting the given numbers yields  $OIP_{3,relay,now} = 0.22 \text{ mW} = -6.4 \text{ dBm}$ . Referred to the input, this is equivalent to  $\text{OIP}_{3,relay} = -11.4 \text{ dBm}$ .

Thus, in summary, the mixer has to be replaced with another one having  $\text{OIP} > -11.4 \text{ dBm}$  in order to improve the linear dynamic range of the receiver by 10 dB.

Problem 3:

A typical design criterion could be that the dispersive broadening can be only 25% of the bit period  $T$ . The dispersive broadening is  $\Delta t \cdot \Delta \lambda$ , where  $\Delta t$  is 16 ps/nm km and  $\Delta \lambda$  denotes the signal bandwidth in [nm] units. We need to relate bandwidth in nm to GHz, and using  $\Delta \lambda / \Delta t = \Delta f$  we find that  $\Delta \lambda = 1 \text{ nm}$  corresponds to  $\Delta f = 125 \text{ GHz}$ , or 1 GHz is 0.008 nm. Then 10 GHz corresponds to 0.08 nm, and 0.32 nm corresponds to 40 GHz.

The allowed fit for 10 GHz is (given our 25% criterion)  $100/4 = 25 \text{ ps}$  and for 40 GHz it is  $25/4 = 6.25 \text{ ps}$ .

This gives  $25/(16 \cdot 0.08) = 19.5 \text{ km}$

and for 40 GHz:  $6.25/(16 \cdot 0.32) = 1.22 \text{ km}$

Finally the scaling:  $L_s \sim \Delta t \cdot \Delta \lambda$  and As scales as  $\sim 1/\Delta f$  and  $\Delta \lambda \sim \Delta f$  we have that  $L_s \sim \Delta f^2$ . This  $19.5/1.22 = 16$  should be a hint as well.

Problem 4:

We have 4 noise sources to calculate: thermal shot, signal-spontaneous and spontaneous-spontaneous beat noise.

We need some data to do this. The received optical power is  $P_r = 10 \text{ dBm}$  (note that the preamp makes the receive optical power higher than the transmit power). The received amplifier noise PSD is  $S_{noise} = N \text{ Gb} / NF = 2.78e-16 \text{ [W/Hz]}$

where  $N = 11$ ,  $G = 100$ , the photon energy  $hv = 1.28e-19 \text{ J}$ , and the noise figure is  $NF = 10^{10.8} = 3.98$ . The photodetector has a responsivity  $R = q/hv = 1.25 \text{ A/W}$ .

The bandwidth  $\Delta f = 40 \text{ GHz}$  and the optical bandwidth is given by the channel spacing to  $\Delta \lambda = 200 \text{ GHz}$ .

This gives:

$$\sigma_{shot}^2 = 4kT ADR = 1.35e-11 \text{ [A}^2]$$

$$\sigma_{sp}^2 = 2q(RP_s)/\Delta f = 1.60e-11 \text{ [A}^2]$$

$$\sigma_{ss}^2 = 4RS_{noise}/\Delta f = 6.95e-8 \text{ [A}^2]$$

$$\sigma_{ss,sp}^2 = 4(RS_{noise})^2 \Delta f = 3.86e-9 \text{ [A}^2]$$

We see that s-s noise is largest, as expected. We can neglect the others and solve with  $kT$  with rate even smaller.

$$\text{Need total gain } G = -20 \text{ dBm} - (-96 \text{ dBm})$$

$$= 76$$

$$F_{NI} = 5 \text{ dB} = F_2 + F_2 - 1 / G$$

$$\Rightarrow F_2 = G_i \cdot (10^{0.5} - F_1) + 1 = 2.8 \approx 4.5 \text{ dB}$$

Dynamic range =  $P_{max} - P_{min,eq} = -79 \text{ dBm} - (-96 \text{ dBm})$

$$= 35 \text{ dB}$$

Assume last amplifier most important  $P_s$

$$DR_p = \left(\frac{P_s}{N_0}\right)^{1/2}$$

$$\Rightarrow P_s = DR_p^{3/2} \cdot Min \cdot G = 4.88 \cdot 10^{-4} \text{ W}$$

$$= -3 \text{ dBm}$$

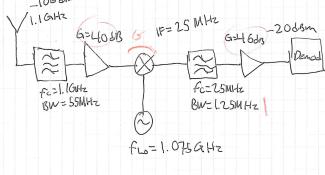
Part 5 Receive from relay

Antenna was  $G = 3 \text{ dBi} \approx 0$   
so  $A_{eff} = \frac{D \lambda^2}{4\pi} = 0.014 \text{ m}^2$

Parabolic antennas

$$d = \sqrt{\frac{4\lambda}{\pi}} = 13 \text{ cm. Size OK,}$$

a bit large for space satellites maybe.



$$G = 8.6 \text{ dB}$$

I=2RP; (assuming  $P_s$  is the average received power)

$$I_p = \sigma_{sp}^2 \cdot \Delta f = (4RS_{noise}RP_s)\Delta f$$

This gives  $Q^2 - P_s \Delta f = 36$  from which we find  $P = 36 \text{ dBm}$ ,  $\Delta f = 4.0 \text{ ms}$ ,  $W = 0.4 \text{ m}$ ,  $= 4 \text{ dBm}$ .

The transmit power is 20 dB lower than this and thus  $-24 \text{ dBm}$ .

(Note: A more accurate but also more complicated result is derived if the sp-sp noise is not neglected. This gives a 2.5 dB higher power of  $-21.5 \text{ dBm}$ .)

## 4 To think about designing systems

### 4.1 Wireless

- Frequency planning
- Explain all assumptions
- Parameter values:  
Antennas:  $e_{rad}, G, T_A, f$   
Filters:  $f_c, B, L$   
Mixers:  $P_3, P_1, F, G$   
Local oscillators:  $f_{LO},$
- RX sensitivity
- Discussion on parameter values
- Dynamic range
- Duplexing

### 4.2 Photonics

- Component choices, for example: external/directly modulated, laser type, detectors, amplifiers, dispersion compensation, SMF/MMF/GRIN, etc.
- Total data capacity:

$$C_{tot} = N_{fibers} \times N_{\lambda-channels} \times R_{symbol} \times Sp.Eff.$$

- Guard bands

Design flow for long-haul fiber system (excerpt from Magnus tips):

1. Assume reasonable channel spacing, modulation format and symbol rate to get the desired bitrate. This gives the required bandwidth (=symbol rate) per channel which is needed for noise calculations.
2. Compensate the dispersion, using reasonable assumptions for DCF (or fiber Bragg gratings) and SMF including parameters. This leads to the total system losses. If you have a coherent receiver the dispersion can be compensated electronically in the receiver. The drawback is the coherent system is more complex, and you must mention this fact as one part of the coherent DSP
3. Assume a reasonable amplifier spacing. Reasonable span losses is around 10-20 dB for ultralong (usually submarine) links and 20-30 dB for shorter terrestrial links. This info together with the total system losses leads to the required number of spans.
4. Use the required Q (for OOK) or SNR (for phase modulation or QAM) to get a transmit power per channel (which you want not much larger than 0 dBm). If too high (e.g. total signal power summed over all channels > 20 dBm) consider either: reducing amplifier span loss (more amplifiers), reducing channel bandwidth, or using FEC and design with a lower Q/SNR.
5. Check all noise sources and verify that s-sp (or LO-sp) dominates.