Δ (2420) 11/2⁺

$$I(J^P) = \frac{3}{2}(\frac{11}{2}^+)$$
 Status: ***

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics **C38** 070001 (2014).

Δ (2420) POLE POSITION

| RFAI | PART |
|------|------|
| REAL | PARI |

| NEAL FAN I | | | | | |
|--|--------------------|----|-------|--------------------------------------|--|
| VALUE (MeV) | DOCUMENT ID | | TECN | COMMENT | |
| 2260 to 2400 (≈ 2330) OUR EST | IMATE | | | | |
| $2454 \pm 4 \pm 11$ | ¹ SVARC | 14 | L+P | $\pi N \rightarrow \pi N$ | |
| 2529 | ARNDT | 06 | DPWA | $\pi N \rightarrow \pi N$, ηN | |
| 2300 | HOEHLER | 93 | ARGD | $\pi N \rightarrow \pi N$ | |
| 2360 ± 100 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ | |
| -2×IMAGINARY PART | | | | | |
| VALUE (MeV) | DOCUMENT ID | | TECN | COMMENT | |
| 350 to 750 (\approx 550) OUR ESTIMA | NTE . | | | | |
| 462± 8±50 | ¹ SVARC | 14 | L + P | $\pi N \rightarrow \pi N$ | |
| 621 | ARNDT | 06 | DPWA | $\pi N \rightarrow \pi N, \eta N$ | |
| 620 | HOEHLER | 93 | ARGD | $\pi N \rightarrow \pi N$ | |
| 420 ± 100 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ | |
| | | | | | |

△(2420) ELASTIC POLE RESIDUE

MODULUS |r|

| VALUE (MeV) | DOCUMENT ID | | TECN | COMMENT |
|---|--------------------|----|------|--|
| 20 to 40 (≈ 30) OUR ESTIMATE | | | | |
| $30 \pm 1 \pm 7$ | ¹ SVARC | 14 | L+P | $\pi N \rightarrow \pi N$ |
| 33 | ARNDT | 06 | DPWA | $\pi N \rightarrow \pi N$, ηN |
| 39 | HOEHLER | 93 | ARGD | $\pi N \rightarrow \pi N$ |
| 18±6 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ |
| PHASE θ | | | | |
| VALUE (°) | DOCUMENT ID | | TECN | COMMENT |
| -60 to 20 (≈ -20) OUR ESTIM | MATE | | | |
| $11\pm \ 1\pm 8$ | ¹ SVARC | 14 | L+P | $\pi N \rightarrow \pi N$ |
| -45 | ARNDT | 06 | DPWA | π N $ ightarrow$ π N, η N |
| -60 | HOEHLER | 93 | ARGD | $\pi N \rightarrow \pi N$ |
| -30 ± 40 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ |

△(2420) BREIT-WIGNER MASS

| VALUE (MeV) | DOCUMENT ID | | TECN | COMMENT |
|---------------------------------|-------------|----|-------------|-----------------------------------|
| 2300 to 2500 (≈ 2420) OUR ESTIN | MATE | | | |
| 2633± 29 | ARNDT | 06 | DPWA | $\pi N \rightarrow \pi N, \eta N$ |
| 2400 ± 125 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ |
| 2416± 17 | HOEHLER | 79 | IPWA | $\pi N \rightarrow \pi N$ |
| | | | | |

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△(2420) BREIT-WIGNER WIDTH

| VALUE (MeV) DOCUMENT ID | TECN | COMMENT |
|---------------------------------|------|-----------------------------------|
| 300 to 500 (≈ 400) OUR ESTIMATE | | |
| 692± 47 ARNDT 06 | DPWA | $\pi N \rightarrow \pi N, \eta N$ |
| 450 ± 150 CUTKOSKY 80 | IPWA | $\pi N \rightarrow \pi N$ |
| 340± 28 HOEHLER 79 | IPWA | $\pi N \rightarrow \pi N$ |

△(2420) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

| | Mode | Fraction (Γ_i/Γ) |
|-----------------------|--------|------------------------------|
| $\overline{\Gamma_1}$ | $N\pi$ | 5–15 % |

△(2420) BRANCHING RATIOS

| $I(N\pi)/I_{total}$ | | | | | l 1/l |
|----------------------|-------------|----|-------------|-----------------------------------|-------|
| VALUE (%) | DOCUMENT ID | | TECN | COMMENT | |
| 5 to 15 OUR ESTIMATE | | | | | |
| 8.5 ± 0.8 | ARNDT | 06 | DPWA | $\pi N \rightarrow \pi N, \eta N$ | |
| 8 ±3 | CUTKOSKY | 80 | IPWA | $\pi N \rightarrow \pi N$ | |
| 8.0 ± 1.5 | HOEHLER | 79 | IPWA | $\pi N \rightarrow \pi N$ | |
| | | | | | |

△(2420) FOOTNOTES

△(2420) REFERENCES

| PDG SVARC | 14 14 | CP C38 070001 PR C89 045205 | K. Olive <i>et al.</i> A. Svarc <i>et al.</i> | (PDG Collab.) |
|--------------|----------|--------------------------------|--|----------------|
| ARNDT | 06 | PR C74 045205 | R.A. Arndt et al. | (GWU) |
| HOEHLER | 93 | π N Newsletter 9 1 | G. Hohler | (KARL) |
| CUTKOSKY | 80 | Toronto Conf. 19 | R.E. Cutkosky et al. | (CMÙ, LBL) IJP |
| Also | | PR D20 2839 | R.E. Cutkosky et al. | (CMU, LBL) |
| HOEHLER | 79 | PDAT 12-1 | G. Hohler <i>et al.</i> | ` (KARLT) IJP |
| Also | | Toronto Conf. 3 | R. Koch | (KARLT) IJP |
| | | | | |

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 $^{^{1}\,\}mathrm{Fit}$ to the amplitudes of HOEHLER 79.