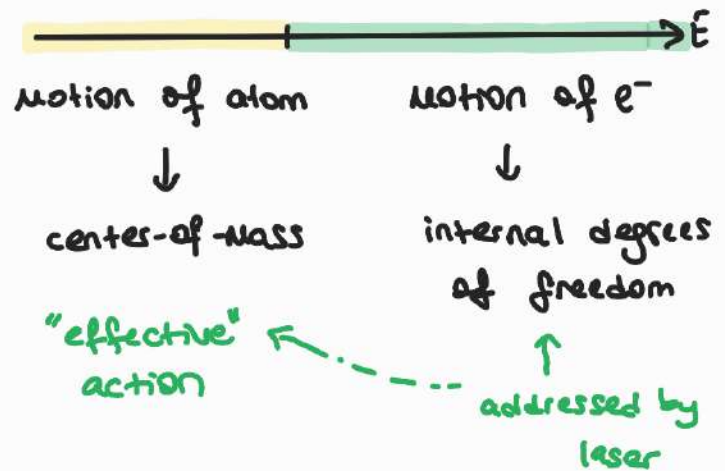


# MECHANICAL EFFECTS OF LIGHT

Idea: . Photon have momentum.

. "Effective" theory



## I. FORMULATION

### 1) Hilbert Space

$$\mathcal{H} = \mathcal{H}_{at} \otimes \mathcal{H}_{field} = \mathcal{H}_{c.o.m} \otimes \mathcal{H}_{internal} \otimes \mathcal{H}_{laser\ mode} \otimes \mathcal{H}_{vac}$$

$$\mathcal{H}_{c.o.m} = \text{Sp} \{ |\vec{p}\rangle, \vec{p} \in \mathbb{R}^3 \} \quad \text{atom in free space}$$

↓  
momentum

$$\mathcal{H}_{internal} = \text{Sp} \{ |n, l, m, j, \dots\rangle, \dots \} \quad \text{energy levels of the atom}$$

$$\mathcal{H}_{internal} = \text{Sp} \{ |n\rangle, n \in \mathbb{N} \} \quad \text{one particular mode addressed by the laser}$$

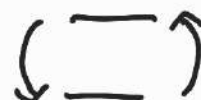
$$\mathcal{H}_{vac} = \mathcal{F} \quad \text{Fock space}$$

### 2) Simplification

. Reduce  $\mathcal{H}_{int}$  to a 2 dim. Hilbert space =  $\text{Sp} \{ |g\rangle, |e\rangle \}$



cycling transition



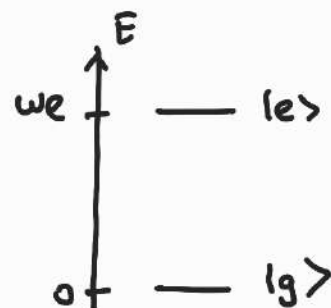
- Laser mode is in a coherent state  $\rightarrow$  classical field
- Vacuum as a reservoir (Markov approximation)

Lindblad Equation

- C.o.m motion is slow compared with internal dynamics.

### 3) Hamiltonian

Acts on  $\mathcal{H}_{\text{c.o.m}} \otimes \mathcal{H}_{\text{int}}$



$$\hat{H} = \frac{\vec{p}^2}{2m} \otimes \hat{\mathbb{I}} + \hat{\mathbb{I}} \otimes \omega_e |e\rangle\langle e| + \frac{\Omega(\vec{r})}{2} \left[ e^{i(\omega_0 t - \vec{k}_L \cdot \vec{r})} \otimes |g\rangle\langle e| + \text{h.c.} \right]$$

key points:  $\Omega$  and  $\phi = \omega_0 t - \vec{k}_L \cdot \vec{r}$  are operators on  $\mathcal{H}_{\text{com}}$

within dipole approximation : "size" of the atom  $\ll \lambda$

the atom "feels"  $\Omega$  and  $\phi$  at the position of its c.o.m.

$$\Omega(\vec{r}) = \langle \Psi_{\text{com}} | \Omega(\vec{r}) | \Psi_{\text{com}} \rangle \text{ for c.o.m in state } |\Psi_{\text{com}}\rangle$$

$$\Omega(\vec{r}) = -\vec{E}(\vec{r}) \cdot \vec{d}_0$$

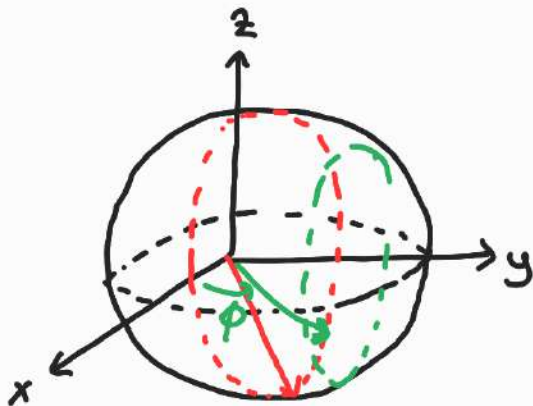
$$\phi(\vec{r}) = \omega_0 t - \vec{k}_L \cdot \vec{r}$$

$\rightarrow$  how c and atom  
bound together

## II. Recoil and Doppler Effects

### 1) Phase Imprinting

Remainder:  $\hat{H} = \frac{\Omega}{2} (e^{-i\phi} |e\rangle\langle g| + \text{h.c.}) + \frac{|\Omega|^2}{2} \hat{n} \hat{\sigma}^z$



Here  $\phi(\vec{r})$

Consider  $|\Psi_{at}\rangle = |\vec{p}\rangle \otimes |g\rangle$

$$\hat{H}_{int} |\Psi_{at}\rangle = \frac{\Omega}{2} \underbrace{(e^{-i\phi(\vec{r})} |\vec{p}\rangle)} \otimes |e\rangle$$

$$\int d^3\vec{r} |\vec{r}\rangle \langle \vec{r}| \vec{r} \times \vec{r} |e^{-i\phi(\vec{r})} |\vec{p}\rangle$$

$$= \int d^3\vec{r} e^{-i\phi(\vec{r})} |\vec{r}\rangle \times \underbrace{|\vec{r}\rangle}_{e^{i\vec{p}\cdot\vec{r}}}$$

$$= \int d^3\vec{r} e^{-i\phi(\vec{r}) + i\vec{p}\cdot\vec{r}} |\vec{r}\rangle$$

Plane wave:  $\phi(\vec{r}) = \omega_0 t - \vec{k}_L \cdot \vec{r}$  then:

$$\hat{H}_{int} |\Psi_{at}\rangle = \frac{\Omega}{2} e^{i\omega_0 t} \int d^3r e^{i(\vec{k}_L \cdot \vec{r} + \vec{p} \cdot \vec{r})} |\vec{r}\rangle$$

$$|\vec{p} + \vec{k}_L\rangle$$

$$\text{So } \hat{H}_{int} = \frac{\Omega}{2} \sum_{\vec{p}} e^{i\omega_0 t} |\vec{p}\rangle \langle \vec{p} + \vec{k}_L, e| + h.c$$

## 2) Shifts

Move to the rest frame of the atom (from the laboratory frame where momentum  $\vec{p}$ )

Unitary transformation:

$$\hat{U}(t) = e^{\frac{i\hat{p}^2 t}{2m}} \quad \hat{H} \rightarrow (i\partial_t \hat{U})\hat{U}^\dagger + \hat{U}\hat{H}\hat{U}^\dagger$$

In this frame:

$$\hat{H} = \frac{\Omega}{2} \sum_{\vec{p}} e^{i\omega_0 t} \hat{U} |\vec{p}\rangle \langle \vec{p} + \vec{k}_L, e| \hat{U}^\dagger + h.c$$

$$e^{\frac{i\hat{p}^2 t}{2m}} |\vec{p}\rangle \langle \vec{p} + \vec{k}_L, e| e^{-\frac{i(\vec{p} + \vec{k}_L)^2 t}{2m}}$$

$$= \frac{\Omega}{2} \sum_{\vec{p}} e^{i(\omega_0 - \frac{\vec{p} \cdot \vec{k}}{m} - \frac{\vec{k}^2}{2m})t} |\vec{p}, g \times \vec{p} + \vec{k}, e| + h.c$$

$$\omega_0 \longrightarrow \omega_0 - \frac{\vec{p} \cdot \vec{k}}{m} - \frac{\vec{k}^2}{2m}$$

$${}^6\text{Li} \sim 0.1 \text{ ms}^{-1}$$

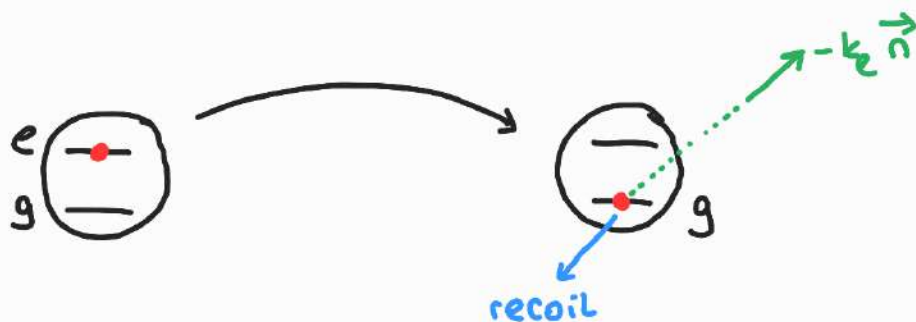
Doppler Shift:  $\frac{\vec{p} \cdot \vec{k}}{m} = \vec{p} \cdot \vec{v}_{\text{rec}}, \quad \vec{v}_{\text{rec}} = \frac{\vec{k}}{m}$

↳ Ex:  ${}^6\text{Li}$  at 300K : 1GHz

Recoil Shift:  $E_R = \frac{\hbar^2 k^2}{2m} = \hbar \cdot 77 \text{ kHz}$  for  ${}^6\text{Li}$

negligible comp with doppler shift

### 3) Jump Operators



$$k_e = \frac{\omega_e}{c}$$

Family of jump operators:  $\hat{L}(\vec{n}) = \sum_{\vec{p}} |\vec{p} + \hbar \vec{n}, g \times \vec{p}, e|$

$$= e^{i \hbar \vec{n} \cdot \vec{\nabla}} \otimes |g \times e|$$



Rate  $\Gamma_{\vec{n}}$  where  $\hat{M}(\vec{n}) = \sqrt{\Gamma_{\vec{n}}} \hat{L}(\vec{n})$   
 $\downarrow$   
 measurement operator

#### 4) Lindblad Equation

$$\dot{\hat{\rho}} = -i [\hat{H}, \hat{\rho}] + \sum_{\vec{n}} \Gamma_{\vec{n}} \left( \hat{L}(\vec{n}) \hat{\rho} \hat{L}(\vec{n})^\dagger - \frac{1}{2} \hat{L}^\dagger \hat{L}(\vec{n}) \hat{\rho} - \frac{1}{2} \hat{\rho} \hat{L}^\dagger \hat{L}(\vec{n}) \right)$$

Note that:  $\hat{L}^\dagger(\vec{n}) \hat{L}(\vec{n}) = |ex\rangle\langle ex|$   
 $\underbrace{e^{-i\cdot} |ex\rangle\langle ex| e^{i\cdot}}_{\text{}} = |ex\rangle\langle ex|$

Introduce  $\Gamma = \sum_{\vec{n}} \Gamma_{\vec{n}}$

see see

$$\dot{\hat{\rho}} = -i [\hat{H}, \hat{\rho}] - \underbrace{\frac{\Gamma}{2} (|ex\rangle\langle ex| \hat{\rho} + \hat{\rho} |ex\rangle\langle ex|)}_{\text{Non-Hermitian "decay" of } |e\rangle} + \underbrace{\sum_{\vec{n}} \hat{L}(\vec{n}) \hat{\rho} \hat{L}^\dagger(\vec{n})}_{\text{kicks upon spontaneous emission}}$$

Non-Hermitian "decay" of  $|e\rangle$

kicks upon spontaneous emission

last term  $\hat{p} = |e\rangle\langle p| \times |p\rangle\langle e|$

Then  $\hat{L}(\vec{n}) \hat{\rho} \hat{L}^\dagger(\vec{n}) = |g, \vec{p} + k\vec{n}\rangle\langle g, \vec{p} + k\vec{n}|$

### III. Semi-classical Forces

#### 1) Further simplification

. The momentum "spread" of the atom is large  $\Delta p \geq k_e, k_L, \dots$

$\Leftrightarrow$  position spread is small  $\Delta R \ll \lambda$

.  $k_e$  is small (momentum exchanges with em field are continuous)

. internal dynamics is much faster than motion of com

Steady state description of the internal state for each position.

$$\Gamma, \Omega \gg \bar{E}_R \rightarrow \text{recoil}$$

$\rightarrow$  Ansatz for the density matrix:  $\hat{\rho} = \hat{\rho}_{\text{com}} \otimes \hat{\rho}_{\text{int}}^{\text{steady state}}$

$$\hat{\rho}_{\text{int}}^{\text{st}}(\vec{R}(t))$$

## 2) Spontaneous Emission

$$\sum_{\vec{n}} \Gamma_{\vec{n}} e^{i k_e \vec{n} \cdot \vec{r}} |g \times e| \hat{p} |e \times g| e^{-i k_e \vec{n} \cdot \vec{r}}$$

Since  $k_e$  is small:  $e^{i k_e \vec{n} \cdot \vec{r}} \sim 1 + i k_e \vec{n} \cdot \vec{r} - \frac{k_e^2}{2} (\vec{n} \cdot \vec{r})^2 + \dots$

So:  $\sum_{\vec{n}} \Gamma_{\vec{n}} |g \times e| \hat{p} |e \times g| + \sum_{\vec{n}} i k_e \vec{n} \cdot \vec{r} |g \times e| \hat{p} |e \times g| + \sum_{\vec{n}} |g \times e| \hat{p} |e \times g| (-i k_e \vec{n} \cdot \vec{r}) + \mathcal{O}(k_e^2)$

Up to  $\mathcal{O}(k_e^2)$ : no net effect of spont. emission of the center of motion!

Now, up to  $\mathcal{O}(k_e^2)$ :

$$\dot{\hat{p}} = -i [\hat{H}, \hat{p}] + \Gamma |g \times e| \hat{p} |e \times g| - \frac{\Gamma}{2} (|e \times e| \hat{p} + \hat{p} |e \times e|)$$

the dissipative part of the Lindblad Eq. is only affecting the internal state!!



### 3) Forces

$$\vec{F} = \frac{d\langle \vec{p} \rangle}{dt} = \frac{d}{dt} \text{Tr}(\hat{\rho} \vec{p}) = \text{Tr}(\dot{\hat{\rho}}, \vec{p})$$

$$\begin{aligned} \text{using cyclicity} \quad \left\{ \begin{aligned} &= -i \text{Tr}([\hat{H}, \hat{\rho}] \cdot \vec{p}) \\ &= -i \text{Tr}(\hat{\rho} \cdot [\vec{p}, \hat{H}]) \end{aligned} \right. \end{aligned}$$

$$\vec{F} = -i \langle [\vec{p}, \hat{H}] \rangle$$

$\hat{H}$  depends on  $\vec{r}$  : move to position representation  
(via  $\Omega(\vec{r})$ ,  $\phi(\vec{r})$ ...)

$$\vec{p} = -i \vec{\nabla}_r$$

$$\text{then } \langle [\vec{p}, \hat{H}] \rangle = -i \langle \vec{\nabla}_r \hat{H} \rangle$$

Proof: for any state  $|\alpha\rangle =$

$$\langle \alpha | \vec{\nabla}_r (\hat{H} | \alpha \rangle) = \langle \alpha | [\vec{\nabla}_r \hat{H}] | \alpha \rangle + \langle \alpha | \hat{H} (\vec{\nabla}_r | \alpha \rangle$$

So,

$$\boxed{\vec{F} = - \langle \vec{\nabla}_r \hat{H} \rangle}$$

For an atom at rest:

$$\hat{H} = \omega_e |e\rangle\langle e| + \frac{\Omega(\vec{r})}{2} \left( e^{i(\omega_0 t - \vec{k}_L \cdot \vec{r})} |g\rangle\langle e| + \text{h.c.} \right)$$

In a frame rotating at  $\omega_0$  (for internal degrees of freedom)

$$\hat{H} = \Delta |e\rangle\langle e| + \frac{\Omega(\vec{r})}{2} \left( e^{-i\vec{k}_L \cdot \vec{r}} |g\rangle\langle e| + \text{h.c.} \right)$$

$$\langle -\vec{\nabla}_r \hat{H} \rangle = - \frac{\nabla \Omega(\vec{r})}{2} \left( \langle e^{-i\vec{k}_L \cdot \vec{r}} |g\rangle\langle e| + \text{h.c.} \rangle \right) \quad \begin{array}{l} \text{internal d.o.f} \\ \downarrow \\ \text{OBE} \\ \text{optical Bloch equation} \end{array}$$

$$+ i k_L \frac{\Omega}{2} \left( \langle e^{-i\vec{k}_L \cdot \vec{r}} |g\rangle\langle e| + \dots \right)$$

$$\vec{F} = \frac{\vec{\nabla} \Omega}{\Omega} \Delta \cdot \frac{s(\vec{r})}{1+s(\vec{r})} + k_L \frac{\vec{r}}{2} \frac{s(\vec{r})}{1+s(\vec{r})}$$

dipole force

$\vec{F}_{\text{dip}}$

$\downarrow$

$$\frac{\Delta}{2} \frac{\vec{\nabla} s}{1+s}$$

radiation pressure

$\vec{F}_{\text{rad}}$

$\downarrow$

$$\vec{k}_L \cdot \vec{r} \cdot \langle \rho_{ee} \rangle_{\text{st}} \rightarrow \text{steady state}$$

$$s(\vec{r}) = \frac{\Omega^2(\vec{r})/2}{\Delta^2 + \Gamma^2/4}$$

## IV Interpretation

### 1) Dipole force

$$\vec{F}_{\text{dip}} = -\vec{\nabla} \left( -\frac{\Delta}{2} \ln(1+s(\vec{r})) \right)$$

dipole force is conservative and  $U_{\text{dip}} = -\frac{\Delta}{2} \ln(1+s)$  <sup>→ potential</sup>

Limiting case:  $\Delta \gg \Omega, \Gamma$

$$s \ll 1 \Rightarrow \langle p_{ee} \rangle \ll 1$$

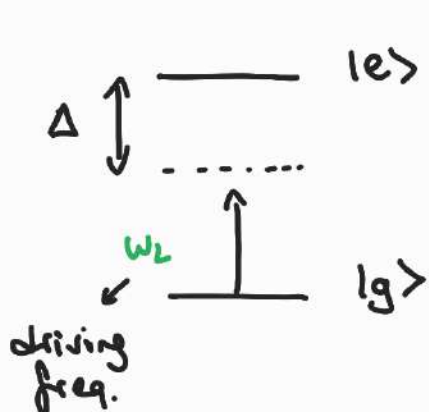
$$\text{Then } U_{\text{dip}} = -\frac{s}{2} \Delta = -\frac{\Omega^2(\vec{r})}{4\Delta}$$

•  $\Omega \propto \mathcal{E}$  E-field amplitude

$U_{\text{dip}} \propto I$  light intensity

•  $U_{\text{dip}}$  is also independent of  $\Gamma$ .

Recovering the dipole potential: 2-level atom in external field.

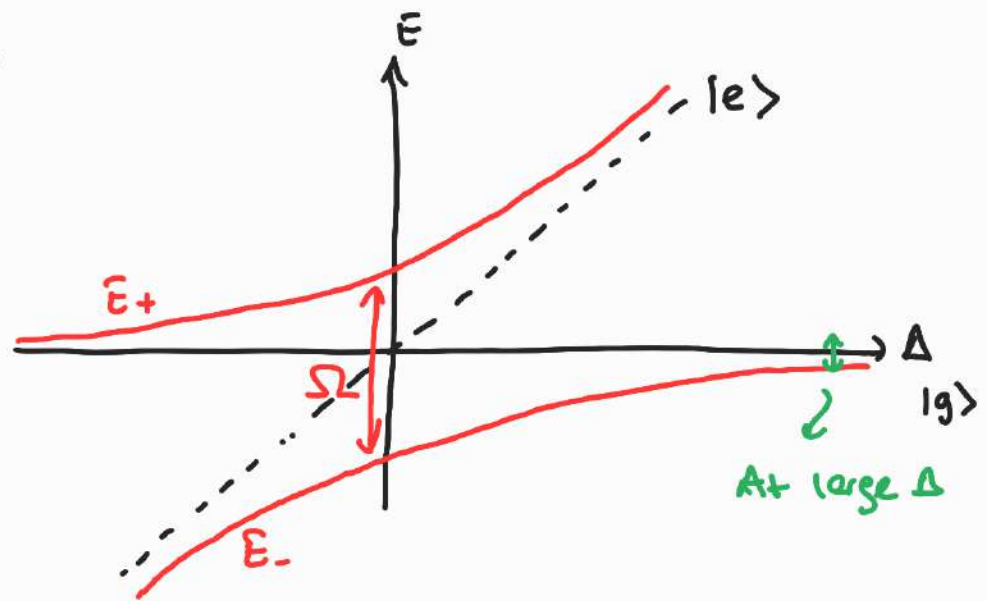


$$\hat{H} = \omega_e |e\rangle\langle e| + \frac{\Omega}{2} (e^{-i\omega_L t} |e\rangle\langle g| + \text{h.c.})$$

In a rot. frame:

$$\hat{H} = \Delta |e\rangle\langle e| + \frac{\Omega}{2} (|e\rangle\langle g| + \text{h.c.})$$

Energy spectrum:



Large  $\Delta$        $E_{\pm} = \frac{1}{2} (\Delta \pm \sqrt{\Delta^2 + \Omega^2})$        $E_- \rightarrow -\frac{\Omega^2}{4\Delta}$

and  $|-\rangle \rightarrow |g\rangle$

Now if  $\Omega(\vec{r}) \rightarrow E_-(\vec{r})$

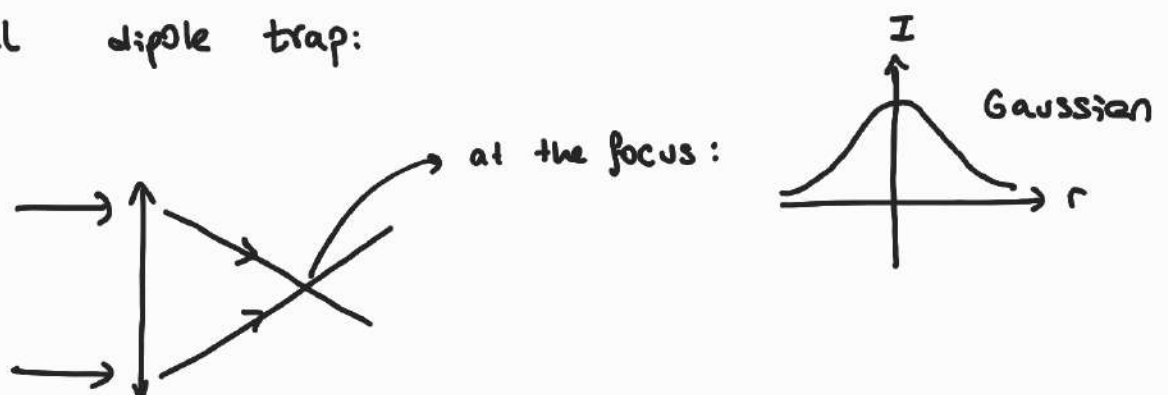
### Remarks

\*  $U_{\text{dip}} \Leftrightarrow$  AC-stark shift

(can also be obtained directly using 2<sup>nd</sup> order perturbation theory.)

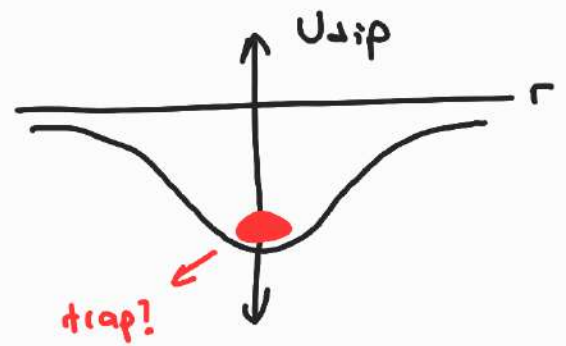
\* Also possible for DC fields:  $\Delta \rightarrow \omega_e$

\* Optical dipole trap:

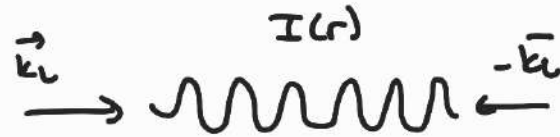


for  $\Delta \geq 0$ : ( $\omega_e \geq \omega_L$ )

"optical tweezers"



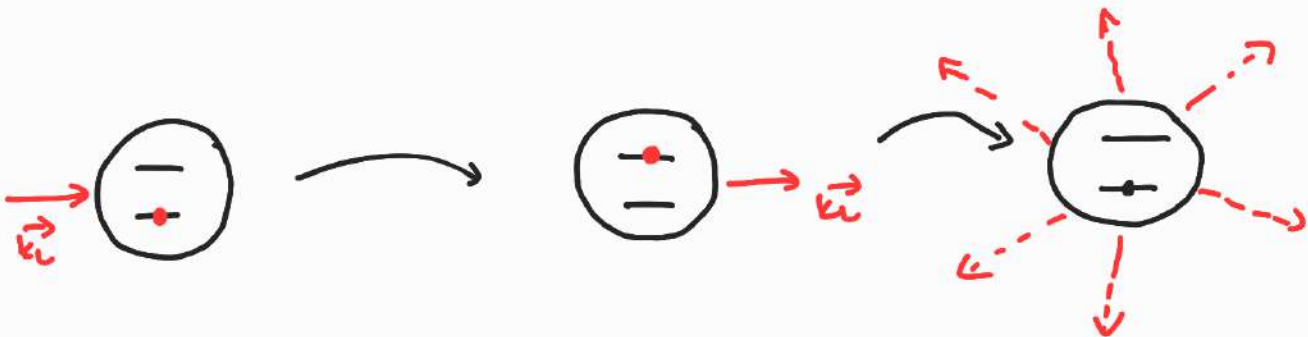
Optical lattices



$$U_{\text{dip}} = U_0 \cos^2(k_L r) \Rightarrow \text{band structure}$$

2) Radiation pressure

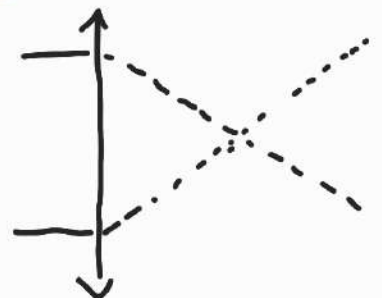
$$\vec{F} = \Gamma \langle p_{ee} \rangle^{\text{st}} \vec{k}$$



Remarks: Intrinsically related to spontaneous emission

$\Leftrightarrow$  incoherent

\*Dipole force: **stimulated emission**





- Radiation pressure is related to photon absorption:

### 3) Doppler cooling

An atom moving with momentum  $\vec{p}$

$$\Delta_0 \rightarrow \Delta_0 - \vec{p} \cdot \vec{v}_{\text{rec}} - E_{\text{rec}} \quad \swarrow \text{negligible}$$

$$\vec{F}_{\text{rad}} = \vec{k}_L \frac{\Gamma \Omega^2}{4} \frac{1}{\Delta_0^2 + \Gamma'^2} \quad \left. \vphantom{\frac{1}{\Delta_0^2 + \Gamma'^2}} \right\} \Gamma' = \text{power broadened linewidth}$$

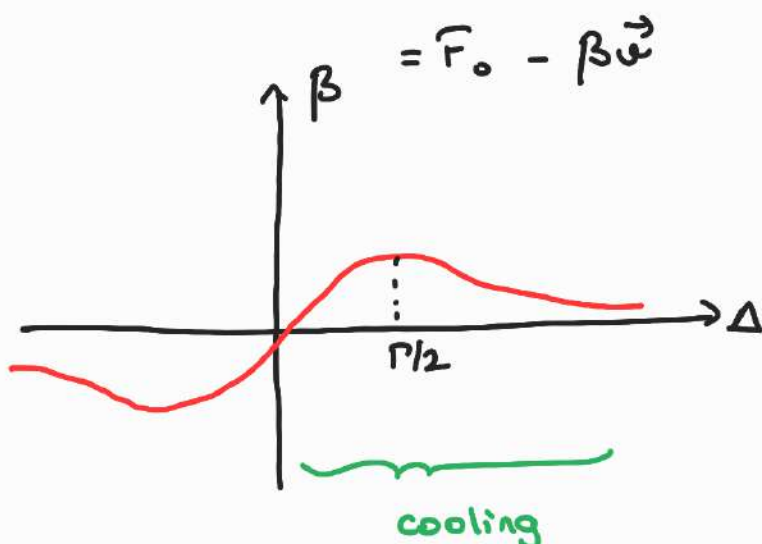
$$= \vec{k}_L \frac{\Gamma \Omega^2}{4} \frac{1}{(\Delta_0 + \vec{p} \cdot \vec{v}_{\text{rec}})^2 + \frac{\Gamma'^2}{2}}$$

First order expansion for Doppler shift:

$$\vec{F}_{\text{rad}} = \vec{k}_L \frac{\Gamma \Omega^2}{4} \left( \frac{1}{\Delta_0^2 + \Gamma'^2} - \vec{p} \cdot \vec{v}_{\text{rec}} \frac{2\Delta_0}{(\Delta_0^2 + \Gamma'^2)^2} + \dots \right)$$

$$\beta = \vec{F}_0 - \beta \vec{v}$$

$\beta$ : friction coefficient



Interpretation: Doppler shift bring the atom closer to resonance when it moves against the beam.



Remarks: to cool in all directions, you need 6 laser beams



#### 4) Limitations and Extensions

Limit to Doppler cooling: terms  $O(k_e^2)$  in the Lindblad equation.

. Random walk in  $k$  (momentum) space  $\langle \Delta p^2 \rangle \sim 2m\hbar\Gamma$

. Limit temperature:  $kT \sim \hbar\Gamma$   $\left( \begin{array}{l} {}^6\text{Li} = 280 \text{ nK} \\ {}^{87}\text{Sr} = 400 \text{ nK} \end{array} \right.$

the finite  $k_e \Rightarrow$  can not reach  $\vec{v}=0$  in free space.

$E_{\text{rec}}$  is the hard limit.

For trapped atoms  $\rightarrow$  reaching the ground state is possible.

Other mechanism to cool in free space: Sisyphus cooling

