

Be(ause  $\vec{\nabla} \times \vec{H} = 0$ , we can let  $\vec{H} = -\vec{\nabla} \times \vec{X}$ Since  $\vec{\nabla} \times (-\vec{\nabla} \times) = 0$ 

If 
$$\hat{H} = -\hat{\nabla} \chi$$
 then  $\hat{B} = -u\hat{\nabla} \chi$ 

by 
$$\vec{\nabla} \cdot \vec{\mathbf{g}} = \vec{\nabla} \cdot (-u \vec{\nabla} \chi) = -u \vec{\nabla} \chi = 0$$

In spherical courd,  $\nabla^2 X = 0$  has solution:

$$X = \sum_{t=0}^{\infty} \sum_{m=-t}^{t} \left( A_{tm} \Gamma^{1} + B_{tm} \Gamma^{(1+1)} \right) Y_{t}^{m} \left( \theta, \phi \right)$$

However problem is azimuthal symmetric, so [m=0]  $\rightarrow \chi = \sum_{i=0}^{\infty} (A_{i} \Gamma^{1} + B_{i} \Gamma^{(i+1)}) P_{i}(\cos\theta)$ 

or 
$$\chi = -\frac{1}{2} \int_{0}^{z} B_{0} \hat{z} \cdot dz' \hat{z} = -\frac{1}{2} B_{0} z$$

$$\chi^{ext} = -\frac{1}{u_0} \beta_0 r \cos \theta$$

as  $r \rightarrow \infty$ ,  $\chi \rightarrow \chi^{ext}$ .

By matching, theres any 
$$t=1$$

By symmetry, don't expect higher variation terms for retition terms for retition terms for retition terms.

then 
$$X^{\text{TD}}(r) = \left(-\frac{1}{46}B_{c}C\left(\frac{r}{C}\right) + B\left(\frac{c}{r}\right)^{2}\right) \cos \theta$$
 for  $r > c$ 

For 
$$A < r < C$$
:  $X = \sum_{l} (a_{l} r^{l} + b_{l} r^{-(l+1)}) P_{l}(os\theta)$ 

Since for r>C, we expect only l=1 term, then we should expect the same behavior since the only difference between A< r< C and r>C is they have different u, but u is homogeneous throughout media. So  $X^{I}$  should have form:

$$\Rightarrow \left( x = \left( ar + br^{-2} \right) \cos \theta = \left( a \left( \frac{r}{A} \right) + b \left( \frac{A}{r} \right)^{2} \right) \cos \theta$$

For r<A: By the same argument, it should have form: x1(r) = (fr+ Yr-) 6.9A but in order to avoid divergence at 1-0, set 8=0 then  $\chi^{I}(\vec{r}) = fr \cos\theta = f f \cos\theta$ Boundary Cond:

Now since  $\vec{\nabla} \cdot \vec{B} = 0 \rightarrow \hat{n} \cdot \vec{B}|_{\vec{F} = C} = \hat{n} \cdot \vec{B}|_{\vec{F} = C}$  $\Rightarrow \hat{\mathbf{n}} \cdot - \mathbf{u}^{\top} \hat{\nabla} \hat{\mathbf{\chi}}_{r=c} = \hat{\mathbf{n}} \cdot - \mathbf{u}^{\top} \hat{\nabla} \hat{\mathbf{\chi}}_{r=c}$ since we have sphere, h= F  $|-u^{II}|_{\Gamma=C} = -u^{II}|_{\Gamma=C}$ and  $\vec{\nabla}_X \vec{H} = 0 \rightarrow \vec{R} \cdot \vec{H} = \vec{dr} \cdot \vec{H}$  $L_{\gamma} - \left( \nabla \chi^{T} \right)_{\mu} \Big|_{\Gamma = C} = -\left( \nabla \chi^{T} \right)_{\mu} \Big|_{\Gamma = C}$ Since problem independent of  $\phi$ , choose dr in  $\hat{\theta}$  direction.  $\left| -\frac{1}{\Gamma} \partial_{\theta} \chi^{\pi} \right|_{\Gamma=C} = -\frac{1}{\Gamma} \partial_{\theta} \chi^{\pi} \Big|_{\Gamma=C}$ then

At 
$$\underline{r} = A$$
  $\Rightarrow u^{1} \partial_{r} \chi^{1} \Big|_{r=A} = u^{1} \partial_{r} \chi^{1} \Big|_{r=A}$ 
 $u(\frac{a}{A} - 2b) \frac{1}{A^{3}} A^{3} = u_{0} \int_{\mathbb{R}^{+}} \chi^{1} \Big|_{r=A}$ 
 $f = \frac{M}{M_{0}} (a - 2b) = (|+\epsilon|)(a - 2b)$ 
 $|+\epsilon| = |+\epsilon|$ 
 $|+\epsilon| =$ 

then 
$$B = a \frac{C}{A} \left( 1 + (\frac{A}{C})^3 \frac{E}{3+2E} \right) + \frac{1}{N_0} B_0 C$$

$$= \frac{B_0 C}{N_0} \left[ \frac{-3 - 3(\frac{A}{C})^3 \frac{E}{3+2E}}{3 + (1 - 2(\frac{A}{C})^3 \frac{E}{3+2E})E} + 1 \right]$$

$$B = \frac{-B_0 C}{N_0} \left[ \frac{(3 + 2E)(\frac{A}{C})^3 \frac{E}{3+2E} - E}{3 + (1 - 2(\frac{A}{C})^3 \frac{E}{3+2E})E} \right]$$
then
$$\chi^1 = \frac{-3B_0 C}{N_0 A} \left[ \frac{(3 + 2E)(\frac{A}{C})^3 \frac{E}{3+2E}}{3 + 2E} \right] \left[ \frac{\Gamma}{A} + (\frac{A}{C})^2 \frac{E}{3+2E} \right] \cos\theta, \quad \Lambda < \Gamma < C$$

$$\chi^{11} = \frac{-3B_0 C}{N_0 A} \left[ \frac{\Gamma}{3} + (1 - 2(\frac{A}{C})^3 \frac{E}{3+2E})E \right] \left[ \frac{\Gamma}{A} + (\frac{A}{C})^2 \frac{E}{3+2E} \right] \cos\theta, \quad \Lambda < \Gamma < C$$

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多几色 stell. Find. B-field.

\$\forall \cdot \vec{\pi} \cdot \vec{\pi} \v 2) a)

For region inerste sphere, r<A

 $\vec{\nabla} \times \vec{B} = 0$ , so lets introduce scalar potential X where  $\vec{B} = -\vec{7}X$ 50 PX (PX) = 0

Then with  $\vec{\nabla} \cdot \vec{B} = -\vec{\nabla} \chi = 0$ 

Since problem has no \$\phi - dependence, \$\chi\$ follows:

Set  $P_{z} = 0$  to avoid singularity at  $r \rightarrow 0$ .

$$\rightarrow \chi = \Sigma \, \alpha_{t} \Gamma^{t} P_{t}(\cos \theta)$$
 for  $r < A$ .

For r>A, we still have  $\vec{J} \times \vec{B} = 0$ ,  $\vec{P} \cdot \vec{B} = 0$ 

But since we want  $X \rightarrow 0$  as  $r \rightarrow \infty$ , let  $[a_1 = 0]$ 

At boundary: 
$$\vec{\nabla} \cdot \vec{B} = \vec{D} \Rightarrow \hat{n} \cdot \vec{B} = \hat{n} \cdot \vec{B}$$

Find I: 
$$\hat{I} = \frac{d\hat{Q}\hat{\theta}}{d\hat{S}} = \frac{d\hat{Q}}{d\hat{S}} = \frac{d\hat{Q}}{$$

Since uniform distribution: 
$$6 = \frac{Q}{4\pi A^2}$$

then 
$$\vec{I} = \frac{Q}{4\pi A^2} A^2 \sin\theta \frac{d\phi d\theta}{dt} \hat{\theta}$$

then 
$$\vec{I} = \frac{Q}{4\pi} \sin\theta \Omega d\theta \hat{\theta}$$

then 
$$(B_{11}^{I} - B_{11}^{I}) A d = 2 + \frac{Q}{411} \sin \Omega d$$

4) 
$$\left( -\frac{1}{2} \partial_{\theta} \chi^{I} + \frac{1}{2} \partial_{\theta} \chi^{I} \right) = u_{\frac{Q}{4\pi A}} \Omega \sin \theta$$

Since we see 
$$\partial\theta \chi \propto \sin\theta$$
, this suggests

$$\chi \sim \cos\theta \propto P_{1}(\cos\theta), i.e. [I=1]$$

$$\Rightarrow so \quad \chi^{I} = \mathcal{A}(\overline{A})\cos\theta \qquad \Gamma < A$$

$$\chi^{I} = \mathcal{B}(\overline{A})^{2}\cos\theta \qquad \Gamma > A$$
Use  $\beta.C. \quad \partial_{\Gamma}\chi^{I}|_{\Gamma = A} = \partial_{\Gamma}\chi^{I}|_{\Gamma = A}$ 

$$\Rightarrow \quad \chi^{I} = -B2\frac{A^{2}}{A^{2}} = -2\beta\frac{A}{A}$$

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 $\vec{B}^{I} = \frac{16}{12\pi} Q \Omega A^{2} \frac{1}{r^{3}} \left( 2 \cos \hat{r} + \sin \hat{\theta} \right)$ 

c) For 
$$M = M_0 \approx = M_0 (\cos \theta \hat{r} - \sin \theta \hat{\theta})$$
 $K_b = M \times \hat{N}$ 
 $= M_0 (\cos \theta \hat{r} - \sin \theta \hat{\theta}) \times \hat{r}$ 
 $\vec{k}_b = M \cos \theta \hat{\theta}$ 
 $\vec{j}_A = \vec{j}_A \vec{M}$ 

$$= M_{o} \stackrel{?}{\nearrow} \times (cos\theta \hat{r} - sin\theta \hat{\theta})$$

$$= + \left( \frac{\partial}{\partial r} (-rsin\theta) - \frac{\partial}{\partial \theta} (cos\theta) \right) M_{o} \stackrel{?}{\phi}$$

$$= \frac{M_{o}}{r} \left( -sin\theta + sin\theta \right) \stackrel{?}{\phi}$$

we have 
$$\vec{\beta} \cdot \vec{B} = 0 \rightarrow \vec{\beta} \cdot \vec{H} = \vec{\beta} \cdot (\vec{a} \cdot \vec{B} - \vec{M}) = -\vec{\beta} \cdot \vec{M}$$

$$\vec{\nabla} \times \vec{B} = \mathcal{U}(\vec{\nabla} \times \vec{M}) = \mathcal{U}(\vec{\nabla} \times \vec{M}) = 0$$

Since 
$$\overrightarrow{\nabla} \times \overrightarrow{H} = 0$$
, introduce  $\overrightarrow{H} = -\overrightarrow{\nabla} X$ 

Then 
$$\vec{\forall} \cdot \vec{H} = - \vec{\forall} \chi = - \vec{\forall} \cdot \vec{M}$$

then 
$$\nabla^2 X = 0$$

By the same argument as we did in part a), we require azimuthal symmetry, non-divergent as r>0 for r<2 and convergence as  $r>\infty$ , then we have

But Boundary Condition is dightly different:

By B.C. 1, we know  $\partial r \times \sim caso$ , which is when 1=1

So 
$$\chi^{I} = A\left(\frac{\Gamma}{\alpha}\right)\cos\theta$$
  $\Gamma < \alpha$ 

$$\chi^{\pi} = B\left(\frac{2}{\Gamma}\right)^2 \cos\theta \qquad \Gamma > \lambda$$

Using B. C. 
$$\widehat{D}$$
:  $\overrightarrow{\vdash} \partial_{\theta} \chi^{1} | = \overrightarrow{\vdash} \partial_{\theta} \chi^{1} |_{\Gamma=2}$ 

We observe  $A = B$ 

Then with B.C.  $\widehat{D}$ :

$$(-A \frac{1}{3} - 2A \frac{1}{3}) \cos \theta = -M_{0} \cos \theta$$

$$then A = \frac{1}{3}M_{0} \times (\frac{1}{3}) \cos \theta \qquad \Gamma < \infty$$

$$\chi^{1} = \frac{1}{3}M_{0} \times (\frac{1}{3}) \cos \theta \qquad \Gamma > \infty$$

Then  $H = \frac{1}{3}\widehat{B} - \widehat{M} \qquad O\Gamma$ 

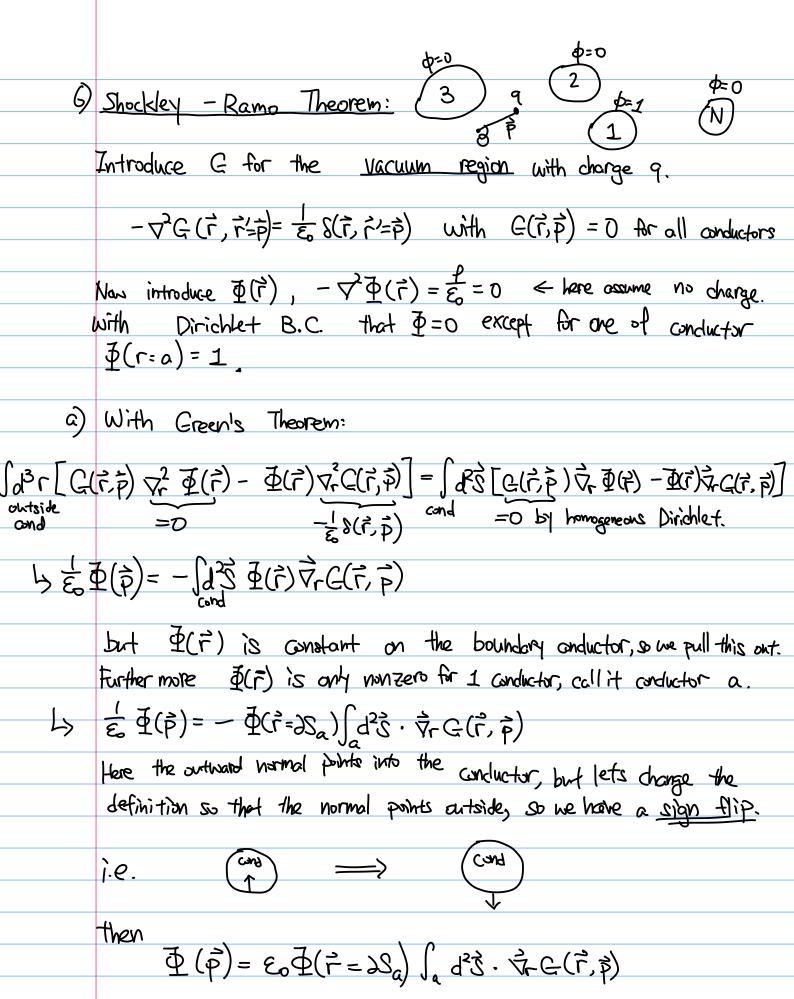
$$\widehat{B}^{1} = \mathcal{U}_{0}(-\overline{\nabla}\chi + \widehat{M})$$

$$= \mathcal{U}_{0}(-\overline{3}M_{0}(\cos \theta \widehat{\Gamma} - \sin \theta \widehat{\theta}) + M_{2}\widehat{2})$$

$$= \mathcal{U}_{0}(-\overline{3}M_{0}(\cos \theta \widehat{\Gamma} - \sin \theta \widehat{\theta}) + M_{2}\widehat{2})$$
Similarly:
$$\widehat{B}^{1} = \mathcal{U}_{0}(-\overline{\nabla}\chi + \widehat{M}) \qquad but \widehat{\Pi} = 0 \text{ outside sphere.}$$

$$= \mathcal{U}_{0}[-\frac{1}{3}M_{0}(\alpha^{2})^{3}(2\cos \theta \widehat{\Gamma} + \sin \theta \widehat{\theta})]$$

$$\widehat{B}^{1} = \frac{\eta_{0}}{3}M_{0}(\alpha^{2})^{3}(2\cos \theta \widehat{\Gamma} + \sin \theta \widehat{\theta}) \qquad for \Gamma > \infty$$



Now multiple q on both sides, and set 
$$\Phi(\vec{r} = 3)=1$$
.

$$9\Phi(\vec{p}) = \Phi(\vec{r}=\lambda S_a)$$
  $9E_0\int_{CM_a}d^2\vec{S}\cdot\vec{\nabla}_r C(\vec{r},\vec{p})$   
Analogus to  $E_0\int_{C}d^2\vec{S}\cdot\vec{\nabla}_r\vec{\Phi} = \int_{C}d^3r(-p)$   
 $= \int_{C}d^3r(-p)$ 

$$9 \overline{\Phi(\vec{p})} = \overline{\Phi(\vec{r} = \partial S_{k})}(-Q)$$

$$= \overline{1}$$

$$9 \overline{\Phi(\vec{p})} = -Q$$

$$D I = \frac{dQ}{dt} = \frac{1}{dt} \left( -9 \, \overline{D}(\overrightarrow{p}) \right)$$

$$\vec{E}(\vec{p}) = -\vec{p} \cdot \vec{p} \cdot$$