

WEEK 3 REVIEW

① Interactions and detection of charged particles

Bethe-Bloch Equation

Ionization energy

loss per cm
unit length

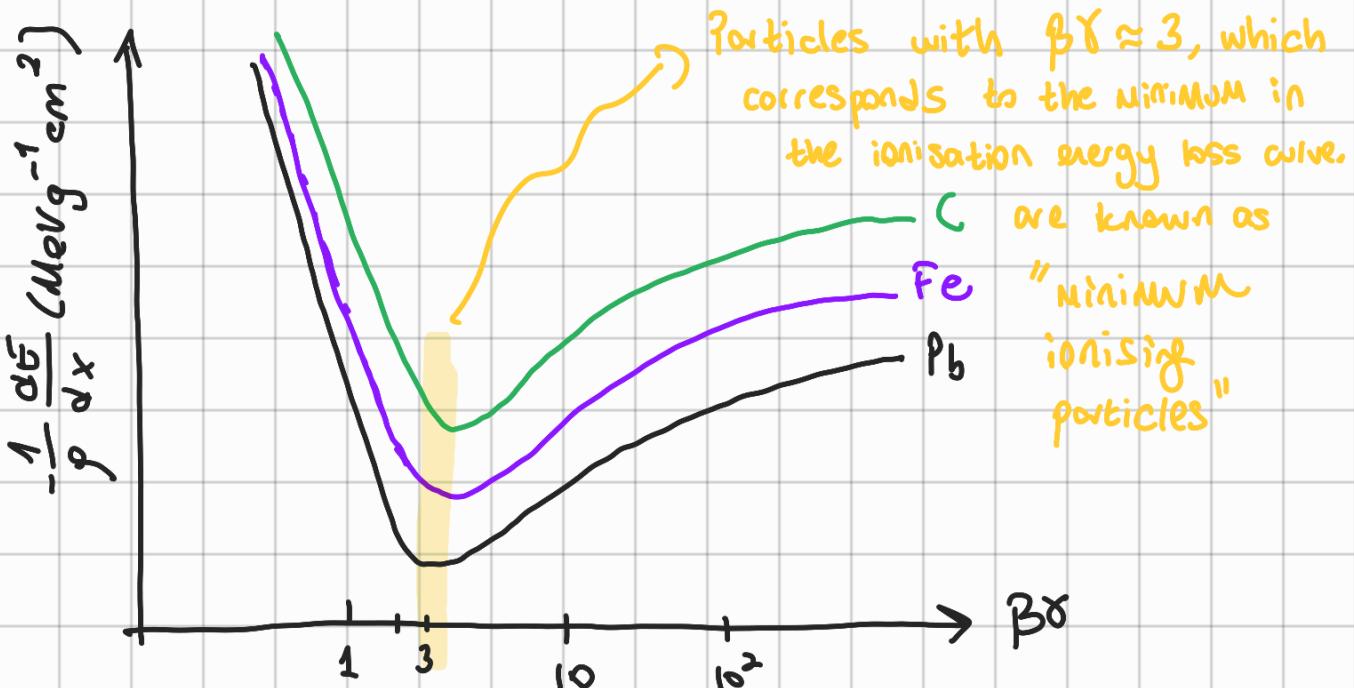
$$\frac{dE}{dx} \approx -\frac{4\pi h^2 c^2 \alpha^2 n Z}{me V^2} \left\{ \ln \left[\frac{2\beta^2 \gamma^2 c^2 N_e}{I_e} \right] - \beta^2 \right\}$$

\Rightarrow effective ionization potential

For modern particle physics where $V \approx c$

In this case, for a given medium $\frac{dE}{dx}$ depends logarithmically on $(\beta\gamma)^2$ where

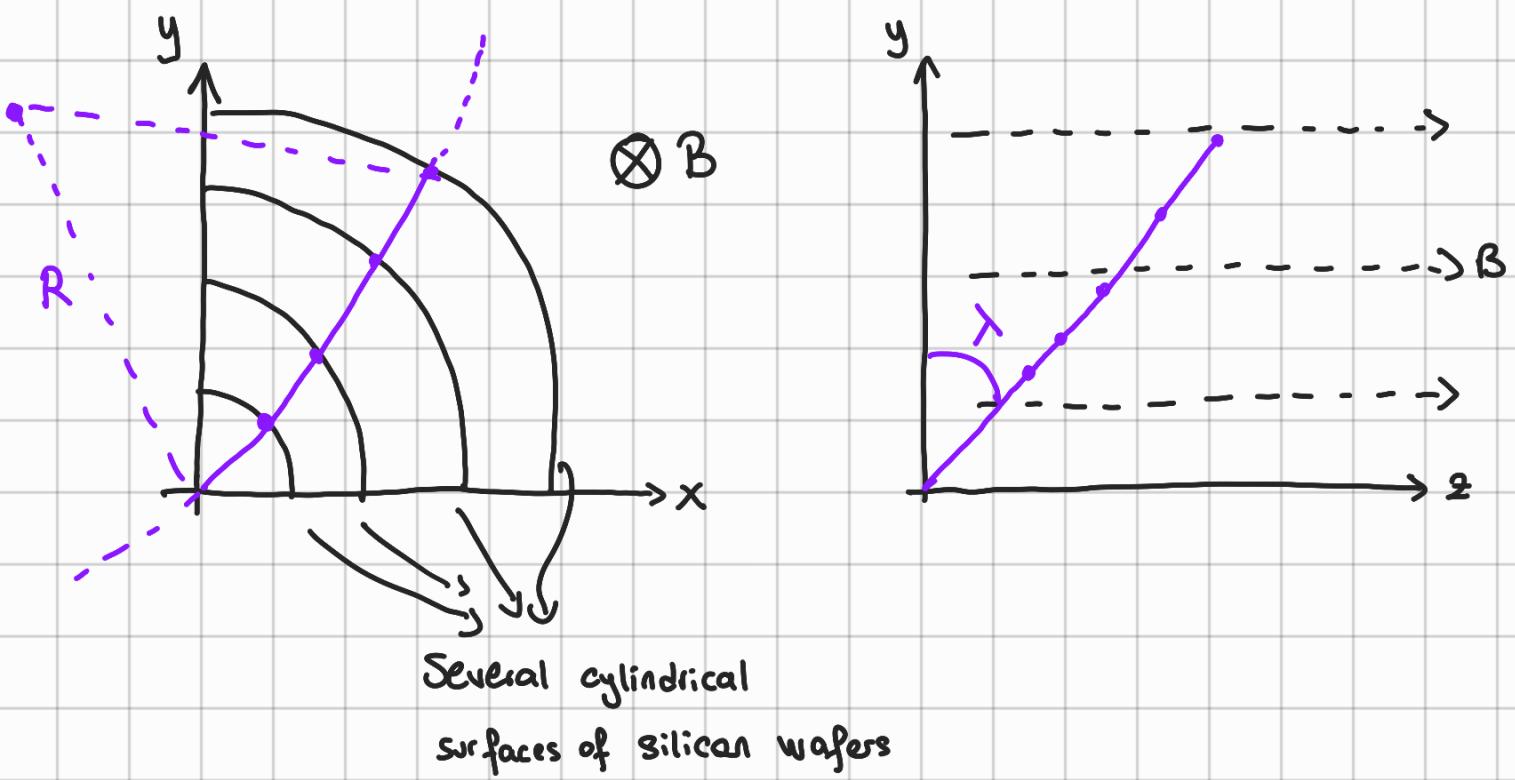
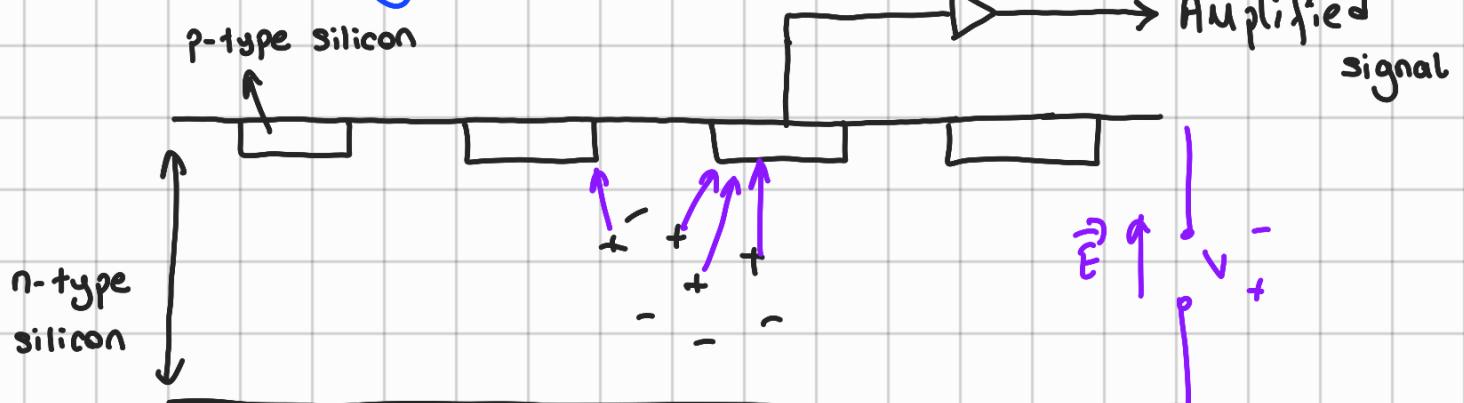
$$\beta\gamma = \frac{V/c}{\sqrt{1-(V/c)^2}} = \frac{P}{mc}$$



$$\frac{1}{g} \frac{dE}{dx} \approx -\frac{4\pi h^2 c^2 \alpha^2}{me V^2} \frac{Z}{Am_u} \left\{ \ln \left[\frac{2\beta^2 \gamma^2 N_e c^2}{I_e} \right] - \beta^2 \right\}$$

atomic mass number atomic mass unit

→ Tracking Detectors



$$p \cos \lambda = 0.3 B R$$

Measuring hits yield λ and R , so that p can be calculated

→ Scintillation Detectors

In an organic scintillator, passage of a charged particle leaves some of the molecules in an excited state

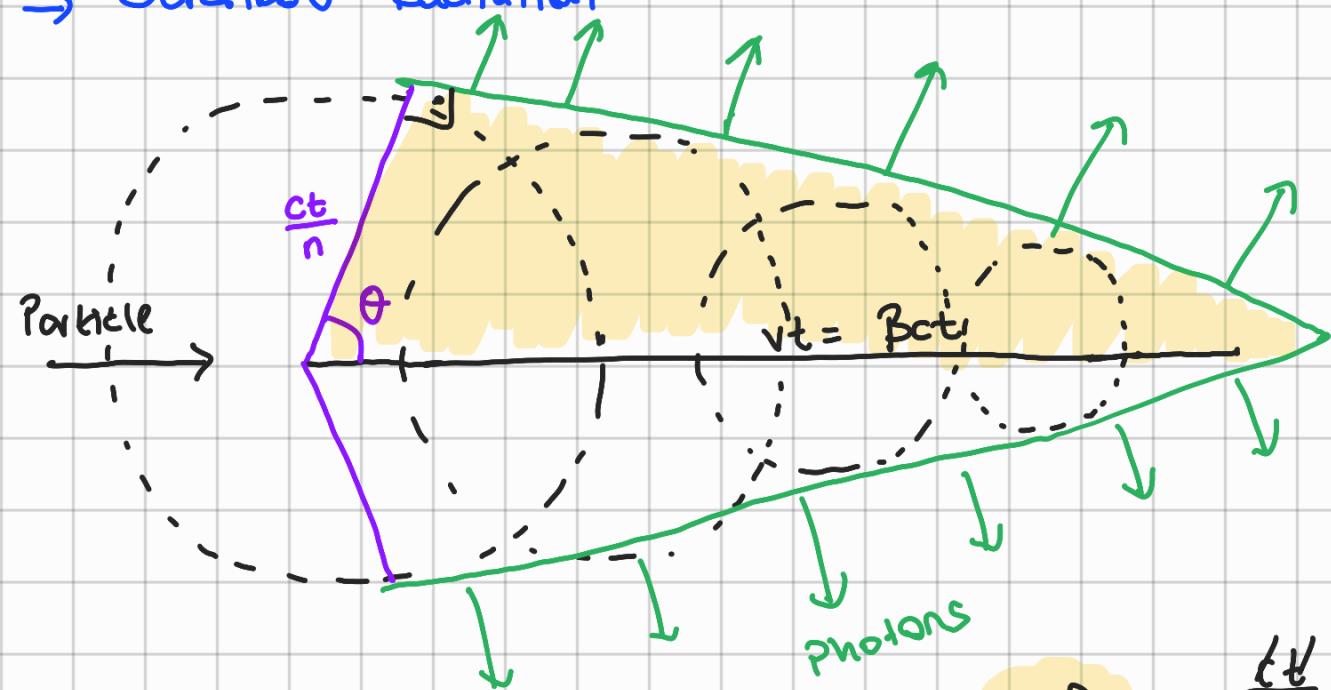
Decay of the excited state \rightsquigarrow emission of UV light

\downarrow
Add fluorescent dyes to the scintillator

\downarrow
Dye absorbs UV light \rightsquigarrow re-emits blue photons

\downarrow
Blue light can be detected by
photomultiplier

→ Čerenkov Radiation



v = Particle's velocity

$\frac{c}{n}$ = Speed of the light in the medium

If $v > c/n$ constructive interference occurs

$$\cos \theta = \frac{\frac{ct}{n}}{\beta \sin \theta} = \frac{1}{\beta n}$$

\checkmark Čerenkov is only emitted only when $\beta > 1/n$

for a relativistic particle

$$p = \gamma M v = \gamma m \beta c \quad |$$
$$\left. \begin{array}{l} p = E \beta \\ \beta = \frac{pc}{E} = \frac{p}{\sqrt{p^2 + m^2 c^2}} \end{array} \right\}$$
$$E = \gamma M c^2 \xrightarrow{c=1} E = \gamma M$$

$$\beta > \frac{1}{n}$$

$$\frac{p}{\sqrt{p^2 + m^2 c^2}} > \frac{1}{n}$$

$$\frac{p^2}{p^2 + m^2 c^2} > \frac{1}{n^2}$$

$$n^2 > \frac{p^2 + m^2 c^2}{p^2} = 1 + \frac{m^2 c^2}{p^2}$$

$$\sqrt{n^2 - 1} > \sqrt{\frac{m^2 c^2}{p^2}}$$

$$(n^2 - 1)^{1/2} > \frac{mc}{p}$$

$$mc < (n^2 - 1)^{1/2} p$$

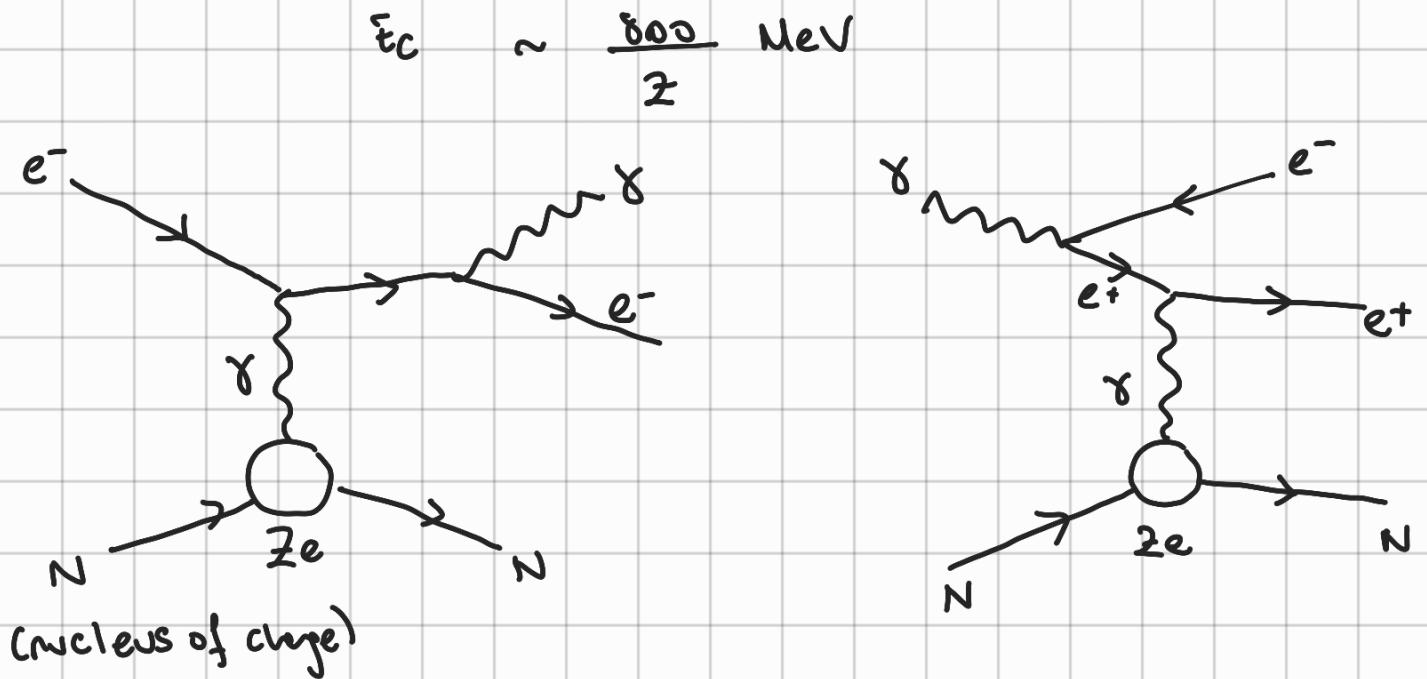
Only particles with this "m" will produce

Čerenkov Radiation

(2) Interactions and detection of electrons and photons

At low energies, the energy loss of electrons is dominated by ionisation.

Energies $> E_c$ (critical energy) "Bremsstrahlung" is the main energy loss mechanism



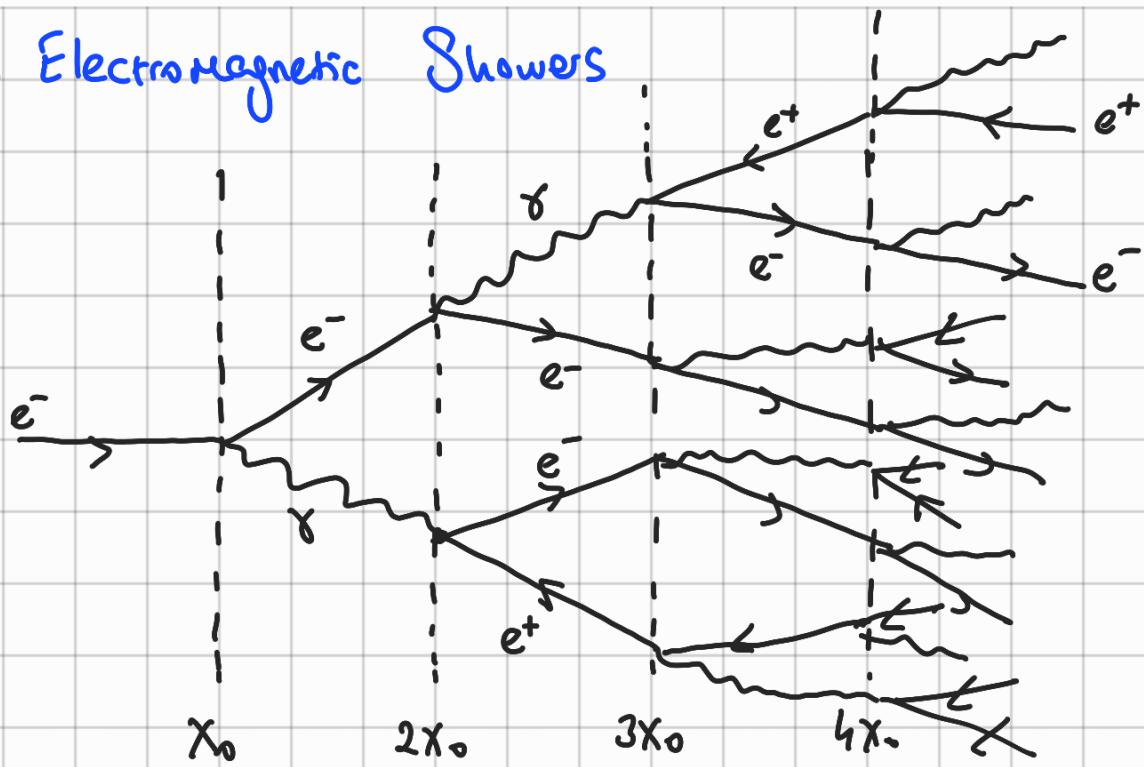
X_0 : The radiation length

(Avg. distance over which the energy of an electron is reduced by bremsstrahlung by a factor $1/e$)

$$X_0 \approx \frac{1}{4 \times n Z^2 r_e^2 \ln(287/2^{12})}$$

atomic number $2.8 \times 10^{-15} \text{ m}$ (classical radius of the e^-)

→ Electromagnetic Showers



$$\langle \bar{v} \rangle \approx \frac{E}{2^x}$$

✓

avg. energy of
the particles after x rad. length

$$x_{\max} = \frac{\ln(E/E_c)}{\ln 2}$$

Collider Experiments

detection
from
missing
momentum

