

# Ph291E Lab 4 – Bessel & Telescope

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## 1 Purpose

Using Bessel's method to determine the focal lengths of two different converging lenses, we will construct a simple refracting telescope using these lenses and determine its magnification both experimentally and theoretically.

## 2 Data

### Approximate Focal Lengths

$$f_1 = 10 \pm 0.3 \text{ cm}$$

$$f_2 = 4.6 \pm 0.6 \text{ cm}$$

## Measurements for Bessel's Method Calculations for Focal Lengths

Table 1: Bessel's Method Data for Lens 1

D (cm)	d (cm)	Random Error D (cm)	Random Error d (cm)	Optical Bench Inst. Error (cm)
43.00	9.85			
43.00	9.85			
43.00	9.25			
43.00	9.20			
43.00	9.60			
43.00	9.75	0	0.12	0.05
<b>Mean D</b>				<b>43.00 cm</b>
<b>Mean d</b>				<b>9.58 cm</b>

Image sharpness uncertainty (for position d):

- Position 1: 0.30 cm
- Position 2: 0.20 cm

Table 2: Bessel's Method Data for Lens 2

D (cm)	d (cm)	Random Error D (cm)	Random Error d (cm)	Optical Bench Inst. Error (cm)
24.95	7.45			
24.95	8.25			
24.95	7.55			
24.95	7.95			
24.95	8.20			
24.95	7.30	0	0.16	0.05
<b>Mean D</b>				<b>24.95 cm</b>
<b>Mean d</b>				<b>7.78 cm</b>

Image sharpness uncertainty (for position d):

- Position 1: 0.05 cm
- Position 2: 0.60 cm

### 3 Calculations

#### Sample Calculations for Distance d (Lens 1)

Mean Calculation:

$$\bar{x} = \frac{1}{n} \sum_{i=1}^n x_i$$

$$\bar{x} = \frac{1}{6}(9.85 + 9.85 + 9.25 + 9.20 + 9.60 + 9.75)$$

$$\bar{x} = 9.5833 \text{ cm}$$

Standard Deviation:

$$S_x = \sqrt{\frac{1}{n-1} \sum_{i=1}^n (x_i - \bar{x})^2}$$

$$S_x = \sqrt{\frac{1}{6-1} [(9.85 - 9.5833)^2 + (9.85 - 9.5833)^2 + \cdots + (9.75 - 9.5833)^2]}$$

$$S_x = 0.29268 \text{ cm}$$

Standard Deviation of the Mean (SDOM):

$$\sigma_{\bar{x}} = \frac{S_x}{\sqrt{n}}$$

$$\sigma_{\bar{x}} = \frac{0.29268}{\sqrt{6}}$$

$$\sigma_{\bar{x}} = 0.11948 \text{ cm}$$

The same procedure was applied to calculate the mean, standard deviation and standard deviation of the mean for the remaining measurements.

#### Bessel's Method Calculations for Lens 1

$$f = \frac{D^2 - d^2}{4D}$$

$$f = \frac{(43.00)^2 - (9.58)^2}{4(43.00)}$$

$$f = 10.22 \text{ cm}$$

### Error Propagation for Bessel's Method (Lens 1)

Since the measurements of  $D$  and  $d$  are independent, we can use the independent error propagation formula:

$$\delta f = \sqrt{\left(\frac{\partial f}{\partial D}\delta D\right)^2 + \left(\frac{\partial f}{\partial d}\delta d\right)^2}$$
$$\delta f = \sqrt{\left(\frac{D^2 + d^2}{4D^2}\delta D\right)^2 + \left(-\frac{d}{2D}\delta d\right)^2}$$

**Chosen uncertainties:**  $\delta D = 0.05$  cm (Instrumental uncertainty, larger than random) and  $\delta d = 0.30$  cm (Image sharpness uncertainty, larger than random and instrumental).

$$\delta f = \sqrt{\left(\frac{(43.00)^2 + (9.58)^2}{4(43.00)^2}(0.05)\right)^2 + \left(-\frac{9.58}{2(43.00)}(0.30)\right)^2}$$
$$\delta f = 0.036 \text{ cm}$$

### Angular Magnification

Where  $f_{obj}$  is the focal length of the lens 1 and  $f_{eye}$  is the focal length of the lens 2.

$$m_\theta = -\frac{f_{obj}}{f_{eye}}$$
$$m_\theta = -\frac{10.216}{5.630}$$
$$m_\theta = -1.814$$

### Error Propagation for Angular Magnification

Since the measurements of  $f_{obj}$  and  $f_{eye}$  are independent, we can use the independent error propagation formula:

$$\delta m_\theta = \sqrt{\left(\frac{\partial m_\theta}{\partial f_{obj}}\delta f_{obj}\right)^2 + \left(\frac{\partial m_\theta}{\partial f_{eye}}\delta f_{eye}\right)^2}$$
$$\delta m_\theta = \sqrt{\left(-\frac{1}{f_{eye}}\delta f_{obj}\right)^2 + \left(\frac{f_{obj}}{f_{eye}^2}\delta f_{eye}\right)^2}$$

**Chosen uncertainties:**  $\delta f_{obj} = 0.036$  cm (calculated) and  $\delta f_{eye} = 0.095$  cm (calculated).

$$\delta m_\theta = \sqrt{\left(-\frac{1}{5.630}(0.036)\right)^2 + \left(\frac{10.216}{(5.630)^2}(0.095)\right)^2}$$

$$\delta m_\theta = 0.031$$

## 4 Results

Focal Lengths:

- Lens 1:  $f_{obj} = 10.22 \pm 0.04$  cm
- Lens 2:  $f_{eye} = 5.63 \pm 0.10$  cm

Angular Magnification:

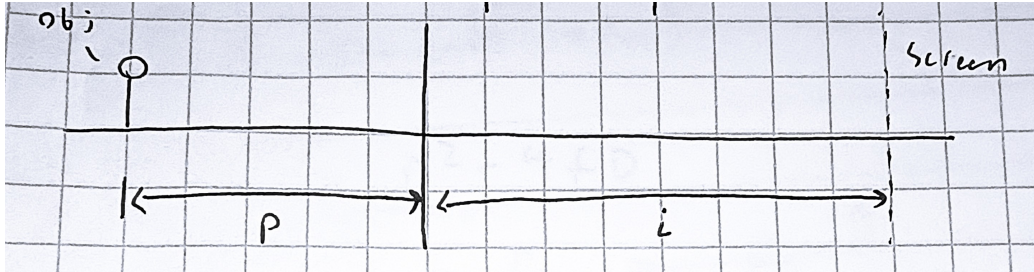
- Theoretical:  $m_\theta = -1.81 \pm 0.03$
- Experimental approximation:  $m_\theta \approx -1.36$

## 5 Conclusion

The focal lengths of the two lenses were determined using Bessel's method to be  $10.22 \pm 0.04$  cm for lens 1 and  $5.63 \pm 0.10$  cm for lens 2. The approximate focal length of lens 1 ( $10 \pm 0.3$  cm) is within the uncertainty range of the calculated focal length, while the approximate focal length of lens 2 ( $4.6 \pm 0.6$  cm) is slightly outside the uncertainty range of the calculated focal length. The theoretical angular magnification of the telescope was calculated to be  $-1.81 \pm 0.03$ , while the experimental approximation yielded a value of  $-1.36$ . This discrepancy could be due to systemic errors such as the lenses not being set up exactly  $f_{obj} + f_{eye}$  apart, and bias in visually comparing the lengths of tape seen through the telescope.

## 6 Answers to questions

### Question 1



Thin lens equation:

$$\frac{1}{f} = \frac{1}{P} + \frac{1}{i}$$

$$\frac{1}{f} = \frac{P+i}{Pi}$$

Since  $D = P + i$  and  $i = D - P$ ,

$$\frac{1}{f} = \frac{D}{P(D - P)}$$

$$f = \frac{P(D - P)}{D} = \frac{PD - P^2}{D}$$

Putting into standard form to find the 2  $P$  values that forms a sharp image:

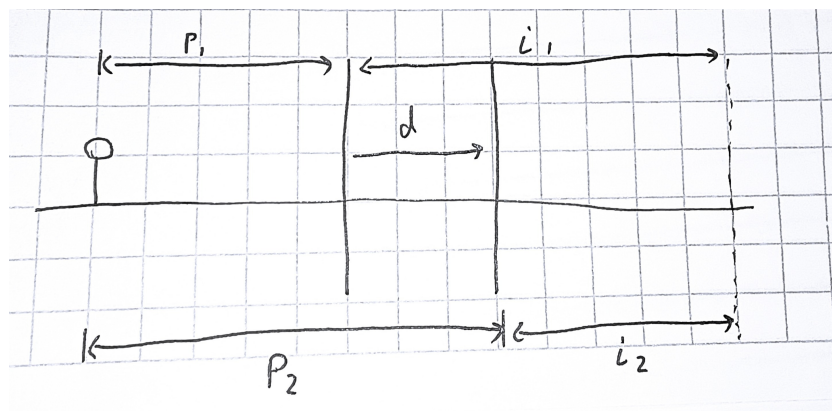
$$P^2 - PD + fD = 0$$

Using the quadratic formula:

$$P = \frac{D \pm \sqrt{D^2 - 4fD}}{2}$$

Since the two values for  $P$  are real and unique, the discriminant must be non-zero:

$$\boxed{D^2 - 4fD > 0 \implies D > 4f}$$



The distance between the two positions of the lens  $P_2 - P_1 = d$  is:

$$d = P_2 - P_1 = \sqrt{D^2 - 4fD}$$

Rearranging for  $f$ :

$$d^2 = D^2 - 4fD$$

$$4fD = D^2 - d^2$$

$$f = \frac{D^2 - d^2}{4D}$$

## Question 2

Thin lens equation:

$$\frac{1}{f} = \frac{1}{P} + \frac{1}{i}$$

$$\frac{1}{f} = \frac{P+i}{Pi}$$

$$f = \frac{Pi}{P+i} = \frac{i}{1+i/P}$$

For the approximate focal length we assumed  $P \rightarrow \infty$ , so  $i/P \rightarrow 0$ :

$$f_{approx} = i$$

However,  $P$  is finite but large, so

$$1 + i/P > 1$$

$$\Rightarrow f = \frac{i}{1 + i/P} < i = f_{approx}$$