$$\Lambda_c(2595)^+$$

$$I(J^P) = 0(\frac{1}{2})$$
 Status: \*\*\*

The  $\Lambda_c^+\pi^+\pi^-$  mode is largely, and perhaps entirely,  $\Sigma_c\pi$ , which is just at threshold; since the  $\Sigma_c$  has  $J^P=1/2^+$ , the  $J^P$  here is almost certainly  $1/2^-$ . This result is in accord with the theoretical expectation that this is the charm counterpart of the strange  $\Lambda(1405)$ .

#### $\Lambda_{c}(2595)^{+}$ MASS

The mass is obtained from the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass-difference measurements below.

VALUE (MeV)

DOCUMENT ID

2592.25 ± 0.28 OUR FIT

# $\Lambda_c(2595)^+ - \Lambda_c^+$ MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID		TECN	COMMENT
305.79 ± 0.24 OUR F	ΊΤ				
$305.79 \pm 0.14 \pm 0.20$	3.5k	AALTONEN	11H	CDF	<i>p</i> <del>p</del> at 1.96 TeV
• • • We do not use	the following	data for averages	, fits,	limits, e	tc. • • •
$305.6 \pm 0.3$		$^{ m 1}$ BLECHMAN	03		Threshold shift
$309.7 \pm 0.9 \pm 0.4$	19	ALBRECHT	97	ARG	$e^+e^-pprox$ 10 GeV
$309.2 \pm 0.7 \pm 0.3$	$14\pm4.5$	FRABETTI	96	E687	$\gamma{ m Be},\ \overline{\it E}_{\gamma}pprox$ 220 GeV
$307.5 \pm 0.4 \pm 1.0$	$112\pm17$	<b>EDWARDS</b>	95	CLE2	$e^+e^-\stackrel{'}{pprox}$ 10.5 GeV

 $<sup>^1</sup>$  BLECHMAN 03 finds that a more sophisticated treatment than a simple Breit-Wigner for the proximity of the threshold of the dominant decay,  $\Sigma_c(2455)\pi$ , lowers the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass difference by 2 or 3 MeV. The analysis of AALTONEN 11H bears this out.

#### $\Lambda_{c}(2595)^{+}$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID		TECN	COMMENT
$2.59 \pm 0.30 \pm 0.47$	3.5k	<sup>2</sup> AALTONEN	11H	CDF	$p\overline{p}$ at 1.96 TeV
• • • We do not use t	he following da	ata for averages, fi	its, lir	nits, etc	. • • •
$2.9 \begin{array}{ccc} +2.9 & +1.8 \\ -2.1 & -1.4 \end{array}$	19	ALBRECHT	97	ARG	$e^+e^-pprox~10~{\rm GeV}$
$3.9 \begin{array}{c} +1.4 & +2.0 \\ -1.2 & -1.0 \end{array}$	$112\pm17$	EDWARDS	95	CLE2	$e^+e^-\approx~10.5~\text{GeV}$
	$(2455)^0 \pi^+$ se	eparately in terms	of a	commor	$ ightarrow$ $\Sigma_c$ (2455) $^{++}$ $\pi^-$ , $\Gamma$ coupling constant $h_2$ d.

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## $\Lambda_c(2595)^+$ DECAY MODES

 $\Lambda_c^+\pi\pi$  and its submode  $\Sigma_c(2455)\pi$  — the latter just barely — are the only strong decays allowed to an excited  $\Lambda_c^+$  having this mass; and the submode seems to dominate.

	Mode	Fraction $(\Gamma_i/\Gamma)$	
$\overline{\Gamma_1}$	$\Lambda_c^+ \pi^+ \pi^-$	[a] —	
$\Gamma_2$	$\Sigma_c(2455)^{++}\pi^-$	24 $\pm$ 7 %	
$\Gamma_3$	$\Sigma_{c}(2455)^{0}\pi^{+}$	24 $\pm$ 7 %	
$\Gamma_4$	$\Lambda_c^+ \pi^+ \pi^-$ 3-body	18 $\pm$ 10 %	
$\Gamma_5$	$\Lambda_c^+ \pi^+ \pi^-$ 3-body $\Lambda_c^+ \pi^0$ $\Lambda_c^+ \gamma$	[b] not seen	
$\Gamma_6$	$\Lambda_c^+ \gamma$	not seen	

- [a] See AALTONEN 11H, Fig. 8, for the calculated ratio of  $\Lambda_c^+\pi^0\pi^0$  and  $\Lambda_c^+\pi^+\pi^-$  partial widths as a function of the  $\Lambda_c(2595)^+-\Lambda_c^+$  mass difference. At our value of the mass difference, the ratio is about 4.
- [b] A test that the isospin is indeed 0, so that the particle is indeed a  $\Lambda_c^+$ .

## $\Lambda_c(2595)^+$ BRANCHING RATIOS

$\Gamma(\Sigma_c(2455)^{++}\pi^-)$	$/\Gamma(\Lambda_c^+\pi^-$	$^{+}\pi^{-})$			$\Gamma_2/\Gamma_1$
VALUE		DOCUMENT ID		TECN	COMMENT
0.36±0.10 OUR AVER	AGE				
$0.37 \pm 0.12 \pm 0.13$		ALBRECHT			$e^+e^-pprox~10~{ m GeV}$
$0.36 \pm 0.09 \pm 0.09$		EDWARDS	95	CLE2	$e^+e^-pprox~10.5~{\rm GeV}$
$\Gamma(\Sigma_c(2455)^0\pi^+)/\Gamma$	$(\Lambda_{a}^{+}\pi^{+}\eta$	r <sup>-</sup> )			$\Gamma_3/\Gamma_1$
VALUE	` ' '	DOCUMENT ID		TECN	COMMENT
0.37±0.10 OUR AVER	AGE			<u> </u>	
$0.29\!\pm\!0.10\!\pm\!0.11$		ALBRECHT	97	ARG	$e^+e^-pprox$ 10 GeV
$0.42\!\pm\!0.09\!\pm\!0.09$		<b>EDWARDS</b>	95	CLE2	$e^+e^-pprox~$ 10.5 GeV
<b>L</b> \	•	, , ,	, .		) $(\Gamma_2 + \Gamma_3)/\Gamma_1$
		DOCUMENT ID			
• • • We do not use the	ne tollowin	g data for average	s, fits	, limits,	etc. • • •
$0.66^{igoplus 0.13}_{-0.16}\!\pm\!0.07$					$e^+e^-pprox$ 10 GeV
>0.51	90	<sup>3</sup> FRABETTI	96	E687	$\gamma\mathrm{Be}$ , $\overline{E}_{\gamma} pprox \;$ 220 GeV
<sup>3</sup> The results of FRA					,
$\Gamma(\Lambda_c^+\pi^0)/\Gamma(\Lambda_c^+\pi^+$	$\pi^-)$				$\Gamma_5/\Gamma_1$
$\Lambda_c^+ \pi^0$ decay is for	orbidden b	y isospin conservat	tion if	this sta	te is in fact a $\Lambda_c$ .
<u>VALUE</u>	<u>CL%</u>	DOCUMENT ID		TECN	COMMENT
<3.53	90	<b>EDWARDS</b>	95	CLE2	$e^+e^-pprox~10.5~{ m GeV}$

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$\Gamma(\Lambda_c^+ \gamma) / \Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID		TECN	COMMENT
<0.98	90	EDWARDS	95	CLE2	$e^+e^-pprox~10.5~{ m GeV}$

# $\Lambda_c$ (2595)<sup>+</sup> REFERENCES

AALTONEN	11H	PR D84 012003	T. Aaltonen et al.	(CDF Collab.)
BLECHMAN	03	PR D67 074033	A.E. Blechman et al.	(JHU, FLOR)
ALBRECHT	97	PL B402 207	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
FRABETTI	96	PL B365 461	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
EDWARDS	95	PRL 74 3331	K.W. Edwards et al.	(CLEO Collab.)

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