

The EMGB solution that will be employed as a thin-shell solution with a normal matter [6] is given by (with  $\Lambda = 0$ )

$$f(r) = 1 + \frac{r^2}{4\alpha} \left( 1 - \sqrt{1 + \frac{8\alpha}{r^4} \left( \frac{2M}{\pi} - \frac{Q^2}{3r^2} \right)} \right) \quad (14)$$

with constants,  $M$  =mass and  $Q$  =charge. For a black hole solution the inner ( $r_-$ ) and event horizons ( $r_+ = r_h$ ) are

$$r_{\pm} = \sqrt{\frac{M}{\pi} - \alpha \pm \left[ \left( \frac{M}{\pi} - \alpha \right)^2 - \frac{Q^2}{3} \right]^{1/2}}. \quad (15)$$

By employing the solution (14) we determine the surface energy-momentum on the thin-shell, which will play the major role in the perturbation. We shall address this problem in the next section.

### III. RADIAL, LINEAR PERTURBATION OF THE THIN-SHELL WORMHOLE WITH NORMAL MATTER

In order to study the radial perturbations of the wormhole we take the throat radius as a function of the proper time, i.e.,  $a = a(\tau)$ . Based on the generalized Birkhoff theorem, for  $r > a(\tau)$  the geometry will be given still by (6). For the metric function  $f(r)$  given in (14) one finds the energy density and pressures as [6]

$$\sigma = -S_{\tau}^{\tau} = -\frac{\Delta}{4\pi} \left[ \frac{3}{a} - \frac{4\alpha}{a^3} (\Delta^2 - 3(1 + \dot{a}^2)) \right], \quad (16)$$

$$S_{\theta}^{\theta} = S_{\phi}^{\phi} = S_{\psi}^{\psi} = p = \frac{1}{4\pi} \left[ \frac{2\Delta}{a} + \frac{\ell}{\Delta} - \frac{4\alpha}{a^2} \left( \ell\Delta - \frac{\ell}{\Delta} (1 + \dot{a}^2) - 2\ddot{a}\Delta \right) \right], \quad (17)$$

where  $\ell = \ddot{a} + f'(a)/2$  and  $\Delta = \sqrt{f(a) + \dot{a}^2}$  in which

$$f(a) = 1 + \frac{a^2}{4\alpha} \left( 1 - \sqrt{1 + \frac{8\alpha}{a^4} \left( \frac{2M}{\pi} - \frac{Q^2}{3a^2} \right)} \right). \quad (18)$$

We note that in our notation a 'dot' denotes derivative with respect to the proper time  $\tau$  and a 'prime' implies differentiation with respect to the argument of the function. By a simple substitution one can show that, the conservation equation

$$\frac{d}{d\tau} (\sigma a^3) + p \frac{d}{d\tau} (a^3) = 0. \quad (19)$$