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NS: Inverse 2 \times 2: For A := [a, b; c, d], A^{-1} := \frac{1}{ad - bc}[d, -b; -c, a] Radial: r\dot{r} = x\dot{x} + y\dot{y}, \ \dot{\theta} = x\dot{x} + y\dot{y}
      \left[\tan^{-1}(y/x)\right]' = \frac{x\dot{y} - \dot{x}y}{x^2 + y^2} Classifications: Node: \lambda_i \in \mathbb{R}, \Pi\lambda_i > 0 Centre: \lambda_i = \pm ib Focus: \lambda_i = a \pm ib
      Hyperbolic: \operatorname{Re}(\lambda) \neq 0 \to \operatorname{hyperbolic}. If all \lambda < 0 for \operatorname{Spec}(Df(x_0)) then A-Stable Invariant Set:
      |\phi_t(S)| \subseteq S Lim Pts: \omega pt. if \lim_{t\to\infty} \phi(x) = p, i.e. flows tend to p. \alpha pt. if \lim_{t\to\infty} \phi(x) = p.
      Attracting Set: A set A \subseteq S if \exists neighbourhood U s.t. \phi(U) \subseteq U \forall t \geq 0, and A = \bigcap_{t>0} \phi(U) Dense
      Orbits: If \forall \epsilon > 0, x \in A with A an attracting set, \exists \tilde{x} \in \Gamma s.t. | x - \tilde{x}| < \epsilon. I.e. a dense orbit goes as close
      as needed to any point within A Attractor: An attracting set with a dense orbit. Lyapunov Stable:
      If \forall \epsilon > 0, \exists \ \delta > 0 s.t. \forall \ x \in B_{\delta}, t \geq 0, \phi(t) \in B_{\delta} (i.e. points stay close within region). Asymptotically
      Stable: If L-Stable and \exists \ \delta > 0 s.t. \phi(x) \to x_0 \forall x \in B_\delta Lyapunov F'n: V(x_0) = 0, V(x) > 0 \forall \ x \neq x_0.
      Then if V < 0 \to A-Stable, or if V \le 0 L-Stable. Stable Manifold: If spectrum of Df(x_0) has k
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      eigvals with positive real parts, and n-k with negative, then \exists an n-k dim manifold tangent to E^s
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      s.t. for all t > 0 \phi(W_{loc}^s) \subseteq W_{loc}^s, and \forall x \in W_{loc}^s \phi(x) \to x_0 as time increases. Repeat for k-dim unstable
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      manifold but for negative time. Then, define e.g. global stable manifold by W^s(x_0) := \bigcup_{t \leq 0} \phi_t(W^s_{loc})
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      Note that we search backwards in time for stable, and forwards for unstable! Centre Manifold: If x_0
      not hyperbolic (0 real part), then E^c is the centre subspace. Then \exists W^c parallel to E^c, of class C^r, and
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      invariant under flow. Want bifurcation at \mu = 0, so with change of variables first find eigercs v_1, v_2.
      Then, construct P := [v_1, v_2] s.t. \vec{x} = P\vec{\xi}. NOTE: first v_i in P is always associated with \text{Re}(\lambda) = 0.
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      Solve for \xi and then expand with \eta = h(\xi, \tilde{\mu}) Alt. Centre Manifold: If vector v_1 \sim E^c = [a, b]^T then
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      we have y = bx/a (e.g. [1,1]^T \to y = x. If bifurcation at \mu = \alpha then have \mu = \tilde{\mu} + \alpha s.t. bifurcation
      when \tilde{\mu}=0. Then have \dot{x}(x,y,\tilde{\mu})=\ldots etc. Next, set up y=h(x,\tilde{\mu})=bx/a+b_1\tilde{\mu}+b_2\tilde{\mu}^2+a_2x^2+c_2\tilde{\mu}x
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      and proceed as usual but at \tilde{\mu}=0, s.t. y is along E^c. Transcritical Bifurcation: Always two points,
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      change type at origin. E.g. \dot{x} = \mu x - x^2 Saddle-Node: E.g. \dot{x} = \mu - x^2 Bifurcation begins to exist
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      at origin. Supercritical: E.g. \dot{x} = \mu x - x^3, where stable \to 2 \times stable and one unstable. Subcritical:
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      E.g. \dot{x} = -\mu x + x^3, where unstable \rightarrow 2 \times unstable and one stable. General co-dim 1: If \dot{x} = f
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      then \dot{x} = \mu f_u + 0.5x^2 f_{xx} + x\mu f_{x\mu} + 0.5\mu^2 f_{\mu\mu}. Generally this is a saddle-node but if f_u = 0 we have
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      \dot{x} = x\mu f_{x\mu} + 0.5x^2 f_{xx}, which is a transcritical. However if flows invariant under x = -x (reflectional
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      symmetry) then \dot{x} = x(\mu f_{x\mu} + \ldots) + x^3(f_{xxx}/6 + \ldots) \to \text{pitchfork}. Saddle-node stable under perturba-
27
      tions! Homoclinic Orbits Sum of roots of cubic = - coeff. of x^2
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      FPDE: Types: 1^{st}: \exists scale s.t. solution found, not so for 2^{nd}. Heat: T = u(T_{\infty} - T_{-\infty}) + T_{-\infty} Oil
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      Spread: Dims: x = x_f + \varepsilon \xi, t = \tau Ground Spread: (1 - s)\phi h_t + Q_x = 0; Q \sim -hh_x, 0 < x_s < x_f. Have
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      h(x_f) = 0, h_t(x_s) = 0, and hh_x|_{x=0,x_f} = 0 (i.e. no flux at centre and front), and h, hh_x cont. at joint.
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      Expansions: Let \xi = z + \epsilon \eta for perturbations Scale: Try x = x_f + \epsilon \xi for groundwater Stefan: S_0 = \epsilon \xi
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      C\left(T_{1}-T_{m}\right)/L, condition = \rho L\dot{s}=kT_{x}|_{s-}^{s+} 1ph Stefan: Bar = T_{h}|liq|_{s}sol|INS. Use T=T_{m}+(T_{1}-T_{m})u
33
      s.t. S_0 u_t = u_{xx}, u = 1 @ x = 0, \{\dot{s} = -u_x, u = 0\} @ x = s, s(0) = 0. Sim. sol is s = \beta \sqrt{t}, f = f(x/\sqrt{t})
      2ph Stefan: Use T = T_m + (T_1 - T_m)u s.t. S_0u_t = u_{xx} @ 0 < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} @ s < x < s, (S_0/\kappa)u_t = u_{xx} 
      1, u = 1 @ x = 0, u_x = 0 @ x = 1, \{\dot{s} = Ku_x|_{s_+} - u_x|_{s_-}, u = 0\} @ x = s, \{s = 0, u = -\theta\} @ x = 0. \text{ Here } u_x = 0, u = 0, u = 0\}
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      \theta := (T_m - T_0)/(T_1 - T_m), \kappa := c_1 k_1/(c_2 k_2), K := k_2/k_1 \text{ Sim. sol is } s = \beta \sqrt{t}, f = f(x/\sqrt{t}) 2-Dim:
37
      U_n = \hat{n} \cdot u = K(u_2)_n - (u_1)_n. If x = f(y,t) then \hat{n} := \nabla(x-f) = [1,-f_y]^T/\sqrt{1+f_y^2} Welding:
38
      Have 0 < s_2 < s_1. Have cold x = a, no flux x = 0. \theta = 1 in liquid. In mush \rho L\theta_t = J^2/\sigma, CoE
      \to \theta \rho L \dot{s} + k T_x \Big|_{s_-}^{s_+} = 0. Have \theta cont. (= 0) at s_1. I.e. we have S_0 u_t = u_{xx} + q, u_x = 0 @ x = 0, u = 0
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      -1 @ x = 1, \theta = 0 @ x = s_1. Also \theta_t = q in mush.
41
      FMM: Integral Constraint If J[y] = \int F dx with \int G dx = C then \tilde{J}[y] = \int F - \lambda G dx Hamilto-
      nian: H:=y'F_{y'}-F\to H'=-F_x. If F=F(y,y) then H=C Hamilton's Eqs. p:=F_{y'},q=y
43
      and so p' = -H_q, q' = H_p Free Boundary: J[y,b] = \int_a^b F(x,y,y')dx where b free. Expand with
      y + \epsilon \eta, b + \epsilon \beta \to J = J_0 + \epsilon \left\{ \int_a^b \eta F_y + \eta' F_{y'} dx + \beta F(b, y(b), y'(b)) \right\} \text{If } y(b) = d \to d = y(b + \epsilon \beta) + \epsilon \eta(b + \epsilon \beta) = 0
      y(b) + \epsilon(\beta y'(b) + \eta(b)) so \eta(b) = -\beta y'(b). IVP on integral so \beta [F - y'F_{y'}]_{x=b} + \int (\ldots) = 0 so F = y'F_{y'} at free boundary. Control: Have \int \xi h_x + \eta h_u dt = 0, \dot{\xi} = \xi f_x + \eta f_u. Sub for \eta, IVP s.t. \frac{d}{dt} \frac{h_u}{f_u} = h_x - f_x \frac{h_u}{f_u}
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47
      and \dot{x}=f Hamiltonian (Control): H:=f\frac{h_u}{f_u}-h s.t. \dot{H}=\frac{h_u}{f_u}f_t-h_t \to \text{autonomous if } h_t=f_t=0.
48
      Fredholm Alt Integ Eqs. For y = f + \int K(x,t)y(t)dt we have ONE (N) has a unique sol y = 0 if f = 0,
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      and adjoint has unique sol, or TWO (H) as sols y_1 \dots y_r iff \forall solutions to H^*, z_i, we have \langle f, z_i \rangle = 0.
      GENERAL CASE: Have y = f + \lambda AG_1 + \lambda BG_2. Solve for system [\alpha_1, \alpha_2; \alpha_3, \alpha_4][A, B]^T = [\gamma_1, \gamma_2]^T
51
      with NONUNIQUE sols for \lambda = \lambda_*. Now for \lambda = \lambda_*, want to solve L^*w = 0 and show this is orthogonal to RHS. First solve [\alpha_1, \alpha_2; \alpha_3, \alpha_4][A, B]^T = [0, 0]^T. Then we have w = \lambda_* A(F(G_1, G_2)).
52
53
      Check if \int fw = 0. If so, return to NONHOM case and solve [\alpha_1, \alpha_2; \alpha_3, \alpha_4]_{\lambda_*}[A, B]^T = [\gamma_1, \gamma_2]^T to get
      B = -\frac{\alpha_1}{\alpha_2}A + \frac{\gamma_1}{\alpha_2}. Sub this into y = f + \lambda_*AG_1 + \lambda_*B(A)G_2. EX: Solve y = 1 - x^2 + \lambda \int (1 - 5x^2t^2)y(t) dt =
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1 - x^2 + \lambda A - 5\lambda Bx^2. \text{ Have } A := \int y_N(t) = \int 1 - x^2 + \lambda A + \dots = \lambda A - \frac{5\lambda}{3} + \frac{2}{3}. \text{ Repeat for } B \text{ s.t.} \\
[1 - \lambda, 5\lambda/3; -\lambda/3, 1 + \lambda][A, B]^T = [2/3, 2/15]^T. \text{ Unique sols if } \lambda \neq \pm \frac{3}{2} \to \text{try when } \lambda_* = \frac{-3}{2}. \text{ Have } L^*w_H = \lambda A - 5\lambda_*Bx^2 \to A := \int \lambda_*A - 5\lambda_*Bx^2, \text{ and } B := \int \dots \text{ Both give consistent results } A = B \text{ so}

w_H = \lambda_* A(1-5x^2). Check that \int w_H(x)(1-x^2) = 0, so we have shown nullspace of adj. orthog. to
RHS. Note that we may also find A, B for adjoint quicker via [1 - \lambda, 5\lambda/3; -\lambda/3, 1 + \lambda]_{\lambda_*}[A, B]^T = [0, 0]^T.
Lastly, return to (N), and having verified \lambda_* permits a solution, solve [1-\lambda, 5\lambda/3; -\lambda/3, 1+\lambda]_{\lambda_*}[A, B]^T =
[2/3,2/15]^T \to A-B=4/15 when \lambda=-3/2. Sub this into y=1-x^2\dots for solution. Fred Diff Eq.
For nonunique sol to exist, need \langle Ly, w \rangle = \langle f, w \rangle \forall ws.t.L^*w = 0
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