Geometry of Image Formation

Pin hole camera: $x = \frac{-fX}{Z}, y = \frac{-fY}{Z}$

Projection of lines
$$\mathbf{X} = \mathbf{A} + \lambda \mathbf{D}$$
: $\lim_{\lambda \to \infty} \mathbf{p} = f \frac{\mathbf{A} + \lambda \mathbf{D}}{A_Z + \lambda D_Z} = f \frac{\mathbf{D}}{D_Z}$

Projection of planes $\mathbf{X} \cdot \mathbf{N} = d$: $\lim_{Z \to \infty} \mathbf{p} \cdot \mathbf{N} = 0 \Rightarrow xN_x + yN_y + fN_z = 0$

Terrestrial perspective:

Ground plane has normal $\mathbf{N} = (0, 1, 0)$ and projects to y = 0. Camera height h_c ; an object of height δY with bottom at $(X, -h_c, Z)$ and top at $(X, \delta Y - h_c, Z)$. The bottom projects to $(fX/Z, -fh_c/Z)$ and top projects to $(fX/Z, f(\delta Y - h_c)/Z)$.

$Geometric\ Transformations$

Pose: the position and orientation of the object with respect to the camera. 6 degrees of freedom.

Shape: the coordinates of the points of an object relative to a coordinate frame on the object; invariant when the object undergoes rotations and translations.

Euclidean Transformations: rotations and translation are isometries or regid body motions that preserves distances between any pair of points.

Orthogonal Transformations:

preserves inner products: $\mathbf{a} \cdot \mathbf{b} = \psi(\mathbf{a}) \cdot \psi(\mathbf{b})$

preserves norms: $\|\mathbf{a}\| = \|\psi(\mathbf{a})\|$

are isometries:
$$(\psi(\mathbf{a}) - \psi(\mathbf{b})) \cdot (\psi(\mathbf{a}) - \psi(\mathbf{b})) = (\mathbf{a} - \mathbf{b}) \cdot (\mathbf{a} - \mathbf{b})$$

Orthogonal matrices:

$$\mathbf{A}' = \mathbf{A}^{-1}, \det(\mathbf{A})^2 = 1 \Rightarrow \det(\mathbf{A}) = +1 \text{ or } -1.$$

$$\begin{bmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{bmatrix} \begin{bmatrix} \cos \theta & \sin \theta \\ \sin \theta & -\cos \theta \end{bmatrix}$$

rotation, det = +1 rotation, det = -1

Group structure of isometries:

 $\psi(\mathbf{a}) = \mathbf{A}\mathbf{a} + \mathbf{t}$, **A** is orthogonal matrix.

$$\psi_1(\mathbf{a}) = \mathbf{A}_1 \mathbf{a} + \mathbf{t}_1, \psi_2(\mathbf{a}) = \mathbf{A}_2 \mathbf{a} + \mathbf{t}_2,$$

$$\psi_1 \circ \psi_2(\mathbf{a}) = (\mathbf{A}_1 \mathbf{A}_2) \mathbf{a} + (\mathbf{A}_1 \mathbf{t}_2 + \mathbf{t}_1),$$

Skew-symmetric matrix: S = -S'

Cross product Matrix:

$$\begin{bmatrix} a_1 \\ a_2 \\ a_3 \end{bmatrix} \wedge \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix} = \begin{bmatrix} a_2b_3 - a_3b_2 \\ a_3b_1 - a_1b_3 \\ a_1b_2 - a_2b_1 \end{bmatrix} \Rightarrow \hat{\mathbf{a}} := \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix}$$

Roderigues Formula:

 $\mathbf{R} = e^{\phi \hat{\mathbf{s}}} = \mathbf{I} + \sin \phi \hat{\mathbf{s}} + (1 - \cos \phi) \hat{\mathbf{s}}^2.$

s is a unit vector along ω and $\phi = ||\omega||t$ is the total amount of rotation.

Affine transformations is a nonsingular linear transformation followed by a translation. $\psi(\mathbf{a}) = \mathbf{A}\mathbf{a} + \mathbf{t} \ (\det \mathbf{A} \neq 0)$ and preserves parallelism and midpoints and does **not** preserve lengths, angles, areas.

Ex. rotation, anisotropic scaling, sheer.

Degrees of freedom:

2D: isometry: 1 + 2 = 3, affine: 4 + 2 = 63D: isometry: 3 + 3 = 6, affine: 9 + 3 = 12

Optical Flow

Motion at 3D world projects to motion in the image. At every point (x, y) in the image we get a 2D vector, corresponding to the motion of the feature located at that point.

egomotion is the optical flow field of a moving observer.

 ${\bf t}$ is translational velocity of camera, ${\boldsymbol \omega}$ is angular velocity of camera

$$\begin{bmatrix} u \\ v \end{bmatrix} = \frac{1}{Z} \begin{bmatrix} -1 & 0 & x \\ 0 & -1 & y \end{bmatrix} \begin{bmatrix} t_x \\ t_y \\ t_z \end{bmatrix} + \begin{bmatrix} xy & -(1+x^2) & y \\ 1+y^2 & -xy & -x \end{bmatrix} \begin{bmatrix} \omega_x \\ \omega_y \\ \omega_z \end{bmatrix}$$

Radiometry of Image Formation

Irradiance:

I(x,y) measures how much light is captured at pixel (x,y). Radiant power per unit area with units W/m^2 , denoted by E.

$$L = \frac{\text{Power}}{(dA\cos\theta)(d\omega)}$$
, units are $Wm^{-2}sr^{-1}$.

Lambertian model: $L = \rho \lambda n \cdot s$